

Encyclopedia of Distances





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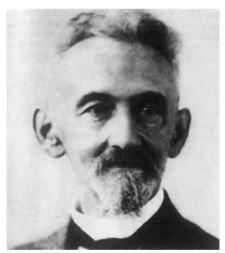
103 years ago, in 1906, Maurice Fréchet submitted his outstanding thesis Sur quelques points du calcul functionnel introducing (within a systematic study of functional operations) the notion of metric space (E-espace, E from écart).

Also, 95 years ago, in 1914, Felix Hausdorff published his famous Grundzüge der Mengenlehre where the theory of topological and metric spaces (metrische Räume) was created.

Let this Encyclopedia be our homage to the memory of these great mathematicians and their lives of dignity through the hard times of the first half of the twentieth century.



Maurice Fréchet (1878–1973) coined in 1906 the concept of écart (semi-metric)



Felix Hausdorff (1868–1942) coined in 1914 the term $metric\ space$

Preface

Encyclopedia of Distances is the result of re-writing and extending our Dictionary of Distances published in 2006 (and put online http://www.sciencedirect.com/science/book/9780444520876) by Elsevier. About a third of the definitions are new, and majority of the remaining ones have been upgraded.

We were motivated by the growing intensity of research on metric spaces and, especially, in distance design for applications. Even if we do not address the practical questions arising during the selection of a "good" distance function, just a sheer listing of the main available distances should be useful for the distance design community.

This Encyclopedia is the first one treating fully the general notion of distance. This broad scope is useful per se, but has limited our options for referencing. We have given an original reference for many definitions but only when it was not too difficult to do so. On the other hand, citing somebody who well developed the notion but was not the original author may induce problems. However, with our data (usually, author name(s) and year), a reader can easily search sources using the Internet.

We found many cases where authors developed very similar distances in different contexts and, clearly, were unaware of it. Such connections are indicated by a simple "cf." in both definitions, without going into priority issues explicitly.

Concerning the style, we have tried to make it a mixture of resource and coffee-table book, with maximal independence of its parts and many cross-references.

PREFACE TO DICTIONARY OF DISTANCES 2006

The concept of *distance* is basic to human experience. In everyday life it usually means some degree of closeness of two physical objects or ideas, i.e., length, time interval, gap, rank difference, coolness or remoteness, while the term *metric* is often used as a standard for a measurement.

But here we consider, except for the last two chapters, the mathematical meaning of those terms, which is an abstraction of measurement.

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The mathematical notions of distance metric (i.e., a function d(x,y) from $X \times X$ to the set of real numbers satisfying to $d(x,y) \ge 0$ with equality only for x = y, d(x,y) = d(y,x), and $d(x,y) \le d(x,z) + d(z,y)$) and of metric space (X,d) were originated a century ago by M. Fréchet (1906) and F. Hausdorff (1914) as a special case of an infinite topological space. The triangle inequality above appears already in Euclid. The infinite metric spaces are usually seen as a generalization of the metric |x-y| on the real numbers. Their main classes are the measurable spaces (add measure) and Banach spaces (add norm and completeness).

However, starting from K. Menger (who, in 1928, introduced metric spaces in Geometry) and L.M. Blumenthal (1953), an explosion of interest in both finite and infinite metric spaces occurred. Another trend is that many mathematical theories, in the process of their generalization, settled on the level of metric space. It is an ongoing process, for example, for Riemannian geometry, Real Analysis, Approximation Theory.

Distance metrics and distances have now become an essential tool in many areas of Mathematics and its applications including Geometry, Probability, Statistics, Coding/Graph Theory, Clustering, Data Analysis, Pattern Recognition, Networks, Engineering, Computer Graphics/Vision, Astronomy, Cosmology, Molecular Biology, and many other areas of science. Devising the most suitable distance metrics and similarities, to quantify the proximity between objects, has become a standard task for many researchers. Especially intense ongoing search for such distances occurs, for example, in Computational Biology, Image Analysis, Speech Recognition, and Information Retrieval.

Often the same distance metric appears independently in several different areas; for example, the edit distance between words, the evolutionary distance in Biology, the Levenstein distance in Coding Theory, and the Hamming+Gap or shuffle-Hamming distance.

This body of knowledge has become too big and disparate to operate in. The numbers of worldwide web entries offered by Google on the topics "distance," "metric space" and "distance metric" approach 300 million (i.e., about 2% of all), 6.5 million and 5.5 million, respectively, not to mention all the printed information outside the Web, or the vast "invisible Web" of searchable databases. However, this vast information on distances is too scattered: the works evaluating distance from some list usually treat very specific areas and are hardly accessible to non-experts.

Therefore many researchers, including us, keep and cherish a collection of distances for use in their areas of science. In view of the growing general need for an accessible interdisciplinary source for a vast multitude of researchers, we have expanded our private collection into this Dictionary. Some additional material was reworked from various encyclopedias, especially Encyclopedia of Mathematics [EM98], MathWorld [Weis99], PlanetMath [PM], and Wikipedia [WFE]. However, the majority of distances are extracted directly from specialist literature.

Preface

Besides distances themselves, we have collected many distance-related notions (especially in Chap. 1) and paradigms, enabling people from applications to get those (arcane for non-specialists) research tools, in ready-to-use fashion. This and the appearance of some distances in different contexts can be a source of new research.

In the time when over-specialization and terminology fences isolate researchers, this Dictionary tries to be "centripetal" and "ecumenical," providing some access and altitude of vision but without taking the route of scientific vulgarization. This attempted balance has defined the structure and style of the Dictionary.

This reference book is a specialized encyclopedic dictionary organized by subject area. It is divided into 29 chapters grouped into seven parts of about the same length. The titles of the parts are purposely approximative: they allow a reader to figure out her/his area of interest and competence. For example, Parts II, III and IV, V require some culture in, respectively, pure and applied Mathematics. Part VII can be read by a layman.

The chapters are thematic lists, by areas of Mathematics or applications, which can be read independently. When necessary, a chapter or a section starts with a short introduction: a field trip with the main concepts. Besides these introductions, the main properties and uses of distances are given, within items, in some instances. We also tried, when it was easy, to trace distances to their originator(s), but the proposed extensive bibliography has a less general ambition: just to provide convenient sources for a quick search.

Each chapter consists of items ordered in a way that hints of connections between them. All item titles and (with majiscules only for proper nouns) selected key terms can be traced in the large Subject Index; they are boldfaced unless the meaning is clear from the context. So, the definitions are easy to locate, by subject, in chapters and/or, by alphabetic order, in the Subject Index.

The introductions and definitions are reader-friendly and generally independent of each other; but they are interconnected, in the three-dimensional HTML manner, by hyperlink-like boldfaced references to similar definitions.

Many nice curiosities appear in this "Who is Who" of distances. Examples of such sundry terms are: ubiquitous Euclidean distance ("as-the-crow-flies"), flower-shop metric (shortest way between two points, visiting a "flower-shop" point first), knight-move metric on a chessboard, Gordian distance of knots, Earth Mover distance, biotope distance, Procrustes distance, lift metric, Post Office metric, Internet hop metric, WWW hyperlink quasi-metric, Moscow metric, and dogkeeper distance.

Besides abstract distances, the distances having physical meaning also appear (especially in Part VI); they range from $1.6 \times 10^{-35} \,\mathrm{m}$ (Planck length) to $4.3 \times 10^{26} \,\mathrm{m}$ (the estimated size of the observable Universe, about $27 \times 10^{60} \,\mathrm{Planck}$ lengths).

The number of distance metrics is infinite, and therefore our Dictionary cannot enumerate all of them. But we were inspired by several successful thematic dictionaries on other infinite lists; for example, on Numbers, x Preface

Integer Sequences, Inequalities, Random Processes, and by atlases of Functions, Groups, Fullerenes, etc. On the other hand, the large scope often forced us to switch to the mode of laconic tutorial.

The target audience consists of all researchers working on some measuring schemes and, to a certain degree, students and a part of the general public interested in science.

We have tried to address, even if incompletely, all scientific uses of the notion of distance. But some distances did not made it to this Dictionary due to space limitations (being too specific and/or complex) or our oversight. In general, the size/interdisciplinarity cut-off, i.e., decision where to stop, was our main headache. We would be grateful to readers who send us their favorite distances missed here. Four pages at the end are reserved for such personal additions.

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Part I Mathematics of Distances

Chapter 1 General Definitions

1.1 Basic definitions

Distance

Let X be a set. A function $d: X \times X \to \mathbb{R}$ is called a **distance** (or **dissimilarity**) on X if, for all $x, y \in X$, there holds:

- 1. $d(x,y) \ge 0$ (non-negativity).
- 2. d(x,y) = d(y,x) (symmetry).
- 3. d(x,x) = 0 (reflexivity).

In Topology, the distance d with d(x,y) = 0 implying x = y is called a **symmetric**. A distance which is a squared metric is called a *quadrance*.

For any distance d, the function D_1 , defined for $x \neq y$ by $D_1(x,y) = d(x,y)+c$, where $c = \max_{x,y,z \in X} (d(x,y)-d(x,z)-d(y,z))$, and D(x,x) = 0, is a **metric**. Also, $D_2(x,y) = d(x,y)^c$ is a metric for sufficiently small $c \geq 0$.

The function $D_3(x,y) = \inf \sum_i d(z_i, z_{i+1})$, where the infimum is taken over all sequences $x = z_0, \ldots, z_{n+1} = y$, is a **semi-metric**.

• Distance space

A distance space (X, d) is a set X equipped with a distance d.

• Similarity

Let X be a set. A function $s: X \times X \to \mathbb{R}$ is called a **similarity** on X if s is non-negative, symmetric, and if $s(x,y) \leq s(x,x)$ holds for all $x,y \in X$, with equality if and only if x = y.

The main transforms used to obtain a distance (dissimilarity) d from a similarity s bounded by 1 from above are: d = 1 - s, $d = \frac{1-s}{s}$, $d = \sqrt{1-s}$, $d = \sqrt{2(1-s^2)}$, $d = \arccos s$, $d = -\ln s$ (cf. Chap. 4).

• Semi-metric

Let X be a set. A function $d: X \times X \to \mathbb{R}$ is called a **semi-metric** (or **écart**) on X if d is non-negative, symmetric, if d(x, x) = 0 for all $x \in X$, and if

$$d(x,y) \le d(x,z) + d(z,y)$$

for all $x, y, z \in X$ (triangle or, sometimes, triangular inequality).

In Topology, it is called a **pseudo-metric**, while the term *semi-metric* is sometimes used for a **symmetric** (a distance d(x, y) with d(x, y) = 0 only if x = y) or for a special case of it; cf. **symmetrizable space** in Chap. 2.

For a semi-metric d, the triangle inequality is equivalent, for each fixed $n \geq 4$, to the following n-gon inequality

$$d(x,y) \le d(x,z_1) + d(z_1,z_2) + \dots + d(z_{n-2},y),$$

for all $x, y, z_1, ..., z_{n-2} \in X$.

For a semi-metric d on X, define an equivalence relation by $x \sim y$ if d(x,y) = 0; equivalent points are equidistant from all other points. Let [x] denote the equivalence class containing x; then D([x],[y]) = d(x,y) is a **metric** on the set $\{[x]: x \in X\}$ of classes.

• Metric

Let X be a set. A function $d: X \times X \to \mathbb{R}$ is called a **metric** on X if, for all $x, y, z \in X$, there holds:

- 1. $d(x,y) \ge 0$ (non-negativity).
- 2. d(x,y) = 0 if and only if x = y (identity of indiscernibles).
- 3. d(x,y) = d(y,x) (symmetry).
- 4. $d(x,y) \le d(x,z) + d(z,y)$ (triangle inequality).

In fact, 1 follows from 3 and 4.

• Metric space

A metric space (X, d) is a set X equipped with a metric d.

A pointed metric space (X, d, x_0) is a metric space (X, d) with a selected base point $x_0 \in X$.

• Metric scheme

A metric scheme is a metric space with an integer-valued metric.

• Extended metric

An **extended metric** is a generalization of the notion of metric: the value ∞ is allowed for a metric d.

• Quasi-distance

Let X be a set. A function $d: X \times X \to \mathbb{R}$ is called a **quasi-distance** on X if d is non-negative, and d(x,x) = 0 holds for all $x \in X$.

In Topology, it is also called a **parametric**.

For a quasi-distance d, the **strong triangle inequality** $d(x,y) \le d(x,z) + d(y,z)$ imply that d is symmetric and so, a semi-metric.

• Quasi-semi-metric

Let X be a set. A function $d: X \times X \to \mathbb{R}$ is called a **quasi-semi-metric** on X if d is non-negative, if d(x,x) = 0 for all $x \in X$, and if

$$d(x,y) \le d(x,z) + d(z,y)$$

for all $x, y, z \in X$ (oriented triangle inequality).

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The set X can be partially ordered by the specialization order: $x \leq y$ if and only if d(x, y) = 0.

A weak quasi-metric is a quasi-semi-metric d on X with weak symmetry, i.e., for all $x, y \in X$ the equality d(x, y) = 0 implies d(y, x) = 0.

An **Albert quasi-metric** is a quasi-semi-metric d on X with weak definiteness, i.e., for all $x, y \in X$ the equality d(x, y) = d(y, x) = 0 implies x = y.

A weightable quasi-semi-metric is a quasi-semi-metric d on X with relaxed symmetry, i.e., for all $x, y, z \in X$

$$d(x,y) + d(y,z) + d(z,x) = d(x,z) + d(z,y) + d(y,x),$$

holds or, equivalently, there exists a weight function $w(x) \in \mathbb{R}$ on X with d(x,y)-d(y,x)=w(y)-w(x) for all $x,y\in X$ (i.e., $d(x,y)+\frac{1}{2}(w(x)-w(y))$ is a semi-metric). If d is a weightable quasi-semi-metric, then d(x,y)+w(x) is a **partial semi-metric** (moreover, **partial metric** if d is an Albert quasi-metric).

• Partial metric

Let X be a set. A non-negative symmetric function $p: X \times X \to \mathbb{R}$ is called a **partial metric** [Matt92] if, for all $x, y, z \in X$, there holds:

- 1. $p(x,x) \le p(x,y)$ (i.e., every **self-distance** p(x,x) is *small*).
- 2. x = y if p(x, x) = p(x, y) = p(y, y) = 0 (T_0 separation axiom).
- 3. $p(x,y) \le p(x,z) + p(z,y) p(z,z)$ (sharp triangle inequality).

If above separation axiom is dropped, the function p is called a **partial semi-metric**. The function p is a partial semi-metric if and only if p(x,y) - p(x,x) is a **weightable quasi-semi-metric** with w(x) = p(x,x).

If above condition $p(x,x) \leq p(x,y)$ is also dropped, the function p is called (Heckmann 1999) a **weak partial semi-metric**.

Cf. distance from measurement in Chap. 3; it is related topologically (Waszkiewicz 2001) to partial metrics.

Sometimes, the term *partial metric* is used when a metric d(x, y) is defined only on a subset of the set of all pairs x, y of points.

• Quasi-metric

Let X be a set. A function $d: X \times X \to \mathbb{R}$ is called a **quasi-metric** on X if $d(x,y) \geq 0$ holds for all $x,y \in X$ with equality if and only if x=y, and

$$d(x,y) \le d(x,z) + d(z,y)$$

for all $x, y, z \in X$ (**oriented triangle inequality**). A *quasi-metric space* (X, d) is a set X equipped with a quasi-metric d.

For any quasi-metric d, the functions $\max\{d(x,y),d(y,x)\}$, $\min\{d(x,y),d(y,x)\}$ and $\frac{1}{2}(d^p(x,y)+d^p(y,x))^{\frac{1}{p}}$ with $p\geq 1$ (usually, p=1 is taken) are **equivalent metrics**.

A non-Archimedean quasi-metric d is a quasi-distance on X which satisfies the following strengthened version of the oriented triangle inequality:

$$d(x,y) \le \max\{d(x,z), d(z,y)\}$$

for all $x, y, z \in X$.

• Near-metric

Let X be a set. A distance d on X is called a **near-metric** (or weak metric) if

$$d(x,y) \le C(d(x,z) + d(z,y))$$

for all $x, y, z \in X$ and some constant C > 1 (C-triangle inequality).

Some recent papers use the term *quasi-triangle inequality* for above inequality and so, *quasi-metric* for the notion of near-metric.

The **power transform** (cf. Chap. 4) $(d(x,y))^{\alpha}$ of any near-metric is a near-metric for any $\alpha > 0$. Also, any near-metric d admits a **bi-Lipschitz mapping** on $(D(x,y))^{\alpha}$ for some semi-metric D on the same set and a positive number α .

A near-metric d on X is called a **Hölder near-metric** if the inequality

$$|d(x,y) - d(x,z)| \le \beta d(y,z)^{\alpha} (d(x,y) + d(x,z))^{1-\alpha}$$

holds for some $\beta > 0, \ 0 < \alpha \le 1$ and all points $x,y,z \in X$. Cf. **Hölder mapping**.

• Coarse-path metric

Let X be a set. A metric d on X is called a **coarse-path metric** if, for a fixed $C \ge 0$ and for every pair of points $x, y \in X$, there exists a sequence $x = x_0, x_1, \ldots, x_t = y$ for which $d(x_{i-1}, x_i) \le C$ for $i = 1, \ldots, t$, and

$$d(x,y) \ge d(x_0,x_1) + d(x_1,x_2) + \dots + d(x_{t-1},x_t) - C,$$

i.e., the weakened triangle inequality $d(x,y) \leq \sum_{i=1}^{t} d(x_{i-1},x_i)$ becomes an equality up to a bounded error.

• Weak ultrametric

A weak ultrametric (or C-pseudo-distance, C-inframetric) d is a distance on X such that for a constant $C \ge 1$ the inequality

$$0 < d(x,y) \leq C \max\{d(x,z),d(z,y)\}$$

holds for all $x, y, z \in X, x \neq y$.

The term **pseudo-distance** is also used, in some applications, for any of a **pseudo-metric**, a **quasi-distance**, a **near-metric**, a distance which can be infinite, a distance with an error, etc.

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• Ultrametric

An ultrametric (or non-Archimedean metric) is (Krasner 1944) a metric d on X which satisfies the following strengthened version of the triangle inequality (Hausdorff 1934), called the ultrametric inequality:

$$d(x,y) \le \max\{d(x,z), d(z,y)\}$$

for all $x, y, z \in X$. So, at least two of d(x, y), d(z, y) and d(x, z) are the same.

A metric d is an ultrametric if and only if its **power transform** (see Chap. 4) d^{α} is a metric for any real positive number α . Any ultrametric satisfies the **four-point inequality**. A metric d is an ultrametric if and only if it is a **Farris transform** (cf. Chap. 4) of a **four-point inequality metric**.

For a finite set X, a symmetric non-negative matrix $A = (A(x,y) : x, y \in X)$ is called *ultrametric* if there exists an ultrametric d on X such that $d(x,y) \leq d(x,z)$ implies $A(x,y) \geq A(x,z)$.

• Robinsonian distance

A distance d on X is called a **Robinsonian distance** (or *monotone distance*) if there exists a total order \leq on X compatible with it, i.e., for $x, y, w, z \in X$,

$$x \leq y \leq w \leq z$$
 implies $d(y, w) \leq d(x, z)$,

or, equivalently, for $x, y, z \in X$,

$$x \leq y \leq z$$
 implies $d(x, y) \leq \max\{d(x, z), d(z, y)\}.$

Any **ultrametric** is a Robinsonian distance.

• Four-point inequality metric

A metric d on X is a **four-point inequality metric** (or **additive metric**) if it satisfies the following strengthened version of the triangle inequality called the **four-point inequality**: for all $x, y, z, u \in X$

$$d(x,y) + d(z,u) \le \max\{d(x,z) + d(y,u), d(x,u) + d(y,z)\}$$

holds. Equivalently, among the three sums d(x,y)+d(z,u), d(x,z)+d(y,u), d(x,u)+d(y,z) the two largest sums are equal.

A metric satisfies the four-point inequality if and only if it is a **tree-like** metric.

Any metric, satisfying the four-point inequality, is a **Ptolemaic metric** and an L_1 -metric (cf. L_p -metric in Chap. 5).

A **bush metric** is a metric for which all four-point inequalities are equalities, i.e., d(x,y) + d(u,z) = d(x,u) + d(y,z) holds for any $u,x,y,z \in X$.

• Relaxed four-point inequality metric

A metric d on X satisfies the **relaxed four-point inequality** if, for all $x, y, z, u \in X$, among the three sums

$$d(x,y) + d(z,u), d(x,z) + d(y,u), d(x,u) + d(y,z)$$

at least two (not necessarily the two largest) are equal.

A metric satisfies the relaxed four-point inequality if and only if it is a relaxed tree-like metric.

• Ptolemaic metric

A **Ptolemaic metric** d is a metric on X which satisfies the **Ptolemaic inequality**

$$d(x,y)d(u,z) \le d(x,u)d(y,z) + d(x,z)d(y,u)$$

(shown by Ptolemy to hold in the Euclidean space) for all $x, y, u, z \in X$.

A Ptolemaic space is a normed vector space (V, ||.||) such that its norm metric ||x - y|| is a Ptolemaic metric. A normed vector space is a Ptolemaic space if and only if it is an **inner product space** (cf. Chap. 5); so, a **Minkowskian metric** (cf. Chap. 6) is Euclidean if and only if it is Ptolemaic.

The involution space $(X \setminus z, d_z)$, where $d_z(x, y) = \frac{d(x, y)}{d(x, z)d(y, z)}$, is a metric space, for any $z \in X$, if and only if d is Ptolemaic [FoSC06].

For any metric d, the metric \sqrt{d} is Ptolemaic [FoSC06].

• δ -hyperbolic metric

Given a number $\delta \geq 0$, a metric d on a set X is called δ -hyperbolic if it satisfies the **Gromov** δ -hyperbolic inequality (another weakening of the four-point inequality): for all $x, y, z, u \in X$

$$d(x,y)+d(z,u)\leq 2\delta+\max\{d(x,z)+d(y,u),d(x,u)+d(y,z)\}$$

holds. A metric space (X, d) is δ -hyperbolic if and only if

$$(x.y)_{x_0} \ge \min\{(x.z)_{x_0}, (y.z)_{x_0}\} - \delta$$

for all $x, y, z \in X$ and for any $x_0 \in X$, where $(x.y)_{x_0} = \frac{1}{2}(d(x_0, x) + d(x_0, y) - d(x, y))$ is the **Gromov product** of the points x and y of X with respect to the base point $x_0 \in X$.

A metric space (X, d) is 0-hyperbolic exactly when d satisfies the **four-point inequality**. Every bounded metric space of diameter D is D-hyperbolic. The n-dimensional hyperbolic space is $\ln 3$ -hyperbolic.

1.1 Basic definitions 9

Every δ -hyperbolic metric space is isometrically embeddable into a **geodesic metric space** (Bonk and Schramm 2000).

• Gromov product similarity

Given a metric space (X, d) with a fixed point $x_0 \in X$, the **Gromov product similarity** (or *Gromov product, covariance*) $(.)_{x_0}$ is a similarity on X, defined by

$$(x.y)_{x_0} = \frac{1}{2}(d(x,x_0) + d(y,x_0) - d(x,y)).$$

If (X,d) is a tree, then $(x.y)_{x_0} = d(x_0, [x,y])$. If (X,d) is a **measure semi-metric space**, i.e., $d(x,y) = \mu(x \triangle y)$ for a Borel measure μ on X, then $(x.y)_{\emptyset} = \mu(x \cap y)$. If d is a **distance of negative type**, i.e., $d(x,y) = d_E^2(x,y)$ for a subset X of an Euclidean space \mathbb{E}^n , then $(x.y)_0$ is the usual *inner product* on \mathbb{E}^n .

Cf. Farris transform metric in Chap. 4.

• Cross difference

Given a metric space (X, d) and quadruple (x, y, z, w) of its points, the **cross difference** is the real number cd defined by

$$cd(x, y, z, w) = d(x, y) + d(z, w) - d(x, z) - d(y, w).$$

For all $x, y, z, w, p \in X$,

$$\frac{1}{2}cd(x,y,z,w) = -(x.y)_p - (z.w)_p + (x.z)_p + (y.w)_p$$

in terms of the **Gromov product similarity**; in particular, it becomes $(x.y)_p$ if y = w = p.

Given a metric space (X, d) and quadruple (x, y, z, w) of its points with $x \neq z$ and $y \neq w$, the **cross-ratio** is the real number cr defined by

$$cr(x,y,z,w) = \frac{d(x,y)d(z,w)}{d(x,z)d(y,w)} \ge 0.$$

• 2k-gonal distance

A 2k-gonal distance d is a distance on X which satisfies the 2k-gonal inequality

$$\sum_{1 \le i < j \le n} b_i b_j d(x_i, x_j) \le 0$$

for all $b \in \mathbb{Z}^n$ with $\sum_{i=1}^n b_i = 0$ and $\sum_{i=1}^n |b_i| = 2k$, and for all distinct elements $x_1, \ldots, x_n \in X$.

• Distance of negative type

A distance of negative type d is a distance on X which is 2k-gonal for any $k \ge 1$, i.e., satisfies the negative type inequality

$$\sum_{1 \le i < j \le n} b_i b_j d(x_i, x_j) \le 0$$

for all $b \in \mathbb{Z}^n$ with $\sum_{i=1}^n b_i = 0$, and for all distinct elements $x_1, \ldots, x_n \in X$.

A distance can be of negative type without being a semi-metric. Cayley proved that a metric d is an L_2 -metric if and only if d^2 is a distance of negative type.

• (2k+1)-gonal distance

A (2k+1)-gonal distance d is a distance on X which satisfies the (2k+1)-gonal inequality

$$\sum_{1 \le i < j \le n} b_i b_j d(x_i, x_j) \le 0$$

for all $b \in \mathbb{Z}^n$ with $\sum_{i=1}^n b_i = 1$ and $\sum_{i=1}^n |b_i| = 2k+1$, and for all distinct elements $x_1, \ldots, x_n \in X$.

The (2k+1)-gonal inequality with k=1 is the usual triangle inequality. The (2k+1)-gonal inequality implies the 2k-gonal inequality.

Hypermetric

A hypermetric d is a distance on X which is (2k + 1)-gonal for any $k \ge 1$, i.e., satisfies the hypermetric inequality

$$\sum_{1 \le i < j \le n} b_i b_j d(x_i, x_j) \le 0$$

for all $b \in \mathbb{Z}^n$ with $\sum_{i=1}^n b_i = 1$, and for all distinct elements $x_1, \ldots, x_n \in X$.

Any hypermetric is a semi-metric, a **distance of negative type** and, moreover, it can be isometrically embedded into some n-sphere \mathbb{S}^n with squared Euclidean distance. Any L_1 -metric (cf. L_p -metric in Chap. 5) is a hypermetric.

• P-metric

A P-metric d is a metric on X with values in [0,1] which satisfies the correlation triangle inequality

$$d(x,y) \le d(x,z) + d(y,z) - d(x,z)d(z,y).$$

The equivalent inequality $(1-d(x,y)) \ge (1-d(x,z))(1-d(z,y))$ expresses that the probability, say, to reach x from y via z is either equal to (1-d(x,z))(1-d(z,y)) (independence of reaching z from x and y from z), or greater than it (positive correlation).

A metric is a *P*-metric if and only if it is a **Schoenberg transform** metric (cf. Chap. 4).

1.2 Main distance-related notions

Metric ball

Given a metric space (X, d), the **metric ball** (or *closed metric ball*) with center $x_0 \in X$ and radius r > 0 is defined by $\overline{B}(x_0, r) = \{x \in X : d(x_0, x) \le r\}$, and the **open metric ball** with center $x_0 \in X$ and radius r > 0 is defined by $B(x_0, r) = \{x \in X : d(x_0, x) < r\}$.

The **metric sphere** with center $x_0 \in X$ and radius r > 0 is defined by $S(x_0, r) = \{x \in X : d(x_0, x) = r\}.$

For the **norm metric** on an *n*-dimensional normed vector space (V, ||.||), the metric ball $\overline{B}^n = \{x \in V : ||x|| \le 1\}$ is called the unit ball, and the set $S^{n-1} = \{x \in V : ||x|| = 1\}$ is called the unit sphere (or unit hypersphere). In a two-dimensional vector space, a metric ball (closed or open) is called a **metric disk** (closed or open, respectively).

• Distance-invariant metric space

A metric space (X, d) is **distance-invariant** if all **metric balls** $\overline{B}(x_0, r) = \{x \in X : d(x_0, x) \leq r\}$ of the same radius have the same number of elements.

• Closed subset of metric space

Given a subset M of a metric space (X,d), a point $x \in X$ is called a *limit* point of M (or accumulation point) if every **open metric ball** $B(x,r) = \{y \in X : d(x,y) < r\}$ contains a point $x' \in M$ with $x' \neq x$. The closure of M, denoted by \overline{M} , is the set M together with all its limit points. The subset M is called **closed** if $M = \overline{M}$.

A closed subset M is **perfect** if every point of M is a limit point of M. Every point of M which is not a limit point of M, is called an *isolated point*. The $interior\ int(M)$ of M is the set of all its isolated points; the $exterior\ ext(M)$ of M is $int(X\backslash M)$ and the $boundary\ \vartheta(M)$ of M is $X\backslash (int(M)\cup ext(M))$.

A subset M is called **topologically discrete** if M = int(M).

• Open subset of metric space

A subset M of a metric space (X,d) is called *open* if, given any point $x \in M$, the **open metric ball** $B(x,r) = \{y \in X : d(x,y) < r\}$ is contained in M for some positive number r. The family of open subsets of a metric space forms a natural topology on it.

An open subset of a metric space is called *clopen* if it is **closed**. An open subset of a metric space is called a *domain* if it is **connected**.

A *door space* is a metric (in general, topological) space in which every subset is either open or closed.

• Connected metric space

A metric space (X, d) is called **connected** if it cannot be partitioned into two non-empty **open** sets (cf. **connected space** in Chap. 2).

(X,d) is distance m locally (path)-connected (extending Holub-Xiong, 2009) if any subspace ($\{y \in X: d(x,y) \in (0,m]\}, d$), $x \in X$, is (path)-connected. A **totally disconnected metric space** is a space in which all connected subsets are \emptyset and one-point sets.

A path-connected metric space is a connected metric space such that any two its points can be joined by an arc (cf. metric curve).

• Cantor connected metric space

A metric space (X,d) is called **Cantor connected** (or pre-connected) if, for any two its points x, y and any $\epsilon > 0$, there exists an ϵ -chain joining them, i.e., a sequence of points $x = z_0, z_1, \ldots, z_{n-1}, z_n = y$ such that $d(z_k, z_{k+1}) \leq \epsilon$ for every $0 \leq k \leq n$. A metric space (X,d) is Cantor connected if and only if it cannot be partitioned into two remote parts A and B, i.e., such that $\inf\{d(x,y): x \in A, y \in B\} > 0$.

The maximal Cantor connected subspaces of a metric space are called its *Cantor connected components*. A **totally Cantor disconnected metric** is the metric of a metric space in which all Cantor connected components are one-point sets.

• Indivisible metric space

A metric space (X, d) is called **indivisible** if it cannot be partitioned into two parts, neither of which contains an isometric copy of (X, d). Any indivisible metric space with $|X| \geq 2$ is infinite, bounded and **totally Cantor disconnected** (Delhomme, Laflamme, Pouzet and Sauer 2007).

A metric space (X, d) is called an **oscillation stable metric space** (Nguyen Van The 2006) if, given any $\epsilon > 0$ and any partition of X into finitely many pieces, the ϵ -neighborhood of one of the pieces includes an isometric copy of (X, d).

• Metric topology

A **metric topology** is a *topology* on X induced by a metric d on X; cf. **equivalent metrics**.

More exactly, given a metric space (X,d), define the *open set* in X as an arbitrary union of (finitely or infinitely many) open metric balls $B(x,r) = \{y \in X : d(x,y) < r\}, x \in X, r \in \mathbb{R}, r > 0$. A *closed set* is defined now as the complement of an open set. The metric topology on (X,d) is defined as the set of all open sets of X. A topological space which can arise in this way from a metric space is called a **metrizable space** (cf. Chap. 2).

Metrization theorems are theorems which give sufficient conditions for a topological space to be metrizable.

On the other hand, the adjective *metric* in several important mathematical terms indicates connection to a measure, rather than distance, for example, *metric Number Theory*, *metric Theory of Functions*, *metric transitivity*.

• Equivalent metrics

Two metrics d_1 and d_2 on a set X are called **equivalent** if they define the same topology on X, i.e., if, for every point $x_0 \in X$, every open metric ball with center at x_0 defined with respect to d_1 , contains an open metric ball with the same center but defined with respect to d_2 , and conversely.

Two metrics d_1 and d_2 are equivalent if and only if, for every $\epsilon > 0$ and every $x \in X$, there exists $\delta > 0$ such that $d_1(x,y) \leq \delta$ implies $d_2(x,y) \leq \epsilon$ and, conversely, $d_2(x,y) \leq \delta$ implies $d_1(x,y) \leq \epsilon$.

All metrics on a finite set are equivalent; they generate the *discrete* topology.

• Closed metric interval

Given two different points $x, y \in X$ of a metric space (X, d), the **closed** metric interval between x and y is the set

$$I(x,y) = \{ z \in X : d(x,y) = d(x,z) + d(z,y) \}.$$

• Underlying graph of a metric space

The **underlying graph** (or *neighborhood graph*) of a metric space (X, d) is a graph with the vertex-set X and xy being an edge if $I(x, y) = \{x, y\}$, i.e., there is no third point $z \in X$, for which d(x, y) = d(x, z) + d(z, y).

• Distance monotone metric space

A metric space (X, d) is called **distance monotone** if any interval I(x, x') is *closed*, i.e., for any $y \in X \setminus I(x, x')$, there exists $x'' \in I(x, x')$ with d(y, x'') > d(x, x').

• Metric triangle

Three distinct points $x, y, z \in X$ of a metric space (X, d) form a **metric triangle** if the **closed metric intervals** I(x, y), I(y, z) and I(z, x) intersect only in the common end points.

• Metric space having collinearity

A metric space (X, d) has **collinearity** if for any $\epsilon > 0$ every its infinite subset M contains three distinct ϵ -collinear (i.e., with $d(x, y) + d(y, z) - d(x, z) \le \epsilon$) points x, y, z.

• Modular metric space

A metric space (X, \bar{d}) is called **modular** if, for any three different points $x, y, z \in X$, there exists a point $u \in I(x, y) \cap I(y, z) \cap I(z, x)$. This should not be confused with **modular distance** in Chap. 10 and **modulus metric** in Chap. 6.

• Median metric space

A metric space (X, d) is called a **median metric space** if, for any three points $x, y, z \in X$, there exists an unique point $u \in I(x, y) \cap I(y, z) \cap I(z, x)$.

Any median metric space is an L_1 -metric; cf. L_p -metric in Chap. 5 and median graph in Chap. 15.

A metric space (X,d) is called an **antimedian metric space** if, for any three points $x,y,z\in X$, there exists a unique point $u\in X$ maximizing d(x,u)+d(y,u)+d(z,u).

• Metric quadrangle

Four different points $x, y, z, u \in X$ of a metric space (X, d) form a **metric quadrangle** if $x, z \in I(y, u)$ and $y, u \in I(x, z)$. Then d(x, y) = d(z, u) and d(x, u) = d(y, z) in such metric quadrangle.

A metric space (X, d) is called *weakly spherical* if, for any three different points $x, y, z \in X$ with $y \in I(x, z)$, there exists $u \in X$ such that x, y, z, u form a metric quadrangle.

• Metric curve

A **metric curve** (or, simply, curve) γ in a metric space (X,d) is a continuous mapping $\gamma: I \to X$ from an interval I of \mathbb{R} into X. A curve is called an **arc** (or **path**, simple curve) if it is injective. A curve $\gamma: [a,b] \to X$ is called a Jordan curve (or simple closed curve) if it does not cross itself, and $\gamma(a) = \gamma(b)$.

The **length of a curve** $\gamma:[a,b]\to X$ is the number $l(\gamma)$ defined by

$$l(\gamma) = \sup \{ \sum_{1 \le i \le n} d(\gamma(t_i), \gamma(t_{i-1})) : n \in \mathbb{N}, a = t_0 < t_1 < \dots < t_n = b \}.$$

A rectifiable curve is a curve with a finite length. A metric space (X,d), where every two points can be joined by a rectifiable curve, is called a **quasi-convex metric space** (or, specifically, C-**quasi-convex metric space**) if there exists a constant $C \geq 1$ such that every pair $x, y \in X$ can be joined by a rectifiable curve of length at most Cd(x,y). If C=1, then this length is equal to d(x,y), i.e., (X,d) is a **geodesic** (or strictly intrinsic) metric space (cf. Chap. 6).

The **metric derivative** of an arc $\gamma: I \to X$ at a limit point t of I is, if it exists,

$$\lim_{s\to 0}\frac{d(\gamma(t+s),\gamma(t))}{|s|}.$$

It generalizes the notion of *speed* to the metric spaces which have not a notion of direction (such as vector spaces).

Geodesic

Given a metric space (X, d), a **geodesic** is a locally shortest **metric curve**, i.e., it is a locally isometric embedding of \mathbb{R} into X; cf. Chap. 6.

A subset S of X is called a **geodesic segment** (or **metric segment**, shortest path, minimizing geodesic) between two distinct points x and y in X, if there exists a segment (closed interval) [a,b] on the real line \mathbb{R} and an isometric embedding $\gamma:[a,b]\to X$, such that $\gamma[a,b]=S$, $\gamma(a)=x$ and $\gamma(b)=y$.

A metric straight line is a geodesic which is minimal between any two of its points; it is an isometric embedding of the whole \mathbb{R} into X. A metric ray and metric great circle are isometric embeddings of, respectively, the half-line $\mathbb{R}_{\geq 0}$ and a circle $S^1(0,r)$ into X.

A **geodesic metric space** (cf. Chap. 6) is a metric space in which any two points are joined by a geodesic segment. If, moreover, the geodesic is unique, the space is called *totally geodesic*. A geodesic metric space is called *geodesically complete* if every geodesic is a subarc of a metric straight line.

Geodesic convexity

Given a **geodesic metric space** (X,d) and a subset $M \subset X$, the set M is called **geodesically convex** (or *convex*) if, for anytwo points of M,

there exists a geodesic segment connecting them which lies entirely in M; the space is called **locally convex** if such a segment exists for any two sufficiently close points in M.

The **injectivity radius** of the set M is the least number r such that, for any two points in M at distance $\langle r \rangle$, there exists exactly one geodesic segment connecting them which lies entirely in M.

The set $M \subset X$ is called a **totally convex metric subspace** of (X, d) if, for any two points of M, any geodesic segment connecting them lies entirely in M. For a given point $x \in X$, the **radius of convexity** is the radius of largest totally convex metric ball with center at x.

• Busemann convexity

A geodesic metric space (X,d) is called Busemann convex (or globally non-positively Busemann curved) if, for any three points $x,y,z \in X$ and $midpoints\ m(x,z)$ and m(y,z) (i.e., $d(x,m(x,z))=d(m(x,z),z)=\frac{1}{2}d(x,z)$ and $d(y,m(y,z))=d(m(y,z),z)=\frac{1}{2}d(y,z)$), there holds

$$d(m(x,z),m(y,z)) \leq \frac{1}{2}d(x,y).$$

Equivalently, the distance $D(c_1, c_2)$ between any geodesic segments $c_1 = [a_1, b_1]$ and $c_2 = [a_2, b_2]$ is a *convex function*; cf. **metric between intervals** in Chap. 10. (A real-valued function f defined on an interval is called *convex* if $f(\lambda x + (1-\lambda)y) \leq \lambda f(x) + (1-\lambda)f(y)$ for any x, y and $\lambda \in (0, 1)$.)

The flat Euclidean strip $\{(x,y) \in \mathbb{R}^2 : 0 < x < 1\}$ is **Gromov hyperbolic** but not Busemann convex. In a complete Busemann convex metric space any two points are joined by a unique geodesic segment. A metric space is **CAT**(0) (cf. Chap. 6) if and only if it is Busemann convex and Ptolemaic (Foertsch, Lytchak and Schroeder 2007).

A geodesic metric space (X, d) is **Busemann locally convex** (Busemann 1948) if the above inequality holds locally. Any geodesic **locally CAT(0)** metric space (cf. Chap. 6) is Busemann locally convex, and any geodesic **CAT(0)** metric space is Busemann convex but not vice versa.

Menger convexity

A metric space (X,d) is called **Menger convex** if, for any different points $x,y \in X$, there exists a third point $z \in X$ for which d(x,y) = d(x,z) + d(z,y), i.e., |I(x,y)| > 2 holds for the **closed metric interval** $I(x,y) = \{z \in X : (x,y) = d(x,z) + d(z,y)\}$. It is called **strictly Menger convex** if such z is unique for all $x,y \in X$.

The **geodesic convexity** implies the Menger convexity. The converse holds for **complete** metric spaces.

A subset $M \subset X$ is a *d-convex set* (Menger 1928) if $I(x,y) \subset M$ for any different points $x,y \in M$. A function $f:M \to \mathbb{R}$ defined on a *d*-convex set $M \subset X$ is a *d*-convex function if for any $z \in I(x,y) \subset M$

$$f(z) \le \frac{d(y,z)}{d(x,y)}f(x) + \frac{d(x,z)}{d(x,y)}f(y).$$

• Midpoint convexity

A metric space (X, d) is called **midpoint convex** (or **having midpoints**, admitting a midpoint map) if, for any different points $x, y \in X$, there exists a third point $m(x, y) \in X$ for which $d(x, m(x, y)) = d(m(x, y), y) = \frac{1}{2}d(x, y)$. Such a point m(x, y) is called a midpoint and the map $m: X \times X \to X$ is called a midpoint map (cf. **midset**); this map is unique if m(x, y) is unique for all $x, y \in X$. For example, the geometric mean \sqrt{xy} is the midpoint map for the metric space $(\mathbb{R}_{>0}, d(x, y)) = |\log x - \log y|$.

A **complete** metric space is a **geodesic** metric space if and only if it is midpoint convex.

A metric space (X, d) is said to have **approximate midpoints** if, for any points $x, y \in X$ and any $\epsilon > 0$, there exists an ϵ -midpoint, i.e., a point $z \in X$ such that $d(x, z) \leq \frac{1}{2}d(x, y) + \epsilon \geq d(z, y)$.

• Ball convexity

A midpoint convex metric space (X, d) is called **ball convex** if

$$d(m(x,y),z) \le \max\{d(x,z),d(y,z)\}\$$

for all $x, y, z \in X$ and any midpoint map m(x, y).

Ball convexity implies that all metric balls are **totally convex** and, in the case of **geodesic** metric space, vice versa. Ball convexity implies also the uniqueness of a midpoint map (geodesics in the case of **complete** metric space).

The metric space $(\mathbb{R}^2, d(x,y) = \sum_{i=1}^2 \sqrt{|x_i - y_i|})$ is not ball convex.

• Distance convexity

A midpoint convex metric space (X, d) is called distance convex if

$$d(m(x,y),z) \leq \frac{1}{2}(d(x,z) + d(y,z)).$$

A **geodesic** metric space is distance convex if and only if the restriction of the distance function $d(x,\cdot)$, $x\in X$, to every geodesic segment is a convex function.

Distance convexity implies **ball convexity** and, in the case of **Busemann convex** metric space, vice versa.

Metric convexity

A metric space (X, d) is called **metrically convex** if, for any different points $x, y \in X$ and any $\lambda \in (0, 1)$, there exists a third point $z = z(x, y, \lambda) \in X$ for which d(x, y) = d(x, z) + d(z, y) and $d(x, z) = \lambda d(x, y)$. Metric convexity implies **Menger convexity**.

The space is called **strictly metrically convex** if such point $z(x, y, \lambda)$ is unique for all $x, y \in X$ and any $\lambda \in (0, 1)$.

A metric space (X, d) is called **strongly metrically convex** if, for any different points $x, y \in X$ and any $\lambda_1, \lambda_2 \in (0, 1)$, there exists a third point $z = z(x, y, \lambda) \in X$ for which $d(z(x, y, \lambda_1), z(x, y, \lambda_2)) = |\lambda_1 - \lambda_2| d(x, y)$.

Strong metric convexity implies metric convexity, and every Menger convex **complete** metric space is strongly metrically convex.

A metric space (X, d) is called **nearly convex** (Mandelkern 1983) if, for any different points $x, y \in X$ and any $\lambda, \mu > 0$ such that $d(x, y) < \lambda + \mu$, there exists a third point $z \in X$ for which $d(x, z) < \lambda$ and $d(z, y) < \mu$, i.e., $z \in B(x, \lambda) \cap B(y, \mu)$. Metric convexity implies near convexity.

• Takahashi convexity

A metric space (X,d) is called **Takahashi convex** if, for any different points $x,y \in X$ and any $\lambda \in (0,1)$, there exists a third point $z = z(x,y,\lambda) \in X$ such that $d(z(x,y,\lambda),u) \leq \lambda d(x,u) + (1-\lambda)d(y,u)$ for all $u \in X$. Any convex subset of a normed space is a Takahashi convex metric space with $z(x,y,\lambda) = \lambda x + (1-\lambda)y$.

A set $M \subset X$ is Takahashi convex if $z(x, y, \lambda) \in M$ for all $x, y \in X$ and any $\lambda \in [0, 1]$. Takahashi has shown in 1970 that, in a Takahashi convex metric space, all metric balls, open metric balls, and arbitrary intersections of Takahashi convex subsets are all Takahashi convex.

Hyperconvexity

A metric space (X, d) is called **hyperconvex** (Aronszajn and Panitchpakdi 1956) if it is **metrically convex** and its metric balls have the *infinite Helly property*, i.e., any family of mutually intersecting closed balls in X has non-empty intersection. A metric space (X, d) is hyperconvex if and only if it is an **injective metric space**.

The spaces l_{∞}^n , l_{∞}^{∞} and l_1^2 are hyperconvex but l_2^{∞} is not.

• Distance matrix

Given a finite metric space $(X = \{x_1, \dots, x_n\}, d)$, its **distance matrix** is the symmetric $n \times n$ matrix $((d_{ij}))$, where $d_{ij} = d(x_i, x_j)$ for any $1 \le i$, $j \le n$.

The probability that a symmetric $n \times n$ matrix, whose diagonal elements are zeros and all other elements are uniformly random real numbers, is a distance matrix is (Mascioni 2005) $\frac{1}{2}$, $\frac{17}{120}$ for n=3,4 and it is within $[1-(0.918)^{n^2},1-(0.707)^{n^2}]$ for n=5.

• Metric cone

The metric cone MET_n is the polyhedral cone in $\mathbb{R}^{\binom{n}{2}}$ of all distance matrices of semi-metrics on the set $V_n = \{1, \ldots, n\}$. Vershik (2004) considers MET_{∞} , i.e., the weakly closed convex cone of infinite distance matrices of semi-metrics on \mathbb{N} .

The **metric fan** is a canonical decomposition MF_n of MET_n into subcones whose faces belong to the fan, and the intersection of any two of them is their common boundary. Two semi-metrics $d, d' \in MET_n$ lie in the same cone of the metric fan if the subdivisions $\delta_d, \delta_{d'}$ of the polyhedron $\delta(n,2) = conv\{e_i + e_j : 1 \le i < j \le n\} \subset \mathbb{R}^n$ are equal. Here a subpolytope P of $\delta(n,2)$ is a cell of the subdivision δ_d if there exists $y \in \mathbb{R}^n$ satisfying $y_i + y_j = d_{ij}$ if $e_i + e_j$ is a vertex of P, and $y_i + y_j > d_{ij}$ otherwise. The complex of bounded faces of the polyhedron dual to δ_d is the **tight span** of the semi-metric d; cf. **combinatorial dimension**.

The term *metric cone* is also used in Bronshtein (1998) for a convex cone equipped with a complete metric compatible with its operations of addition and multiplication by non-negative numbers.

• Cayley-Menger matrix

Given a finite metric space $(X = \{x_1, \dots, x_n\}, d)$, its **Cayley–Menger matrix** is the symmetric $(n + 1) \times (n + 1)$ matrix

$$CM(X,d) = \begin{pmatrix} 0 & e \\ e^T & D \end{pmatrix},$$

where $D = ((d^2(x_i, x_j)))$ and e is the n-vector all components of which are 1.

The determinant of CM(X,d) is called the Cayley-Menger determinant. If (X,d) is a metric subspace of the Euclidean space \mathbb{E}^{n-1} , then CM(X,d) is $(-1)^n 2^{n-1} ((n-1)!)^2$ times squared (n-1)-dimensional volume of the convex hull of X in \mathbb{R}^{n-1} .

• Gram matrix

Given elements v_1, \ldots, v_k of a Euclidean space, their **Gram matrix** is the symmetric $k \times k$ matrix

$$G(v_1, \ldots, v_k) = ((\langle v_i, v_j \rangle))$$

of pairwise inner products of v_1, \ldots, v_k .

A $k \times k$ matrix is positive-semi-definite if and only if it is a Gram matrix. A $k \times k$ matrix is positive-definite if and only if it is a Gram matrix with linearly independent defining vectors.

We have $G(v_1, \ldots v_k) = \frac{1}{2}((d_E^2(v_0, v_i) + d_E^2(v_0, v_j) - d_E^2(v_i, v_j)))$, i.e., the inner product \langle, \rangle is the **Gromov product similarity** of the **squared Euclidean distance** d_E^2 . A $k \times k$ matrix $((d_E^2(v_i, v_j)))$ defines a **distance of negative type** on $\{1, \ldots, k\}$; all such $k \times k$ matrices form the (non-polyhedral) closed convex cone of all such distances on a k-set.

The determinant of a Gram matrix is called the *Gram determinant*; it is equal to the square of the k-dimensional volume of the parallelotope constructed on $v_1, \ldots v_k$.

Midset

Given a metric space (X, d) and distinct $y, z \in X$, the **midset** (or *bisector*) of points y and z is the set $M = \{x \in X : d(x, y) = d(x, z)\}$ of *midpoints* x.

A metric space is said to have the n-points midset property if, for every pair of its points, the midset has exactly n points. The 1-point midset property mean uniqueness of the midpoint map (cf. midpoint convexity).

• Distance k-sector

Given a metric space (X, d) and disjoint subsets $Y, Z \subset X$, the bisector of Y and Z is the set $M = \{x \in X : \inf_{y \in Y} d(x, y) = \inf_{z \in Z} d(x, z)\}.$

The **distance** k-sector of Y and Z is the sequence M_1, \ldots, M_{k-1} of subsets of X such that M_i , for any $1 \le i \le k-1$, is the bisector of sets M_{i-1}

and M_{i+1} , where $Y = M_0$ and $Z = M_k$. Asano, Matousek and Tokuyama (2006) considered the distance k-sector on the Euclidean plane (\mathbb{R}^2, l_2); for compact sets Y and Z, the sets M_1, \ldots, M_{k-1} are curves partitioning the plane into k parts.

• Metric basis

Given a metric space (X, d), a subset $M \subset X$ is called a **metric basis** (or set of uniqueness) of X if d(x, s) = d(y, s) for all $s \in M$ implies x = y. For $x \in X$, the numbers $d(x, s), s \in M$, are called **metric coordinates** of x.

1.3 Metric numerical invariants

• Metric density

A metric space (X, d) is called **metrically dense** (or, specifically, μ -dense) if (Tukia and Väisälä 1980) there exist numbers λ_1, λ_2 with $0 \le \lambda_1 \le \lambda_2 \le 1$ such that, for every pair of points $x, y \in X$, there exists a point $z \in X$ with $\lambda_1 d(x, y) \le d(x, z) \le \lambda_2 d(x, y)$. In this case, $\mu = (\frac{1+\lambda_2}{\lambda_1(1-\lambda_2)})^2$.

The quantity $\inf\{\mu\}$, where (X, d) is μ -dense, is the coefficient of metric density of (X, d). For the middle-third Cantor set on the interval [0, 1], this coefficient is 12.25 (Ibragimov 2002).

• Metric entropy

Given $\epsilon > 0$, the **metric entropy** (or ϵ -entropy) $H_{\epsilon}(M, X)$ of a subset $M \subset X$ of a metric space (X, d), is defined (Kolmogorov and Tihomirov 1956) by

$$H_{\epsilon}(M, X) = \log_2 CAP_{\epsilon}(M, X),$$

where the function $CAP_{\epsilon}(M, X)$ of $\epsilon > 0$, called the **capacity of metric** space (M, d), is the smallest number of points in an ϵ -net (or ϵ -covering, ϵ -approximation) for the metric space (M, d), i.e., a set of points such that the union of open ϵ -balls, centered at those points, covers M.

The notion of metric entropy for a **dynamical system** is one of the most important invariants in Ergodic Theory.

• Metric dimension

For a metric space (X, d) and any real number $\epsilon > 0$, let $N_X(\epsilon)$ be the minimal number of sets with diameter at most ϵ which are needed in order to cover X (cf. **metric entropy**). The number $\lim_{\epsilon \to 0} \frac{\ln N(\epsilon)}{\ln((\epsilon)^{-1})}$ (if it exists) is called the **metric dimension** (or **Minkowski–Bouligand dimension**, Minkowski dimension, packing dimension, box-counting dimension) of X.

If the limit above does not exist, then the following notions of dimension are considered:

1. The number $\underline{\lim}_{\epsilon \to 0} \frac{\ln N(\epsilon)}{\ln((\epsilon^{-1})}$ is called the **lower Minkowski dimension** (or lower metric dimension, lower box dimension, Pontryagin–Snirelman dimension);

2. The number $\overline{\lim}_{\epsilon \to 0} \frac{\ln N(\epsilon)}{\ln((\epsilon^{-1})}$ is called the **Kolmogorov–Tihomirov dimension** (or *upper metric dimension*, *entropy dimension*, *upper box dimension*).

See below examples of other, less prominent, notions of metric dimension occurring in the mathematical literature:

- 1. The (basis) metric dimension (or location number) of a metric space is the minimum cardinality of its **metric basis**. The partition dimension (Chartrand, Salevi and Zhang 1998) is the minimum cardinality of its resolving partition, i.e., an ordered partition S_1, \ldots, S_k of the space such that no two points have, for $1 \le i \le k$, the same minimal distances to the set S_i . So, partition dimension is at most basis metric dimension plus 1.
- 2. The (equilateral) metric dimension of a metric space is the maximum cardinality of its equidistant subset, i.e., such that any two of its distinct points are at the same distance. For a normed space, this dimension is equal to the maximum number of translates of its unit ball that pairwise touch.
- 3. For any c > 1, the (normed space) metric dimension $\dim_c(X)$ of a finite metric space (X,d) is the least dimension of a real normed space (V,||.||) such that there is an embedding $f: X \to V$ with $\frac{1}{c}d(x,y) \le ||f(x) f(y)|| \le d(x,y)$.
- 4. The (Euclidean) metric dimension of a finite metric space (X, d) is the least dimension n of a Euclidean space \mathbb{E}^n such that (X, f(d)) is its metric subspace, where the minimum is taken over all continuous monotone increasing functions f(t) of $t \geq 0$.
- 5. The dimensionality of a metric space is $\frac{\mu^2}{2\sigma^2}$, where μ and σ^2 are the mean and variance of its histogram of distance values; this notion is used in Information Retrieval for proximity searching. The term dimensionality is also used for the minimal dimension, if it is finite, of Euclidean space in which a given metric space embeds isometrically.

• Volume of finite metric space

Given a metric space (X, d) with $|X| = k < \infty$, its **volume** (Feige 2000) is the maximal (k-1)-dimensional volume of the simplex with vertices $\{f(x): x \in X\}$ over all **short mappings** $f: (X, d) \to (\mathbb{R}^{k-1}, l_2)$. The volume coincides with the metric for k = 2. It is monotonically increasing and continuous in the metric d.

• Rank of metric space

The **Minkowski rank of metric space** (X, d) is the maximal dimension of a normed vector space (V, ||.||) such that there is an isometric embedding $(V, ||.||) \to (X, d)$.

The **Euclidean rank of metric space** (X, d) is the maximal dimension of a *flat* in it, that is of a Euclidean space \mathbb{E}^n such that there is an isometric embedding $\mathbb{E}^n \to (X, d)$.

The quasi-Euclidean rank of metric space (X, d) is the maximal dimension of a *quasi-flat* in it, that is, of a Euclidean space \mathbb{E}^n such that there is a quasi-isometry $\mathbb{E}^n \to (X, d)$. Every Gromov hyperbolic metric space has this rank 1.

• Hausdorff dimension

For a metric space (X,d) and any real p,q>0, let $M_p^q(X)=\inf \sum_{i=1}^{+\infty} (diam(A_i))^p$, where the infimum is taken over all countable coverings $\{A_i\}_i$ of X with the diameter of A_i less than q. The **Hausdorff dimension** (or **fractal dimension**, Hausdorff-Besicovitch dimension, capacity dimension) $dim_{Haus}(X,d)$ of (X,d) is defined by

$$\inf\{p: \lim_{q\to 0} M_p^q(X) = 0\}.$$

Any countable metric space has Hausdorff dimension 0; the Hausdorff dimension of the Euclidean space \mathbb{E}^n is equal to n.

For each **totally bounded** metric space, its Hausdorff dimension is bounded from above by its **metric dimension** and from below by its **topological dimension**.

• Topological dimension

For any compact metric space (X, d) its **topological dimension** (or **Lebesgue covering dimension**) is defined by

$$\inf_{d'} \{ \dim_{Haus}(X, d') \},$$

where d' is any metric on X topologically equivalent to d, and dim_{Haus} is the **Hausdorff dimension**.

This dimension does not exceed also the **Assouad-Nagata dimension** of (X, d).

In general, the **topological dimension** of a topological space X is the smallest integer n such that, for any finite open covering of X, there exists a finite open sub-covering (i.e., a refinement of it) with no point of X belonging to more than n+1 elements.

Fractal

For a metric space, its **topological dimension** does not exceed its **Hausdorff dimension**. A **fractal** is a metric space for which this inequality is strict. (Originally, Mandelbrot defined a fractal as a point set with non-integer Hausdorff dimension.) For example, the *Cantor set*, seen as a compact metric subspace of $(\mathbb{R}, d(x,y) = |x-y|)$ has the Hausdorff dimension $\frac{\ln 2}{\ln 3}$; cf. another **Cantor metric** on it in Chaps. 11 and 18. Another classical fractal, the *Sierpinski carpet* of $[0,1] \times [0,1]$ is a **complete geodesic** metric subspace of $(\mathbb{R}^2, d(x,y) = ||x-y||_1)$.

The term *fractal* is used also, more generally, for *self-similar* (i.e., roughly, looking similar at any scale) object (usually, a subset of \mathbb{R}^n). Cf. scale invariance in Chap. 29.

• Doubling dimension

The **doubling dimension** of a metric space (X,d) is the smallest integer n (or ∞ if such n does not exist) such that every metric ball (or, say, a set of finite diameter) can be covered by a family of at most 2^n metric balls (respectively, sets) of half the diameter.

If (X, d) has finite doubling dimension (or, equivalently, finite **Assouad-Nagata dimension**), then d is called a **doubling metric**.

• Assouad-Nagata dimension

The Assouad-Nagata dimension $dim_{AN}(X,d)$ of a metric space (X,d) is the smallest integer n (or ∞ if such n does not exist) for which there exists a constant C > 0 such that, for all s > 0, there exists a covering of X by its subsets of diameter $\leq Cs$ with every subset of X of diameter $\leq s$ meeting $\leq n+1$ elements of covering.

Replacing "for all s>0" in above definition by "for s>0 sufficiently large" or by "for s>0 sufficiently small," gives microscopic $mi-dim_{AN}(X,d)$ and macroscopic $ma-dim_{AN}(X,d)$ Assouad-Nagata dimensions, respectively. Then (Brodskiy, Dudak, Higes and Mitra 2006) $mi-dim_{AN}(X,d)=dim_{AN}(X,\min\{d,1\})$ and $ma-dim_{AN}(X,d)=dim_{AN}(X,\max\{d,1\})$ (here $\max\{d(x,y),1\}$ means 0 for x=y).

In general, the Assouad-Nagata dimension is not preserved under quasiisometry but it is preserved (Lang and Schlichenmaier 2004) under quasisymmetric mapping.

• Vol'berg-Konyagin dimension

The Vol'berg–Konyagin dimension of a metric space (X, d) is the smallest constant C > 1 (or ∞ if such C does not exist) for which X carries a doubling measure, i.e., a Borel measure μ such that

$$\mu(\overline{B}(x,2r)) \leq C\mu(\overline{B}(x,r))$$

for all $x \in X$ and r > 0.

A metric space (X, d) carries a doubling measure if and only if d is a **doubling metric**, and any complete doubling metric carries a doubling measure.

The **Karger–Ruhl constant** of a metric space (X, d) is the smallest constant c > 1 (or ∞ if such c does not exist) such that

$$|\overline{B}(x,2r)| \le c|\overline{B}(x,r)|$$

for all $x \in X$ and r > 0.

If c is finite, then the **doubling dimension** of (X, d) is at most 4c.

Hyperbolic dimension

A metric space (X, d) is called an (R, N)-large-scale doubling if there exist a number R > 0 and integer N > 0 such that every ball of radius $r \ge R$ in (X, d) can be covered by N balls of radius $\frac{r}{2}$.

The **hyperbolic dimension** hypdim(X,d) of a metric space (X,d) (Buyalo and Schroeder 2004) is the smallest integer n such that, for every r > 0, there exist a real number R > 0, an integer N > 0 and a covering of X with the following properties:

- 1. Every ball of radius r meets at most n+1 elements of covering.
- 2. The covering is an (R, N)-large-scale doubling, and any finite union of its elements is an (R', N)-large-scale doubling for some R' > 0.

The hyperbolic dimension is 0 if (X, d) is a large-scale doubling, and it is n if (X, d) is the n-dimensional hyperbolic space.

Also, $hypdim(X,d) \leq asdim(X,d)$ since the **asymptotic dimension** asdim(X,d) corresponds to the case N=1 in the definition of hypdim(X,d).

The hyperbolic dimension is preserved under a quasi-isometry.

• Asymptotic dimension

The **asymptotic dimension** asdim(X,d) of a metric space (X,d) (Gromov 1993) is the smallest integer n such that, for every r > 0, there exist a constant D = D(r) and a covering of X by its subsets of diameter at most D such that every ball of radius r meets at most n + 1 elements of the covering.

The asymptotic dimension is preserved under a quasi-isometry.

• Width dimension

Let (X,d) be a compact metric space. For a given number $\epsilon > 0$, the **width** dimension $Widim_{\epsilon}(X,d)$ of (X,d) is (Gromov 1999) the minimum integer n such that there exist an n-dimensional polyhedron P and a continuous map $f: X \to P$ (called an ϵ -embedding) with $Diam f^{-1}(y) \leq \epsilon$ for all $y \in P$.

 $\lim_{\epsilon \to 0} Widim_{\epsilon}(X, d)$ is the **topological dimension** of (X, d). Thus the width dimension of (X, d) is its macroscopic dimension at the scale $\geq \epsilon$.

Godsil–MaKay dimension

We say that a metric space (X,d) has **Godsil–McKay dimension** $n \geq 0$ if there exists an element $x_0 \in X$ and two positive constants c and C such that the inequality $ck^n \leq |\{x \in X : d(x,x_0) \leq k\}| \leq Ck^n$ holds for every integer $k \geq 0$. This notion was introduced in [GoMc80] for the **path metric** of a countable locally finite graph. It was proved there that, if the group \mathbb{Z}^n acts faithfully and with a finite number of orbits on the vertices of the graph, then this dimension is equal to n.

• Length of metric space

The **Fremlin length** of a metric space (X, d) is the one-dimensional Hausdorff outer measure on X.

The **Hejcman length** lng(M) of a subset $M \subset X$ of a metric space (X,d) is $\sup\{lng(M'): M' \subset M, |M'| < \infty\}$. Here $lng(\emptyset) = 0$ and, for a finite subset $M' \subset X$, $lng(M') = \min \sum_{i=1}^{n} d(x_{i-1}, x_i)$ over all sequences x_0, \ldots, x_n such that $\{x_i : i = 0, 1, \ldots, n\} = M'$.

The **Schechtman length** of a finite metric space (X, d) is $\inf \sqrt{\sum_{i=1}^n a_i^2}$ over all sequences a_1, \ldots, a_n of positive numbers such that there exists a sequence X_0, \ldots, X_n of partitions of X with following properties:

- 1. $X_0 = \{X\}$ and $X_n = \{\{x\} : x \in X\}$.
- 2. X_i refines X_{i-1} for $i = 1, \ldots, n$.
- 3. For $i=1,\ldots,n$ and $B,C\subset A\in X_{i-1}$ with $B,C\in X_i$, there exists a one-to-one map f from B onto C such that $d(x,f(x))\leq a_i$ for all $x\in B$.

• Roundness of metric space

The **roundness of a metric space** (X, d) is the supremum of all q such that

$$d(x_1, x_2)^q + d(y_1, y_2)^q \le d(x_1, y_1)^q + d(x_1, y_2)^q + d(x_2, y_1)^q + d(x_2, y_2)^q$$

for any four points $x_1, x_2, y_1, y_2 \in X$.

Every metric space has roundness ≥ 1 ; it is ≤ 2 if the space has **approximate midpoints**. The roundness of L_p -space is p if $1 \leq p \leq 2$.

The generalized roundness of a metric space (X, d) is (Enflo 1969) the supremum of all q such that, for any $2k \geq 4$ points $x_i, y_i \in X$ with $1 \leq i \leq k$,

$$\sum_{1 \le i < j \le k} (d(x_i, x_j)^q + d(y_i, y_j)^q) \le \sum_{1 \le i, j \le k} d(x_i, y_j)^q.$$

So, the generalized roundness is the supremum of q such that the **power transform** (cf. Chap. 4) d^q is 2k-gonal distance.

Every **CAT**(0) **space** (cf. Chap. 6) has roundness 2, but some of them have generalized roundness 0 (Lafont and Prassidis 2006).

Type of metric space

The **Enflo type** of a metric space (X,d) is p if there exists a constant $1 \leq C < \infty$ such that, for every $n \in \mathbb{N}$ and every function $f: \{-1,1\}^n \to X$, $\sum_{\epsilon \in \{-1,1\}^n} d^p(f(\epsilon),f(-\epsilon))$ is at most $C^p \sum_{j=1}^n \sum_{\epsilon \in \{-1,1\}^n} d^p(f(\epsilon_1,\ldots,\epsilon_{j-1},\epsilon_j,\epsilon_{j+1},\ldots,\epsilon_n))$, $f(\epsilon_1,\ldots,\epsilon_{j-1},-\epsilon_j,\epsilon_{j+1},\ldots,\epsilon_n)$.

A Banach space (V, ||.||) of Enflo type p has Rademacher type p, i.e., for every $x_1, \ldots, x_n \in V$,

$$\sum_{\epsilon \in \{-1,1\}^n} ||\sum_{j=1}^n \epsilon_j x_j||^p \le C^p \sum_{j=1}^n ||x_j||^p.$$

Given a metric space (X,d), a symmetric Markov chain on X is a Markov chain $\{Z_l\}_{l=0}^{\infty}$ on a state space $\{x_1,\ldots,x_m\}\subset X$ with a symmetrical transition $m\times m$ matrix $((a_{ij}))$, such that $P(Z_{l+1}=x_j:Z_l=x_i)=a_{ij}$

and $P(Z_0 = x_i) = \frac{1}{m}$ for all integers $1 \le i, j \le m$ and $l \ge 0$. A metric space (X, d) has **Markov type** p (Ball 1992) if $\sup_T M_p(X, T) < \infty$ where $M_p(X, T)$ is the smallest constant C > 0 such that the inequality

$$\mathbb{E}d^p(Z_T, Z_0) \le TC^p \mathbb{E}d^p(Z_1, Z_0)$$

holds for every symmetric Markov chain $\{Z_l\}_{l=0}^{\infty}$ on X holds, in terms of expected value (mean) $\mathbb{E}[X] = \sum_x xp(x)$ of the discrete random variable X.

A metric space of Markov type p has Enflo type p.

• Strength of metric space

Given a finite metric space (X,d) with s different non-zero values of $d_{ij} = d(i,j)$, its **strength** is the largest number t such that, for any integers $p,q \geq 0$ with $p+q \leq t$, there is a polynomial $f_{pq}(s)$ of degree at most $\min\{p,q\}$ such that $((d_{ij}^{2p}))((d_{ij}^{2q})) = ((f_{pq}(d_{ij}^2)))$.

• Polynomial metric space

Let (X,d) be a metric space with a finite diameter D and a finite normalized measure μ_X . Let the Hilbert space $L_2(X,d)$ of complex-valued functions decompose into a countable (when X is infinite) or a finite (with D+1 members when X is finite) direct sum of mutually orthogonal subspaces $L_2(X,d) = V_0 \oplus V_1 \oplus \ldots$

Then (X, d) is a **polynomial metric space** if there exists an ordering of the spaces V_0, V_1, \ldots such that, for $i = 0, 1, \ldots$, there exist zonal spherical functions, i.e., real polynomials $Q_i(t)$ of degree i such that

$$Q_i(t(d(x,y))) = \frac{1}{r_i} \sum_{j=1}^{r_i} v_{ij}(x) \overline{v_{ij}(y)}$$

for all $x, y \in X$, where r_i is the dimension of V_i , $\{v_{ii}(x) : 1 \le j \le r_i\}$ is an orthonormal basis of V_i , and t(d) is a continuous decreasing real function such that t(0) = 1 and t(D) = -1. The zonal spherical functions constitute an orthogonal system of polynomials with respect to some weight w(t).

The finite polynomial metric spaces are also called $(P \ and \ Q)$ -polynomial association schemes; cf. distance-regular graph in Chap. 15.

The infinite polynomial metric spaces are the *compact connected two-point homogeneous spaces*; Wang (1952) classified them as the Euclidean unit spheres, the real, complex, quaternionic projective spaces or the Cayley projective line and plane.

• Growth rate of metric space

Let (X,d) be a **distance-invariant metric space**, i.e., all **metric balls** $\overline{B}(x,n) = \{y \in X : d(x,y) \leq n\}$ of the same radius have the same number of elements. The **growth rate of a metric space** (X,d) is the function $f(n) = |\overline{B}(x,n)|$.

(X,d) is a metric space of polynomial growth if there are some positive constants k,C such that $f(n) \leq Cn^k$ for all $n \geq 0$. It has exponential growth if there is a constant C > 1 such that $f(n) > C^n$ for all $n \geq 0$. Cf. graph of polynomial growth, including the group case, in Chap. 15.

For a **metrically discrete metric space** (X, d) (i.e., with $a = \inf_{x,y \in X, x \neq y} d(x,y) > 0$), its *growth rate* was defined also (Gordon, Linial and Rabinovich 1998) by

$$\max_{x \in X, r \ge 2} \frac{\log |\overline{B}(x, ar)|}{\log r}.$$

• Rendez-vous number

Given a metric space (X,d), its **rendez-vous number** (or *Gross number*, *magic number*) is a positive real number r(X,d) (if it exists), defined by the property that for each integer n and all (not necessarily distinct) $x_1, \ldots, x_n \in X$ there exists a point $x \in X$ such that

$$r(X,d) = \frac{1}{n} \sum_{i=1}^{n} d(x_i, x).$$

If the number r(X,d) exists, then it is said that (X,d) has the **average distance property**. Every compact connected metric space has this property. The *unit ball* $\{x \in V : ||x|| \le 1\}$ of a **Banach space** (V,||.||) has the rendez-vous number 1.

• Wiener polynomial

Given a finite subset M of a metric space (X,d) and a parameter q, the **Wiener polynomial** of M is

$$W(M;q) = \frac{1}{2} \sum_{x,y \in M: x \neq y} q^{d(x,y)}.$$

It is a generating function for the **distance distribution** (cf. a very similar notion in Chap. 16) of M, i.e., the coefficient of q^i in W(M;q) is the number of unordered pairs $x, y \in M$ having d(x, y) = i.

The number $W'(M;1) = \frac{1}{2} \sum_{x,y \in M} d(x,y)$, in the case when d is the **path metric** of a graph with vertex-set M, is called *Wiener index*; cf. chemical distance in Chap. 24. The degree distance of this graph is (Dobrynin and Kochetova 1994) $\frac{1}{2} \sum_{x,y \in M} d(x,y)(r(x)+r(y))$, where r(z) is the degree of the vertex $z \in M$.

The average distance of M is the number $\frac{1}{|M|(|M|-1)} \sum_{x,y \in M} d(x,y)$.

• s-energy

Given a finite subset M of a metric space (X, d), the s-energy of M is the number

$$\sum_{x,y\in M, x\neq y} \frac{1}{d^s(x,y)} \text{ and } \sum_{x,y\in M, x\neq y} \log \frac{1}{d(x,y)} = -\prod_{x,y\in M, x\neq y} d(x,y),$$

for $s \neq 0$ and s = 0, respectively. The (-s)-energy with s > 0 is also called the (unnormalized) s-moment of M.

The discrete Riesz s-energy is s-energy for Euclidean distance d and s > 0.

A 1-median and a center of mass of M are points $x_1^*, x_2^* \in X$ minimizing the functionals $\sum_{y \in M} d(x_1, y)$ and $\sum_{y \in M} d^2(x_2, y)$, respectively.

In general, given a completely monotonic (i.e., $(-1)^k f^k \geq 0$ for any k) function $f \in \mathbb{C}^{\infty}$, the f-potential energy of a finite subset M of a metric space (X,d) is $\sum_{x,y\in M, x\neq y} f(d^2(x,y))$. The metric subspace (M,d) is called (Cohn and Kumar 2007) universally optimal if it minimizes, among subspaces (M',d) with |M'| = |M|, f-potential energy for any such function f.

Given an ordered subset $M = \{x_1, \ldots, x_n\}$ of a metric space (X, d) (usually l_2^3), its Lennard-Jones potential energy is $\sum_{i=1}^{n-1} \sum_{j=i+1}^{n} (d(x_i, x_j)^{-12} - 2d(x_i, x_j)^{-6})$.

• Transfinite diameter

The *n*-th diameter $D_n(M)$ and the *n*-th Chebyshev constant $C_n(M)$ of a set $M \subseteq X$ in a metric space (X, d) are defined (Fekete 1923, for the complex plane \mathbb{C}) as

$$D_n(M) = \sup_{x_1, \dots, x_n \in M} \prod_{i \neq j} d(x_i, x_j)^{\frac{1}{n(n-1)}} \text{ and }$$

$$C_n(M) = \inf_{x \in X} \sup_{x_1, \dots, x_n \in M} \prod_{j=1}^n d(x_j, x_j)^{\frac{1}{n}}.$$

The number $\log D_n(M)$ (the supremum of average distance) is called the n-extent of M. The numbers $D_n(M), C_n(M)$ come from the geometric mean averaging; they also come as the limit case $s \to 0$ of the s-moment $\sum_{i \neq j} d(x_i, x_j)^s$ averaging.

The transfinite diameter (or ∞ -th diameter) and the ∞ -th Chebyshev constant $C_{\infty}(M)$ of M are defined as

$$D_{\infty}(M) = \lim_{n \to \infty} D_n(M)$$
 and $C_{\infty}(M) = \lim_{n \to \infty} C_n(M)$;

these limits existing since $\{D_n(M)\}$ and $\{C_n(M)\}$ are non-increasing sequences of non-negative real numbers. Define $D_{\infty}(\emptyset) = 0$. The transfinite diameter of a compact subset of \mathbb{C} is its *capacity*; for a segment in \mathbb{C} , it is $\frac{1}{4}$ of its length.

• Metric diameter

The **metric diameter** (or **diameter**, width) diam(M) of a set $M \subseteq X$ in a metric space (X, d) is defined by

$$\sup_{x,y\in M} d(x,y).$$

The **diameter graph** of M has, as vertices, all points $x \in M$ with d(x,y) = diam(M) for some $y \in M$; it has, as edges, all pairs of its vertices at distance diam(M) in (X,d).

A metric space (X, d) is called an **antipodal metric space** (or *diametrical metric space*) if, for any $x \in X$, there exists the *antipode*, i.e., a unique $x' \in X$ such that the interval I(x, x') is X.

In a metric space endowed with a measure, one says that the *isodiametric* inequality holds if the metric balls maximize the measure among all sets with given diameter. It holds for the volume in Euclidean space but not, for example, for the **Heisenberg metric** on the *Heisenberg group* (cf. Chap. 10).

The k-diameter of a finite metric space (X, d) is (Chung, Delorme and Sole 1999) $\max_{K\subseteq X: |K|=k} \min_{x,y\in K: x\neq y} d(x,y)$; cf. minimum distance in Chap. 16.

Given a property $P \subseteq X \times X$ of a pair (K, K') of subsets of a finite metric space (X, d), the **conditional diameter** (or P-diameter, Balbuena, Carmona, Fábrega and Fiol 1996) is $\max_{(K,K')\in P} \min_{(x,y)\in K\times K'} d(x,y)$. It is diam(X,d) if $P=\{(K,K')\in X\times X: |K|=|K'|=1\}$. When (X,d) models an interconnection network, the P-diameter corresponds to the maximum delay of the messages interchanged between any pair of clusters of nodes, K and K', satisfying a given property P of interest.

• Metric spread

Given a metric space (X, d), let M be a **bounded** (i.e., with finite diameter A) and **metrically discrete** (i.e., the infimum $a = \inf_{x,y \in M, x \neq y} d(x,y) > 0$) subset of X.

The **metric spread** (or aspect ratio, distance ratio, normalized diameter) of M is the ratio $\frac{A}{a}$.

Eccentricity

Given a bounded metric space (X, d), the **eccentricity** (or *Koenig number*) of a point $x \in X$ is the number $e(x) = \max_{y \in X} d(x, y)$.

The numbers $\max_{x \in X} e(x)$ and $\min_{x \in X} e(x)$ are called the **diameter** and the **radius** of (X, d), respectively. For finite |X|, the average eccentricity is $\frac{1}{|X|} \sum_{x \in X} e(x)$.

The sets $\{x \in X : e(x) \le e(z) \text{ for any } z \in X\}$, $\{x \in X : e(x) \ge e(z) \text{ for any } z \in X\}$ and $\{x \in X : \sum_{y \in X} d(x,y) \le \sum_{y \in X} d(z,y) \text{ for any } z \in X\}$ are called, respectively, the **metric center** (or *eccentricity center*, *center*), **metric antimedian** (or *periphery*) and the **metric median** (or *distance center*) of (X, d).

• Radii of metric space

Given a bounded metric space (X,d) and a set $M\subseteq X$, the **metric radius** (or **radius**) r(M) of M is the infimum of radii of metric balls which contain M, i.e., the number $\inf_{x\in M}\sup_{y\in M}d(x,y)$. Then $\frac{diam(M)}{2}\leq r(M)\leq diam(M)$, where diam(M) is the **diameter** of the set M, with r(M)=diam(M) in any **equidistant metric space** and $r(M)=\frac{diam(M)}{2}$ in any **injective metric space**. Some authors define the radius to be the number $\frac{diam(M)}{2}$.

The covering radius of a set $M \subset X$ is $\max_{x \in X} \min_{y \in M} d(x, y)$, i.e., the smallest number R such that the open metric balls of radius R with centers at the elements of M cover X. It is also called the **directed Hausdorff distance** from X to M. The set M is called an ϵ -covering if its covering radius does not exceed ϵ . Given a positive number m, a **minimax distance design of size** m is a m-subset of X having smallest covering radius.

The **packing radius** of a set $M \subset X$ is the largest r such that the open metric balls of radius r with centers at the elements of M are pairwise disjoint, i.e., $\min_{x \in M} \min_{y \in M \setminus \{x\}} d(x,y) > 2r$. The set M is called an ϵ -packing if its packing radius is no less than ϵ . Given a positive number m, a **maximum distance design of size** m is an m-subset of X having largest packing radius.

The size of the smallest ϵ -covering is at most the size of the largest $\frac{\epsilon}{2}$ -packing. An $\frac{\epsilon}{2}$ -packing M is non-extendible if $M \cup \{x\}$ is not an $\frac{\epsilon}{2}$ -packing for every $x \in X \setminus M$, i.e., M is also an ϵ -net.

• Congruence order of metric space

A metric space (X, d) has **congruence order** n if every finite metric space which is not **isometrically embeddable** in (X, d) has a subspace with at most n points which is not isometrically embeddable in (X, d).

For example, the congruence order of l_2^n is n+3 (Menger 1928); it is 4 for the **path metric** of a tree.

• Chromatic numbers of metric space

Given a metric space (X,d) and a set D of positive real numbers, the D-chromatic number of (X,d) is the standard chromatic number of the D-distance graph of (X,d), i.e., the graph with the vertex-set X and the edge-set $\{xy:d(x,y)\in D\}$. Usually, (X,d) is an l_p -space and $D=\{1\}$ (Benda-Perles chromatic number) or $D=[1-\epsilon,1+\epsilon]$ (the chromatic number of the ϵ -unit distance graph). Rosenfeld conjectured that the D-chromatic number of \mathbb{R}^2 is ∞ if D is the set of odd integers.

For a metric space (X, d), the **polychromatic number** is the minimum number of colors needed to color all the points $x \in X$ so that, for each color class C_i , there is a distance d_i such that no two points of C_i are at distance d_i .

For any integer t > 0, the t-distance chromatic number of a metric space (X, d) is the minimum number of colors needed to color all the points $x \in X$ so that any two points whose distance is $\leq t$ have distinct colors.

For any integer t > 0, the t-th Babai number of (X, d) is the minimum number of colors needed to color all the points $x \in X$ so that, for any set D of positive distances with $|D| \le t$, any two points whose distance belongs to D have distinct colors.

• Steiner ratio

Given a metric space (X, d) and a finite subset $V \subset X$, consider the complete weighted graph G = (V, E) with the vertex-set V and edge-weights d(x, y) for all $x, y \in V$.

A spanning tree T in G is a subset of |V|-1 edges forming a tree on V with the weight d(T) equal to the sum of weights of its edges. Let MST_V be a minimum spanning tree in G, i.e., a spanning tree in G with the minimal weight $d(MST_V)$.

A minimum Steiner tree on V is a tree SMT_V such that its vertex-set is a subset of X containing V, and $d(SMT_V) = \inf_{M \subset X: V \subset M} d(MST_M)$.

The **Steiner ratio** St(X,d) of the metric space (X,d) is defined by

$$\inf_{V \subset X} \frac{d(SMT_V)}{d(MST_V)}.$$

For any metric space (X, d) we have $\frac{1}{2} \leq St(X, d) \leq 1$. For the l_2 -metric (cf. L_p -metric in Chap. 5) on \mathbb{R}^2 , it is equal to $\frac{\sqrt{3}}{2}$, while for the l_1 -metric on \mathbb{R}^2 it is equal to $\frac{2}{3}$.

Cf. arc routing problems in Chap. 15.

1.4 Metric mappings

• Distance function

A distance function (or ray function) is a continuous function on a metric space (X,d) (usually, on an Euclidean space \mathbb{E}^n) $f:X\to\mathbb{R}_{\geq 0}$ which is homogeneous, i.e., f(tx)=tf(x) for all $t\geq 0$ and all $x\in X$.

A distance function f is called *symmetric* if f(x) = f(-x), *positive* if f(x) > 0 for all $x \neq 0$, and *convex* if $f(x+y) \leq f(x) + f(y)$ with f(0) = 0. If $X = \mathbb{E}^n$, the set $\{x \in \mathbb{R}^n : f(x) < 1\}$ is called a *star body*, it corresponds to a unique distance function. The star body is bounded if f is positive, it is symmetric about the origin if f is symmetric, and it is convex if f is a **convex distance function**.

In Topology, the term distance function is often used for distance.

Convex distance function

Given a compact convex region $B \subset \mathbb{R}^n$ which contains the origin in its interior, the **convex distance function** (or **gauge**, *Minkowski distance function*) $d_B(x,y)$ is the quasi-metric on \mathbb{R}^n defined, for $x \neq y$, by

$$\inf\{\alpha > 0 : y - x \in \alpha B\}.$$

It is also defined, equivalently, as $\frac{||y-x||_2}{||z-x||_2}$, where z is the unique point of the boundary $\partial(x+B)$ hit by the ray from x through y. Then $B=\{x\in\mathbb{R}^n:d_B(0,x)\leq 1\}$ with equality only for $x\in\partial B$. A convex distance function is called polyhedral if B is a polytope, tetrahedral if it is a tetrahedron and so on.

If B is centrally-symmetric with respect to the origin, then d_B is a **Minkowskian metric** (cf. Chap. 6) whose unit ball is B.

• Element of best approximation

Given a metric space (X, d) and a subset $M \subset X$, an element $u_0 \in M$ is called an **element of best approximation** to a given element $x \in X$ if $d(x, u_0) = \inf_{u \in M} d(x, u)$, i.e., if $d(x, u_0)$ is the **point-set distance** d(x, M).

A metric projection (or operator of best approximation, nearest point map) is a multi-valued mapping associating to each element $x \in X$ the set of elements of best approximation from the set M (cf. distance map).

A Chebyshev set in a metric space (X, d) is a subset $M \subset X$ containing a unique element of best approximation for every $x \in X$.

A subset $M \subset X$ is called a **semi-Chebyshev set** if the number of such elements is at most one, and a **proximinal set** if this number is at least one.

The Chebyshev radius of the set M is $\inf_{x \in X} \sup_{y \in M} d(x, y)$, and a Chebyshev center of M is an element $x_0 \in X$ realizing this infimum.

• Distance map

Given a metric space (X, d) and a subset $M \subset X$, the **distance map** is a function $f_M : X \to \mathbb{R}_{\geq 0}$, where $f_M(x) = \inf_{u \in M} d(x, u)$ is the **point-set distance** d(x, M) (cf. **metric projection**).

If the boundary B(M) of the set M is defined, then the **signed distance** function g_M is defined by $g_M(x) = -\inf_{u \in B(M)} d(x, u)$ for $x \in M$, and $g_M(x) = \inf_{u \in B(M)} d(x, u)$ otherwise. If M is a (closed and orientable) manifold in \mathbb{R}^n , then g_M is the solution of the *eikonal equation* $|\nabla g| = 1$ for its gradient ∇ .

If $X = \mathbb{R}^n$ and, for every $x \in X$, there is unique element u(x) with d(x, M) = d(x, u(x)) (i.e., M is a **Chebyshev set**), then ||x - u(x)|| is called a **vector distance function**.

Distance maps are used in Robot Motion (M being the set of obstacle points) and, especially, in Image Processing (M being the set of all or only boundary pixels of the image). For $X = \mathbb{R}^2$, the graph $\{(x, f_M(x)) : x \in X\}$ of d(x, M) is called the *Voronoi surface* of M.

Isometry

Given metric spaces (X, d_X) and (Y, d_Y) , a function $f: X \to Y$ is called an **isometric embedding** of X into Y if it is injective and the equality $d_Y(f(x), f(y)) = d_X(x, y)$ holds for all $x, y \in X$.

An **isometry** (or *congruence mapping*) is a bijective isometric embedding. Two metric spaces are called **isometric** (or *isometrically isomorphic*) if there exists an isometry between them.

A property of metric spaces which is invariant with respect to isometries (completeness, boundedness, etc.) is called a **metric property** (or *metric invariant*).

A path isometry (or *arcwise isometry*) is a mapping from X into Y (not necessarily bijective) preserving lengths of curves.

• Rigid motion of metric space

A **rigid motion** (or, simply, **motion**) of a metric space (X, d) is an **isometry** of (X, d) onto itself.

For a motion f, the **displacement function** $d_f(x)$ is d(x, f(x)). The motion f is called *semisimple* if $\inf_{x \in X} d_f(x) = d(x_0, f(x_0))$ for some $x_0 \in X$, and *parabolic* otherwise. A semisimple motion is called *elliptic* if $\inf_{x \in X} d_f(x) = 0$, and *axial* (or *hyperbolic*) otherwise. A motion is called a *Clifford translation* if the displacement function $d_f(x)$ is a constant for all $x \in X$.

• Symmetric metric space

A metric space (X,d) is called **symmetric** if, for any point $p \in X$, there exists a *symmetry* relative to that point, i.e., a **motion** f_p of this metric space such that $f_p(f_p(x)) = x$ for all $x \in X$, and p is an isolated fixed point of f_p .

• Homogeneous metric space

A metric space is called **homogeneous** (or *highly transitive*, *ultrahomogeneous*) if any isometry between two of its finite subspaces extends to the whole space.

A metric space is called *point-homogeneous* if, for any two points of it, there exists a motion mapping one of the points to the other. In general, a *homogeneous space* is a set together with a given transitive group of *symmetries*.

A metric space (X,d) is called (Grünbaum–Kelly) a **metrically homogeneous metric space** if $\{d(x,z):z\in X\}=\{d(y,z):z\in X\}$ for any $x,y\in X$.

• Dilation

Given a metric space (X,d) and a positive real number r, a function $f: X \to X$ is called a **dilation** if d(f(x), f(y)) = rd(x, y) holds for any $x, y \in X$.

• Metric cone structure

Given a **pointed metric space** (X, d, x_0) (i.e., a space (X, d) with a fixed point $x_0 \in X$), a **metric cone structure** on it is a (pointwise) continuous family f_t $(t \in \mathbb{R}_{>0})$ of **dilations** of X, leaving the point x_0 invariant, such that $d(f_t(x), f_t(y)) = td(x, y)$ for all x, y and $f_t \circ f_s = f_{ts}$.

A Banach space has such a structure for the dilations $f_t(x) = tx$ $(t \in \mathbb{R}_{>0})$. The *Euclidean cone over a metric space* (cf. **cone over metric space** in Chap. 9) is another example. Cf. also **cone metric** in Chap. 3.

A cone over a topological space (X, τ) (the base of the cone) is the quotient space $(X \times [0,1])/(X \times \{0\})$ obtained from the product $X \times [0,1]$ by collapsing the subspace $X \times \{0\}$ to a point v (the vertex of the cone).

The **tangent metric cone** over a metric space (X, d) at a point x_0 is (for all dilatations tX = (X, td)) the closure of $\cup_{t>0} tX$, i.e., of $\lim_{t\to\infty} tX$ taken in the pointed Gromov–Hausdorff topology (cf. **Gromov–Hausdorff metric**).

The asymptotic metric cone over (X, d) is its tangent metric cone "at infinity," i.e., $\bigcap_{t>0} tX = \lim_{t\to 0} tX$. Cf. boundary of metric space in Chap. 6.

• Metric fibration

Given a **complete** metric space (X, d), two subsets M_1 and M_2 of X are called *equidistant* if for each $x \in M_1$ there exists $y \in M_2$ with d(x, y) being equal to the **Hausdorff metric** between the sets M_1 and M_2 . A **metric fibration** of (X, d) is a partition \mathcal{F} of X into isometric mutually equidistant closed sets.

The quotient metric space X/\mathcal{F} inherits a natural metric for which the **distance map** is a **submetry**.

• Paradoxical metric space

Given a metric space (X, d) and an equivalence relation on the subsets of X, the space (X, d) is called **paradoxical metric space** if X can be decomposed into two disjoint sets M_1 , M_2 so that M_1 , M_2 and X are pairwise equivalent.

Deuber, Simonovitz and Sós (1995) introduced this idea for wobbling equivalent subsets $M_1, M_2 \subset X$, i.e., there is a bijective wobbling (a mapping $f: M_1 \to M_2$ with bounded $\sup_{x \in X} d(x, f(x))$). For example, (\mathbb{R}^2, l_2) is paradoxical for wobbling equivalence but not for isometry equivalence.

• Homeomorphic metric spaces

Two metric spaces (X, d_X) and (Y, d_Y) are called **homeomorphic** (or topologically isomorphic) if there exists a homeomorphism from X to Y, i.e., a bijective function $f: X \to Y$ such that f and f^{-1} are continuous (the preimage of every open set in Y is open in X).

Two metric spaces (X, d_X) and (Y, d_Y) are called uniformly isomorphic if there exists a bijective function $f: X \to Y$ such that f and f^{-1} are uniformly continuous functions. (A function g is uniformly continuous if, for any $\epsilon > 0$, there exists $\delta > 0$ such that, for any $x, y \in X$, the inequality $d_X(x,y) < \delta$ implies that $d_Y(g(x), f(y)) < \epsilon$; a continuous function is uniformly continuous if X is compact.)

• Möbius mapping

Given a metric space (X, d) and quadruple (x, y, z, w) of its distinct points, the **cross-ratio** is the positive number defined by

$$cr((x, y, z, w), d) = \frac{d(x, y)d(z, w)}{d(x, z)d(y, w)}.$$

Given metric spaces (X, d_X) and (Y, d_Y) , a **homeomorphism** $f: X \to Y$ is called a **Möbius mapping** if, for every quadruple (x, y, z, w) of distinct points of X,

$$cr((x, y, z, w), d_X) = cr((f(x), f(y), f(z), f(w)), d_Y).$$

A homeomorphism $f: X \to Y$ is called a **quasi-Möbius mapping** (Väisälä 1984) if there exists a homeomorphism $\tau: [0, \infty) \to [0, \infty)$ such that, for every quadruple (x, y, z, w) of distinct points of X,

$$cr((f(x), f(y), f(z), f(w)), d_Y) \le \tau(cr((x, y, z, w), d_X)).$$

• Quasi-symmetric mapping

Given metric spaces (X, d_X) and (Y, d_Y) , a **homeomorphism** $f: X \to Y$ is called a **quasi-symmetric mapping** (Tukia and Väisälä 1980) if there exists a homeomorphism $\tau: [0, \infty) \to [0, \infty)$ such that, for every triple (x, y, z) of distinct points of X,

$$\frac{d_Y(f(x), f(y))}{d_Y(f(x), f(z))} \le \tau \frac{d_X(x, y)}{d_X(x, z)}.$$

Quasi-symmetric mappings are **quasi-Möbius**, and quasi-Möbius mappings between bounded metric spaces are quasi-symmetric. In the case $f: \mathbb{R}^n \to \mathbb{R}^n$, quasi-symmetric mappings are exactly the same as **quasi-conformal mappings**.

• Conformal metric mapping

Given metric spaces (X, d_X) and (Y, d_Y) , which are domains in \mathbb{R}^n , a **homeomorphism** $f: X \to Y$ is called a **conformal metric mapping** if, for any non-isolated point $x \in X$, the limit $\lim_{y \to x} \frac{d_Y(f(x), f(y))}{d(x, y)}$ exists, is finite and positive.

A homeomorphism $f:X\to Y$ is called a **quasi-conformal mapping** (or, specifically, C-quasi-conformal mapping) if there exists a constant C such that

$$\lim_{r \to 0} \sup \frac{\max\{d_Y(f(x), f(y)) : d_X(x, y) \le r\}}{\min\{d_Y(f(x), f(y)) : d_X(x, y) \ge r\}} \le C$$

for each $x \in X$. The smallest such constant C is called the **conformal** dilation.

The **conformal dimension** of a metric space (X, d) (Pansu 1989) is the infimum of **Hausdorff dimension** over all quasi-conformal mappings of (X, d) into some metric space. For the middle-third Cantor set on [0, 1], it is 0 but, for any of its quasi-conformal images, it is positive.

• Hölder mapping

Let $c, \alpha \geq 0$ be constants. Given metric spaces (X, d_X) and (Y, d_Y) , a function $f: X \to Y$ is called **Hölder mapping** (or α -Hölder mapping if the constant α should be mentioned) if for all $x, y \in X$

$$d_Y(f(x), f(y)) \le c(d_X(x, y))^{\alpha}.$$

A 1-Hölder mapping is a **Lipschitz mapping**; 0-Hölder mapping means that the metric d_Y is bounded.

• Lipschitz mapping

Let c be a positive constant. Given metric spaces (X, d_X) and (Y, d_Y) , a function $f: X \to Y$ is called a **Lipschitz mapping** (or c-Lipschitz mapping if the constant c should be mentioned) if for all $x, y \in X$

$$d_Y(f(x), f(y)) \le cd_X(x, y).$$

A c-Lipschitz mapping is called a **short mapping** if c = 1, and is called a **contraction** if c < 1.

• Bi-Lipschitz mapping

Let c > 1 be a positive constant. Given metric spaces (X, d_X) and (Y, d_Y) , a function $f: X \to Y$ is called a **bi-Lipschitz mapping** (or c-bi-Lipschitz mapping, **c-embedding**) if there exists a positive real number r such that, for any $x, y \in X$, we have the following inequalities:

$$rd_X(x,y) \le d_Y(f(x),f(y)) \le crd_X(x,y).$$

Every bi-Lipschitz mapping is a quasi-symmetric mapping.

The smallest c for which f is a c-bi-Lipschitz mapping is called the **distortion** of f. Bourgain proved that every k-point metric space c-embeds into a Euclidean space with distortion $O(\ln k)$. Gromov's distortion for curves is the maximum ratio of arclength to chord length.

Two metrics d_1 and d_2 on X are called **bi-Lipschitz equivalent metrics** if there are positive constants c and C such that $cd_1(x,y) \leq d_2(x,y) \leq Cd_1(x,y)$ for all $x,y \in X$, i.e., the identity mapping is a bi-Lipschitz mapping from (X,d_1) into (X,d_2) . Bi-Lipschitz equivalent metrics are **equivalent**, i.e., generate the same topology but, for example, equivalent L_1 -metric and L_2 -metric (cf. L_p -metric in Chap. 5) on $\mathbb R$ are not bi-Lipschitz equivalent.

A bi-Lipschitz mapping $f: X \to Y$ is a **c-isomorphism** $f: X \to f(X)$.

• c-isomorphism of metric spaces

Given two metric spaces (X, d_X) and (Y, d_Y) , the Lipschitz norm $||.||_{Lip}$ on the set of all injective mappings $f: X \to Y$ is defined by

$$||f||_{Lip} = \sup_{x,y \in X, x \neq y} \frac{d_Y(f(x), f(y))}{d_X(x, y)}.$$

Two metric spaces X and Y are called c-isomorphic if there exists an injective mapping $f: X \to Y$ such that $||f||_{Lip}||f^{-1}||_{Lip} \le c$.

• Metric Ramsey number

For a given class \mathcal{M} of metric spaces (usually, l_p -spaces), an integer $n \geq 1$, and a real number $c \geq 1$, the **metric Ramsey number** (or *c-metric Ramsey number*) $R_{\mathcal{M}}(c,n)$ is the largest integer m such that every n-point metric space has a subspace of size m that c-embeds into a member of \mathcal{M} (see [BLMN05]).

The Ramsey number R_n is the minimal number of vertices of a complete graph such that any coloring of the edges with n colors produces a monochromatic triangle. The following metric analog of R_n was considered in [Masc04]. Let D_n be the least number of points a finite metric space must contain in order to contain an equilateral triangle, i.e., to have equilateral metric dimension greater than 2.

• Uniform metric mapping

Given metric spaces (X, d_X) and (Y, d_Y) , a function $f: X \to Y$ is called a **uniform metric mapping** if there are two non-decreasing functions g_1 and g_2 from $\mathbb{R}_{\geq 0}$ to itself with $\lim_{r\to\infty} g_i(r) = \infty$ for i = 1, 2, such that the inequality

$$g_1(d_X(x,y)) \le d_Y(f(x), f(y)) \le g_2(d_X(x,y))$$

holds for all $x, y \in X$.

A **bi-Lipschitz mapping** is a uniform metric mapping with linear functions g_1 and g_2 .

• Metric compression

Given metric spaces (X, d_X) (unbounded) and (Y, d_Y) , a function $f: X \to Y$ is a large scale Lipschitz mapping if, for some $c > 0, D \ge 0$ and all $x, y \in X$,

$$d_Y(f(x), f(y)) \le cd_X(x, y) + D.$$

The compression of such a mapping f is $\rho_f(r) = \inf_{d_X(x,y) \geq r} d_Y(f(x), f(y))$.

The **metric compression** of (X, d_X) in (Y, d_Y) is defined by

$$R(X,Y) = \sup_{f} \{ \underline{\lim}_{r \to \infty} \frac{\log \max\{\rho_f(r), 1\}}{\log r} \},$$

where supremum is over all large scale Lipschitz mappings f.

The main interesting case, when (Y, d_Y) is a Hilbert space and (X, d_X) is a (finitely generated discrete) group with **word metric**, was considered by Guentner and Kaminker in 2004. Then R(X,Y) = 0 if there is no **uniform metric mapping** from (X, d_X) to (Y, d_Y) and R(X,Y) = 1 for free groups (even if there is no **quasi-isometry**). Arzhantzeva, Guba and Sapir (2006) found groups with $\frac{1}{2} \leq R(X,Y) \leq \frac{3}{4}$.

• Quasi-isometry

Given metric spaces (X, d_X) and (Y, d_Y) , a function $f: X \to Y$ is called a **quasi-isometry** (or (C, c)-quasi-isometry) if there exist real numbers C > 1 and c > 0 such that

$$C^{-1}d_X(x,y) - c \le d_Y(f(x), f(y)) \le Cd_X(x,y) + c,$$

and $Y = \bigcup_{x \in X} B_{d_Y}(f(x), c)$, i.e., for every point $y \in Y$, there exists a point $x \in X$ such that $d_Y(y, f(x)) < \frac{c}{2}$. Quasi-isometry is an equivalence relation on metric spaces; it is a bi-Lipschitz equivalence up to small distances.

A quasi-isometry with C=1 is called a **coarse isometry** (or rough isometry, almost isometry, Hausdorff approximation).

Cf. quasi-Euclidean rank of a metric space.

• Coarse embedding

Given metric spaces (X, d_X) and (Y, d_Y) , a function $f: X \to Y$ is called a **coarse embedding** if there exist non-decreasing functions $\rho_1, \rho_2: [0, \infty) \to [0, \infty)$ such that $\rho_1(d_X(x, y)) \leq d_Y(f(x), f(y)) \leq \rho_2(d_X(x, y))$ for all $x, y \in X$, and $\lim_{t \to \infty} \rho_1(t) = +\infty$.

Metrics d_1 and d_2 on X are called **coarsely equivalent metrics** if there exist non-decreasing functions $f, g : [0, \infty) \to [0, \infty)$ such that $d_1 \leq f(d_2)$ and $d_2 \leq g(d_1)$.

• Contraction

Given metric spaces (X, d_X) and (Y, d_Y) , a function $f: X \to Y$ is called a **contraction** if the inequality

$$d_Y(f(x), f(y)) \le cd_X(x, y)$$

holds for all $x, y \in X$ and some real number $c, 0 \le c < 1$.

Every contraction is a **contractive mapping** (but not necessarily the other way around) and it is uniformly continuous. *Banach Fixed Point Theorem* (or *Contraction Principle*): every contraction from a **complete** metric space into itself has a unique fixed point.

• Contractive mapping

Given metric spaces (X, d_X) and (Y, d_Y) , a function $f: X \to Y$ is called a **contractive mapping** (or *strictly short mapping*) if, for all different points $x, y \in X$,

$$d_Y(f(x), f(y)) < d_X(x, y).$$

Every contractive mapping from a **compact** metric space into itself has a unique fixed point.

A function $f: X \to Y$ is called a **non-contractive mapping** (or dominating mapping) if, for all different $x, y \in X$,

$$d_Y(f(x), f(y)) \ge d_X(x, y).$$

Every non-contractive bijection from a **totally bounded** metric space onto itself is an **isometry**.

• Short mapping

Given metric spaces (X, d_X) and (Y, d_Y) , a function $f : X \to Y$ is called a **short mapping** (or 1-Lipschitz mapping, non-expansive mapping, metric mapping semi-contraction) if the inequality

$$d_Y(f(x), f(y)) \le d_X(x, y)$$

holds for all $x, y \in X$.

A **submetry** is a short mapping such that the image of any metric ball is a metric ball of the same radius.

The set of short mappings $f: X \to Y$ for bounded metric spaces (X, d_X) and (Y, d_Y) is a metric space under **uniform metric** $\sup\{d_Y(f(x), g(x)): x \in X\}$.

Two subsets A and B of a metric space (X,d) are called (Gowers 2000) **similar** if there exist short mappings $f: A \to X$, $g: B \to X$ and a small $\epsilon > 0$ such that every point of A is within ϵ of some point of B, every point of B is within ϵ of some point of A, and $|d(x, g(f(x))) - d(y, f(g(y)))| \le \epsilon$ for every $x \in A$ and $y \in B$.

• Category of metric spaces

A category Ψ consists of a class $Ob\Psi$, whose elements are called *objects of* the category, and a class $Mor\Psi$, elements of which are called *morphisms* of the category. These classes have to satisfy the following conditions:

- 1. To each ordered pair of objects A, B is associated a set H(A,B) of morphisms.
- 2. Each morphism belongs to only one set H(A, B).
- 3. The composition $f \cdot g$ of two morphisms $f : A \to B$, $g : C \to D$ is defined if B = C in which case it belongs to H(A, D).
- 4. The composition of morphisms is associative.
- 5. Each set H(A, A) contains, as an *identity*, a morphism id_A such that $f \cdot id_A = f$ and $id_A \cdot g = g$ for any morphisms $f : X \to A$ and $g : A \to Y$.

The **category of metric spaces**, denoted by Met (see [Isbe64]), is a category which has metric spaces as objects and **short mappings** as morphisms. A unique **injective envelope** exists in this category for every one of its objects; it can be identified with its **tight span**. In Met, the *monomorphisms* are injective short mappings, and *isomorphisms* are **isometries**. Met is a subcategory of the category which has metric spaces as objects and **Lipschitz mappings** as morphisms.

• Injective metric space

A metric space (X, d) is called **injective** if, for every isometric embedding $f: X \to X'$ of (X, d) into another metric space (X', d'), there exists a **short mapping** f' from X' into X with $f' \cdot f = id_X$, i.e., X is a **retract** of X'. Equivalently, X is an **absolute retract**, i.e., a retract of every metric space into which it embeds isometrically. A metric space (X, d) is injective if and only if it is **hyperconvex**.

Examples of injective metric spaces include l_1^2 -space, l_{∞}^n -space, any **real** tree and the **tight span** of a metric space.

• Injective envelope

The notion of **injective envelope** (or **metric envelope**) is a generalization of the notion of **Cauchy completion**. Given a metric space (X, d), it can be embedded isometrically into an **injective metric space** (\hat{X}, \hat{d}) ; given any such isometric embedding $f: X \to \hat{X}$, there exists a unique smallest injective subspace $(\overline{X}, \overline{d})$ of (\hat{X}, \hat{d}) containing f(X) which is called **injective envelope** of X. It is isometrically identified with the **tight span** of (X, d).

A metric space coincides with its injective envelope if and only if it is an injective metric space.

• Tight extension

An extension (X',d') of a metric space (X,d) is called a **tight extension** if, for every semi-metric d'' on X' satisfying the conditions $d''(x_1,x_2) = d(x_1,x_2)$ for all $x_1,x_2 \in X$, and $d''(y_1,y_2) \leq d'(y_1,y_2)$ for any $y_1,y_2 \in X'$, one has $d''(y_1,y_2) = d'(y_1,y_2)$ for all $y_1,y_2 \in X'$.

The **tight span** is the *universal tight extension* of X, i.e., it contains, up to canonical isometries, every tight extension of X, and it has no proper tight extension itself.

• Tight span

Given a metric space (X,d) of finite diameter, consider the set $\mathbb{R}^X = \{f: X \to \mathbb{R}\}$. The **tight span** T(X,d) of (X,d) is defined as the set $T(X,d) = \{f \in \mathbb{R}^X : f(x) = \sup_{y \in X} (d(x,y) - f(y)) \text{ for all } x \in X\}$, endowed with the metric induced on T(X,d) by the $\sup_{x \in X} |f(x)|$.

The set X can be identified with the set $\{h_x \in T(X,d) : h_x(y) = d(y,x)\}$ or, equivalently, with the set $T^0(X,d) = \{f \in T(X,d) : 0 \in f(X)\}$. The **injective envelope** $(\overline{X},\overline{d})$ of X is isometrically identified with the tight span T(X,d) by

$$\overline{X} \to T(X,d), \ \overline{x} \to h_{\overline{x}} \in T(X,d): h_{\overline{x}}(y) = \overline{d}(f(y),\overline{x}).$$

The tight span T(X,d) of a finite metric space is the metric space $(T(X), D(f,g) = \max |f(x) - g(x)|)$, where T(X) is the set of functions $f: X \to \mathbb{R}$ such that for any $x, y \in X$, $f(x) + f(y) \ge d(x,y)$ and, for each $x \in X$, there exists $y \in X$ with f(x) + f(y) = d(x,y). The mapping of any x into the function $f_x(y) = d(x,y)$ gives an isometric embedding of (X,d) into T(X,d). For example, if $X = \{x_1, x_2\}$, then T(X,d) is the interval of length $d(x_1, x_2)$.

The tight span of a metric space (X, d) of finite diameter can be considered as a polytopal complex of bounded faces of the polyhedron

$$\{y \in \mathbb{R}^n_{\geq 0} : y_i + y_j \geq d(x_i, x_j) \text{ for } 1 \leq i < j \leq n\}$$

if, for example, $X = \{x_1, \dots, x_n\}$. The dimension of this complex is called (Dress 1984) the **combinatorial dimension** of (X, d).

• Real tree

A metric space (X, d) is called (Tits 1977) a **real tree** (or \mathbb{R} -tree) if, for all $x, y \in X$, there exists a unique **arc** from x to y, and this arc is a **geodesic segment**. So, an \mathbb{R} -tree is a (uniquely) arcwise connected metric space in which each arc is isometric to a subarc of \mathbb{R} . A real tree is also called a **metric tree**, not to be confused with a **metric tree** in Data Analysis (cf. Chap. 17).

A metric space (X, d) is a real tree if and only if it is **path-connected** and Gromov 0-hyperbolic (i.e., satisfies the four-point inequality).

Real trees are exactly **tree-like** metric spaces which are **geodesic**; they are **injective** metric spaces among tree-like spaces. Tree-like metric spaces are by definition metric subspaces of real trees.

If (X, d) is a finite metric space, then the **tight span** T(X, d) is a real tree and can be viewed as an edge-weighted graph-theoretical tree.

A metric space is a complete real tree if and only if it is **hyperconvex** and any two points are joined by a **metric segment**.

The plane \mathbb{R}^2 with the **Paris metric** or **lift metric** (cf. Chap. 19) are examples of \mathbb{R} -tree.

1.5 General distances

• Discrete metric

Given a set X, the **discrete metric** (or **trivial metric**, **sorting distance**) is a metric on X, defined by d(x,y) = 1 for all distinct $x,y \in X$ and d(x,x) = 0. Cf. the more general notion of a (metrically or topologically) **discrete metric space**.

• Indiscrete semi-metric

Given a set X, the **indiscrete semi-metric** d is a semi-metric on X, defined by d(x, y) = 0 for all $x, y \in X$.

• Equidistant metric

Given a set X and a positive real number t, the **equidistant metric** d is a metric on X, defined by d(x,y) = t for all distinct $x,y \in X$ (and d(x,x) = 0).

• (1,2) - B-metric

Given a set X, the (1,2)-B-metric d is a metric on X such that, for any $x \in X$, the number of points $y \in X$ with d(x,y)=1 is at most B, and all other distances are equal to 2. The (1,2)-B-metric is the **truncated** metric of a graph with maximal vertex degree B.

• Permutation metric

Given a finite set X, a metric d on it is called a **permutation metric** (or *linear arrangement metric*) if there exists a bijection $\omega: X \to \{1, \ldots, |X|\}$ such that for all $x, y \in X$

$$d(x,y) = |\omega(x) - \omega(y)|.$$

Given an integer $n \geq 1$, the **line metric on** $\{1,\ldots,n\}$ is defined by |x-y| for any $1 \leq x,y \leq n$. Even, Naor, Rao and Schieber (2000) defined more general **spreading metric**, i.e., any metric d on $\{1,\ldots,n\}$ such that $\sum_{y \in M} d(x,y) \geq \frac{|M|(|M|+2)}{4}$ for any $1 \leq x \leq n$ and $M \subseteq \{1,\ldots,n\} \setminus \{x\}$ with |M| > 2.

• Induced metric

Given a metric space (X, d) and a subset $X' \subset X$, an **induced metric** is the restriction d' of d to X'. A metric space (X', d') is called a **metric subspace** of (X, d), and the metric space (X, d) is called a **metric extension** of (X', d').

Katětov mapping

Given a metric space (X, d), the mapping $f: X \to \mathbb{R}$ is a **Katětov mapping** if

$$|f(x) - f(y)| \le d(x, y) \le f(x) + f(y)$$

for any $x, y \in X$, i.e., setting d(x, z) = f(x) defines a one-point **metric** extension $(X \cup \{z\}, d)$ of (X, d).

The set E(X) of Katětov mappings on X endowed with distance $D(f,g) = \sup_{x \in X} |f(x) - g(x)|$ is a complete metric space; (X,d) embeds isometrically in it via the *Kuratowski mapping* $x \to d(x,.)$, with unique extension of each isometry of X to one of E(X).

• Dominating metric

Given metrics d and d_1 on a set X, d_1 dominates d if $d_1(x, y) \ge d(x, y)$ for all $x, y \in X$. Cf. non-contractive mapping (or dominating mapping).

• Metric transform

A **metric transform** is a distance obtained as a function of a given metric (cf. Chap. 4).

• Complete metric

Given a metric space (X, d), a sequence $\{x_n\}$, $x_n \in X$, is said to have convergence to $x^* \in X$ if $\lim_{n\to\infty} d(x_n, x^*) = 0$, i.e., for any $\epsilon > 0$, there exists $n_0 \in \mathbb{N}$ such that $d(x_n, x^*) < \epsilon$ for any $n > n_0$.

A sequence $\{x_n\}_n$, $x_n \in X$, is called a *Cauchy sequence* if, for any $\epsilon > 0$, there exists $n_0 \in \mathbb{N}$ such that $d(x_n, x_m) < \epsilon$ for any $m, n > n_0$.

A metric space (X, d) is called a **complete metric space** if every Cauchy sequence in it converges. In this case the metric d is called a **complete metric**. An example of incomplete metric space is $(\mathbb{N}, d(m, n) = \frac{|m-n|}{mn})$.

• Cauchy completion

Given a metric space (X, d), its **Cauchy completion** is a metric space (X^*, d^*) on the set X^* of all equivalence classes of *Cauchy sequences*, where the sequence $\{x_n\}_n$ is called *equivalent to* $\{y_n\}_n$ if $\lim_{n\to\infty} d(x_n, y_n) = 0$. The metric d^* is defined by

$$d^*(x^*, y^*) = \lim_{n \to \infty} d(x_n, y_n)$$

for any $x^*, y^* \in X^*$, where $\{x_n\}_n$ (respectively, $\{y_n\}_n$) is any element in the equivalence class x^* (respectively, y^*).

The Cauchy completion (X^*, d^*) is a unique, up to isometry, **complete** metric space, into which the metric space (X, d) embeds as a *dense* metric subspace.

The Cauchy completion of the metric space $(\mathbb{Q}, |x-y|)$ of rational numbers is the real line $(\mathbb{R}, |x-y|)$. A **Banach space** is the Cauchy completion of a normed vector space (V, ||.||) with the **norm metric** ||x-y||. A **Hilbert space** corresponds to the case an inner product norm $||x|| = \sqrt{\langle x, x \rangle}$.

• Perfect metric space

A complete metric space (X, d) is called **perfect** if every point $x \in X$ is a limit point, i.e., $|B(x, r) = \{y \in X : d(x, y) < r\}| > 1$ holds for any r > 0.

Every non-empty perfect **totally disconnected** compact metric space is **homeomorphic** to the *Cantor set* with the natural metric |x-y|. The totally disconnected countable metric space $(\mathbb{Q}, |x-y|)$ of rational numbers also consists only of limit points but it is not complete and not **locally compact**.

Every proper metric ball of radius r in a metric space has diameter at most 2r. Given a number $0 < c \le 1$, a metric space is called a c-uniformly perfect metric space if this diameter is at least 2cr. Cf. radius of metric space.

• Metrically discrete metric space

A metric space (X, d) is called **metrically discrete** (or *uniformly discrete*) if there exists a number r > 0 such that $B(x, r) = \{y \in X : d(x, y) < r\} = \{x\}$ for every $x \in X$.

(X,d) is a **topologically discrete metric space** (or a discrete metric space) if the underlying topological space is **discrete**, i.e., each point $x \in X$ is an isolated point: there exists a number r(x) > 0 such that $B(x,r(x)) = \{x\}$. For $X = \{\frac{1}{n} : n = 1,2,3,\dots\}$, the metric space (X,|x-y|) is topologically but not metrically discrete. Cf. **translation discrete metric** in Chap. 10.

Alternatively, a metric space (X, d) is called *discrete* if any of the following holds:

1. (Burdyuk and Burdyuk 1991) it has a proper *isolated subset*, i.e., $M \subset X$ with $\inf\{d(x,y): x \in M, y \notin M\} > 0$ (any such space admits a unique decomposition into *continuous*, i.e., non-discrete, components).

2. (Lebedeva, Sergienko and Soltan 1984) for any two distinct points $x, y \in X$, there exists a point z of the **closed metric interval** I(x, y) with $I(x, z) = \{x, z\}$.

3. a stronger property holds: for any two distinct points $x, y \in X$, every sequence of points z_1, z_2, \ldots with $z_k \in I(x, y)$ but $z_{k+1} \in I(x, z_k) \setminus \{z_k\}$ for $k = 1, 2, \ldots$ is a finite sequence.

• Bounded metric space

A metric (moreover, a distance) d on a set X is called **bounded** if there exists a constant C > 0 such that $d(x, y) \le C$ for any $x, y \in X$.

For example, given a metric d on X, the metric D on X, defined by $D(x,y)=\frac{d(x,y)}{1+d(x,y)}$, is bounded with C=1.

A metric space (X, d) with a bounded metric d is called a **bounded** metric space.

• Totally bounded metric space

A metric space (X,d) is called **totally bounded** if, for every $\epsilon > 0$, there exists a finite ϵ -net, i.e., a finite subset $M \subset X$ with the **point-set distance** $d(x,M) < \epsilon$ for any $x \in X$ (cf. **totally bounded space** in Chap. 2).

Every totally bounded metric space is **bounded** and **separable**.

A metric space is totally bounded if and only if its **Cauchy completion** is a **compact** metric space.

• Separable metric space

A metric space is called **separable** if it contains a countable *dense* subset, i.e., some countable subset with which all its elements can be approached.

A metric space is separable if and only if it is **second-countable**, and if and only if it is **Lindelöf**.

• Metric compactum

A metric compactum (or compact metric space) is a metric space in which every sequence has a *Cauchy subsequence*, and those subsequences are convergent. A metric space is compact if and only if it is **totally bounded** and **complete**.

Every bounded and closed subset of a Euclidean space is compact. Every finite metric space is compact. Every compact metric space is **second-countable**.

• Proper metric space

A metric space is called **proper** (or *finitely compact*, or having the *Heine-Borel property*) if every closed metric ball in it is compact. Every proper metric space is **complete**.

• UC metric space

A metric space is called a **UC metric space** (or *Atsuji space*) if any continuous function from it into an arbitrary metric space is *uniformly continuous*.

Every **metric compactum** is a UC metric space. Every UC metric space is **complete**.

• Polish space

A Polish space is a complete separable metric space. A metric space is called a Souslin space if it is a continuous image of a Polish space.

A metric triple (or mm-space) is a Polish space (X, d) with a Borel probability measure μ , i.e., a non-negative real function μ on the Borel sigma-algebra \mathcal{F} of X with the following properties: $\mu(\emptyset) = 0$, $\mu(X) = 1$, and $\mu(\cup_n A_n) = \sum_n \mu(A_n)$ for any finite or countable collection of pairwise disjoint sets $A_n \in \mathcal{F}$.

Given a topological space (X, τ) , a sigma-algebra on X is a collection \mathcal{F} of subsets of X with the following properties: $\emptyset \in \mathcal{F}$, $X \setminus U \in \mathcal{F}$ for $U \in \mathcal{F}$, and $\bigcup_n A_n \in \mathcal{F}$ for a finite or countable collection $\{A_n\}_n$, $A_n \in \mathcal{F}$.

The sigma-algebra on X which is related to the topology of X, i.e., consists of all open and closed sets of X, is called a *Borel sigma-algebra* of X. Any metric space is a *Borel space*, i.e., a set equipped with a Borel sigma-algebra.

• Norm metric

Given a normed vector space (V, ||.||), the **norm metric** on V is defined by

$$||x-y||$$
.

The metric space (V, ||x - y||) is called a **Banach space** if it is **complete**. Examples of norm metrics are l_p - and L_p -metrics, in particular, the **Euclidean metric**.

Any metric space (X, d) admits an isometric embedding into a Banach space B such that its convex hull is dense in B (cf. Monge–Kantorovich metric); (X, d) is a linearly rigid metric space if such embedding is unique up to isometry.

• Path metric

Given a connected graph G = (V, E), its **path metric** (or *graphic metric*) d_{path} is a metric on V, defined as the length (i.e., the number of edges) of a shortest path connecting two given vertices x and y from V (cf. Chap. 15).

• Editing metric

Given a finite set X and a finite set \mathcal{O} of (unary) editing operations on X, the **editing metric** on X is the **path metric** of the graph with the vertex-set X and xy being an edge if y can be obtained from x by one of the operations from \mathcal{O} .

• Gallery metric

A chamber system is a set X (whose elements are referred to as chambers) equipped with n equivalence relations \sim_i , $1 \leq i \leq n$. A gallery is a sequence of chambers x_1, \ldots, x_m such that $x_i \sim_j x_{i+1}$ for every i and some j depending on i.

The **gallery metric** is an **extended metric** on X which is the length of the shortest gallery connecting x and $y \in X$ (and is equal to ∞ if there is no connecting gallery). The gallery metric is the (extended) **path metric** of the graph with the vertex-set X and xy being an edge if $x \sim_i y$ for some $1 \le i \le n$.

• Riemannian metric

Given a connected n-dimensional smooth $manifold M^n$, its **Riemannian** metric is a collection of positive-definite symmetric bilinear forms $((g_{ij}))$ on the tangent spaces of M^n which varies smoothly from point to point.

The length of a curve γ on M^n is expressed as $\int_{\gamma} \sqrt{\sum_{i,j} g_{ij} dx_i dx_j}$, and the **intrinsic metric** on M^n , sometimes also called the **Riemannian distance**, is the infimum of lengths of curves connecting any two given points $x, y \in M^n$. Cf. Chap. 7.

• Linearly additive metric

A linearly additive metric (or projective metric) d is a continuous metric on \mathbb{R}^n which satisfies the condition

$$d(x,z) = d(x,y) + d(y,z)$$

for any collinear points x, y, z lying in that order on a common line. The Hilbert fourth problem asked in 1900 to classify such metrics; it is solved only for dimension n = 2 [Amba76]. Cf. Chap. 6.

Every **norm metric** on \mathbb{R}^n is linearly additive. Every linearly additive metric on \mathbb{R}^2 is a **hypermetric**.

• Product metric

Given a finite or countable number n of metric spaces (X_1, d_1) , (X_2, d_2) , ..., (X_n, d_n) , the **product metric** is a metric on the Cartesian product $X_1 \times X_2 \times \cdots \times X_n = \{x = (x_1, x_2, \dots, x_n) : x_1 \in X_1, \dots, x_n \in X_n\}$, defined as a function of d_1, \dots, d_n (cf. Chap. 4).

• Hamming metric

The **Hamming metric** d_H is a metric on \mathbb{R}^n , defined (Hamming 1950) by

$$|\{i: 1 \le i \le n, x_i \ne y_i\}|.$$

On binary vectors $x, y \in \{0, 1\}^n$ the Hamming metric and the l_1 -metric (cf. L_p -metric in Chap. 5) coincide; they are equal to $|I(x)\Delta I(y)| = |I(x)\backslash I(y)| + |I(y)\backslash I(x)|$, where $I(z) = \{1 \le i \le n : z_i = 1\}$. In fact, $\max\{|I(x)\backslash I(y)|, |I(y)\backslash I(x)|\}$ is also a metric.

• Lee metric

Given $m, n \in \mathbb{N}$, $m \geq 2$, the **Lee metric** d_{Lee} is a metric on $\mathbb{Z}_m^n = \{0, 1, \ldots, m-1\}^n$, defined (Lee 1958) by

$$\sum_{1\leq i\leq n}\min\{|x_i-y_i|,m-|x_i-y_i|\}.$$

The metric space (Z_m^n, d_{Lee}) is a discrete analog of the *elliptic space*.

The Lee metric coincides with the Hamming metric d_H if m=2 or m=3. The metric spaces (Z_4^n,d_{Lee}) and $Z_2^{2n},d_H)$ are isometric. The Lee metric is applied for phase modulation while the Hamming metric is used in case of orthogonal modulation.

Cf. absolute summation distance and generalized Lee metric in Chap. 16.

• Symmetric difference metric

Given a measure space $(\Omega, \mathcal{A}, \mu)$, the symmetric difference semimetric (or measure semi-metric) d_{\triangle} is a semi-metric on the set $\mathcal{A}_{\mu} = \{A \in \mathcal{A} : \mu(A) < \infty\}$, defined by

$$\mu(A\triangle B),$$

where $A \triangle B = (A \cup B) \setminus (A \cap B)$ is the *symmetric difference* of the sets A and $B \in \mathcal{A}_{\mu}$.

The value $d_{\triangle}(A, B) = 0$ if and only if $\mu(A \triangle B) = 0$, i.e., A and B are equal almost everywhere. Identifying two sets $A, B \in \mathcal{A}_{\mu}$ if $\mu(A \triangle B) = 0$, we obtain the symmetric difference metric (or Fréchet–Nikodym–Aronszyan distance, measure metric).

If μ is the cardinality measure, i.e., $\mu(A) = |A|$ is the number of elements in A, then $d_{\triangle}(A, B) = |A \triangle B|$. In this case $|A \triangle B| = 0$ if and only if A = B. The **Johnson distance** between k-sets A and B is $\frac{|A \triangle B|}{2} = k - |A \cap B|$. The symmetric difference metric between ordered q-partitions $A = (A_1, \ldots, A_q)$ and $B = (B_1, \ldots, B_q)$ of a finite set is $\sum_{i=1}^q |A_i \triangle B_i|$. Cf. metrics between partitions in Chap. 10.

• Enomoto-Katona metric

Given a finite set X and an integer k, $2k \leq |X|$, the **Enomoto–Katona metric** is the distance between unordered pairs (X_1, X_2) and (Y_1, Y_2) of disjoint k-subsets of X, defined by

$$\min\{|X_1 \setminus Y_1| + |X_2 \setminus Y_2|, |X_1 \setminus Y_2| + |X_2 \setminus Y_1|\}.$$

• Steinhaus distance

Given a measure space $(\Omega, \mathcal{A}, \mu)$, the **Steinhaus distance** d_{St} is a semimetric on the set $\mathcal{A}_{\mu} = \{A \in \mathcal{A} : \mu(A) < \infty\}$, defined by

$$\frac{\mu(A \triangle B)}{\mu(A \cup B)} = 1 - \frac{\mu(A \cap B)}{\mu(A \cup B)}$$

if $\mu(A \cup B) > 0$ (and is equal to 0 if $\mu(A) = \mu(B) = 0$). It becomes a metric on the set of equivalence classes of elements from \mathcal{A}_{μ} ; here $A, B \in \mathcal{A}_{\mu}$ are called *equivalent* if $\mu(A \triangle B) = 0$.

The biotope distance (or Tanimoto distance, Marczewski–Steinhaus distance) $\frac{|A\triangle B|}{|A\cup B|}$ is the special case of Steinhaus distance obtained for the *cardinality measure* $\mu(A) = |A|$ for finite sets (cf. also generalized biotope transform metric in Chap. 4).

• Point-set distance

Given a metric space (X, d), the **point-set distance** d(x, A) between a point $x \in X$ and a subset A of X is defined as

$$\inf_{y \in A} d(x, y).$$

For any $x, y \in X$ and for any non-empty subset A of X, we have the following version of the triangle inequality: $d(x, A) \leq d(x, y) + d(y, A)$ (cf. distance map).

For a given point-measure $\mu(x)$ on X and a penalty function p, an **optimal quantizer** is a set $B \subset X$ such that $\int p(d(x,B))d\mu(x)$ is as small as possible.

• Set-set distance

Given a metric space (X, d), the **set-set distance** between two subsets A and B of X is defined by

$$\inf_{x \in A, y \in B} d(x, y).$$

This distance can be 0 even for disjoint sets, for example, for the intervals (1,2), (2,3) on \mathbb{R} . The sets A and B are positively separated if their set-set distance is positive.

In Data Analysis, the set-set distance between clusters is called the single linkage, while $\sup_{x \in A, y \in B} d(x, y)$ is called the **complete linkage**.

Matching distance

Given a metric space (X, d), the **matching distance** (or *multiset-multiset distance*) between two multisets A and B in X is defined by

$$\inf_{\phi} \max_{x \in A} d(x, \phi(x)),$$

where ϕ runs over all bijections between A and B, as multisets. Cf. **metrics** between multisets.

The matching distance in d'Amico, Frosini and Landi (2006) is, roughly, the case when d is the L_{∞} -metric on cornerpoints of the size functions $f: \{(x,y) \in \mathbb{R}^2 : x < y\} \to \mathbb{N}$.

The matching distance is not related to the **perfect matching distance** in Chap. 15 nor to the **non-linear elastic matching distance** in Chap. 21.

• Hausdorff metric

Given a metric space (X,d), the **Hausdorff metric** (or two-sided Hausdorff distance) d_{Haus} is a metric on the family \mathcal{F} of all compact subsets of X, defined by

$$\max\{d_{dHaus}(A,B), d_{dHaus}(B,A)\},\$$

where $d_{dHaus}(A, B) = \max_{x \in A} \min_{y \in B} d(x, y)$ is the **directed Hausdorff distance** (or *one-sided Hausdorff distance*) from A to B. In other words, $d_{Haus}(A, B)$ is the minimal number ϵ (called also the **Blaschke distance**) such that closed ϵ -neighborhood of A contains B and a closed ϵ -neighborhood of B contains A. Then $d_{Haus}(A, B)$ is equal to

$$\sup_{x \in X} |d(x, A) - d(x, B)|,$$

where $d(x, A) = \min_{y \in A} d(x, y)$ is the **point-set distance**. The Hausdorff metric is not a **norm metric**.

If the above definition is extended for non-compact closed subsets A and B of X, then $d_{Haus}(A, B)$ can be infinite, i.e., it becomes an **extended** metric.

For not necessarily closed subsets A and B of X, the **Hausdorff semi-**metric between them is defined as the Hausdorff metric between their closures. If X is finite, d_{Haus} is a metric on the class of all subsets of X.

• L_p -Hausdorff distance

Given a finite metric space (X, d), the L_p -Hausdorff distance [Badd92] between two subsets A and B of X is defined by

$$(\sum_{x \in X} |d(x,A) - d(x,B)|^p)^{\frac{1}{p}},$$

where d(x, A) is the **point-set distance**. The usual **Hausdorff metric** corresponds to the case $p = \infty$.

• Generalized G-Hausdorff metric

Given a group (G, \cdot, e) acting on a metric space (X, d), the **generalized** G-Hausdorff metric between two closed bounded subsets A and B of X is defined by

$$\min_{g_1, g_2 \in G} d_{Haus}(g_1(A), g_2(B)),$$

where d_{Haus} is the **Hausdorff metric**. If d(g(x), g(y)) = d(x, y) for any $g \in G$ (i.e., if the metric d is *left-invariant* with respect of G), then above metric is equal to $\min_{g \in G} d_{Haus}(A, g(B))$.

• Gromov-Hausdorff metric

The **Gromov–Hausdorff metric** is a metric on the set of all *isometry classes* of compact metric spaces, defined by

$$\inf d_{Haus}(f(X), g(Y))$$

for any two classes X^* and Y^* with the representatives X and Y, respectively, where d_{Haus} is the **Hausdorff metric**, and the minimum is taken over all metric spaces M and all isometric embeddings $f: X \to M, g: Y \to M$. The corresponding metric space is called the Gromov-Hausdorff space.

The **Hausdorff–Lipschitz distance** between isometry classes of compact metric spaces X and Y is defined by

$$\inf\{d_{GH}(X, X_1) + d_L(X_1, Y_1) + d_{GH}(Y, Y_1)\},\$$

where d_{GH} is the Gromov–Hausdorff metric, d_L is the **Lipschitz metric**, and the minimum is taken over all (isometry classes of compact) metric spaces X_1, Y_1 .

• Fréchet metric

Let (X,d) be a metric space. Consider a set \mathcal{F} of all continuous mappings $f:A\to X,\ g:B\to X,\ldots$, where A,B,\ldots are subsets of \mathbb{R}^n , homeomorphic to $[0,1]^n$ for a fixed dimension $n\in\mathbb{N}$.

The Fréchet semi-metric d_F is a semi-metric on \mathcal{F} , defined by

$$\inf_{\sigma} \sup_{x \in A} d(f(x), g(\sigma(x))),$$

where the infimum is taken over all orientation preserving homeomorphisms $\sigma: A \to B$. It becomes the **Fréchet metric** on the set of equivalence classes $f^* = \{g: d_F(g,f) = 0\}$. Cf. the **Fréchet surface metric** in Chap. 8.

• Banach-Mazur distance

The Banach–Mazur distance d_{BM} between two Banach spaces V and W is

$$\ln\inf_{T}||T||\cdot||T^{-1}||,$$

where the infimum is taken over all isomorphisms $T: V \to W$.

It can also be written as $\ln d(V,W)$, where the number d(V,W) is the smallest positive $d \geq 1$ such that $\overline{B}_W^n \subset T(\overline{B}_V^n) \subset d\overline{B}_W^n$ for some linear invertible transformation $T: V \to W$. Here $\overline{B}_V^n = \{x \in V: ||x||_V \leq 1\}$ and $\overline{B}_W^n = \{x \in W; ||x||_W \leq 1\}$ are the *unit balls* of the normed spaces $(V, ||.||_V)$ and $(W, ||.||_W)$, respectively.

One has $d_{BM}(V, W) = 0$ if and only if V and W are isometric, and d_{BM} becomes a metric on the set X^n of all equivalence classes of n-dimensional normed spaces, where $V \sim W$ if they are isometric. The pair (X^n, d_{BM}) is a compact metric space which is called the **Banach–Mazur compactum**.

The Gluskin–Khrabrov distance (or modified Banach–Mazur distance) is defined by

$$\inf\{||T||_{X\to Y}:|detT|=1\}\cdot\inf\{||T||_{Y\to X}:|detT|=1\}.$$

 ${\bf Tomczak-Jaegermann\ distance}\ ({\rm or}\ \textit{weak\ Banach-Mazur\ distance})$ is defined by

$$\max\{\overline{\gamma}_Y(id_X), \overline{\gamma}_X(id_Y)\},\$$

where id is the identity map and, for an operator $U: X \to Y$, $\overline{\gamma}_Z(U)$ denotes $\inf \sum ||W_k|| ||V_k||$. Here the infimum is taken over all representations $U = \sum W_k V_k$ for $W_k: X \to Z$ and $V_k: Z \to Y$. This distance never exceeds the corresponding Banach–Mazur distance.

• Kadets distance

The gap (or opening) between two closed subspaces X and Y of a Banach space (V, ||.||) is defined by

$$gap(X, Y) = \max\{\delta(X, Y), \delta(Y, X)\},\$$

where $\delta(X,Y) = \sup\{\inf_{y \in Y} ||x-y|| : x \in X, ||x|| = 1\}$ (cf. gap distance in Chap. 12 and gap metric in Chap. 18).

The **Kadets distance** between two Banach spaces V and W is a semi-metric, defined (Kadets 1975) by

$$\inf_{Z,f,g} gap(B_{f(V)}, B_{g(W)}),$$

where the infimum is taken over all Banach spaces Z and all linear isometric embeddings $f: V \to Z$ and $g: W \to Z$; here $B_{f(V)}$ and $B_{g(W)}$ are the unit metric balls of Banach spaces f(V) and g(W), respectively.

The non-linear analogue of the Kadets distance is the following Gromov-Hausdorff distance between Banach spaces U and W:

$$\inf_{Z,f,g} d_{Haus}(f(B_V),g(B_W)),$$

where the infimum is taken over all metric spaces Z and all isometric embeddings $f: V \to Z$ and $g: W \to Z$; here d_{Haus} is the **Hausdorff metric**.

The **Kadets path distance** between Banach spaces V and W is defined (Ostrovskii 2000) as the infimum of the length (with respect to the Kadets distance) of all curves joining V and W (and is equal to ∞ if there is no such curve).

• Lipschitz distance

Given $\alpha \geq 0$ and two metric spaces (X, d_X) , (Y, d_Y) , the α -Hölder norm $||.||_{Hol}$ on the set of all injective functions $f: X \to Y$ is defined by

$$||f||_{Hol} = \sup_{x,y \in X, x \neq y} \frac{d_Y(f(x), f(y))}{d_X(x, y)^{\alpha}}.$$

The Lipschitz norm $||.||_{Lip}$ is the case $\alpha = 1$ of $||.||_{Hol}$.

The **Lipschitz distance** between metric spaces (X, d_X) and (Y, d_Y) is defined by

$$\ln \inf_{f} ||f||_{Lip} \cdot ||f^{-1}||_{Lip},$$

where the infimum is taken over all bijective functions $f:X \to Y$. Equivalently, it is the infimum of numbers $\ln a$ such that there exists a bijective **bi-Lipschitz mapping** between (X, d_X) and (Y, d_Y) with constants $\exp(-a)$, $\exp(a)$.

It becomes a metric – **Lipschitz metric** – on the set of all isometry classes of compact metric spaces. Cf. **Hausdorff–Lipschitz distance**.

This distance is an analog to the **Banach–Mazur distance** and, in the case of finite-dimensional real Banach spaces, coincides with it.

It coincides also with the **Hilbert projective metric** on *non-negative* projective spaces, obtained by starting with $\mathbb{R}^n_{>0}$ and identifying any point x with cx, c > 0.

• Lipschitz distance between measures

Given a compact metric space (X,d), the Lipschitz semi-norm $||.||_{Lip}$ on the set of all functions $f:X\to\mathbb{R}$ is defined by $||f||_{Lip}=\sup_{x,y\in X,x\neq y}\frac{|f(x)-f(y)|}{d(x,y)}$.

The Lipschitz distance between measures μ and ν on X is defined by

$$\sup_{||f||_{Lip} \le 1} \int f d(\mu - \nu).$$

If μ and ν are probability measures, then it is the **Kantorovich–Mallows–Monge–Wasserstein metric**.

An analog of the Lipschitz distance between measures for the *state* space of unital C^* -algebra is the **Connes metric**.

• Barycentric metric space

Given a metric space (X, d), let $(B(X), ||\mu - \nu||_{TV})$ be the metric space, where B(X) is the set of all regular Borel probability measures on X with bounded support, and $||\mu - \nu||_{TV}$ is the **total variation norm distance** $\int_X |p(\mu) - p(\nu)| d\lambda$. Here $p(\mu)$ and $p(\nu)$ are the density functions of measures μ and ν , respectively, with respect to the σ -finite measure $\frac{\mu + \nu}{2}$.

A metric space (X, d) is **barycentric** if there exist a constant $\beta > 0$ and a surjection $f: B(X) \to X$ such that the inequality

$$d(f(\mu), f(\nu)) \le \beta diam(supp(\mu + \nu))||\mu - \nu||_{TV}$$

holds for any measures $\mu, \nu \in B(X)$.

Any Banach space (X, d = ||x - y||) is a barycentric metric space with the smallest β being 1 and the map $f(\mu)$ being the usual center of mass $\int_{Y} x d\mu(x)$.

Any Hadamard space (i.e., a complete **CAT**(0) space, cf. Chap. 6) is barycentric with the smallest β being 1 and the map $f(\mu)$ being the unique minimizer of the function $g(y) = \int_X d^2(x,y) d\mu(x)$ on X.

• Metrics between multisets

A multiset (or bag) on a set S is a mapping $m: S \to \mathbb{Z}_{\geq 0}$, where m(x) represents the "multiplicity" of $x \in S$. Multisets are good models for multi-attribute objects as, say, all symbols in a string, all words in a document, criminal records, etc.

A multiset m is finite if S and all m(x) are finite; the *complement* of a finite multiset m is the multiset $\overline{m}: S \to \mathbb{Z}_{>0}$, where $\overline{m}(x) =$

 $\max_{y\in S} m(y) - m(x)$. Given two multisets m_1 and m_2 , denote by $m_1 \cup m_2$, $m_1 \cap m_2$, $m_1 \setminus m_2$ and $m_1 \Delta m_2$ the multisets on S, defined, for any $x \in S$, by $m_1 \cup m_2(x) = \max\{m_1(x), m_2(x)\}$, $m_1 \cap m_2(x) = \min\{m_1(x), m_2(x)\}$, $m_1 \setminus m_2(x) = \max\{0, m_1(x) - m_2(x)\}$ and $m_1 \Delta m_2(x) = |m_1(x) - m_2(x)|$, respectively. Also, $m_1 \subseteq m_2$ denotes that $m_1(x) \leq m_2(x)$ for all $x \in S$.

The measure $\mu(m)$ of a multiset m may be defined, for instance, as a linear combination $\mu(m) = \sum_{x \in S} \lambda(x) m(x)$ with $\lambda(x) \geq 0$. In particular, the number |m| of elements in the multiset m, $\sum_{x \in S} m(x)$, is its counting measure.

For any measure $\mu(m) \in \mathbb{R}_{\geq 0}$, Petrovsky (2003) proposed several **metrics between multisets** m_1 and m_2 including $d_1(m_1, m_2) = \mu(m_1 \Delta m_2)$ and $d_2(m_1, m_2) = \frac{\mu(m_1 \Delta m_2)}{\mu(m_1 \cup m_2)}$ (with $d_2(\emptyset, \emptyset) = 0$ by definition). Cf. symmetric difference metric and Steinhaus distance.

Among examples of other metrics between multisets are **matching** distance in Chap. 1, metric space of roots in Chap. 12, μ -metric in Chap. 15 and, in Chap. 11, bag distance $\max\{|m_1 \backslash m_2|, |m_2 \backslash m_1|\}$ and q-gram similarity.

• Metrics between fuzzy sets

A fuzzy subset of a set S is a mapping $\mu: S \to [0,1]$, where $\mu(x)$ represents the "degree of membership" of $x \in S$. It is an ordinary (crisp) if all $\mu(x)$ are 0 or 1. Fuzzy sets are good models for gray scale images (cf. gray scale images distances in Chap. 21), random objects and objects with non-sharp boundaries.

Bhutani and Rosenfeld (2003) introduced the following two metrics between two fuzzy subsets μ and ν of a finite set S. The **diff-dissimilarity** is a metric (a fuzzy generalization of **Hamming metric**), defined by

$$d(\mu, \nu) = \sum_{x \in S} |\mu(x) - \nu(x)|.$$

The **perm-dissimilarity** is a semi-metric, defined by

$$\min\{d(\mu,p(\nu))\},$$

where the minimum is taken over all permutations p of S.

The Chaudhuri–Rosenfeld metric (1996) between two fuzzy sets μ and ν with crisp points (i.e., the sets $\{x \in S : \mu(x) = 1\}$ and $\{x \in S : \nu(x) = 1\}$ are non-empty) is an **extended metric**, defined by

$$\int_0^1 2t d_{Haus}(\{x \in S : \mu(x) \ge t\}, \{x \in S : \nu(x) \ge t\}) dt,$$

where d_{Haus} is the **Hausdorff metric**.

A fuzzy number is a fuzzy subset μ of the real line \mathbb{R} such that the level set $\{x \in \mathbb{R} : \mu(x) \geq t\}$ is convex for every $t \in [0,1]$. The sendograph of a fuzzy set μ is the set

$$send(\mu) = \{(x, t) \in S \times [0, 1] : \mu(x) > 0, \ \mu(x) \ge t\}.$$

The **sendograph metric** (Kloeden 1980) between two fuzzy numbers μ , ν with crisp points and compact sendographs is the **Hausdorff metric**

$$\max\{\sup_{a=(x,t)\in send(\mu)}d(a,send(\nu)),\sup_{b=(x',t')\in send(\nu)}d(b,send(\mu))\},$$

where d(a, b) = d((x, t), (x', t')) is a **box metric** (cf. Chap. 4) $\max\{|x - x'|, |t - t'|\}$.

The t-cut of a fuzzy set μ is the set $A_{\mu}(t) = \{x \in S : \mu(x) \ge t\}$.

The Klement–Puri–Ralesku metric (1988) between two fuzzy numbers μ , ν is

$$\int_0^1 d_{Haus}(A_{\mu}(t), A_{\nu}(t))dt,$$

where $d_{Haus}(A_{\mu}(t), A_{\nu}(t))$ is the **Hausdorff metric**

$$\max \{ \sup_{x \in A_{\mu}(t)} \inf_{y \in A_{\nu}(t)} |x - y|, \sup_{x \in A_{\nu}(t)} \inf_{x \in A_{\mu}(t)} |x - y| \}.$$

Several other Hausdorff-like metrics on some families of fuzzy sets were proposed by Boxer in 1997, Fan in 1998 and Brass in 2002; Brass also argued the non-existence of a "good" such metric.

If q is a quasi-metric on [0,1] and S is a finite set, then $Q(\mu,\nu) = \sup_{x \in S} q(\mu(x),\nu(x))$ is a quasi-metric on fuzzy subsets of S.

Cf. fuzzy Hamming distance in Chap. 11 and, in Chap. 23, fuzzy set distance and fuzzy polynucleotide metric. Cf. fuzzy metric spaces in Chap. 3 for fuzzy-valued generalizations of metrics and, for example, [Bloc99] for a survey.

• Metrics between intuitionistic fuzzy sets

An intuitionistic fuzzy subset of a set S is (Atanassov 1999) an ordered pair of mappings $\mu, \nu : \to [0,1]$ with $0 \le \mu(x) + \nu(x) \le 1$ for all $x \in S$, representing the "degree of membership" and the "degree of non-membership" of $x \in S$, respectively. It is an ordinary fuzzy subset if $\mu(x) + \nu(x) = 1$ for all $x \in S$.

Given two intuitionistic fuzzy subsets $(\mu(x), \nu(x))$ and $(\mu'(x), \nu'(x))$ of a finite set $S = \{x_1, \ldots, x_n\}$, their **Atanassov distances** (1999) are:

$$\frac{1}{2}\sum_{i=1}^{n}(|\mu(x_i)-\mu'(x_i)|+|\nu(x_i)-\nu'(x_i)|)$$
 (Hamming distance) and

$$\sqrt{\frac{1}{2}\sum_{i=1}^{n}((\mu(x_i)-\mu'(x_i))^2+(\nu(x_i)-\nu'(x_i)^2)}$$
 (Euclidean distance).

Their Grzegorzewski distances (2004) are:

$$\sum_{i=1}^{n} \max\{|\mu(x_i) - \mu'(x_i)|, |\nu(x_i) - \nu'(x_i)|\}$$
 (Hamming distance) and

$$\sqrt{\sum_{i=1}^{n} \max\{(\mu(x_i) - \mu'(x_i))^2, (\nu(x_i) - \nu'(x_i))^2\}}$$
 (Euclidean distance).

The normalized versions – dividing the above four sums by n – were proposed also.

Szmidt and Kacprzyk (1997) proposed modification of the above, adding $\pi(x) - \pi'(x)$, where $\pi(x)$ is the third mapping $1 - \mu(x) - \nu(x)$.

An interval-valued fuzzy subset of a set S is a mapping $\mu : \to [I]$, where [I] is the set of closed intervals $[a^-, a^+] \subseteq [0, 1]$. Let $\mu(x) = [\mu^-(x), \mu^+(x)]$, where $0 \le \mu^-(x) \le \mu^+(x) \le 1$ and an interval-valued fuzzy subset is an ordered pair of mappings μ^- and μ^+ . This notion is very close to the above intuitionistic one; so, the above distance can easily be adapted. For example, $\sum_{i=1}^n \max\{|\mu^-(x_i) - \mu'^-(x_i)|, |\mu^+(x_i) - \mu'^+(x_i)|\}$ is a Hamming distance between interval-valued fuzzy subsets (μ^-, μ^+) and (μ'^-, μ'^+) .

• Compact quantum metric space

Let V be a normed space (or, more generally, a **locally convex** topological vector space, and let V' be its **continuous dual space**, i.e., the set of all continuous linear functionals f on V. The weak* topology (or Gelfand topology) on V' is defined as the weakest (i.e., with the fewest open sets) topology on V' such that, for every $x \in V$, the map $F_x : V' \to \mathbb{R}$ defined by $F_x(f) = f(x)$ for all $f \in V'$, remains continuous.

An order-unit space is a partially ordered real (complex) vector space (A, \preceq) with a distinguished element e, called an order unit, which satisfies the following properties:

- 1. For any $a \in A$, there exists $r \in \mathbb{R}$ with $a \leq re$.
- 2. If $a \in A$ and $a \leq re$ for all positive $r \in \mathbb{R}$, then $a \leq 0$ (Archimedean property).

The main example of an order-unit space is the vector space of all self-adjoint elements in a $unital\ C^*$ -algebra with the identity element being the order unit. Here a C^* -algebra is a $Banach\ algebra$ over $\mathbb C$ equipped with a special involution. It is called unital if it has a unit (multiplicative identity element); such C^* -algebras are also called, roughly, $compact\ non-commutative\ topological\ spaces$.

The typical example of a unital C^* -algebra is the complex algebra of linear operators on a complex **Hilbert space** which is topologically closed in the norm topology of operators, and is closed under the operation of taking adjoints of operators.

The state space of an order-unit space (A, \leq, e) is the set $S(A) = \{f \in A'_+ : ||f|| = 1\}$ of states, i.e., continuous linear functionals f with ||f|| = f(e) = 1.

Rieffel's **compact quantum metric space** is a pair $(A, ||.||_{Lip})$, where (A, \leq, e) is an order-unit space, and $||.||_{Lip}$ is a semi-norm on A (with values in $[0, +\infty]$), called *Lipschitz semi-norm*, which satisfies the following conditions:

- 1. For $a \in A$, $||a||_{Lip} = 0$ holds if and only if $a \in \mathbb{R}e$.
- 2. The metric $d_{Lip}(f,g) = \sup_{a \in A: ||a||_{Lip} \le 1} |f(a) g(a)|$ generates on the state space S(A) its weak* topology.

So, one has a usual metric space $(S(A), d_{Lip})$. If the order-unit space (A, \leq, e) is a C^* -algebra, then d_{Lip} is the **Connes metric**, and if, moreover, the C^* -algebra is non-commutative, the metric space $(S(A), d_{Lip})$ is called a **non-commutative metric space**.

The expression quantum metric space comes from the belief, by many experts in Quantum Gravity and String Theory, that the Planck-scale geometry of space-time is similar to one coming from such non-commutative C^* -algebras.

For example, Non-commutative Field Theory supposes that, on sufficiently small (quantum) distances, the spatial coordinates do not commute, i.e., it is impossible to measure exactly the position of a particle with respect to more than one axis.

• Dynamical system

A (deterministic) **dynamical system** is a tuple (T, X, f) consisting of a metric space (X, d), called the *phase space*, a *time set* $T \subseteq \mathbb{R}$, and a continuous function $f: T \times X \to X$, called the *evolution law*. The system is *discrete* (or *cascade*) if $T = \{0, 1, 2...\}$; it is *continuous* (or *real*, *flow*) if T is an open interval in \mathbb{R} .

The dynamical systems are studied in Control Theory in the context of stability of systems; Chaos Theory considers the systems with maximal possible instability.

A discrete dynamical system is defined by a self-map $f: X \to X$. For any $x \in X$, its *orbit* (or *trajectory*) is the sequence $\{f^n(x)\}_n$; here $f^n(x) = f(f^{n-1}(x))$ with $f^0(x) = x$. The orbit of $x \in X$ is called *periodic* if $f^n(x) = x$ for some n > 0.

A pair $(x, y) \in X \times X$ is called *proximal* if $\underline{\lim}_{n \to \infty} d(f^n(x), f^n(y)) = 0$, and *distal* otherwise. The system is called *distal* if any pair (x, y) of distinct points is distal.

The dynamical system is called *expansive* if there exists a constant D > 0 such that the inequality $d(f^n(x), f^n(y)) \ge D$ holds for any distinct $x, y \in X$ and some n.

An attractor is a closed subset A of X such that there exists an open neighborhood U of A with the property that $\lim_{n\to\infty} d(f^n(b), A) = 0$ for every $b \in U$, i.e., A attracts all nearby orbits. Here $d(x, A) = \inf_{y \in A} d(x, y)$ is the **point-set distance**.

If for large n and small r there exists a number α such that

$$C(X,n,r) = \frac{|\{(i,j): d(f^i(x),f^j(x)) \leq r, 1 \leq i,j \leq n\}|}{n^2} \sim r^{\alpha},$$

then α is called (Grassberger, Hentschel and Procaccia 1983) the *correlation dimension*.

• Universal metric space

A metric space (U,d) is called **universal** for a collection \mathcal{M} of metric spaces if any metric space (M,d_M) from \mathcal{M} is isometrically embeddable in (U,d), i.e., there exists a mapping $f:M\to U$ which satisfies $d_M(x,y)=d(f(x),f(y))$ for any $x,y\in M$. Some examples follow.

Every separable metric space (X, d) isometrically embeds (Fréchet 1909) in (a non-separable) **Banach space** l_{∞}^{∞} . In fact, $d(x, y) = \sup_{i} |d(x, a_{i}) - d(y, a_{i})|$, where $(a_{1}, \ldots, a_{i}, \ldots)$ is a dense countable subset of X.

Every metric space isometrically embeds (Kuratowski 1935) in the **Banach space** $L^{\infty}(X)$ of bounded functions $f: X \to \mathbb{R}$ with the norm $\sup_{x \in X} |f(x)|$.

The **Urysohn space** is a **homogeneous** complete separable space which is the universal metric space for all separable metric spaces.

The **Hilbert cube** is the universal metric space for the class of metric spaces with a countable base.

The **graphic** metric space of the **Random graph** (Rado 1964; the vertex-set consists of all prime numbers $p \equiv 1 \pmod{4}$ with pq being an edge if p is a quadratic residue modulo q) is the universal metric space for any finite or countable metric space with distances 0, 1 and 2 only. It is a discrete analog of the Urysohn space.

There exists a metric d on \mathbb{R} , inducing the usual (interval) topology, such that (\mathbb{R},d) is a universal metric space for all finite metric spaces (Holsztynski 1978). The Banach space l_{∞}^n is a universal metric space for all metric spaces (X,d) with $|X| \leq n+2$ (Wolfe 1967). The Euclidean space \mathbb{E}^n is a universal metric space for all ultrametric spaces (X,d) with $|X| \leq n+1$; the space of all finite functions $f(t) : \mathbb{R}_{\geq 0} \to \mathbb{R}$ equipped with the metric $d(f,g) = \sup\{t : f(t) \neq g(t)\}$ is a universal metric space for all ultrametric spaces (A. Lemin and V. Lemin 1996).

The universality can be defined also for mappings, other than isometric embeddings, of metric spaces, say, bi-Lipschitz embedding, etc. For example, any compact metric space is a continuous image of the **Cantor set** with the natural metric |x-y| inherited from \mathbb{R} , and any complete separable metric space is a continuous image of the space of irrational numbers.

• Constructive metric space

A constructive metric space is a pair (X, d), where X is some set of constructive objects (usually, words over an alphabet), and d is an algorithm converting any pair of elements of X into a constructive real number d(x, y) such that d becomes a metric on X.

• Effective metric space

Let $\{x_n\}_{n\in\mathbb{N}}$ be a sequence of elements from a given **complete** metric space (X,d) such that the set $\{x_n:n\in\mathbb{N}\}$ is *dense* in (X,d). Let $\mathcal{N}(m,n,k)$ be the *Cantor number* of a triple $(n,m,k)\in\mathbb{N}^3$, and let $\{q_k\}_{k\in\mathbb{N}}$ be a fixed total standard numbering of the set \mathbb{Q} of rational numbers.

The triple $(X, d, \{x_n\}_{n \in \mathbb{N}})$ is called an **effective metric space** [Hemm02] if the set $\{\mathcal{N}(n, m, k) : d(x_m, x_n) < q_k\}$ is recursively enumerable. It is an adaptation of Weihrauch's notion of **computable metric space** (or **recursive metric space**).

Chapter 2 Topological Spaces

A topological space (X, τ) is a set X with a topology τ , i.e., a collection of subsets of X with the following properties:

- 1. $X \in \tau$, $\emptyset \in \tau$.
- 2. If $A, B \in \tau$, then $A \cap B \in \tau$.
- 3. For any collection $\{A_{\alpha}\}_{\alpha}$, if all $A_{\alpha} \in \tau$, then $\cup_{\alpha} A_{\alpha} \in \tau$.

The sets in τ are called *open sets*, and their complements are called *closed sets*. A *base* of the topology τ is a collection of open sets such that every open set is a union of sets in the base. The coarsest topology has two open sets, the empty set and X, and is called the *trivial topology* (or *indiscrete topology*). The finest topology contains all subsets as open sets, and is called the *discrete topology*.

In a metric space (X,d) define the *open ball* as the set $B(x,r) = \{y \in X : d(x,y) < r\}$, where $x \in X$ (the *center* of the ball), and $r \in \mathbb{R}, r > 0$ (the *radius* of the ball). A subset of X which is the union of (finitely or infinitely many) open balls, is called an *open set*. Equivalently, a subset U of X is called *open* if, given any point $x \in U$, there exists a real number $\epsilon > 0$ such that, for any point $y \in X$ with $d(x,y) < \epsilon, y \in U$.

Any metric space is a topological space, the topology (metric topology, topology induced by the metric d) being the set of all open sets. The metric topology is always T_4 (see below a list of topological spaces). A topological space which can arise in this way from a metric space, is called a metrizable space.

A quasi-pseudo-metric topology is a topology on X induced similarly by a quasi-semi-metric d on X, using the set of open d-balls B(x,r) as the base. In particular, quasi-metric topology and pseudo-metric topology are the terms used in Topology for the case of, respectively, quasi-metric and semi-metric d. In general, those topologies are not T_0 .

Given a topological space (X, τ) , a neighborhood of a point $x \in X$ is a set containing an open set which in turn contains x. The closure of a subset of a topological space is the smallest closed set which contains it. An open cover of X is a collection \mathcal{L} of open sets, the union of which is X; its subcover is a cover \mathcal{K} such that every member of \mathcal{K} is a member of \mathcal{L} ; its refinement

is a cover \mathcal{K} , where every member of \mathcal{K} is a subset of some member of \mathcal{L} . A collection of subsets of X is called *locally finite* if every point of X has a neighborhood which meets only finitely many of these subsets.

A subset $A \subset X$ is called *dense* if it has non-empty intersection with every non-empty open set or, equivalently, if the only closed set containing it is X. In a metric space (X,d), a *dense set* is a subset $A \subset X$ such that, for any $x \in X$ and any $\epsilon > 0$, there exists $y \in A$, satisfying $d(x,y) < \epsilon$. A *local base* of a point $x \in X$ is a collection \mathcal{U} of neighborhoods of x such that every neighborhood of x contains some member of \mathcal{U} .

A function from one topological space to another is called *continuous* if the preimage of every open set is open. Roughly, given $x \in X$, all points close to x map to points close to f(x). A function f from one metric space (X, d_X) to another metric space (Y, d_Y) is *continuous* at the point $c \in X$ if, for any positive real number ϵ , there exists a positive real number δ such that all $x \in X$ satisfying $d_X(x, c) < \delta$ will also satisfy $d_Y(f(x), f(c)) < \epsilon$; the function is continuous on an interval I if it is continuous at any point of I.

The following classes of topological spaces (up to T_4) include any metric space:

• T_0 -space

A T_0 -space (or $Kolmogorov\ space$) is a topological space (X, τ) fulfilling the T_0 -separation axiom: for every two points $x, y \in X$ there exists an open set U such that $x \in U$ and $y \notin U$, or $y \in U$ and $x \notin U$ (every two points are topologically distinguishable).

• T_1 -space

A T_1 -space (or accessible space) is a topological space (X, τ) fulfilling the T_1 -separation axiom: for every two points $x, y \in X$ there exist two open sets U and V such that $x \in U$, $y \notin U$, and $y \in V$, $x \notin V$ (every two points are separated). T_1 -spaces are always T_0 .

• T_2 -space

A T_2 -space (or Hausdorff space, separated space) is a topological space (X, τ) fulfilling the T_2 -axiom: every two points $x, y \in X$ have disjoint neighborhoods. Thus, (X, τ) is Hausdorff if and only if it is both T_0 and preregular, i.e., any two topologically distinguishable points in it are separated by neighbourhoods. T_2 -spaces are always T_1 .

• Regular space

A regular space is a topological space in which every neighborhood of a point contains a closed neighborhood of the same point.

• T_3 -space

A T_3 -space (or Vietoris space, regular Hausdorff space) is a topological space which is T_1 and regular.

Completely regular space

A completely regular space (or *Tychonoff space*) is a **Hausdorff space** (X, τ) in which any closed set A and any $x \notin A$ are functionally separated.

Two subsets A and B of X are functionally separated if there exists a continuous function $f: X \to [0,1]$ such that f(x) = 0 for any $x \in A$, and f(y) = 1 for any $y \in B$.

• Perfectly normal space

A **perfectly normal space** is a topological space (X, τ) in which any two disjoint closed subsets of X are functionally separated.

• Normal space

A **normal space** is a topological space in which, for any two disjoint closed sets A and B, there exist two disjoint open sets U and V such that $A \subset U$, and $B \subset V$.

• T_4 -space

A T_4 -space (or *Tietze space*, normal Hausdorff space) is a topological space which is T_1 and normal. Any metric space (X, d) is a T_4 -space.

• Completely normal space

A **completely normal space** is a topological space in which any two separated sets have disjoint neighborhoods.

Sets A and B are *separated* in X if each is disjoint from the other's closure.

• T_5 -space

A T_5 -space (or completely normal Hausdorff space) is a topological space which is **completely normal** and T_1 . T_5 -spaces are always T_4 .

• T_6 -space

A T_6 -space (or perfectly normal Hausdorff space) is a topological space which is T_1 and **perfectly normal**. T_6 -spaces are always T_5 .

• Moore space

A Moore space is a regular space with a development.

A development is a sequence $\{\mathcal{U}_n\}_n$ of open covers such that, for every $x \in X$ and every open set A containing x, there exists n such that $St(x,\mathcal{U}_n) = \bigcup \{U \in \mathcal{U}_n : x \in U\} \subset A$, i.e., $\{St(x,\mathcal{U}_n)\}_n$ is a neighborhood base at x.

• Separable space

A **separable space** is a topological space which has a countable dense subset.

• Lindelöf space

A Lindelöf space is a topological space in which every open cover has a countable subcover.

• First-countable space

A topological space is called **first-countable** if every point has a countable local base. Any metric space is first-countable.

• Second-countable space

A topological space is called **second-countable** if its topology has a countable base. Second-countable spaces are always **separable**, **first-countable**, and **Lindelöf**.

For metric spaces the properties of being second-countable, **separable**, and **Lindelöf** are all equivalent.

The Euclidean space \mathbb{E}^n with its usual topology is second-countable.

• Baire space

A Baire space is a topological space in which every intersection of countably many dense open sets is dense. Every complete metric space is a Baire space. Every locally compact T_2 -space (hence, every manifold) is a Baire space.

• Alexandrov space

An **Alexandrov space** is a topological space in which every intersection of arbitrarily many open sets is open.

A topological space is called a P-space if every G_{δ} -set (i.e., the intersection of countably many open sets) is open.

A topological space (X, τ) is called a Q-space if every subset $A \subset X$ is a G_{δ} -set.

• Connected space

A topological space (X, τ) is called **connected** if it is not the union of a pair of disjoint non-empty open sets. In this case the set X is called a *connected set*.

A topological space (X, τ) is called **locally connected** if every point $x \in X$ has a local base consisting of connected sets.

A topological space (X, τ) is called **path-connected** (or 0-connected) if for every points $x, y \in X$ there is a path γ from x to y, i.e., a continuous function $\gamma : [0, 1] \to X$ with $\gamma(x) = 0, \gamma(y) = 1$.

A topological space (X, τ) is called **simply connected** (or 1-connected) if it consists of one piece, and has no circle-shaped "holes" or "handles" or, equivalently, if every continuous curve of X is contractible, i.e., can be reduced to one of its points by a continuous deformation.

• Paracompact space

A topological space is called **paracompact** if every open cover of it has an open locally finite refinement. Every metric (moreover, **metrizable**) space is paracompact.

• Totally bounded space

A topological space (X, τ) is called **totally bounded** (or *pre-compact*) if it can be covered by finitely many subsets of any fixed size.

A metric space (X, d) is a **totally bounded metric space** if, for every real number r > 0, there exist finitely many open balls of radius r, whose union is equal to X.

Compact space

A topological space (X, τ) is called **compact** if every open cover of X has a finite subcover.

Compact spaces are always Lindelöf, totally bounded, and paracompact. A metric space is compact if and only if it is complete and totally bounded. A subset of a Euclidean space \mathbb{E}^n is compact if and only if it is closed and bounded.

There exist a number of topological properties which are equivalent to compactness in metric spaces, but are nonequivalent in general topological spaces. Thus, a metric space is compact if and only if it is a *sequentially*

compact space (every sequence has a convergent subsequence), or a countably compact space (every countable open cover has a finite subcover), or a pseudo-compact space (every real-valued continuous function on the space is bounded), or a weakly countably compact space (i.e., every infinite subset has an accumulation point).

Continuum

A continuum is a compact connected T_2 -space.

• Locally compact space

A topological space is called **locally compact** if every point has a local base consisting of compact neighborhoods. The Euclidean spaces \mathbb{E}^n and the spaces \mathbb{Q}_p of *p-adic numbers* are locally compact.

A topological space (X, τ) is called a k-space if, for any compact set $Y \subset X$ and $A \subset X$, the set A is closed whenever $A \cap Y$ is closed. The k-spaces are precisely quotient images of locally compact spaces.

• Locally convex space

A topological vector space is a real (complex) vector space V which is a T_2 -space with continuous vector addition and scalar multiplication. It is a uniform space (cf. Chap. 3).

A locally convex space is a topological vector space whose topology has a base, where each member is a convex balanced absorbent set. A subset A of V is called convex if, for all $x, y \in A$ and all $t \in [0, 1]$, the point $tx + (1-t)y \in A$, i.e., every point on the line segment connecting x and y belongs to A. A subset A is balanced if it contains the line segment between x and -x for every $x \in A$; A is absorbent if, for every $x \in V$, there exist t > 0 such that $tx \in A$.

The locally convex spaces are precisely vector spaces with topology induced by a family $\{||.||_{\alpha}\}$ of semi-norms such that x=0 if $||x||_{\alpha}=0$ for every α .

Any metric space (V, ||x - y||) on a real (complex) vector space V with a **norm metric** ||x - y|| is a locally convex space; each point of V has a local base consisting of convex sets. Every L_p with 0 is an example of a vector space which is not locally convex.

• Fréchet space

A **Fréchet space** is a **locally convex space** (V, τ) which is complete as a **uniform space** and whose topology is defined using a countable set of semi-norms $||.||_1, \ldots, ||.||_n, \ldots$, i.e., a subset $U \subset V$ is open in (V, τ) if, for every $u \in U$, there exist $\epsilon > 0$ and $N \ge 1$ with $\{v \in V : ||u-v||_i < \epsilon \text{ if } i \le N\} \subset U$.

A Fréchet space is precisely a locally convex **F-space** (cf. Chap. 5). Its topology can be induced by a **translation invariant metric** and it is a complete and **metrizable space** with respect to this topology. But this topology may be induced by many such metrics; so, there is no natural notion of distance between points of a Fréchet space.

Every Banach space is a Fréchet space.

• Countably-normed space

A countably-normed space is a locally convex space (V, τ) whose topology is defined using a countable set of *compatible norms* $||.||_1, ..., ||.||_n, ...$ It means that, if a sequence $\{x_n\}_n$ of elements of V that is fundamental in the norms $||.||_i$ and $||.||_j$ converges to zero in one of these norms, then it also converges in the other. A countably-normed space is a **metrizable space**, and its metric can be defined by

$$\sum_{n=1}^{\infty} \frac{1}{2^n} \frac{||x-y||_n}{1+||x-y||_n}.$$

• Metrizable space

A topological space is called **metrizable** if it is homeomorphic to a metric space, i.e., X admits a metric d such that the set of open d-balls $\{B(x,r): r>0\}$ forms a neighborhood base at each point $x\in X$.

Metrizable spaces are always **paracompact** T_2 -spaces (hence, normal and **completely regular**), and **first-countable**.

A topological space is called **locally metrizable** if every point in it has a metrizable neighborhood.

A topological space (X, τ) is called **submetrizable** if there exists a metrizable topology τ' on X which is coarser than τ .

A topological space (X, τ) is called **protometrizable** if it is paracompact and has an *orthobase*, i.e., a base \mathcal{B} such that, for $\mathcal{B}' \subset \mathcal{B}$, either $\cap \mathcal{B}'$ is open, or \mathcal{B}' is a local base at the unique point in $\cap \mathcal{B}'$.

Some examples of other direct generalizations of metrizable spaces follow.

A sequential space is a quotient image of a metrizable space.

Morita's M-space is a topological space (X, τ) from which there exists a continuous map f onto a metrizable topological space (Y, τ') such that f is closed and $f^{-1}(y)$ is countably compact for each $y \in Y$.

Ceder's M_1 -space is a topological space (X, τ) having a σ -closure-preserving base (metrizable spaces have σ -locally finite bases).

Okuyama's σ -space is a topological space (X, τ) having a σ -locally finite net, i.e., a collection \mathcal{U} of subsets of X such that, given of a point $x \in U$ with U open, there exists $U' \in \mathcal{U}$ with $x \in U' \subset U$ (a base is a net consisting of open sets). Every compact subset of a σ -space is metrizable.

Michael's **cosmic space** is a topological space (X, τ) having a countable net (equivalently, a Lindelöf σ -space). It is exactly a continuous image of a separable metric space. A T_2 -space is called **analytic** if it is a continuous image of a complete separable metric space; it is called a **Lusin space** if, moreover, the image is one-to-one.

• Quasi-metrizable space

A topological space (X, τ) is called a **quasi-metrizable space** if X admits a quasi-metric d such that the set of open d-balls $\{B(x, r) : r > 0\}$ forms a neighborhood base at each point $x \in X$.

A more general γ -space is a topological space admitting a γ -metric d (i.e., a function $d: X \times X \to \mathbb{R}_{\geq 0}$ with $d(x, z_n) \to 0$ whenever $d(x, y_n) \to 0$ and $d(y_n, z_n) \to 0$) such that the set of open forward d-balls $\{B(x, r) : r > 0\}$ forms a neighborhood base at each point $x \in X$.

The Sorgenfrey line is the topological space (\mathbb{R}, τ) defined by the base $\{[a,b): a,b \in \mathbb{R}, a < b\}$. It is not metrizable but it is a first-countable separable and paracompact T_5 -space; neither it is second-countable, nor locally compact or locally connected. However, the Sorgenfrey line is quasimetrizable by the **Sorgenfrey quasi-metric** (cf. Chap. 12) defined as y-x if y > x, and 1 otherwise.

• Symmetrizable space

A topological space (X, τ) is called **symmetrizable** (and τ is called the **distance topology**) if there is a **symmetric** d on X (i.e., a distance $d: X \times X \to \mathbb{R}_{\geq 0}$ with d(x,y) = 0 implying x = y) such that a subset $U \subset X$ is open if and only if, for each $x \in U$, there exists $\epsilon > 0$ with $B(x,\epsilon) = \{y \in X : d(x,y) < \epsilon\} \subset U$. In other words, a subset $H \subset X$ is closed if and only if $d(x,H) = \inf_y \{d(x,y) : y \in H\} > 0$ for each $x \in X \setminus U$. A symmetrizable space is **metrizable** if and only if it is a Morita's M-space.

In Topology, the term **semi-metrizable space** refers to a topological space (X,τ) admitting a symmetric d such that, for each $x\in X$, the family $\{B(x,\epsilon):\epsilon>0\}$ of balls forms a (not necessarily open) neighborhood base at x. In other words, a point x is in the closure of a set H if and only if d(x,H)=0. A topological space is semi-metrizable if and only if it is symmetrizable and **first-countable**. Also, a symmetrizable space is semi-metrizable if and only if it is a $Fr\acute{e}chet-Urysohn$ space (or E-space), i.e., for any subset A and for any point x of its closure, there is a sequence in A converging to x.

• Hyperspace

A hyperspace of a topological space (X, τ) is a topological space on the set CL(X) of all non-empty closed (or, moreover, compact) subsets of X. The topology of a hyperspace of X is called a hypertopology. Examples of such a hit-and-miss topology are the Vietoris topology, and the Fell topology. Examples of such a weak hyperspace topology are the Hausdorff metric topology, and the Wijsman topology.

Discrete topological space

A topological space (X, τ) is **discrete** if τ is the discrete topology (the finest topology on X), i.e., containing all subsets of X as open sets. Equivalently, it does not contain any *limit point*, i.e., it consists only of *isolated points*.

Indiscrete topological space

A topological space (X, τ) is **indiscrete** if τ is the *indiscrete topology* (the coarsest topology on X), i.e., having only two open sets, \emptyset and X. It can be considered as the semi-metric space (X, d) with the **indiscrete semi-metric**: d(x, y) = 0 for any $x, y \in X$.

• Extended topology

Consider a set X and a map $cl: P(X) \to P(X)$, where P(X) is the set of all subsets of X. The set cl(A) (for $A \subset X$), its dual set $int(A) = X \setminus cl(X \setminus A)$ and the map $N: X \to P(X)$ with $N(x) = \{A \subset X : x \in int(A)\}$ are called the *closure*, *interior* and *neighborhood* map, respectively. So, $x \in cl(A)$ is equivalent to $X \setminus A \in P(X) \setminus N(x)$. A subset $A \subset X$ is *closed* if A = cl(A) and *open* if A = int(A). Consider the following possible properties of cl; they are meant to hold for all $A, B \in P(X)$:

- 1. $cl(\emptyset) = \emptyset$.
- 2. $A \subseteq B$ implies $cl(A) \subseteq cl(B)$ (isotony).
- 3. $A \subseteq cl(A)$ (enlarging).
- 4. $cl(A \cup B) = cl(A) \cup cl(B)$ (linearity, and, in fact, 4 implies 2).
- 5. cl(cl(A)) = cl(A) (idempotency).

The pair (X, cl) satisfying 1 is called an **extended topology** if 2 holds, a **Brissaud space** (Brissaud 1974) if 3 holds, a **neighborhood space** (Hammer 1964) if 2 and 3 hold, a **Smyth space** (Smyth 1995) if 4 holds, a **pretopology** (Čech 1966) if 3 and 4 hold, and a **closure space** (Soltan 1984) if 2, 3 and 5 hold.

(X, cl) is the usual topology, in closure terms, if 1, 3, 4 and 5 hold.

Chapter 3

Generalizations of Metric Spaces

Some immediate generalizations of the notion of metric, for example, **quasimetric**, **near-metric**, **extended metric**, were defined in Chap. 1. Here we give some generalizations in the direction of Topology, Probability, Algebra, etc.

3.1 m-metrics

• m-hemi-metric

Let X be a set. A function $d: X^{m+1} \to \mathbb{R}$ is called m-hemi-metric if:

- 1. d is non-negative, i.e., $d(x_1, \ldots, x_{n+1}) \ge 0$ for all $x_1, \ldots, x_{n+1} \in X$.
- 2. d is totally symmetric, i.e., satisfies $d(x_1, \ldots, x_{m+1}) = d(x_{\pi(1)}, \ldots, x_{\pi(m+1)})$ for all $x_1, \ldots, x_{m+1} \in X$ and for any permutation π of $\{1, \ldots, m+1\}$.
- 3. d is zero conditioned, i.e., $d(x_1, \ldots, x_{m+1}) = 0$ if and only if x_1, \ldots, x_{m+1} are not pairwise distinct.
- 4. For all $x_1, \ldots, x_{m+2} \in X$, d satisfies the m-simplex inequality:

$$d(x_1,\ldots,x_{m+1}) \le \sum_{i=1}^{m+1} d(x_1,\ldots,x_{i-1},x_{i+1},\ldots,x_{m+2}).$$

• 2-metric

Let X be a set. A function $d: X^3 \to \mathbb{R}$ is called a 2-metric if d is non-negative, totally symmetric, zero conditioned, and satisfies the **tetra-hedron inequality**

$$d(x_1, x_2, x_3) \le d(x_4, x_2, x_3) + d(x_1, x_4, x_3) + d(x_1, x_2, x_4).$$

It is the most important case m=2 of the m-hemi-metric.

• (m, s)-super-metric

Let X be a set, and let s be a positive real number. A function $d: X^{m+1} \to \mathbb{R}$ is called (m, s)-super-metric [DeDu03] if d is non-negative, totally symmetric, zero conditioned, and satisfies t the (m, s)-simplex inequality:

$$sd(x_1,\ldots,x_{m+1}) \le \sum_{i=1}^{m+1} d(x_1,\ldots,x_{i-1},x_{i+1},\ldots,x_{m+2}).$$

An (m, s)-super-metric is an m-hemi-metric if $s \ge 1$.

3.2 Indefinite metrics

• Indefinite metric

An **indefinite metric** (or G-metric) on a real (complex) vector space V is a bilinear (in the complex case, sesquilinear) form G on V, i.e., a function $G: V \times V \to \mathbb{R}$ (\mathbb{C}), such that, for any $x, y, z \in V$ and for any scalars α, β , we have the following properties: $G(\alpha x + \beta y, z) = \alpha G(x, z) + \beta G(y, z)$, and $G(x, \alpha y + \beta z) = \overline{\alpha} G(x, y) + \overline{\beta} G(x, z)$, where $\overline{\alpha} = \overline{a + bi} = a - bi$ denotes the complex conjugation.

If a positive-definite form G is symmetric, then it is an *inner product* on V, and one can use it to canonically introduce a *norm* and the corresponding **norm metric** on V. In the case of a general form G, there is neither a norm, nor a metric canonically related to G, and the term **indefinite metric** only recalls the close relation of such forms with certain metrics in vector spaces (cf. Chaps. 7 and 26).

The pair (V, G) is called a *space with an indefinite metric*. A finite-dimensional space with an indefinite metric is called a *bilinear metric space*. A **Hilbert space** H, endowed with a continuous G-metric, is called a *Hilbert space with an indefinite metric*. The most important example of such space is a J-space.

A subspace L in a space (V, G) with an indefinite metric is called a *positive subspace*, negative subspace, or neutral subspace, depending on whether G(x,x) > 0, G(x,x) < 0, or G(x,x) = 0 for all $x \in L$.

• Hermitian G-metric

A Hermitian G-metric is an indefinite metric G^H on a complex vector space V such that, for all $x, y \in V$, we have the equality

$$G^H(x,y) = \overline{G^H(y,x)},$$

where $\overline{\alpha} = \overline{a + bi} = a - bi$ denotes the *complex conjugation*.

• Regular G-metric

A regular G-metric is a continuous indefinite metric G on a Hilbert space H over \mathbb{C} , generated by an invertible $Hermitian\ operator\ T$ by the formula

$$G(x,y) = \langle T(x), y \rangle,$$

where \langle , \rangle is the *inner product* on H.

A Hermitian operator on a Hilbert space H is a linear operator T on H, defined on a domain D(T) of H such that $\langle T(x), y \rangle = \langle x, T(y) \rangle$ for any $x, y \in D(T)$. A bounded Hermitian operator is either defined on the whole of H, or can be so extended by continuity, and then $T = T^*$. On a finite-dimensional space a Hermitian operator can be described by a Hermitian $matrix((a_{ij})) = ((\overline{a}_{ji}))$.

• J-metric

A *J*-metric is a continuous indefinite metric G on a Hilbert space H over \mathbb{C} , defined by a certain *Hermitian involution* J on H by the formula

$$G(x,y) = \langle J(x), y \rangle,$$

where \langle , \rangle is the *inner product* on H.

An involution is a mapping H onto H whose square is the identity mapping. The involution J may be represented as $J = P_+ - P_-$, where P_+ and P_- are orthogonal projections in H, and $P_+ + P_- = H$. The rank of indefiniteness of the J-metric is defined as min $\{\dim P_+, \dim P_-\}$.

The space (H, G) is called a *J-space*. A *J-space* with finite rank of indefiniteness is called a *Pontryagin space*.

3.3 Topological generalizations

• Metametric space

A metametric space (Väisälä 2003) is a pair (X,d), where X is a set, and d is a non-negative symmetric function $d: X \times X \to \mathbb{R}$ such that d(x,y) = 0 implies x = y and triangle inequality $d(x,y) \le d(x,z) + d(z,y)$ holds for all $x,y,z \in X$. A metametric space is metrizable: the metametric d defines the same topology as the metric d' defined by d'(x,x) = 0 and d'(x,y) = d(x,y) if $x \neq y$. A metametric d induces a Hausdorff topology with the usual definition of a $ball\ B(x_0,r) = \{x \in X : d(x_0,x) < r\}$.

• Resemblance

Let X be a set. A function $d: X \times X \to \mathbb{R}$ is called a **resemblance** on X if d is *symmetric* and if, for all $x, y \in X$, either $d(x, x) \leq d(x, y)$ (in which case d is called a **forward resemblance**), or $d(x, x) \geq d(x, y)$ (in which case d is called a **backward resemblance**).

Every resemblance d induces a *strict partial order* \prec on the set of all unordered pairs of elements of X by defining $\{x,y\} \prec \{u,v\}$ if and only if d(x,y) < d(u,v).

For any backward resemblance d, the forward resemblance -d induces the same partial order.

• w-distance

Given a metric space (X,d), a w-distance on X (Kada, Suzuki and Takahashi 1996) is a non-negative function $p: X \times X \to \mathbb{R}$ which satisfies the following conditions:

- 1. $p(x,z) \le p(x,y) + p(y,z)$ for all $x, y, z \in X$.
- 2. For any $x \in X$, the function $p(x,.): X \to \mathbb{R}$ is lower semicontinuous, i.e., if a sequence $\{y_n\}_n$ in X converges to $y \in X$, then $p(x,y) \leq \liminf_{n \to \infty} p(x,y_n)$.
- 3. For any $\epsilon > 0$, there exists $\delta > 0$ such that $p(z, x) \leq \delta$ and $p(z, y) \leq \delta$ imply $d(x, y) \leq \epsilon$, for each $x, y, z \in X$.

• τ -distance space

A τ -distance space is a pair (X, f), where X is a topological space and f is an Aamri-Moutawakil's τ -distance on X, i.e., a non-negative function $f: X \times X \to \mathbb{R}$ such that, for any $x \in X$ and any neighborhood U of x, there exists $\epsilon > 0$ with $\{y \in X : f(x,y) < \epsilon\} \subset U$.

Any distance space (X,d) is a τ -distance space for the topology τ_f defined as follows: $A \in \tau_f$ if, for any $x \in X$, there exists $\epsilon > 0$ with $\{y \in X : f(x,y) < \epsilon\} \subset A$. However, there exist non-metrizable τ -distance spaces. A τ -distance f(x,y) need be neither symmetric, nor vanishing for x = y; for example, $e^{|x-y|}$ is a τ -distance on $X = \mathbb{R}$ with usual topology.

• Proximity space

A **proximity space** (Efremovich 1936) is a set X with a binary relation δ on the *power set* P(X) of all of its subsets which satisfies the following conditions:

- 1. $A\delta B$ if and only if $B\delta A$ (symmetry).
- 2. $A\delta(B \cup C)$ if and only if $A\delta B$ or $A\delta C$ (additivity).
- 3. $A\delta A$ if and only if $A \neq \emptyset$ (reflexivity).

The relation δ defines a **proximity** (or *proximity structure*) on X. If $A\delta B$ fails, the sets A and B are called *remote sets*.

Every metric space (X, d) is a proximity space: define $A\delta B$ if and only if $d(A, B) = \inf_{x \in A, y \in B} d(x, y) = 0$.

Every proximity on X induces a (**completely regular**) topology on X by defining the *closure operator* $cl: P(X) \to P(X)$ on the set of all subsets of X as $cl(A) = \{x \in X : \{x\}\delta A\}$.

• Uniform space

A uniform space is a topological space (with additional structure) providing a generalization of metric space, based on **set-set distance**.

A uniform space (Weil 1937) is a set X with an uniformity (or uniform structure) \mathcal{U} – a non-empty collection of subsets of $X \times X$, called entourages, with the following properties:

- 1. Every subset of $X \times X$ which contains a set of \mathcal{U} belongs to \mathcal{U} .
- 2. Every finite intersection of sets of \mathcal{U} belongs to \mathcal{U} .
- 3. Every set $V \in \mathcal{U}$ contains the diagonal, i.e., the set $\{(x,x) : x \in X\} \subset X \times X$.
- 4. If V belongs to \mathcal{U} , then the set $\{(y,x):(x,y)\in V\}$ belongs to \mathcal{U} .
- 5. If V belongs to \mathcal{U} , then there exists $V' \in \mathcal{U}$ such that $(x,z) \in V$ whenever $(x,y),(y,z) \in V'$.

Every metric space (X, d) is a uniform space. An entourage in (X, d) is a subset of $X \times X$ which contains the set $V_{\epsilon} = \{(x, y) \in X \times X : d(x, y) < \epsilon\}$ for some positive real number ϵ . Other basic example of uniform space are topological groups.

Every uniform space (X, \mathcal{U}) generates a topology consisting of all sets $A \subset X$ such that, for any $x \in A$, there is a set $V \in \mathcal{U}$ with $\{y : (x, y) \in V\} \subset A$.

Every uniformity induces a **proximity** σ where $A\sigma B$ if and only if $A\times B$ has non-empty intersection with any entourage.

A topological space admits a uniform structure inducing its topology if only if the topology is **completely regular** (cf. Chap. 2) and, also, if only if it is a *gauge space*, i.e., the topology is defined by a family of semi-metrics.

• Nearness space

A nearness space (Herrich 1974) is a set X with a nearness structure, i.e., a non-empty collection \mathcal{U} of families of subsets of X, called near families, with the following properties:

- 1. Each family refining a near family is near.
- 2. Every family with non-empty intersection is near.
- 3. V is near if $\{cl(A): A \in V\}$ is near, where cl(A) is $\{x \in X: \{\{x\}, A\} \in \mathcal{U}\}$.
- 4. \emptyset is near, while the set of all subsets of X is not.
- 5. If $\{A \cup B : A \in \mathcal{F}_1, B \in \mathcal{F}_2\}$ is near family, then so is \mathcal{F}_1 or \mathcal{F}_2 .

The uniform spaces are precisely paracompact nearness spaces.

Approach space

An approach space is a topological space providing a generalization of metric space, based on **point-set distance**.

An approach space (Lowen 1989) is a pair (X, D), where X is a set and D is a **point-set distance**, i.e., a function $X \times P(X) \to [0, \infty]$ (where P(X) is the set of all subsets of X) satisfying, for all $x \in X$ and all $A, B \in P(X)$, the following conditions:

- 1. $D(x, \{x\}) = 0$.
- 2. $D(x, \{\emptyset\}) = \infty$.

- 3. $D(x, A \cup B) = \min\{D(x, A), D(x, B)\}.$
- 4. $D(x, A) \leq D(x, A^{\epsilon}) + \epsilon$ for any $\epsilon \in [0, \infty]$, where $A^{\epsilon} = \{x : D(x, A) \leq \epsilon\}$ is the " ϵ -ball" with center x.

Every metric space (X, d) (moreover, any extended quasi-semi-metric space) is an approach space with D(x, A) being the usual point-set distance $\min_{y \in A} d(x, y)$.

Given a locally compact separable metric space (X,d) and the family \mathcal{F} of its non-empty closed subsets, the **Baddeley–Molchanov distance** function gives a tool for another generalization. It is a function $D: X \times \mathcal{F} \to \mathbb{R}$ which is lower semi-continuous with respect to its first argument, measurable with respect to the second, and satisfies the following two conditions: $F = \{x \in X : D(x, F) \leq 0\}$ for $F \in \mathcal{F}$, and $D(x, F_1) \geq D(x, F_2)$ for $x \in X$, whenever $F_1, F_2 \in \mathcal{F}$ and $F_1 \subset F_2$.

The additional conditions $D(x, \{y\}) = D(y, \{x\})$, and $D(x, F) \leq D(x, \{y\}) + D(y, F)$ for all $x, y \in X$ and every $F \in \mathcal{F}$, provide analogs of symmetry and the triangle inequality. The case D(x, F) = d(x, F) corresponds to the usual point-set distance for the metric space (X, d); the case D(x, F) = d(x, F) for $x \in X \setminus F$ and $D(x, F) = -d(x, X \setminus F)$ for $x \in X$ corresponds to the **signed distance function** in Chap. 1.

• Metric bornology

Given a topological space X, a bornology of X is any family \mathcal{A} of proper subsets A of X such that the following conditions hold:

- 1. $\bigcup_{A \in \mathcal{A}} A = X$.
- 2. \mathcal{A} is an *ideal*, i.e., contains all subsets and finite unions of its members. The family \mathcal{A} is a **metric bornology** [Beer99] if, moreover:
- 3. A contains a countable base.
- 4. For any $A \in \mathcal{A}$ there exists $A' \in \mathcal{A}$ such that the closure of A coincides with the interior of A'.

The metric bornology is called trivial if \mathcal{A} is the set P(X) of all subsets of X; such a metric bornology corresponds to the family of bounded sets of some bounded metric. For any non-compact **metrizable** topological space X, there exists an unbounded metric compatible with this topology. A nontrivial metric bornology on such a space X corresponds to the family of bounded subsets with respect to some such unbounded metric. A noncompact metrizable topological space X admits uncountably many nontrivial metric bornologies.

3.4 Beyond numbers

• Probabilistic metric space

A notion of **probabilistic metric space** is a generalization of the notion of metric space (see, for example, [ScSk83]) in two ways: distances become

probability distributions, and the sum in the triangle inequality becomes a triangle operation.

Formally, let A be the set of all probability distribution functions, whose support lies in $[0,\infty]$. For any $a \in [0,\infty]$ define step functions $\epsilon_a \in A$ by $\epsilon_a(x) = 1$ if x > a or $x = \infty$, and $\epsilon_a(x) = 0$, otherwise. The functions in A are ordered by defining $F \leq G$ to mean $F(x) \leq G(x)$ for all $x \geq 0$; the minimal element is ϵ_0 . A commutative and associative operation τ on A is called a **triangle function** if $\tau(F, \epsilon_0) = F$ for any $F \in A$ and $\tau(E, F) \leq \tau(G, H)$ whenever $E \leq G$, $F \leq H$. The semi-group (A, τ) generalizes the group $(\mathbb{R}, +)$.

A **probabilistic metric space** is a triple (X, D, τ) , where X is a set, D is a function $X \times X \to A$, and τ is a triangle function, such that for any $p, q, r \in X$:

- 1. $D(p,q) = \epsilon_0$ if and only if p = q.
- 2. D(p,q) = D(q,p).
- 3. $D(p,r) \ge \tau(D(p,q), D(q,r))$.

For any metric space (X, d) and any triangle function τ , such that $\tau(\epsilon_a, \epsilon_b) \geq \epsilon_{a+b}$ for all $a, b \geq 0$, the triple $(X, D = \epsilon_{d(x,y)}, \tau)$ is a probabilistic metric space.

For any $x \ge 0$, the value D(p,q) at x can be interpreted as "the probability that the distance between p and q is less than x;" this was approach of Menger, who proposed in 1942 the original version, statistical metric space, of this notion.

A probabilistic metric space is called a Wald space if the triangle function is a convolution, i.e., of the form $\tau_x(E,F) = \int_{\mathbb{R}} E(x-t) dF(t)$.

A probabilistic metric space is called a **generalized Menger space** if the triangle function has form $\tau_x(E,F) = \sup_{u+v=x} T(E(u),F(v))$ for a *t-norm* T, i.e., such a commutative and associative operation on [0,1] that T(a,1)=a, T(0,0)=0 and $T(c,d)\geq T(a,b)$ whenever $c\geq a, d\geq b$.

• Fuzzy metric spaces

A fuzzy subset of a set S is a mapping $\mu: S \to [0,1]$, where $\mu(x)$ represents the "degree of membership" of $x \in S$.

A continuous t-norm is a binary commutative and associative continuous operation T on [0,1], such that T(a,1)=a and $T(c,d)\geq T(a,b)$ whenever $c\geq a, d\geq b$.

A KM fuzzy metric space (Kramosil and Michalek 1975) is a pair $(X, (\mu, T))$, where X is a non-empty set and a fuzzy metric (μ, T) is a pair comprising a continuous t-norm T and a fuzzy set $\mu: X^2 \times \mathbb{R}_{\geq 0} \to [0, 1]$, such that for $x, y, z \in X$ and $s, t \geq 0$:

- 1. $\mu(x, y, 0) = 0$.
- 2. $\mu(x, y, t) = 1$ if and only if x = y, t > 0.
- 3. $\mu(x, y, t) = \mu(y, x, t)$.
- 4. $T(\mu(x, y, t), \mu(y, z, s)) \le \mu(x, z, t + s)$.
- 5. The function $\mu(x,y,\cdot):\mathbb{R}_{\geq 0}\to[0,1]$ is left continuous.

A KM fuzzy metric space is called also a fuzzy Menger space since by defining $D_t(p,q) = \mu(p,q,t)$ one gets a generalized Menger space. The following modification of the above notion, using a stronger form of metric fuzziness, can be seen as a generalized Menger space with $D_t(p,q)$ positive and continuous on $\mathbb{R}_{>0}$ for all p,q.

A **GV** fuzzy metric space (George and Veeramani 1994) is a pair $(X, (\mu, T))$, where X is a non-empty set, and a fuzzy metric (μ, T) is a pair comprising a continuous t-norm T and a fuzzy set $\mu: X^2 \times \mathbb{R}_{>0} \to [0, 1]$, such that for $x, y, z \in X$ and s, t > 0:

- 1. $\mu(x, y, t) > 0$.
- 2. $\mu(x, y, t) = 1$ if and only if x = y.
- 3. $\mu(x, y, t) = \mu(y, x, t)$.
- 4. $T(\mu(x, y, t), \mu(y, z, s)) \le \mu(x, z, t + s)$.
- 5. The function $\mu(x,y,\cdot):\mathbb{R}_{>0}\to[0,1]$ is continuous.

An example of a GV fuzzy metric space can be obtained from any metric space (X,d) by defining T(a,b)=b-ab and $\mu(x,y,t)=\frac{t}{t+d(x,y)}$. Conversely, any GV fuzzy metric space (and also any KM fuzzy metric space) generates a metrizable topology.

A fuzzy number is a fuzzy set $\mu: \mathbb{R} \to [0,1]$ which is $normal\ (\{x \in \mathbb{R} : \mu(x) = 1\} \neq \emptyset)$, $convex\ (\mu(tx + (1-t)y) \geq \min\{\mu(x), \mu(y)\}$ for every $x, y \in \mathbb{R}$ and $t \in [0,1]$) and $upper\ semicontinuous$ (at each point x_0 , the values $\mu(x)$ for x near x_0 are either close to $\mu(x_0)$ or less than $\mu(x_0)$). Denote the set of all fuzzy numbers which are non-negative, i.e., $\mu(x) = 0$ for all x < 0, by G. The additive and multiplicative identities of fuzzy numbers are denoted by 0 and 0, respectively. The $level\ set\ [\mu]_t = \{x: \mu(x) \geq t\}$ of a fuzzy number μ is a closed interval.

Given a non-empty set X and a mapping $d: X^2 \to G$, let the mappings $L, R: [0,1]^2 \to [0,1]$ be symmetric and non-decreasing in both arguments and satisfy L(0,0) = 0, R(1,1) = 1. For all $x, y \in X$ and $t \in (0,1]$, let $[d(x,y)]_t = [\lambda_t(x,y), \rho_t(x,y)]$.

A KS fuzzy metric space (Kaleva and Seikkala 1984) is a quadruple (X, d, L, R) with fuzzy metric d, if for all $x, y, z \in X$:

- 1. $d(x,y) = \tilde{0}$ if and only if x = y.
- 2. d(x,y) = d(y,x).
- 3. $d(x,y)(s+t) \ge L(d(x,z)(s),d(z,y)(t))$ whenever $s \le \lambda_1(x,z), t \le \lambda_1(z,y)$, and $s+t \le \lambda_1(x,y)$.
- 4. $d(x,y)(s+t) \leq R(d(x,z)(s),d(z,y)(t))$ whenever $s \geq \lambda_1(x,z), t \geq \lambda_1(z,y)$, and $s+t \geq \lambda_1(x,y)$.

The following functions are some frequently used choices for L and R:

$$\max\{a+b-1,0\}, ab, \min\{a,b\}, \max\{a,b\}, a+b-ab, \min\{a+b,1\}.$$

3.4 Beyond numbers 75

Several other notions of fuzzy metric space were proposed, including those by Erceg (1979), Deng (1982), and Voxman (1998), Xu and Li (2001), Tran and Duckstein (2002), C. Chakraborty and D. Chakraborty (2006). Cf. also metrics between fuzzy sets, fuzzy Hamming distance, gray-scale image distances and fuzzy polynucleotide metric in Chaps. 1, 11, 21 and 23, respectively.

• Interval-valued metric space

Let $I(\mathbb{R}_{>0})$ denote the set of closed intervals of $\mathbb{R}_{>0}$.

An **interval-valued metric space** (Coppola and Pacelli 2006) is a pair $((X, \leq), \Delta)$, where (X, \leq) is a partially ordered set and Δ is an interval-valued mapping $\Delta: X \times X \to I(\mathbb{R}_{>0})$, such that for every $x, y, z \in X$:

- 1. $\Delta(x, x) \star [0, 1] = \Delta(x, x)$.
- 2. $\Delta(x,y) = \Delta(y,x)$.
- 3. $\Delta(x,y) \Delta(z,z) \leq \Delta(x,z) + \Delta(z,y)$.
- 4. $\Delta(x,y) \Delta(x,y) \leq \Delta(x,x) + \Delta(y,y)$.
- 5. $x \le x'$ and $y \le y'$ imply $\Delta(x, y) \subseteq \Delta(x', y')$.
- 6. $\Delta(x,y) = 0$ if and only if x = y and x,y are atoms (minimal elements of (X, \leq)).

Here the following interval arithmetic rules hold: $[u, v] \leq [u', v']$ if and only if $u \leq u'$,

$$[u, v] \star [u', v'] = [\min\{uu', uv', vu', vv'\}, \max\{uu', uv', vu', vv'\}],$$

 $[u, v] + [u', v'] = [u + u', v + v'] \text{ and } [u, v] - [u', v'] = [u - u', v - v'].$
Cf. metric between intervals in Chap. 10.

The usual metric spaces coincide with above spaces in which all $x \in X$ are atoms.

• Generalized metric

Let X be a set. Let $(G, +, \leq)$ be an ordered semi-group (not necessarily commutative) having a least element 0. A function $d: X \times X \to G$ is called a **generalized metric** if the following conditions hold:

- 1. d(x, y) = 0 if and only if x = y.
- 2. $d(x,y) \le d(x,z) + d(z,y)$ for all $x, y \in X$.
- 3. $d(x,y) = \overline{d(y,x)}$, where $\overline{\alpha}$ is a fixed order-preserving involution of G.

The pair (X, d) is called a **generalized metric space**.

If the condition 2 and "only if" in 1 above are dropped, we obtain a generalized distance d, and a generalized distance space (X, d).

• Cone metric

Let C be a proper cone in a real Banach space W, i.e., C is closed, $C \neq \emptyset$, the interior of C is not equal to $\{0\}$ and:

- 1. If $x, y \in C$ and $a, b \in \mathbb{R}_{\geq 0}$, then $ax + by \in C$.
- 2. If $x \in C$ and $-x \in C$, then x = 0.

Define a partial ordering (W, \leq) on W by letting $x \leq y$ if $y - x \in C$. The following variation of **generalized metric** was defined in Huang and Zhang (2007). Given a set X, a **cone metric** is a mapping $d: X \times X \to (W, \leq)$ such that:

- 1. $d(x,y) \ge 0$ with equality if and only if x = y.
- 2. d(x,y) = d(y,x) for all $x, y \in X$.
- 3. $d(x,y) \leq d(x,z) + d(z,y)$ for all $x,y \in X$.

• Distance on building

A Coxeter group is a group $(W,\cdot,1)$ generated by the elements $\{w_1,\ldots,w_n:(w_iw_j)^{m_{ij}}=1,1\leq i,j\leq n\}$. Here $M=((m_{ij}))$ is a Coxeter matrix, i.e., an arbitrary symmetric $n\times n$ matrix with $m_{ii}=1$, and the other values are positive integers or ∞ .

The length l(x) of $x \in W$ is the smallest number of generators w_1, \ldots, w_n needed to represent x.

Let X be a set, and let $(W, \cdot, 1)$ be a Coxeter group. The pair (X, d) is called a *building* over $(W, \cdot, 1)$ if the function $d: X \times X \to W$, called a **distance on building**, has the following properties:

- 1. d(x,y) = 1 if and only if x = y.
- 2. $d(y,x) = (d(x,y))^{-1}$.
- 3. The relation \sim_i , defined by $x \sim_i y$ if d(x, y) = 1 or w_i , is an equivalence relation.
- 4. Given $x \in X$ and an equivalence class C of \sim_i , there exists a unique $y \in C$ such that d(x,y) is shortest (i.e., of smallest length), and $d(x,y') = d(x,y)w_i$ for any $y' \in C$, $y' \neq y$.

The gallery distance on building d' is a usual metric on X, defined by l(d(x,y)). The distance d' is the **path metric** in the graph with the vertex-set X and xy being an edge if $d(x,y) = w_i$ for some $1 \le i \le n$. The gallery distance on building is a special case of a **gallery metric** (of chamber system X).

• Boolean metric space

A Boolean algebra (or Boolean lattice) is a distributive lattice (B, \vee, \wedge) admitting a least element 0 and greatest element 1 such that every $x \in B$ has a complement \overline{x} with $x \vee \overline{x} = 1$ and $x \wedge \overline{x} = 0$.

Let X be a set, and let (B, \vee, \wedge) be a Boolean algebra. The pair (X, d) is called a **Boolean metric space** over B if the function $d: X \times X \to B$ has the following properties:

- 1. d(x,y) = 0 if and only if x = y.
- 2. $d(x,y) \le d(x,z) \lor d(z,y)$ for all $x,y,z \in X$.

• Space over algebra

A space over algebra is a metric space with a differential-geometric structure, whose points can be provided with coordinates from some *algebra* (usually, an associative algebra with identity).

A module over an algebra is a generalization of a vector space over a field, and its definition can be obtained from the definition of a vector space by replacing the field by an associative algebra with identity. An affine space over an algebra is a similar generalization of an affine space over a field. In affine spaces over algebras one can specify a Hermitian metric, while in the case of commutative algebras even a quadratic metric can be given. To do this one defines in a unital module a scalar product $\langle x, y \rangle$, in the first case with the property $\langle x, y \rangle = J(\langle y, x \rangle)$, where J is an involution of the algebra, and in the second case with the property $\langle y, x \rangle = \langle x, y \rangle$.

The n-dimensional projective space over an algebra is defined as the variety of one-dimensional submodules of an (n+1)-dimensional unital module over this algebra. The introduction of a scalar product $\langle x,y\rangle$ in a unital module makes it possible to define a Hermitian metric in a projective space constructed by means of this module or, in the case of a commutative algebra, quadratic elliptic and hyperbolic metrics. The metric invariant of the points of these spaces is the cross-ratio $W = \langle x,x\rangle^{-1}\langle x,y\rangle\langle y,y\rangle^{-1}\langle y,x\rangle$. If W is a real number, then the invariant w, for which $W = \cos^2 w$, is called the **distance** between x and y in the space over algebra.

• Partially ordered distance

Let X be a set. Let (G, \leq) be a partially ordered set with a least element g_0 . A **partially ordered distance** is a function $d: X \times X \to G$ such that, for any $x, y \in X$, $d(x, y) = g_0$ if and only if x = y.

A **generalized ultrametric** (Priess-Crampe and Ribenboim 1993) is a symmetric (i.e., d(x,y) = d(y,x)) partially ordered distance, such that $d(z,x) \leq g$ and $d(z,y) \leq g$ imply $d(x,y) \leq g$ for any $x,y,z \in X$ and $g \in G$.

Suppose from now that $G' = G \setminus \{g_0\}$ is non-empty and, for any $g_1, g_2 \in G'$, there exists $g_3 \in G'$ such that $g_3 \leq g_1$ and $g_3 \leq g_2$. Consider the following possible properties:

- 1. For any $g_1 \in G'$, there exists $g_2 \in G'$ such that, for any $x, y \in X$, from $d(x, y) \leq g_2$ it follows that $d(y, x) \leq g_1$.
- 2. For any $g_1 \in G'$, there exist $g_2, g_3 \in G'$ such that, for any $x, y, z \in X$, from $d(x, y) \leq g_2$ and $d(y, z) \leq g_3$ it follows that $d(x, z) \leq g_1$.
- 3. For any $g_1 \in G'$, there exists $g_2 \in G'$ such that, for any $x, y, z \in X$, from $d(x, y) \leq g_2$ and $d(y, z) \leq g_2$ it follows that $d(y, x) \leq g_1$.
- 4. G' has no first element.
- 5. d(x,y) = d(y,x) for any $x,y \in X$.
- 6. For any $g_1 \in G'$, there exists $g_2 \in G'$ such that, for any $x, y, z \in X$, from $d(x, y) <^* g_2$ and $d(y, z) <^* g_2$ it follows that $d(x, z) <^* g_1$; here $p <^* q$ means that either p < q, or p is not comparable to q.
- 7. The order relation < is a total ordering of G.

In terms of above properties, d is called: the **Appert partially ordered** distance if 1 and 2 hold; the **Golmez partially ordered** distance of

first type if 4, 5, and 6 hold; the Golmez partially ordered distance of second type if 3, 4, and 5 hold; the Kurepa–Fréchet distance if 3, 4, 5, and 7 hold.

In fact, the case $G = \mathbb{R}_{\geq 0}$ of the Kurepa–Fréchet distance corresponds to the **Fréchet** V-space, i.e., a pair (X,d), where X is a set, and d(x,y) is a non-negative symmetric function $d: X \times X \to \mathbb{R}$ (voisinage of points x and y) such that d(x,y) = 0 if and only if x = y, and there exists a non-negative function $f: \mathbb{R} \to \mathbb{R}$ with $\lim_{t\to 0} f(t) = 0$ with the following property: for all $x, y, z \in X$ and all positive r, the inequality $\max\{d(x,y), d(y,z)\} \leq r$ implies $d(x,z) \leq f(r)$.

• Distance from measurement

This notion is an analog of distance on domains in Computer Science; it was developed in [Mart00].

A dcpo is a partially ordered set (D, \preceq) , in which every directed subset $S \subset D$ (i.e., $S \neq \emptyset$ and any pair $x, y \in S$ is bounded: there is $z \in S$ with $x, y \preceq z$) has a supremum $\sqcup S$, i.e., the least of such upper bounds z. For $x, y \in D$, y is an approximation of x if, for all directed subsets $S \subset D$, $x \preceq \sqcup S$ implies $y \preceq s$ for some $s \in S$.

A dcpo (D, \preceq) is *continuous* if for all $x \in D$ the set of all approximations of x is directed and x is its supremum. A *domain* is a continuous dcpo (D, \preceq) such that for all $x, y \in D$ there is $z \in D$ with $z \preceq x, y$. A *Scott domain* is a continuous dcpo (D, \preceq) with least element, in which any bounded pair $x, y \in D$ has a supremum.

A measurement is a mapping $\mu: D \to \mathbb{R}_{\geq 0}$ between dcpo (D, \preceq) and dcpo $(\mathbb{R}_{\geq 0}, \preceq)$, where $\mathbb{R}_{\geq 0}$ is ordered as $x \preceq y$ if $y \leq x$, such that:

- 1. $x \leq y$ implies $\mu(x) \leq \mu(y)$.
- 2. $\mu(\sqcup S) = \sqcup(\{\mu(s): s \in S\})$ for every directed subset $S \subset D$.
- 3. For all $x \in D$ with $\mu(x) = 0$ and all sequences $(x_n), n \to \infty$, of approximations of x with $\lim_{n\to\infty} \mu(x_n) = \mu(x)$, one has $\sqcup (\cup_{n=1}^{\infty} \{x_n\}) = x$.

Given a measurement μ , the **distance from measurement** is a mapping $d: D \times D \to \mathbb{R}_{\geq 0}$ given by

$$d(x,y) = \inf\{\mu(z) : z \text{ approximates } x,y\} = \inf\{\mu(z) : z \leq x,y\}.$$

One has $d(x,x) \leq \mu(x)$. The function d(x,y) is a metric on the set $\{x \in D : \mu(x) = 0\}$ if μ satisfies the following **measurement triangle inequality**: for all bounded pairs $x,y \in D$, there is an element $z \leq x,y$ such that $\mu(z) \leq \mu(x) + \mu(y)$.

Waszkiewicz (2001) found topological connections between topologies coming from a distance from measurement and from a **partial metric** defined in Chap. 1.

Chapter 4 Metric Transforms

There are many ways to obtain new distances (metrics) from given distances (metrics). Metric transforms give new distances as a functions of given metrics (or given distances) on the same set X. A metric so obtained is called a **transform metric**. We give some important examples of transform metrics in Sect. 4.1.

Given a metric on a set X, one can construct a new metric on an extension of X; similarly, given a collection of metrics on sets X_1, \ldots, X_n , one can obtain a new metric on an extension of X_1, \ldots, X_n . Examples of such operations are given in Sect. 4.2.

Given a metric on X, there are many distances on other structures connected with X, for example, on the set of all subsets of X. The main distances of this kind are considered in Sect. 4.3.

4.1 Metrics on the same set

• Metric transform

A **metric transform** is a distance on a set X, obtained as a function of given metrics (or given distances) on X.

In particular, given a continuous monotone increasing function f(x) of $x \geq 0$ with f(0) = 0, called the *scale*, and a distance space (X, d), one obtains another distance space (X, d_f) , called a **scale metric transform** of X, defining $d_f(x, y) = f(d(x, y))$. For every finite distance space (X, d), there exists a scale f, such that (X, d_f) is a metric subspace of a Euclidean space \mathbb{E}^n .

If (X, d) is a metric space and f is a continuous differentiable strictly increasing scale with f(0) = 0 and non-increasing f', then (X, d_f) is a metric space (cf. functional transform metric).

The metric d is an **ultrametric** if and only if f(d) is a metric for every non-decreasing function $f: \mathbb{R}_{\geq 0} \to \mathbb{R}_{\geq 0}$.

• Transform metric

A transform metric is a metric on a set X which is a metric transform, i.e., is obtained as a function of a given metric (or given metrics) on X. In

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particular, transform metrics can be obtained from a given metric d (or given metrics d_1 and d_2) on X by any of the following operations (here t > 0):

- 1. td(x,y) (t-scaled metric, or dilated metric, similar metric)
- 2. $\min\{t, d(x, y)\}\ (t$ -truncated metric)
- 3. $\max\{t, d(x, y)\}\$ for $x \neq y\ (t$ -uniformly discrete metric)
- 4. d(x,y) + t for $x \neq y$ (t-translated metric)
- $5. \quad \frac{d(x,y)}{1+d(x,y)}$
- 6. $d^p(x,y) = \frac{2d(x,y)}{d(x,p)+d(y,p)+d(x,y)}$, where p is an fixed element of X (biotope transform metric, or Steinhaus transform metric)
- 7. $\max\{d_1(x,y), d_2(x,y)\}$
- 8. $\alpha d_1(x,y) + \beta d_2(x,y)$, where $\alpha,\beta > 0$ (cf. **metric cone** in Chap. 1)

• Generalized biotope transform metric

For a given metric d on a set X and a closed set $M \subset X$, the **generalized** biotope transform metric d^M on X is defined by

$$d^M(x,y) = \frac{2d(x,y)}{d(x,y) + \inf_{z \in M} (d(x,z) + d(y,z))}.$$

In fact, $d^M(x,y)$ and its 1-truncation $\min\{1,d^M(x,y)\}$ are both metrics. The **biotope transform metric** is $d^M(x,y)$ with M consisting only of a point, say, p; the **Steinhaus distance** from Chap. 1 is the case $d(x,y)=\mu(x\triangle y)$ with $p\neq\emptyset$ and the **biotope distance** from Chap. 23 is its subcase $d(x,y)=\mu(x\triangle y)=|x\triangle y|$.

• Metric-preserving function

A function $f: \mathbb{R}_{\geq 0} \to \mathbb{R}_{\geq 0}$ with $f^{-1}(0) = \{0\}$ is a **metric-preserving** function if, for each metric space (X, d), the **metric transform**

$$d_f(x,y) = f(d(x,y))$$

is a metric on X; cf. [Cora99]. In this case d_f is called a **functional transform metric**. For example, αd ($\alpha > 0$), $d^{\alpha}(0 < \alpha \le 1)$, $\ln(1+d)$, arcsinh d, arcs

The superposition, sum and maximum of two metric-preserving functions are metric-preserving. If f is subadditive, i.e., $f(x+y) \leq f(x) + f(y)$ for all $x,y \geq 0$, and non-decreasing, then it is metric-preserving. But, for example, the function $f(x) = \frac{x+2}{x+1}$, for x>0, and f(0)=0, is decreasing and metric-preserving. If f is concave, i.e., $f(\frac{x+y}{2}) \geq \frac{f(x)+f(y)}{2}$ for all $x,y \geq 0$, then it is metric-preserving. In particular, a twice differentiable function $f: \mathbb{R}_{\geq 0} \to \mathbb{R}_{\geq 0}$ such that f(0)=0, f'(x)>0 for all $x\geq 0$, and $f''(x)\leq 0$ for all $x\geq 0$, is metric-preserving.

If f is metric-preserving, then it is subadditive.

The function f is **strongly metric-preserving function** if d and f(d(x,y)) are **equivalent metrics** on X, for each metric space (X,d). A metric-preserving function is strongly metric-preserving if and only if it is continuous at 0.

• Power transform metric

Let $0 < \alpha \le 1$. Given a metric space (X, d), the **power transform metric** (or *snowflake transform metric*) is a **functional transform metric** on X, defined by

$$(d(x,y))^{\alpha}$$
.

The distance $d(x,y) = (\sum_{1}^{n} |x_i - y_i|^p)^{\frac{1}{p}}$ with $0 is not a metric on <math>\mathbb{R}^n$, but its power transform $(d(x,y)^{\alpha})$ is a metric.

For a given metric d on X and any $\alpha > 1$, the function d^{α} is, in general, only a distance on X. It is a metric, for any positive α , if and only if d is an **ultrametric**.

A metric d is a **doubling metric** if and only if (Assouad 1983) the power transform metric d^{α} admits a **bi-Lipschitz embedding** in some Euclidean space for every $0 < \alpha < 1$ (cf. Chap. 1 for definitions).

• Schoenberg transform metric

Let $\lambda > 0$. Given a metric space (X, d), the **Schoenberg transform** metric is a functional transform metric on X, defined by

$$1 - e^{-\lambda d(x,y)}.$$

The Schoenberg transform metrics are exactly *P*-metrics (cf. Chap. 1).

• Pullback metric

Given two metric spaces (X, d_X) , (Y, d_Y) and an injective mapping $g: X \to Y$, the **pullback metric** (of (Y, d_Y) by g) on X is defined by

$$d_Y(g(x), g(y)).$$

If (X, d_X) coincides with (Y, d_Y) , then the pullback metric is called a *g*-transform metric.

• Internal metric

Given a metric space (X,d) in which every pair of points x,y is joined by a rectifiable curve, the **internal metric** (or **inner metric**, induced **intrinsic metric**, **interior metric**) D is a **transform metric** on X, obtained from d as the infimum of the lengths of all rectifiable curves connecting two given points x and $y \in X$.

The metric d on X is called an **intrinsic metric** (or **length metric**, cf. Chap. 6) if it coincides with its internal metric.

• Farris transform metric

Given a metric space (X, d) and a point $z \in X$, the **Farris transform** is a metric transform D_z on $X \setminus \{z\}$ defined by $D_z(x, x) = 0$ and, for different $x, y \in X \setminus \{z\}$, by

$$D_z(x,y) = C - (x.y)_z,$$

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where C is a positive constant, and $(x.y)_z = \frac{1}{2}(d(x,z) + d(y,z) - d(x,y))$ is the **Gromov product** (cf. Chap. 1). It is a metric if $C \ge \max_{x \in X \setminus \{z\}} d(x,z)$; in fact, there exists a number $C_0 \in (\max_{x,y \in X \setminus \{z\}, x \ne y} (x.y)_z, \max_{x \in X \setminus \{z\}} d(x,z)]$ such that it is a metric if and only if $C \ge C_0$. The Farris transform is an **ultrametric** if and only if d satisfies the **four-point inequality**. In Phylogenetics, where it was applied first, the term Farris transform is used for the function d(x,y) - d(x,z) - d(y,z).

• Involution transform metric

Given a metric space (X, d) and a point $z \in X$, the **involution transform** metric is a metric transform d_z on $X \setminus \{z\}$ defined by

$$d_z(x,y) = \frac{d(x,y)}{d(x,z)d(y,z)}.$$

It is a metric for any $z \in X$, if and only if d is a **Ptolemaic metric** [FoSC06].

4.2 Metrics on set extensions

• Extension distances

If d is a metric on $V_n = \{1, ..., n\}$, and $\alpha \in \mathbb{R}, \alpha > 0$, then the following extension distances (see, for example, [DeLa97]) are used.

The gate extension distance $gat = gat_{\alpha}^{d}$ is a metric on $V_{n+1} = \{1, \ldots, n+1\}$, defined by the following conditions:

- 1. $gat(1, n + 1) = \alpha$.
- 2. $gat(i, n + 1) = \alpha + d(1, i)$ if $2 \le i \le n$.
- 3. gat(i, j) = d(i, j) if $1 \le i < j \le n$.

The distance gat_0^d is called the **gate** 0-extension or, simply, 0-extension of d.

If $\alpha \ge \max_{2 \le i \le n} d(1, i)$, then the **antipodal extension distance** $ant = ant_{\alpha}^{d}$ is a distance on V_{n+1} , defined by the following conditions:

- 1. $ant(1, n + 1) = \alpha$.
- 2. $ant(i, n + 1) = \alpha d(1, i)$ if $2 \le i \le n$.
- 3. ant(i, j) = d(i, j) if $1 \le i < j \le n$.

If $\alpha \ge \max_{1 \le i,j \le n} d(i,j)$, then the **full antipodal extension distance** $Ant = Ant_{\alpha}^d$ is a distance on $V_{2n} = \{1, \dots, 2n\}$, defined by the following conditions:

- 1. $Ant(i, n+i) = \alpha \text{ if } 1 \leq i \leq n.$
- 2. $Ant(i, n+j) = \alpha d(i, j)$ if $1 \le i \ne j \le n$.
- 3. $Ant(i, j) = d(i, j) \text{ if } 1 \le i \ne j \le n.$
- 4. Ant(n+i, n+j) = d(i, j) if $1 \le i \ne j \le n$.

It is obtained by applying the antipodal extension operation iteratively ntimes, starting from d.

The spherical extension distance $sph = sph_{\alpha}^{d}$ is a metric on V_{n+1} , defined by the following conditions:

- 1. $sph(i, n + 1) = \alpha \text{ if } 1 \le i \le n.$ 2. sph(i, j) = d(i, j) if $1 \le i < j \le n$.
- 1-sum distance

Let d_1 be a distance on a set X_1 , let d_2 be a distance on a set X_2 , and suppose that $X_1 \cap X_2 = \{x_0\}$. The 1-sum distance of d_1 and d_2 is the distance d on $X_1 \cup X_2$, defined by the following conditions:

$$d(x,y) = \begin{cases} d_1(x,y), & \text{if} & x,y \in X_1, \\ d_2(x,y), & \text{if} & x,y \in X_2, \\ d(x,x_0) + d(x_0,y), & \text{if} & x \in X_1, y \in X_2. \end{cases}$$

In Graph Theory, the 1-sum distance is a **path metric**, corresponding to the clique 1-sum operation for graphs.

• Disjoint union metric

Given a family $(X_t, d_t), t \in T$, of metric spaces, the **disjoint union met**ric is an extended metric on the set $\bigcup_t X_t \times \{t\}$, defined by

$$d((x, t_1), (y, t_2)) = d_t(x, y)$$

for $t_1 = t_2$, and $d((x, t_1), (y, t_2)) = \infty$ otherwise.

• Product metric

Given finite or countable number n of metric spaces (X_1, d_1) , (X_2, d_2) , \ldots , (X_n, d_n) , the **product metric** is a metric on the Cartesian product $X_1 \times X_2 \times \cdots \times X_n = \{x = (x_1, x_2, \dots, x_n) : x_1 \in X_1, \dots, x_n \in X_n\},\$ defined as a function of d_1, \ldots, d_n . The simplest finite product metrics are defined by:

- 1. $\sum_{i=1}^{n} d_i(x_i, y_i)$ 2. $(\sum_{i=1}^{n} d_i^p(x_i, y_i))^{\frac{1}{p}}, 1
 3. <math display="block">\max_{1 \le i \le n} d_i(x_i, y_i)$ 4. $\sum_{i=1}^{n} \frac{1}{2^i} \frac{d_i(x_i, y_i)}{1 + d_i(x_i, y_i)}$

The last metric is **bounded** and can be extended to the product of countably many metric spaces.

If $X_1 = \cdots = X_n = \mathbb{R}$, and $d_1 = \cdots = d_n = d$, where d(x, y) = |x - y| is the **natural metric** on \mathbb{R} , all product metrics above induce the Euclidean topology on the n-dimensional space \mathbb{R}^n . They do not coincide with the Euclidean metric on \mathbb{R}^n , but they are equivalent to it. In particular, the set \mathbb{R}^n with the Euclidean metric can be considered as the Cartesian product $\mathbb{R} \times \cdots \times \mathbb{R}$ of n copies of the real line (\mathbb{R}, d) with the product metric, defined by $\sqrt{\sum_{i=1}^n d^2(x_i, y_i)}$.

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• Box metric

Let (X, d) be a metric space and I the unit interval of \mathbb{R} . The **box metric** is the **product metric** d' on the Cartesian product $X \times I$ defined by

$$d'((x_1, t_1), (x_2, t_2)) = \max(d(x_1, x_2), |t_1 - t_2|).$$

Cf. unrelated **bounded box metric** in Chap. 18.

• Fréchet product metric

Let (X,d) be a metric space with a **bounded** metric d. Let $X^{\infty} = X \times \cdots \times X \dots = \{x = (x_1, \dots, x_n, \dots) : x_1 \in X_1, \dots, x_n \in X_n, \dots \}$ be the countable Cartesian product space of X.

The Fréchet product metric is a product metric on X^{∞} , defined by

$$\sum_{n=1}^{\infty} A_n d(x_n, y_n),$$

where $\sum_{n=1}^{\infty} A_n$ is any convergent series of positive terms. Usually, $A_n = \frac{1}{2n}$ is used.

A metric (sometimes called the *Fréchet metric*) on the set of all sequences $\{x_n\}_n$ of real (complex) numbers, defined by

$$\sum_{n=1}^{\infty} A_n \frac{|x_n - y_n|}{1 + |x_n - y_n|},$$

where $\sum_{n=1}^{\infty} A_n$ is any convergent series of positive terms, is a Fréchet product metric of countably many copies of \mathbb{R} (\mathbb{C}). Usually, $A_n = \frac{1}{n!}$ or $A_n = \frac{1}{2^n}$ are used.

• Hilbert cube metric

The Hilbert cube I^{\aleph_0} is the Cartesian product of countable many copies of the interval [0,1], equipped with the metric

$$\sum_{i=1}^{\infty} 2^{-i} |x_i - y_i|$$

(cf. **Fréchet infinite metric product**). It also can be identified up to homeomorphisms with the compact metric space formed by all sequences $\{x_n\}_n$ of real numbers such that $0 \le x_n \le \frac{1}{n}$, where the metric is defined as $\sqrt{\sum_{n=1}^{\infty} (x_n - y_n)^2}$.

• Hamming cube

Given integers $n \geq 1$ and $q \geq 2$, the Hamming space H(n,q) is the set of all n-tuples over an alphabet of size q (say, the Cartesian product of n copies of the set $\{0,1,\ldots,q-1\}$), equipped with the **Hamming metric** (cf. Chap. 1), i.e., the distance between two n-tuples is the number of coordinates where they differ. The **Hamming cube** is the Hamming space H(n,2).

The infinite Hamming cube $H(\infty, 2)$ is the set of all infinite strings over the alphabet $\{0, 1\}$ containing only finitely many 1s, equipped with the Hamming metric.

• Cameron-Tarzi cube

Given integers $n \geq 1$ and $q \geq 2$, the normalized Hamming space $H_n(q)$ is the set of all n-tuples over an alphabet of size q, equipped with the **Hamming metric** divided by n. Clearly, there are isometric embeddings

$$H_1(q) \to H_2(q) \to H_4(q) \to H_8(q) \to \dots$$

Let H(q) denote the **Cauchy completion** (cf. Chap. 1) of the union (denote it by $H_{\omega}(q)$) of all metric spaces $H_n(q)$ with $n \geq 1$. This metric space was introduced in [CaTa08]. Call H(2) the **Cameron–Tarzi cube**.

It is shown in [CaTa08] that $H_{\omega}(2)$ is the **word metric** space (cf. Chap. 10) of the *countable Nim group*, i.e., the elementary Abelian 2-group of all natural numbers under bitwise addition modulo 2 of the number expressions in base 2. The Cameron–Tarzi cube is also the word metric space of an Abelian group.

• Warped product metric

Let (X, d_X) and (Y, d_Y) be two complete **length spaces** (cf. Chap. 1), and let $f: X \to \mathbb{R}$ be a positive continuous function. Given a curve $\gamma: [a, b] \to X \times Y$, consider its projections $\gamma_1: [a, b] \to X$ and $\gamma_2: [a, b] \to Y$ to X and Y, and define the length of γ by the formula $\int_a^b \sqrt{|\gamma_1'|^2(t) + f^2(\gamma_1(t))|\gamma_2'|^2(t)}dt$.

The warped product metric is a metric on $X \times Y$, defined as the infimum of lengths of all rectifiable curves connecting two given points in $X \times Y$ (see [BuIv01]).

4.3 Metrics on other sets

Given a metric space (X,d), one can construct several distances between some subsets of X. The main such distances are: the **point-set distance** $d(x,A) = \inf_{y \in A} d(x,y)$ between a point $x \in X$ and a subset $A \subset X$, the **set-set distance** $\inf_{x \in A, y \in B} d(x,y)$ between two subsets A and B of X, and the **Hausdorff metric** between compact subsets of X, which are considered in Chap. 1. In this section we list some other distances of this kind.

• Line-line distance

The line-line distance (or vertical distance between lines) is the set-set distance in \mathbb{E}^3 between two skew lines, i.e., two straight lines

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that do not lie in a plane. It is the length of the segment of their common perpendicular whose end points lie on the lines. For l_1 and l_2 with equations l_1 : x = p + qt, $t \in \mathbb{R}$, and l_2 : x = r + st, $t \in \mathbb{R}$, the distance is given by

$$\frac{|\langle r - p, q \times s \rangle|}{||q \times s||_2},$$

where \times is the *cross product* on \mathbb{E}^3 , \langle , \rangle is the *inner product* on \mathbb{E}^3 , and $||.||_2$ is the Euclidean norm. For $x = (q_1, q_2, q_3)$, $s = (s_1, s_2, s_3)$, one has $q \times s = (q_2s_3 - q_3s_2, q_3s_1 - q_1s_3, q_1s_2 - q_2s_1)$.

• Point-line distance

The **point-line distance** is the **point-set distance** between a point and a line.

In \mathbb{E}^2 , the distance between a point $z = (z_1, z_2)$ and a line $l: ax_1 + bx_2 + c = 0$ is given by

$$\frac{|az_1+bz_2+c|}{\sqrt{a^2+b^2}}.$$

In \mathbb{E}^3 , the distance between a point z and a line l: x = p + qt, $t \in \mathbb{R}$, is given by

$$\frac{||q\times(p-z)||_2}{||q||_2},$$

where \times is the *cross product* on \mathbb{E}^3 , and $||.||_2$ is the Euclidean norm.

• Point-plane distance

The **point-plane distance** is the **point-set distance** in \mathbb{E}^3 between a point and a plane. The distance between a point $z = (z_1, z_2, z_3)$ and a plane α : $ax_1 + bx_2 + cx_3 + d = 0$ is given by

$$\frac{|az_1 + bz_2 + cz_3 + d|}{\sqrt{a^2 + b^2 + c^2}}.$$

• Prime number distance

The **prime number distance** is the **point-set distance** in $(\mathbb{N}, |n-m|)$ between a number $n \in \mathbb{N}$ and the set of prime numbers $P \subset \mathbb{N}$. It is the absolute difference between n and the nearest prime number.

• Distance up to nearest integer

The distance up to nearest integer is the point-set distance in $(\mathbb{R}, |x-y|)$ between a number $x \in \mathbb{R}$ and the set of integers $\mathbb{Z} \subset \mathbb{R}$, i.e., $\min_{n \in \mathbb{Z}} |x-n|$.

• Busemann metric of sets

Given a metric space (X, d), the **Busemann metric of sets** (see [Buse55]) is a metric on the set of all non-empty closed subsets of X, defined by

$$\sup_{x \in X} |d(x, A) - d(x, B)|e^{-d(p, x)},$$

where p is a fixed point of X, and $d(x, A) = \min_{y \in A} d(x, y)$ is the **point-set distance**.

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Instead of the weighting factor $e^{-d(p,x)}$, one can take any distance transform function which decreases fast enough (cf. L_p -Hausdorff distance in Chap. 1, and the list of variations of the Hausdorff metric in Chap. 21).

• Quotient semi-metric

Given an **extended metric space** (X,d) (i.e., a possibly infinite metric) and an equivalence relation \sim on X, the **quotient semi-metric** is a semi-metric on the set $\overline{X} = X/\sim$ of equivalence classes defined, for any $\overline{x}, \overline{y} \in \overline{X}$, by

$$\overline{d}(\overline{x}, \overline{y}) = \inf_{m \in \mathbb{N}} \sum_{i=1}^{m} d(x_i, y_i),$$

where the infimum is taken over all sequences $x_1, y_1, x_2, y_2, \ldots, x_m, y_m$ with $x_1 \in \overline{x}$, $y_m \in \overline{y}$, and $y_i \sim x_{i+1}$ for $i = 1, 2, \ldots, m-1$. One has $\overline{d}(\overline{x}, \overline{y}) \leq d(x, y)$ for all $x, y \in X$, and \overline{d} is the biggest semi-metric on \overline{X} with this property.

Chapter 5

Metrics on Normed Structures

In this chapter we consider a special class of metrics, defined on some *normed structures*, as the norm of the difference between two given elements. This structure can be a group (with a *group norm*), a vector space (with a *vector norm* or, simply, a *norm*), a vector lattice (with a *Riesz norm*), a field (with a *valuation*), etc.

• Group norm metric

A group norm metric is a metric on a group (G, +, 0), defined by

$$||x + (-y)|| = ||x - y||,$$

where ||.|| is a group norm on G, i.e., a function $||.||: G \to \mathbb{R}$ such that, for all $x, y \in G$, we have the following properties:

- 1. $||x|| \ge 0$, with ||x|| = 0 if and only if x = 0.
- 2. ||x|| = ||-x||.
- 3. $||x+y|| \le ||x|| + ||y||$ (triangle inequality).

Any group norm metric d is **right-invariant**, i.e., d(x,y) = d(x+z,y+z) for any $x,y,z \in G$. Conversely, any right-invariant (as well as any **left-invariant**, and, in particular, any **bi-invariant**) metric d on G is a group norm metric, since one can define a group norm on G by ||x|| = d(x,0).

• F-norm metric

A vector space (or linear space) over a field \mathbb{F} is a set V equipped with operations of vector addition $+: V \times V \to V$ and scalar multiplication $\cdot: \mathbb{F} \times V \to V$ such that (V, +, 0) forms an Abelian group (where $0 \in V$ is the zero vector), and, for all vectors $x, y \in V$ and any scalars $a, b \in \mathbb{F}$, we have the following properties: $1 \cdot x = x$ (where 1 is the multiplicative unit of \mathbb{F}), $(ab) \cdot x = a \cdot (b \cdot x), (a + b) \cdot x = a \cdot x + b \cdot x$, and $a \cdot (x + y) = a \cdot x + a \cdot y$.

A vector space over the field $\mathbb R$ of real numbers is called a *real vector space*. A vector space over the field $\mathbb C$ of complex numbers is called *complex vector space*.

A F-norm metric is a metric on a real (complex) vector space V, defined by

$$||x-y||_F$$

where $||.||_F$ is an F-norm on V, i.e., a function $||.||_F : V \to \mathbb{R}$ such that, for all $x, y \in V$ and for any scalar a with |a| = 1, we have the following properties:

- 1. $||x||_F \ge 0$, with $||x||_F = 0$ if and only if x = 0.
- 2. $||ax||_F = ||x||_F$.
- 3. $||x+y||_F \le ||x||_F + ||y||_F$ (triangle inequality).

An *F*-norm is called *p*-homogeneous if $||ax||_F = |a|^p||x||_F$ for any scalar *a*. Any *F*-norm metric *d* is a **translation invariant metric**, i.e., d(x,y) = d(x+z,y+z) for all $x,y,z \in V$. Conversely, if *d* is a translation invariant metric on *V*, then $||x||_F = d(x,0)$ is an *F*-norm on *V*.

• F^* -metric

An F^* -metric is an F-norm metric $||x-y||_F$ on a real (complex) vector space V such that the operations of scalar multiplication and vector addition are *continuous* with respect to $||.||_F$. Thus $||.||_F$ is a function $||.||_F:V\to\mathbb{R}$ such that, for all $x,y,x_n\in V$ and for all scalars a,a_n , we have the following properties:

- 1. $||x||_F \ge 0$, with $||x||_F = 0$ if and only if x = 0.
- 2. $||ax||_F = ||x||_F$ for all a with |a| = 1.
- 3. $||x+y||_F \le ||x||_F + ||y||_F$.
- 4. $||a_n x||_F \to 0 \text{ if } a_n \to 0.$
- 5. $||ax_n||_F \to 0 \text{ if } x_n \to 0.$
- 6. $||a_n x_n||_F \to 0 \text{ if } a_n \to 0, x_n \to 0.$

The metric space $(V, ||x - y||_F)$ with an F^* -metric is called a nF^* -space. Equivalently, an F^* -space is a metric space (V, d) with a **translation** invariant metric d such that the operation of scalar multiplication and vector addition are continuous with respect to this metric.

A complete F^* -space is called an F-space. A locally convex F-space is known as a **Fréchet space** (cf. Chap. 2) in Functional Analysis.

A modular space is an F^* -space $(V, ||.||_F)$ in which the F-norm $||.||_F$ is defined by

$$||x||_F = \inf\{\lambda > 0 : \rho\left(\frac{x}{\lambda}\right) < \lambda\},$$

and ρ is a metrizing modular on V, i.e., a function $\rho: V \to [0, \infty]$ such that, for all $x, y, x_n \in V$ and for all scalars a, a_n , we have the following properties:

- 1. $\rho(x) = 0$ if and only if x = 0.
- 2. $\rho(ax) = \rho(x)$ implies |a| = 1.
- 3. $\rho(ax+by) \leq \rho(x) + \rho(y)$ implies $a, b \geq 0, a+b=1$.
- 4. $\rho(a_n x) \to 0$ if $a_n \to 0$ and $\rho(x) < \infty$.
- 5. $\rho(ax_n) \to 0$ if $\rho(x_n) \to 0$ (metrizing property).
- 6. For any $x \in V$, there exists k > 0 such that $\rho(kx) < \infty$.

• Norm metric

A **norm metric** is a metric on a real (complex) vector space V, defined by

$$||x-y||$$

where ||.|| is a *norm* on V, i.e., a function $||.||:V\to\mathbb{R}$ such that, for all $x,y\in V$ and for any scalar a, we have the following properties:

- 1. $||x|| \ge 0$, with ||x|| = 0 if and only if x = 0.
- 2. ||ax|| = |a|||x||.

1936) by

3. $||x+y|| \le ||x|| + ||y||$ (triangle inequality).

Therefore, a norm ||.|| is a 1-homogeneous F-norm. The vector space (V, ||.||) is called a normed vector space or, simply, normed space.

Any metric space can be embedded isometrically in some normed vector space as a closed linearly independent subset. Every finite-dimensional normed space is **complete**, and all norms on it are equivalent.

In general, the norm ||.|| is equivalent (Maligranda 2008) to the norm

$$||x||_{u,p} = (||x + ||x|| \cdot u||^p + ||x - ||x|| \cdot u||^p)^{\frac{1}{p}},$$

introduced, for any $u \in V$ and $p \ge 1$, by Odell and Schlumprecht in 1998. The **norm-angular distance** between x and y is defined (Clarkson

$$d(x,y) = ||\frac{x}{||x||} - \frac{y}{||y||}||.$$

The following sharpening of the triangle inequality (Maligranda 2003) holds:

$$\frac{||x-y||-|||x||-||y|||}{\min\{||x||,||y||\}} \leq d(x,y) \leq \frac{||x-y||+|||x||-||y|||}{\max\{||x||,||y||\}}, \text{ i.e.,}$$

$$(2 - d(x, -y)) \min\{||x||, ||y||\} \le ||x|| + ||y|| - ||x + y|| \le (2 - d(x, -y)) \max\{||x||, ||y||\}.$$

Dragomir, 2004, call $|\int_a^b f(x)dx| \le \int_a^b f|(x)|dx$ continuous triangle inequality.

• Reverse triangle inequality

The triangle inequality $||x+y|| \le ||x|| + ||y||$ in a normed space (V, ||.||) is equivalent to the following inequality, for any $x_1, \ldots, x_n \in V$ with $n \ge 2$:

$$||\sum_{i=1}^{n} x_i|| \le \sum_{i=1}^{n} ||x_i||.$$

If in the normed space (V, ||.||) for some $C \ge 1$ one has

$$C||\sum_{i=1}^{n} x_i|| \ge \sum_{i=1}^{n} ||x_i||,$$

then this inequality is called the **reverse triangle inequality**. This term is used, sometimes, also for the **inverse triangle inequality** (cf. **kinematic metric** in Chap. 26) and for the eventual inequality $Cd(x,z) \ge d(x,y) + d(y,z)$ with $C \ge 1$ in a metric space (X,d).

The triangle inequality $||x+y|| \le ||x|| + ||y||$, for any $x, y \in V$, in a normed space (V, ||.||) is, for any number q > 1, equivalent (Belbachir, Mirzavaziri and Moslenian 2005) to the following inequality:

$$||x+y||^q \le 2^{q-1}(||x||^q + ||y||^q).$$

The parallelogram inequality $||x+y||^2 \le 2(||x||^2 + ||y||^2)$ is the case q=2 of above.

• Semi-norm semi-metric

A **semi-norm semi-metric** is a semi-metric on a real (complex) vector space V, defined by

$$||x-y||$$
,

where ||.|| is a *semi-norm* (or *pseudonorm*) on V, i.e., a function $||.||:V\to\mathbb{R}$ such that, for all $x,y\in V$ and for any scalar a, we have the following properties:

- 1. $||x|| \ge 0$, with ||0|| = 0.
- 2. ||ax|| = |a|||x||.
- 3. $||x+y|| \le ||x|| + ||y||$ (triangle inequality).

The vector space (V, ||.||) is called a *semi-normed vector space*. Many normed vector spaces, in particular, **Banach spaces**, are defined as the quotient space by the subspace of elements of semi-norm zero.

A quasi-normed space is a vector space V, on which a quasi-norm is given. A quasi-norm on V is a non-negative function $||.||:V\to\mathbb{R}$ which satisfies the same axioms as a norm, except for the triangle inequality which is replaced by the weaker requirement: there exists a constant C>0 such that, for all $x,y\in V$, the following C-triangle inequality holds:

$$||x + y|| \le C(||x|| + ||y||)$$

(cf. near-metric in Chap. 1). An example of a quasi-normed space, that is not normed, is the Lebesgue space $L_p(\Omega)$ with 0 in which a quasi-norm is defined by

$$||f|| = (\int_{\Omega} |f(x)|^p dx)^{1/p}, f \in L_p(\Omega).$$

• Banach space

A Banach space (or B-space) is a complete metric space (V, ||x-y||)on a vector space V with a norm metric ||x-y||. Equivalently, it is the complete normed space (V, ||.||). In this case, the norm ||.|| on V is called the Banach norm. Some examples of Banach spaces are:

- 1. l_p^n -spaces, l_p^∞ -spaces, $1 \le p \le \infty, n \in \mathbb{N}$ 2. The space C of convergent numerical sequences with the norm ||x||=1 $\sup_{n} |x_n|$
- 3. The space C_0 of numerical sequences which converge to zero with the norm $||x|| = \max_n |x_n|$
- 4. The space $C^p_{[a,b]}$, $1 \le p \le \infty$, of continuous functions on [a,b] with the L_p -norm $||f||_p = (\int_a^b |f(t)|^p dt)^{\frac{1}{p}}$
- 5. The space C_K of continuous functions on a compactum K with the norm $||f|| = \max_{t \in K} |f(t)|$
- 6. The space $(C_{[a,b]})^n$ of functions on [a,b] with continuous derivatives up to and including the order n with the norm $||f||_n =$ $\sum_{k=0}^{n} \max_{a \leq t \leq b} |f^{(k)}(t)|$ 7. The space $C^{n}[I^{m}]$ of all functions defined in an m-dimensional cube that
- are continuously differentiable up to and including the order n with the norm of uniform boundedness in all derivatives of order at most n
- 8. The space $M_{[a,b]}$ of bounded measurable functions on [a,b] with the norm

$$||f|| = ess \sup_{a < t < b} |f(t)| = \inf_{e, \mu(e) = 0} \sup_{t \in [a, b] \setminus e} |f(t)|$$

- 9. The space $A(\Delta)$ of functions analytic in the open unit disk $\Delta = \{z \in A(\Delta)\}$ $\mathbb{C}:|z|<1$ and continuous in the closed disk $\overline{\Delta}$ with the norm ||f||= $\max_{z \in \overline{\Lambda}} |f(z)|$
- 10. The Lebesgue spaces $L_p(\Omega)$, $1 \le p \le \infty$
- 11. The Sobolev spaces $W^{k,p}(\Omega)$, $\Omega \subset \mathbb{R}^n$, $1 \leq p \leq \infty$, of functions f on Ω such that f and its derivatives, up to some order k, have a finite L_p -norm, with the norm $||f||_{k,p} = \sum_{i=0}^k ||f^{(i)}||_p$
- 12. The Bohr space AP of almost periodic functions with the norm

$$||f|| = \sup_{-\infty < t < +\infty} |f(t)|$$

A finite-dimensional real Banach space is called a *Minkowskian space*. A norm metric of a Minkowskian space is called a Minkowskian **metric** (cf. Chap. 6). In particular, any l_p -metric is a Minkowskian metric.

All n-dimensional Banach spaces are pairwise isomorphic; the set of such spaces becomes compact if one introduces the Banach-Mazur distance by $d_{BM}(V, W) = \ln \inf_{T} ||T|| \cdot ||T^{-1}||$, where the infimum is taken over all operators which realize an isomorphism $T: V \to W$.

• l_p -metric

The l_p -metric d_{l_p} , $1 \leq p \leq \infty$, is a norm metric on \mathbb{R}^n (or on \mathbb{C}^n), defined by

$$||x-y||_p$$

where the l_p -norm $||.||_p$ is defined by

$$||x||_p = (\sum_{i=1}^n |x_i|^p)^{\frac{1}{p}}.$$

For $p=\infty$, we obtain $||x||_{\infty}=\lim_{p\to\infty}\sqrt[p]{\sum_{i=1}^n|x_i|^p}=\max_{1\leq i\leq n}|x_i|$. The metric space (\mathbb{R}^n,d_{l_p}) is abbreviated as l_p^n and is called l_p^n -space. The l_p -metric, $1\leq p\leq \infty$, on the set of all sequences $x=\{x_n\}_{n=1}^\infty$ of real (complex) numbers, for which the sum $\sum_{i=1}^\infty |x_i|^p$ (for $p=\infty$, the sum $\sum_{i=1}^{\infty} |x_i|$) is finite, is defined by

$$\left(\sum_{i=1}^{\infty}|x_i-y_i|^p\right)^{\frac{1}{p}}.$$

For $p = \infty$, we obtain $\max_{i \ge 1} |x_i - y_i|$. This metric space is abbreviated as l_p^{∞} and is called l_p^{∞} -space.

Most important are l_1 -, l_2 - and l_{∞} -metrics; the l_2 -metric on \mathbb{R}^n is also called the **Euclidean metric**. The l_2 -metric on the set of all sequences $\{x_n\}_n$ of real (complex) numbers, for which $\sum_{i=1}^{\infty}|x_i|^2<\infty$, is also known as the Hilbert metric. On \mathbb{R} all l_p -metrics coincide with the natural metric |x-y|.

Among l_p -metrics, only l_1 - and l_{∞} -metrics are crystalline metrics, i.e., metrics having polygonal unit balls.

Euclidean metric

The Euclidean metric (or Pythagorean distance, as-the-crow-flies **distance**, beeline distance) d_E is the metric on \mathbb{R}^n , defined by

$$||x - y||_2 = \sqrt{(x_1 - y_1)^2 + \dots + (x_n - y_n)^2}.$$

It is the ordinary l_2 -metric on \mathbb{R}^n . The metric space (\mathbb{R}^n, d_E) is abbreviated as \mathbb{E}^n and is called **Euclidean space** (or real Euclidean space). Sometimes, the expression "Euclidean space" stands for the case n=3, as opposed to the Euclidean plane for the case n=2. The Euclidean line (or real line) is obtained for n=1, i.e., it is the metric space $(\mathbb{R},|x-y|)$ with the **natural metric** |x-y| (cf. Chap. 12).

In fact, \mathbb{E}^n is an inner product space (and even a Hilbert space), i.e., $d_E(x,y) = ||x-y||_2 = \sqrt{\langle x-y,x-y\rangle}$, where $\langle x,y\rangle$ is the inner product on \mathbb{R}^n which is given in a suitably chosen (Cartesian) coordinate system by the formula $\langle x,y\rangle=\sum_{i=1}^n x_iy_i$. In a standard coordinate system

one has $\langle x, y \rangle = \sum_{i,j} g_{ij} x_i y_j$, where $g_{ij} = \langle e_i, e_j \rangle$, and the **metric tensor** $((g_{ij}))$ is a positive-definite symmetric $n \times n$ matrix.

In general, a Euclidean space is defined as a space, the properties of which are described by the axioms of *Euclidean Geometry*.

• Norm-related metrics on \mathbb{R}^n

On the vector space \mathbb{R}^n , there are many well-known metrics related to a given norm ||.|| on \mathbb{R}^n , especially, to the Euclidean norm $||.||_2$. Some examples are given below:

1. The British Rail metric (cf. Chap. 19), defined by

$$||x|| + ||y||,$$

for $x \neq y$ (and is equal to 0, otherwise).

2. The radar screen metric (cf. Chap. 19), defined by

$$\min\{1, ||x - y||\}.$$

3. The (p,q)-relative metric (cf. Chap. 19), defined by

$$\frac{||x-y||_2}{\left(\frac{1}{2}(||x||_2^p+||y||_2^p)\right)^{\frac{q}{p}}},$$

for x or $y \neq 0$ (and equal to 0, otherwise), where $0 < q \leq 1$, and $p \geq \max\{1-q, \frac{2-q}{3}\}$. For q=1 and any $1 \leq p < \infty$, one obtains the **p-relative metric**; for q=1 and $p=\infty$, one obtains the **relative metric** (cf. Chap. 19).

4. The M-relative metric (cf. Chap. 19), defined by

$$\frac{||x-y||_2}{f(||x||_2) \cdot f(||y||_2)},$$

for x or $y \neq 0$, where $f:[0,\infty) \to (0,\infty)$ is a convex increasing function such that $\frac{f(x)}{x}$ is decreasing for x>0. In particular, the distance $\frac{||x-y||_2}{\sqrt[p]{1+||x||_2^p}\sqrt[p]{1+||y||_2^p}}$ is a metric on \mathbb{R}^n if and only if $p\geq 1$. A similar metric on $\mathbb{R}^n\setminus\{0\}$ can be defined by $\frac{||x-y||_2}{||x||_2\cdot||y||_2}$.

The last two constructions can be used for any Ptolemaic space (V, ||.||).

• Unitary metric

The unitary metric (or complex Euclidean metric) is the l_2 -metric on \mathbb{C}^n , defined by

$$||x-y||_2 = \sqrt{|x_1 - y_1|^2 + \dots + |x_n - y_n|^2}.$$

The metric space $(\mathbb{C}^n, ||x-y||_2)$ is called the unitary space (or complex Euclidean space). For n=1, we obtain the complex plane (or Argand plane), i.e., the metric space $(\mathbb{C}, |z-u|)$ with the **complex modulus** metric |z-u|; here $|z|=|z_1+iz_2|=\sqrt{z_1^2+z_2^2}$ is the complex modulus (cf. also quaternion metric in Chap. 12).

• L_p -metric

An L_p -metric d_{L_p} , $1 \le p \le \infty$, is a norm metric on $L_p(\Omega, \mathcal{A}, \mu)$, defined by

$$||f-g||_p$$

for any $f, g \in L_p(\Omega, \mathcal{A}, \mu)$. The metric space $(L_p(\Omega, \mathcal{A}, \mu), d_{L_p})$ is called the L_p -space (or Lebesgue space).

Here Ω is a set, and \mathcal{A} is n σ -algebra of subsets of Ω , i.e., a collection of subsets of Ω satisfying the following properties:

- 1. $\Omega \in \mathcal{A}$.
- 2. If $A \in \mathcal{A}$, then $\Omega \backslash A \in \mathcal{A}$.
- 3. If $A = \bigcup_{i=1}^{\infty} A_i$ with $A_i \in \mathcal{A}$, then $A \in \mathcal{A}$.

A function $\mu: \mathcal{A} \to \mathbb{R}_{\geq 0}$ is called a *measure* on \mathcal{A} if it is *additive*, i.e., $\mu(\cup_{i\geq 1}A_i) = \sum_{i\geq 1}\mu(A_i)$ for all pairwise disjoint sets $A_i \in \mathcal{A}$, and satisfies $\mu(\emptyset) = 0$. A *measure space* is a triple $(\Omega, \mathcal{A}, \mu)$.

Given a function $f: \Omega \to \mathbb{R}(\mathbb{C})$, its L_p -norm is defined by

$$||f||_p = \left(\int_{\Omega} |f(\omega)|^p \mu(d\omega)\right)^{\frac{1}{p}}.$$

Let $L_p(\Omega, \mathcal{A}, \mu) = L_p(\Omega)$ denote the set of all functions $f: \Omega \to \mathbb{R}$ (\mathbb{C}) such that $||f||_p < \infty$. Strictly speaking, $L_p(\Omega, \mathcal{A}, \mu)$ consists of equivalence classes of functions, where two functions are *equivalent* if they are equal almost everywhere, i.e., the set on which they differ has measure zero. The set $L_{\infty}(\Omega, \mathcal{A}, \mu)$ is the set of equivalence classes of measurable functions $f: \Omega \to \mathbb{R}$ (\mathbb{C}) whose absolute values are bounded almost everywhere.

The most classical example of an L_p -metric is d_{L_p} on the set $L_p(\Omega, \mathcal{A}, \mu)$, where Ω is the open interval (0,1), \mathcal{A} is the Borel σ -algebra on (0,1), and μ is the Lebesgue measure. This metric space is abbreviated by $L_p(0,1)$ and is called $L_p(0,1)$ -space.

In the same way, one can define the L_p -metric on the set $C_{[a,b]}$ of all real (complex) continuous functions on [a,b]: $d_{L_p}(f,g)=(\int_a^b|f(x)-g(x)|^pdx)^{\frac{1}{p}}$. For $p=\infty$, $d_{L_\infty}(f,g)=\max_{a\leq x\leq b}|f(x)-g(x)|$. This metric space is abbreviated by $C_{[a,b]}^p$ and is called $C_{[a,b]}^p$ -space.

If $\Omega = \mathbb{N}$, $A = 2^{\Omega}$ is the collection of all subsets of Ω , and μ is the cardinality measure (i.e., $\mu(A) = |A|$ if A is a finite subset of Ω , and $\mu(A) = \infty$, otherwise), then the metric space $(L_p(\Omega, 2^{\Omega}, |.|), d_{L_p})$ coincides with the space l_p^{∞} .

If $\Omega = V_n$ is a set of cardinality n, $\mathcal{A} = 2^{V_n}$, and μ is the cardinality measure, then the metric space $(L_p(V_n, 2^{V_n}, |.|), d_{L_p})$ coincides with the space l_p^n .

• Dual metrics

The l_p -metric and the l_q -metric, $1 < p, q < \infty$, are called **dual** if 1/p + 1/q = 1.

In general, when dealing with a normed vector space $(V, ||.||_V)$, one is interested in the continuous linear functionals from V into the base field $(\mathbb{R} \text{ or } \mathbb{C})$. These functionals form a **Banach space** $(V', ||.||_{V'})$, called the continuous dual of V. The norm $||.||_{V'}$ on V' is defined by $||T||_{V'} = \sup_{||x||_V \le 1} |T(x)|$.

The continuous dual for the metric space l_p^n (l_p^∞) is l_q^n (l_q^∞ , respectively). The continuous dual of l_1^n (l_1^∞) is l_∞^n (l_∞^∞ , respectively). The continuous duals of the Banach spaces C (consisting of all convergent sequences, with l_∞ -metric) and C_0 (consisting of the sequences converging to zero, with l_∞ -metric) are both naturally identified with l_1^∞ .

• Inner product space

An **inner product space** (or *pre-Hilbert space*) is a metric space (V, ||x - y||) on a real (complex) vector space V with an *inner product* $\langle x, y \rangle$ such that the norm metric ||x - y|| is constructed using the *inner product norm* $||x|| = \sqrt{\langle x, x \rangle}$.

An inner product \langle, \rangle on a real (complex) vector space V is a symmetric bilinear (in the complex case, sesquilinear) form on V, i.e., a function $\langle, \rangle: V \times V \longrightarrow \mathbb{R}$ (\mathbb{C}) such that, for all $x,y,z \in V$ and for all scalars α,β , we have the following properties:

- 1. $\langle x, x \rangle \geq 0$, with $\langle x, x \rangle = 0$ if and only if x = 0.
- 2. $\langle x, y \rangle = \overline{\langle y, x \rangle}$, where the bar denotes *complex conjugation*.
- 3. $\langle \alpha x + \beta y, z \rangle = \alpha \langle x, z \rangle + \beta \langle y, z \rangle$.

For a complex vector space, an inner product is called also a *Hermitian inner product*, and the corresponding metric space is called a *Hermitian inner product space*.

A norm ||.|| in a normed space (V, ||.||) is generated by an inner product if and only if, for all $x, y \in V$, we have: $||x+y||^2 + ||x-y||^2 = 2(||x||^2 + ||y||^2)$.

• Hilbert space

A Hilbert space is an inner product space which, as a metric space, is **complete**. More precisely, a Hilbert space is a complete metric space (H, ||x-y||) on a real (complex) vector space H with an inner product \langle , \rangle such that the norm metric ||x-y|| is constructed using the inner product norm $||x|| = \sqrt{\langle x, x \rangle}$. Any Hilbert space is a **Banach space**.

An example of a Hilbert space is the set of all sequences $x = \{x_n\}_n$ of real (complex) numbers such that $\sum_{i=1}^{\infty} |x_i|^2$ converges, with the **Hilbert metric** defined by

$$(\sum_{i=1}^{\infty} (x_i - y_i)^2)^{\frac{1}{2}}.$$

Other examples of Hilbert spaces are any L_2 -space, and any finite-dimensional inner product space. In particular, any Euclidean space is a Hilbert space.

A direct product of two **Hilbert spaces** is called a *Liouville space* (or *line space*, *extended Hilbert space*).

Given an infinite cardinal number τ and a set A of the cardinality τ , let \mathbb{R}_a , $a \in A$, be the copies of \mathbb{R} . Let $H(A) = \{\{x_a\} \in \prod_{a \in A} \mathbb{R}_a : \sum_a x_a^2 < \infty\}$; then H(A) with the metric, defined for $\{x_a\}, \{y_a\} \in H(A)$ as

$$(\sum_{a \in A} (x_a - y_a)^2)^{\frac{1}{2}},$$

is called the **generalized Hilbert space** of weight τ .

• Erdös space

The **Erdös space** (or rational Hilbert space) is the metric subspace of l_2 consisting of all vectors in l_2 the coordinates of which are all rational. It has topological dimension 1 and is not complete. Erdös space is **homeomorphic** to its countable infinite power, and every non-empty open subset of it is homeomorphic to whole space.

The **complete Erdös space** (or *irrational Hilbert space*) is the complete metric subspace of l_2 consisting of all vectors in l_2 the coordinates of which are all irrational.

• Riesz norm metric

A Riesz space (or vector lattice) is a partially ordered vector space (V_{Ri}, \preceq) in which the following conditions hold:

- 1. The vector space structure and the partial order structure are compatible, i.e., from $x \leq y$ it follows that $x + z \leq y + z$, and from x > 0, $a \in \mathbb{R}$, a > 0 it follows that ax > 0.
- 2. For any two elements $x, y \in V_{Ri}$, there exist the join $x \vee y \in V_{Ri}$ and $meet \ x \wedge y \in V_{Ri}$ (cf. Chap. 10).

The **Riesz norm metric** is a norm metric on V_{Ri} defined by

$$||x-y||_{Bi}$$

where $||.||_{Ri}$ is a *Riesz norm* on V_{Ri} , i.e., a norm such that, for any $x, y \in V_{Ri}$, the inequality $|x| \leq |y|$, where $|x| = (-x) \vee (x)$, implies $||x||_{Ri} \leq ||y||_{Ri}$.

The space $(V_{Ri}, ||.||_{Ri})$ is called a *normed Riesz space*. In the case of completeness, it is called a *Banach lattice*.

• Banach-Mazur compactum

The Banach–Mazur distance d_{BM} between two *n*-dimensional normed spaces $(V, ||.||_V)$ and $(W, ||.||_W)$ is defined by

$$\ln \inf_{T} ||T|| \cdot ||T^{-1}||,$$

where the infimum is taken over all isomorphisms $T:V\to W$. It is a metric on the set X^n of all equivalence classes of n-dimensional normed spaces, where $V\sim W$ if and only if they are isometric. Then the pair (X^n,d_{BM}) is a compact metric space which is called the **Banach–Mazur compactum**.

• Quotient metric

Given a normed space $(V, ||.||_V)$ with a norm $||.||_V$ and a closed subspace W of V, let $(V/W, ||.||_{V/W})$ be the normed space of cosets $x+W=\{x+w: w\in W\}, x\in V$, with the quotient norm $||x+W||_{V/W}=\inf_{w\in W}||x+w||_V$.

The quotient metric is a norm metric on V/W defined by

$$||(x+W) - (y+W)||_{V/W}.$$

• Tensor norm metric

Given normed spaces $(V, ||.||_V)$ and $(W, ||.||_W)$, a norm $||.||_{\otimes}$ on the tensor product $V \otimes W$ is called tensor norm (or cross norm) if $||x \otimes y||_{\otimes} = ||x||_V ||y||_W$ for all decomposable tensors $x \otimes y$.

The **tensor product metric** is a norm metric on $V \otimes W$ defined by

$$||z-t||_{\otimes}$$
.

For any $z \in V \otimes W$, $z = \sum_j x_j \otimes y_j$, $x_j \in V$, $y_j \in W$, the projective norm (or π -norm) of z is defined by $||z||_{pr} = \inf \sum_j ||x_j||_V ||y_j||_W$, where the infimum is taken over all representations of z as a sum of decomposable vectors. It is the largest tensor norm on $V \otimes W$.

• Valuation metric

A valuation metric is a metric on a field \mathbb{F} defined by

$$||x-y||$$
,

where ||.|| is a *valuation* on \mathbb{F} , i.e., a function $||.||: \mathbb{F} \to \mathbb{R}$ such that, for all $x, y \in \mathbb{F}$, we have the following properties:

- 1. $||x|| \ge 0$, with ||x|| = 0 if and only if x = 0.
- 2. $||xy|| = ||x|| \, ||y||$.
- 3. $||x+y|| \le ||x|| + ||y||$ (triangle inequality).

If $||x+y|| \le \max\{||x||, ||y||\}$, the valuation ||.|| is called *non-Archimedean*. In this case, the valuation metric is an **ultrametric**. The simplest valuation is the *trivial valuation* $||.||_{tr}$: $||0||_{tr} = 0$, and $||x||_{tr} = 1$ for $x \in \mathbb{F} \setminus \{0\}$. It is non-Archimedean.

There are different definitions of valuation in Mathematics. Thus, the function $\nu : \mathbb{F} \to \mathbb{R} \cup \{\infty\}$ is called a valuation if $\nu(x) \geq 0$, $\nu(0) = \infty$, $\nu(xy) = \nu(x) + \nu(y)$, and $\nu(x+y) \geq \min\{\nu(x), \nu(y)\}$ for all $x, y \in \mathbb{F}$. The valuation ||.|| can be obtained from the function ν by the formula $||x|| = \alpha^{\nu(x)}$ for some fixed $0 < \alpha < 1$ (cf. **p-adic metric** in Chap. 12).

The $K\ddot{u}rsch\ddot{a}k$ valuation $|.|_{Krs}$ is a function $|.|_{Krs}: \mathbb{F} \to \mathbb{R}$ such that $|x|_{Krs} \geq 0$, $|x|_{Krs} = 0$ if and only if x = 0, $|xy|_{Krs} = |x|_{Krs}|y|_{Krs}$, and $|x+y|_{Krs} \leq C \max\{|x|_{Krs},|y|_{Krs}\}$ for all $x,y\in\mathbb{F}$ and for some positive constant C, called the constant of valuation. If $C\leq 2$, one obtains the ordinary valuation ||.|| which is non-Archimedean if $C\leq 1$. In general, any $|.|_{Krs}$ is equivalent to some ||.||, i.e., $|.|_{Krs}^p = ||.||$ for some p>0.

Finally, given an ordered group (G,\cdot,e,\leq) equipped with zero, the Krull valuation is a function $|.|:\mathbb{F}\to G$ such that |x|=0 if and only if x=0, |xy|=|x||y|, and $|x+y|\leq \max\{|x|,|y|\}$ for any $x,y\in\mathbb{F}$. It is a generalization of the definition of non-Archimedean valuation ||.|| (cf. generalized metric in Chap. 3).

• Power series metric

Let \mathbb{F} be an arbitrary algebraic field, and let $\mathbb{F}\langle x^{-1}\rangle$ be the field of power series of the form $w = \alpha_{-m}x^m + \cdots + \alpha_0 + \alpha_1x^{-1} + \cdots$, $\alpha_i \in \mathbb{F}$. Given l > 1, a non-Archimedean valuation ||.|| on $\mathbb{F}\langle x^{-1}\rangle$ is defined by

$$||w|| = \begin{cases} l^m, & \text{if } w \neq 0, \\ 0, & \text{if } w = 0. \end{cases}$$

The **power series metric** is the **valuation metric** ||w-v|| on $\mathbb{F}\langle x^{-1}\rangle$.

Part II Geometry and Distances

Chapter 6 Distances in Geometry

Geometry arose as the field of knowledge dealing with spatial relationships. It was one of the two fields of pre-modern Mathematics, the other being the study of numbers. Earliest known evidence of abstract representation – ochre rocks marked with cross hatches and lines to create a consistent complex geometric motif, dated about 70,000 BC – were found in Blombos Cave, South Africa. In modern times, geometric concepts have been generalized to a high level of abstraction and complexity.

6.1 Geodesic Geometry

In Mathematics, the notion of "geodesic" is a generalization of the notion of "straight line" to curved spaces. This term is taken from *Geodesy*, the science of measuring the size and shape of the Earth.

Given a metric space (X,d), a **metric curve** γ is a continuous function $\gamma: I \to X$, where I is an *interval* (i.e., non-empty connected subset) of \mathbb{R} . If γ is r times continuously differentiable, it is called a *regular curve* of class C^r ; if $r = \infty$, γ is called a *smooth curve*.

In general, a curve may cross itself. A curve is called a *simple curve* (or *arc*, *path*) if it does not cross itself, i.e., if it is injective. A curve $\gamma : [a, b] \to X$ is called a *Jordan curve* (or *simple closed curve*) if it does not cross itself, and $\gamma(a) = \gamma(b)$.

The length (which may be equal to ∞) $l(\gamma)$ of a curve $\gamma:[a,b]\to X$ is defined by $\sup\sum_{i=1}^n d(\gamma(t_{i-1}),\gamma(t_i))$, where the supremum is taken over all finite decompositions $a=t_0< t_1< ...< t_n=b, n\in \mathbb{N},$ of [a,b]. A curve with finite length is called rectifiable. For each regular curve $\gamma:[a,b]\to X$ define the natural parameter s of γ by $s=s(t)=l(\gamma|_{[a,t]})$, where $l(\gamma|_{[a,t]})$ is the length of the part of γ corresponding to the interval [a,t]. A curve with this natural parametrization $\gamma=\gamma(s)$ is called **of unit speed**, (or parameterized by arc length, normalized); in this parametrization, for any $t_1,t_2\in I$, one has $l(\gamma|_{[t_1,t_2]})=|t_2-t_1|$, and $l(\gamma)=|b-a|$.

The length of any curve $\gamma:[a,b]\to X$ is at least the distance between its end points: $l(\gamma)\geq d(\gamma(a),\gamma(b))$. The curve γ , for which $l(\gamma)=d(\gamma(a),\gamma(b))$, is called the **geodesic segment** (or *shortest path*) from $x=\gamma(a)$ to $y=\gamma(b)$, and denoted by [x,y]. Thus, a geodesic segment is a shortest join of its endpoints; it is an isometric embedding of [a,b] in X. In general, geodesic segments need not exist, except for a trivial case when the segment consists of one point only. A geodesic segment joining two points need not be unique.

A **geodesic** (cf. Chap. 1) is a curve which extends indefinitely in both directions and behaves locally like a segment, i.e., is everywhere locally a distance minimizer. More exactly, a curve $\gamma : \mathbb{R} \to X$, given in the natural parametrization, is called a *geodesic* if, for any $t \in \mathbb{R}$, there exists a *neighborhood* U of t such that, for any $t_1, t_2 \in U$, we have $d(\gamma(t_1), \gamma(t_2)) = |t_1 - t_2|$. Thus, any geodesic is a locally isometric embedding of the whole of \mathbb{R} in X.

A geodesic is called a **metric straight line** if the equality $d(\gamma(t_1), \gamma(t_2)) = |t_1 - t_2|$ holds for all $t_1, t_2 \in \mathbb{R}$. Such a geodesic is an isometric embedding of the whole real line \mathbb{R} in X. A geodesic is called a **metric great circle** if it is an isometric embedding of a circle $S^1(0,r)$ in X. In general, geodesics need not exist.

• Geodesic metric space

A metric space (X, d) is called **geodesic** if any two points in X can be joined by a **geodesic segment**, i.e., for any two points $x, y \in X$, there is an isometry from the segment [0, d(x, y)] into X. Any complete *Riemannian space* and any Banach space is a geodesic metric space.

A metric space (X, d) is called a **locally geodesic metric space** if any two sufficiently close points in X can be joined by a geodesic segment; it is called D-geodesic if any two points at distance < D can be joined by a geodesic segment.

• Geodesic distance

The **geodesic distance** (or **shortest path distance**) is the length of a **geodesic segment** (i.e., a *shortest path*) between two points.

• Intrinsic metric

Given a metric space (X, d) in which every two points are joined by a rectifiable curve, the **internal metric** (cf. Chap. 4) D on X is defined as the infimum of the lengths of all rectifiable curves, connecting two given points $x, y \in X$.

The metric d on X is called the **intrinsic metric** (or **length metric**) if it coincides with its internal metric D. A metric space with the intrinsic metric is called a **length space** (or **path metric space**, inner metric space, intrinsic space).

If, moreover, any pair x, y of points can be joined by a curve of length d(x, y), the intrinsic metric d is called *strictly intrinsic*, and the length space (X, d) is a **geodesic** metric space.

A complete metric space (X, d) is a length space if and only if it is **having approximate midpoints**, i.e., for any points $x, y \in X$ and for any $\epsilon > 0$, there exists a third point $z \in X$ with $d(x, z), d(y, z) \leq \frac{1}{2}d(x, y) + \epsilon$.

Any complete locally compact length space is a **proper** geodesic metric space.

• G-space

A G-space (or space of geodesics) is a metric space (X, d) with the geometry characterized by the fact that extensions of geodesics, defined as locally shortest lines, are unique. Such geometry is a generalization of $Hilbert\ Geometry$ (see [Buse55]).

More exactly, a G-space (X, d) is defined by the following conditions:

- 1. It is **proper** (or *finitely compact*), i.e., all metric balls are compact.
- 2. It is **Menger-convex**, i.e., for any different $x, y \in X$, there exists a third point $z \in X$, $z \neq x, y$, such that d(x, z) + d(z, y) = d(x, y).
- 3. It is *locally extendable*, i.e., for any $a \in X$, there exists r > 0 such that, for any distinct points x, y in the ball B(a, r), there exists z distinct from x and y such that d(x, y) + d(y, z) = d(x, z).
- 4. It is uniquely extendable, i.e., if in 3 above two points z_1 and z_2 were found, so that $d(y, z_1) = d(y, z_2)$, then $z_1 = z_2$.

The existence of geodesic segments is ensured by finite compactness and Menger-convexity: any two points of a finitely compact Menger-convex set X can be joined by a geodesic segment in X. The existence of geodesics is ensured by the axiom of local prolongation: if a finitely compact Menger-convex set X is locally extendable, then there exists a geodesic containing a given segment. Finally, the uniqueness of prolongation ensures the assumption of Differential Geometry that a *line element* determines a geodesic uniquely.

All Riemannian and Finsler spaces are G-spaces. A one-dimensional G-space is a metric straight line or a metric great circle. Any two-dimensional G-space is a topological manifold.

Every G-space is a *chord space*, i.e., a metric space with a set distinguished geodesic segments such that any two points are joined by a unique such segment (see [BuPh87]).

• Desarguesian space

A **Desarguesian space** is a G-space (X, d) in which the role of geodesics is played by ordinary straight lines. Thus, X may be topologically mapped into a *projective space* $\mathbb{R}P^n$ so that each geodesic of X is mapped into a straight line of $\mathbb{R}P^n$. Any X mapped into $\mathbb{R}P^n$ must either cover all of $\mathbb{R}P^n$ and, in such a case, the geodesics of X are all metric great circles of the same length, or X may be considered as an open convex subset of an affine space A^n .

A space (X,d) of geodesics is a Desarguesian space if and only if the following conditions hold:

- 1. The geodesic passing through two different points is unique.
- 2. For dimension n = 2, both the direct and the converse *Desargues theorems* are valid and, for dimension n > 2, any three points in X lie in one plane.

Among *Riemannian spaces*, the only Desarguesian spaces are Euclidean, *hyperbolic*, and *elliptic spaces*. An example of the non-Riemannian Desarguesian space is the *Minkowskian space* which can be regarded as the prototype of all non-Riemannian spaces, including *Finsler spaces*.

• G-space of elliptic type

A G-space of elliptic type is a G-space in which the geodesic through two points is unique, and all geodesics are the metric great circles of the same length.

Every G-space such that there is unique geodesic through each given pair of points is either a G-space of elliptic type, or a **straight** G-space.

• Straight G-space

A straight G-space is a G-space in which extension of a geodesic is possible globally, so that any segment of the geodesic remains a shortest path. In other words, for any two points $x, y \in X$, there is a unique geodesic segment joining x to y, and a unique metric straight line containing x and y.

Any geodesic in a straight G-space is a metric straight line, and is uniquely determined by any two of its points. Any two-dimensional straight G-space is homeomorphic to the plane.

All simply-connected Riemannian spaces of non-positive curvature (including Euclidean and hyperbolic spaces), Hilbert geometries, and Teichmüller spaces of compact Riemann surfaces of genus g > 1 (when metrized by the **Teichmüller metric**) are straight G-spaces.

• Gromov hyperbolic metric space

A metric space (X, d) is called **Gromov hyperbolic** if it is **geodesic** and δ -hyperbolic for some $\delta \geq 0$.

Any complete simply connected Riemannian space of sectional curvature $k \leq -a^2$ is a Gromov hyperbolic metric space with $\delta = \frac{\ln 3}{a}$. An important class of Gromov hyperbolic metric spaces are the hyperbolic groups, i.e., finitely generated groups whose **word metric** is δ -hyperbolic for some $\delta \geq 0$. A metric space is a **real tree** exactly when it is a Gromov hyperbolic metric space with $\delta = 0$.

A geodesic metric space (X,d) is δ -hyperbolic if and only if it is Rips 4δ -hyperbolic, i.e., each of its $geodesic\ triangles$ (the union of three $geodesic\ segments\ [x,y],\ [x,z],\ [y,z])$ is 4δ -thin (or 4δ -slim): every side of the triangle is contained in the 4δ -neighborhood of the other two sides (a 4δ -neighborhood of a subset $A \subset X$ is the set $\{b \in X : \inf_{a \in A} d(b,a) < 4\delta\}$).

Every $CAT(\kappa)$ space with $\kappa < 0$ is Gromov hyperbolic. Every Euclidean space \mathbb{E}^n is a CAT(0) space; it is Gromov hyperbolic only for n = 1.

• CAT(κ) space

Let (X,d) be a metric space. Let M^2 be a simply connected twodimensional Riemannian manifold of constant curvature κ , i.e., the 2-sphere S_{κ}^2 with $\kappa > 0$, the Euclidean plane \mathbb{E}^2 with $\kappa = 0$, or the hyperbolic plane H_{κ}^2 with $\kappa < 0$. Let D_{κ} denote the diameter of M^2 , i.e., $D_{\kappa} = \frac{\pi}{\sqrt{\kappa}}$ if $\kappa > 0$, and $D_{\kappa} = \infty$ if $\kappa \leq 0$. A triangle T in X consists of three points in X together with three geodesic segments joining them pairwise; the segments are called the sides of the triangle. For a triangle $T \subset X$, a comparison triangle for T in M^2 is a triangle $T' \subset M^2$ together with a map f_T which sends each side of T isometrically onto a side of T'. A triangle T is said to satisfy the Gromov $\mathbf{CAT}(\kappa)$ inequality (for Cartan, Alexandrov and Toponogov) if, for every $x, y \in T$, we have

$$d(x,y) \le d_{M^2}(f_T(x), f_T(y)),$$

where f_T is the map associated to a comparison triangle for T in M^2 . So, the geodesic triangle T is at least as "thin" as its comparison triangle in M^2 .

The metric space (X, d) is a $CAT(\kappa)$ space if it is D_{κ} -geodesic (i.e., any two points at distance $< D_{\kappa}$ can be joined by a geodesic segment), and all triangles T with perimeter $< 2D_{\kappa}$ satisfy the $CAT(\kappa)$ inequality.

Every $CAT(\kappa_1)$ space is a $CAT(\kappa_2)$ space if $\kappa_1 < \kappa_2$. Every **real tree** is a $CAT(-\infty)$ space, i.e., is a $CAT(\kappa_1)$ space for all $\kappa \in \mathbb{R}$.

An Alexandrov space with curvature bounded from above by κ (or locally CAT(κ) space) is a metric space (X,d) in which every point $p \in X$ has a neighborhood U such that any two points $x,y \in U$ are connected by a geodesic segment, and the CAT(κ) inequality holds for any $x,y,z \in U$. A Riemannian manifold is locally CAT(κ) if and only if its sectional curvature is at most κ .

An Alexandrov space with curvature bounded from below by κ is a metric space (X,d) in which every point $p \in X$ has a neighborhood U such that any two points $x,y \in U$ are connected by a geodesic segment, and the reverse $CAT(\kappa)$ inequality

$$d(x,y) \ge d_{M^2}(f_T(x), f_T(y))$$

holds for any $x, y, z \in U$, where f_T is the map associated to a comparison triangle for T in M^2 .

The above two definitions differ only by the sign (≤ 0 or ≥ 0) of $d(x,y)-d_{M^2}(f_T(x),f_T(y))$. In the case $\kappa=0$, the above spaces are called **non-positively curved** and **non-negatively curved** metric spaces, respectively; they differ also by the sign of

$$2d^{2}(z, m(x, y)) - (d^{2}(z, x) + d^{2}(z, y) + \frac{1}{2}d^{2}(x, y))$$

(≤ 0 or ≥ 0 , respectively) where again x, y, z are any three points in a neighborhood U for each $p \in X$, and m(x, y) is the midpoint of the **metric** interval I(x, y).

In a CAT(0) space, any two points are connected by a unique geodesic segment, and the distance is a convex function. Any CAT(0) space is **Busemann convex** and **Ptolemaic** (cf. Chap. 1) and vice versa.

Euclidean spaces, hyperbolic spaces, and trees are CAT(0) spaces. Complete CAT(0) spaces are called also *Hadamard spaces*.

• Bruhat-Tits metric space

A metric space (X, d) satisfies **semi-parallelogram law** (or Bruhat–Tits CN inequality) if for any $x, y \in X$, there is a point m(x, y) that satisfies

$$2d^{2}(z, m(x, y)) - (d^{2}(z, x) + d^{2}(z, y) + \frac{1}{2}d^{2}(x, y)) \le 0.$$

In fact, the point m(x, y) is the unique *midpoint* between x and y (cf. **midpoint convexity** in Chap. 1).

A geodesic space is a CAT(0) space if and only if it satisfies above inequality.

The usual vector parallelogram law $||u-v||^2 + ||u+v||^2 = 2||u||^2 + 2||v||^2$, characterizing norms induced by inner products, is equivalent to the semi-parallelogram law with the inequality replaced by an equality.

A Bruhat-Tits metric space is a complete metric space satisfying the semi-parallelogram law.

Boundary of metric space

There are many notions of the **boundary** ∂X of a metric space (X, d). We give below some of the most general among them. Usually, if (X, d) is locally compact, $X \cup \partial X$ is its *compactification*:

1. **Ideal boundary**. Given a geodesic metric space (X, d), let γ^1 and γ^2 be two **metric rays**, i.e., geodesics with isometry of $\mathbb{R}_{\geq 0}$ into X. These rays are called *equivalent* if the **Hausdorff distance** between them (associated with the metric d) is finite, i.e., if $\sup_{t\geq 0} d(\gamma^1(t), \gamma^2(t)) < \infty$.

The boundary at infinity (or ideal boundary) of (X, d) is the set $\partial_{\infty}X$ of equivalence classes γ_{∞} of all metric rays. Cf. metric cone structure, asymptotic metric cone in Chap. 1.

If (X, d) is a complete CAT(0) space, then the **Tits metric** (or asymptotic angle of divergence) on $\partial_{\infty}X$ is defined by

$$2\arcsin\left(\frac{\rho}{2}\right)$$

for all $\gamma_{\infty}^1, \gamma_{\infty}^2 \in \partial_{\infty} X$, where $\rho = \lim_{t \to +\infty} \frac{1}{t} d(\gamma^1(t), \gamma^2(t))$. The set $\partial_{\infty} X$ equipped with the Tits metric is called the **Tits boundary** of X. If (X, d, x_0) is a pointed complete CAT(-1) space, then the **Bourdon metric** (or **visual distance**) on $\partial_{\infty} X$ is defined by

$$\rho^{-}(x.y)$$

for any distinct points $x, y \in \partial_{\infty} X$, where (x.y) denotes the **Gromov product** $(x.y)_{x_0}$.

The visual sphere of (X,d) at a point $x_0 \in X$ is the set of equivalence classes of all metric rays emanating from x_0 .

2. **Gromov boundary**. Given a pointed metric space (X, d, x_0) , the **Gromov boundary** of it (as generalized by Buckley and Kokkendorff 2005, from the case of the Gromov hyperbolic space) is the set $\partial_G X$ of equivalence classes of *Gromov sequences*. A sequence $x = \{x_n\}_n$ in X is a *Gromov sequence* if the Gromov product $(x_i.x_j)_{x_0} \to \infty$ as $i, j \to \infty$. Two Gromov sequences x and y are equivalent if there is a finite chain of Gromov sequences x^k , $0 \le k \le k'$, such that $x = x^0, y = x^{k'}$, and $\lim_{i,j\to\infty}\inf(x_i^{k-1}.x_j^k) = \infty$ for $0 \le k \le k'$.

In a **proper** geodesic Gromov hyperbolic space (X, d), the visual sphere does not depends on the base point x_0 and is naturally isomorphic to its *Gromov boundary* $\partial_G X$, which can be identified with $\partial_{\infty} X$.

- 3. g-boundary. Denote by $\overline{X_d}$ the metric completion of (X, d) and, viewing X as a subset of $\overline{X_d}$, denote by ∂X_d the difference $\overline{X_d} \backslash X$. Let (X, l, x_0) be a pointed unbounded **length space**, i.e., its metric coincides with the **internal metric** l of (X, d). Given a measurable function $g: \mathbb{R}_{\geq 0} \to \mathbb{R}_{\geq 0}$, the g-boundary of (X, d, x_0) (as generalized by Buckley and Kokkendorff 2005, from spherical and Floyd boundaries) is $\partial_g X = \partial X_\sigma \backslash \partial X_l$, where $\sigma(x,y) = \inf \int_{\gamma} g(z) dl(z)$ for all $x, y \in X$ (here the infimum is taken over all metric segments $\gamma = [x, y]$).
- 4. Hotchkiss boundary. Given a pointed proper Busemann convex metric space (X, d, x_0) , the Hotchkiss boundary of it is the set $\partial_H(X, x_0)$ of isometries $f: \mathbb{R}_{\geq 0} \to X$ with $f(0) = x_0$. The boundaries $\partial_H^{x_0}X$ and $\partial_H^{x_1}X$ are homeomorphic for distinct $x_0, x_1 \in X$. In a Gromov hyperbolic space, $\partial_H^{x_0}X$ is homeomorphic to the Gromov boundary $\partial_G X$.
- 5. Metric boundary. Given a pointed metric space (X,d,x_0) and an unbounded subset S of $\mathbb{R}_{\geq 0}$, a ray $\gamma:S\to X$ is called a weakly geodesic ray if, for every $x\in X$ and every $\epsilon>0$, there is an integer N such that $|d(\gamma(t),\gamma(0))-t|<\epsilon$, and $|d(\gamma(t),x)-d(\gamma(s),x)-(t-s)|<\epsilon$ for all $s,t\in T$ with $s,t\geq N$. Let $\mathcal{G}(X,d)$ be the commutative unital C^* -algebra with the norm $||.||_{\infty}$, generated by the (bounded, continuous) functions which vanish at infinity, the constant functions, and the functions of the form $g_y(x)=d(x,x_0)-d(x,y)$; cf. quantum metric space for definitions. The Rieffel's metric boundary $\partial_R X$ of (X,d) is the difference $\overline{X}^d\backslash X$, where \overline{X}^d is the metric compactification of (X,d), i.e., the maximum ideal space (the set of pure states) of this C^* -algebra.

For a proper metric space (X, d) with a countable base, the boundary $\partial_R X$ consists of the limits $\lim_{t\to\infty} f(\gamma(t))$ for every weakly geodesic ray γ and every function f from the above C^* -algebra (Rieffel 2002).

• Projectively flat metric space

A metric space, in which geodesics are defined, is called **projectively flat** if it locally admits a *geodesic mapping* (or *projective mapping*), i.e., a mapping preserving geodesics into an Euclidean space. Cf. Euclidean **rank of metric space** in Chap. 1; similar terms are: *affinely flat*, *conformally flat*, etc.

A Riemannian space is projectively flat if and only if it has constant (sectional) curvature.

6.2 Projective Geometry

Projective Geometry is a branch of Geometry dealing with the properties and invariants of geometric figures under projection. Affine Geometry, Similarity (or Metric) Geometry and Euclidean Geometry are subsets of Projective Geometry of increasing complexity. The main invariants of Projective, Affine, Metric, Euclidean Geometry are, respectively, cross-ratio, parallelism (and relative distances), angles (and relative distances), absolute distances.

An *n*-dimensional projective space $\mathbb{F}P^n$ is the space of one-dimensional vector subspaces of a given (n+1)-dimensional vector space V over a field \mathbb{F} . The basic construction is to form the set of equivalence classes of non-zero vectors in V under the relation of scalar proportionality. This idea goes back to mathematical descriptions of perspective. The use of a basis of V allows the introduction of homogeneous coordinates of a point in $\mathbb{K}P^n$ which are usually written as $(x_1:x_2:\ldots:x_n:x_{n+1})$ – a vector of length n+1, other than $(0:0:0:\ldots:0)$. Two sets of coordinates that are proportional denote the same point of the projective space. Any point of projective space which can be represented as $(x_1:x_2:\ldots:x_n:0)$ is called a point at infinity. The part of a projective space $\mathbb{K}P^n$ not "at infinity," i.e., the set of points of the projective space which can be represented as $(x_1:x_2:\ldots:x_n:1)$, is an n-dimensional affine space A^n .

The notation $\mathbb{R}P^n$ denotes the real projective space of dimension n, i.e., the space of one-dimensional vector subspaces of \mathbb{R}^{n+1} . The notation $\mathbb{C}P^n$ denotes the complex projective space of complex dimension n. The projective space $\mathbb{R}P^n$ carries a natural structure of a compact smooth n-dimensional manifold. It can be viewed as the space of lines through the zero element 0 of \mathbb{R}^{n+1} (i.e., as a ray space). It can be viewed also as the set \mathbb{R}^n , considered as an affine space, together with its points at infinity. Also it can be seen as the set of points of an n-dimensional sphere in \mathbb{R}^{n+1} with identified diametrically-opposite points.

The projective points, projective straight lines, projective planes,..., projective hyperplanes of $\mathbb{K}P^n$ are one-dimensional, two-dimensional, three-dimensional,..., n-dimensional subspaces of V, respectively. Any two projective straight lines in a projective plane have one and only one common point. A projective transformation (or collineation, projectivity) is a bijection

of a projective space onto itself, preserving collinearity (the property of points to be on one line) in both directions. Any projective transformation is a composition of a pair of *perspective projections*. Projective transformations do not preserve sizes or angles but do preserve *type* (that is, points remain points, and lines remain lines), *incidence* (that is, whether a point lies on a line), and *cross-ratio*.

Here, given four collinear points $x,y,z,t\in\mathbb{F}P^n$, their cross-ratio is defined by $(x,y,z,t)=\frac{(x-z)(y-t)}{(y-z)(x-t)}$, where $\frac{x-z}{x-t}$ denotes the ratio $\frac{f(x)-f(z)}{f(x)-f(t)}$ for some affine bijection f from the straight line $l_{x,y}$ through the points x and y onto \mathbb{K} . Given four projective straight lines l_x,l_y,l_z,l_t , containing points x,y,z,t, respectively, and passing through a given point, their cross-ratio, defined by $(l_x,l_y,l_z,l_t)=\frac{\sin(l_x,l_z)\sin(l_y,l_t)}{\sin(l_y,l_z)\sin(l_x,l_t)}$, coincides with (x,y,z,t). The cross-ratio of four complex numbers x,y,z,t is given by $(x,y,z,t)=\frac{(x-z)(y-t)}{(y-z)(x-t)}$. It is real if and only if the four numbers are either collinear or concyclic.

• Projective metric

Given a convex subset D of a projective space $\mathbb{R}P^n$, the **projective metric** d is a metric on D such that shortest paths with respect to this metric are parts of or entire projective straight lines. It is assumed that the following conditions hold:

- 1. D does not belong to a hyperplane.
- 2. For any three non-collinear points $x, y, z \in D$, the triangle inequality holds in the strict sense: d(x, y) < d(x, z) + d(z, y).
- 3. If x, y are different points in D, then the intersection of the straight line $l_{x,y}$ through x and y with D is either all of $l_{x,y}$, and forms a **metric great circle**, or is obtained from $l_{x,y}$ by discarding some segment (which can be reduced to a point), and forms a **metric straight line**.

The metric space (D,d) is called a **projective metric space** (cf. **projectively flat space**). The problem of determining all projective metrics constitutes the *fourth problem of Hilbert*; it has been solved only for dimension n=2. In fact, given a smooth measure on the space of hyperplanes in $\mathbb{R}P^n$, define the distance between any two points $x,y\in\mathbb{R}P^n$ as one-half the measure of all hyperplanes intersecting the line segment joining x and y. The obtained metric is projective; it is the Busemann's construction of projective metrics. For n=2, Ambartzumian [Amba76] proved that all projective metrics can be obtained from the Busemann's construction.

In a projective metric space there cannot simultaneously be both types of straight lines: they are either all metric straight lines, or they are all metric great circles of the same length (Hamel's theorem). Spaces of the first kind are called open. They coincide with subspaces of an affine space; the geometry of open projective metric spaces is a Hilbert Geometry. Hyperbolic Geometry is a Hilbert Geometry in which there exist reflections at all straight lines. Thus, the set D has Hyperbolic Geometry if and only if it is the interior of an ellipsoid. The geometry of open projective metric spaces

whose subsets coincide with all of affine space, is a *Minkowski Geometry*. Euclidean Geometry is a Hilbert Geometry and a Minkowski Geometry, simultaneously. Spaces of the second kind are called *closed*; they coincide with the whole of $\mathbb{R}P^n$. Elliptic Geometry is the geometry of a projective metric space of the second kind.

• Strip projective metric

The strip projective metric [BuKe53] is a projective metric on the strip $St = \{x \in \mathbb{R}^2 : -\pi/2 < x_2 < \pi/2\}$ defined by

$$\sqrt{(x_1-y_1)^2+(x_2-y_2)^2}+|\tan x_2-\tan y_2|.$$

Note, that St with the ordinary Euclidean metric $\sqrt{(x_1 - y_1)^2 + (x_2 - y_2)^2}$ is not a projective metric space.

• Half-plane projective metric

The half-plane projective metric [BuKe53] is a projective metric on $\mathbb{R}^2_+ = \{x \in \mathbb{R}^2 : x_2 > 0\}$ defined by

$$\sqrt{(x_1-y_1)^2+(x_2-y_2)^2}+\left|\frac{1}{x_2}-\frac{1}{y_2}\right|.$$

• Hilbert projective metric

Given a set H, the **Hilbert projective metric** h is a **complete projective metric** on H. It means that H contains, together with two arbitrary distinct points x and y, also the points z and t for which h(x,z)+h(z,y)=h(x,y), h(x,y)+h(y,t)=h(x,t), and that H is homeomorphic to a *convex* set in an n-dimensional affine space A^n , the geodesics in H being mapped to straight lines of A^n . The metric space (H,h) is called the *Hilbert projective space*, and the geometry of a Hilbert projective space is called *Hilbert Geometry*.

Formally, let D be a non-empty convex open set in A^n with the boundary ∂D not containing two proper coplanar but non-collinear segments (ordinarily the boundary of D is a strictly convex closed curve, and D is its interior). Let $x, y \in D$ be located on a straight line which intersects ∂D at z and t, z is on the side of y, and t is on the side of x. Then the Hilbert metric h on D is defined by

$$\frac{r}{2}\ln(x,y,z,t),$$

where (x, y, z, t) is the *cross-ratio* of x, y, z, t, and r is a fixed positive constant.

The metric space (D, h) is a **straight** G-space. If D is an ellipsoid, then h is the **hyperbolic metric**, and defines Hyperbolic Geometry on D. On the unit disk $\Delta = \{z \in \mathbb{C} : |z| < 1\}$ the metric h coincides with

the Cayley-Klein-Hilbert metric. If n = 1, the metric h makes D isometric to the Euclidean line.

If ∂D contains coplanar but non-collinear segments, a projective metric on D can be given by h(x,y) + d(x,y), where d is any **Minkowskian metric** (usually, the Euclidean metric).

• Minkowskian metric

The Minkowskian metric (or Minkowski-Hölder distance) is the norm metric of a finite-dimensional real Banach space.

Formally, let \mathbb{R}^n be an *n*-dimensional real vector space, let K be a *symmetric convex body* in \mathbb{R}^n , i.e., an open neighborhood of the origin which is bounded, convex, and *symmetric* $(x \in K \text{ if and only if } -x \in K)$. Then the *Minkowski functional* $||.||_K : \mathbb{R}^n \to [0, \infty)$ defined by

$$||x||_K = \inf\{\alpha > 0 : \frac{x}{\alpha} \in \partial K\}$$

is a *norm* on \mathbb{R}^n , and the Minkowskian metric m on \mathbb{R}^n is defined by

$$||x-y||_K$$
.

The metric space (\mathbb{R}^n, m) is called $Minkowskian \, space$. It can be considered as an n-dimensional affine space A^n with a metric m in which the role of the $unit \, ball$ is played by a given centrally-symmetric convex body. The geometry of a Minkowskian space is called Minkowski Geometry. For a strictly convex symmetric body the Minkowskian metric is a **projective metric**, and (\mathbb{R}^n, m) is a G-straight space. A Minkowski Geometry is Euclidean if and only if its $unit \, sphere$ is an ellipsoid.

The Minkowskian metric m is proportional to the Euclidean metric d_E on every given line l, i.e., $m(x,y) = \phi(l)d_E(x,y)$. Thus, the Minkowskian metric can be considered as a metric which is defined in the whole affine space A^n and has the property that the affine ratio $\frac{ac}{ab}$ of any three collinear points a, b, c (cf. Sect. 6.3) is equal to their distance ratio $\frac{m(a,c)}{m(a,b)}$.

• Busemann metric

The **Busemann metric** [Buse55] is a metric on the real *n*-dimensional projective space $\mathbb{R}P^n$ defined by

$$\min \left\{ \sum_{i=1}^{n+1} \left| \frac{x_i}{||x||} - \frac{y_i}{||y||} \right|, \sum_{i=1}^{n+1} \left| \frac{x_i}{||x||} + \frac{y_i}{||y||} \right| \right\}$$

for any
$$x = (x_1 : ... : x_{n+1}), y = (y_1 : ... : y_{n+1}) \in \mathbb{R}P^n$$
, where $||x|| = \sqrt{\sum_{i=1}^{n+1} x_i^2}$.

• Flag metric

Given an *n*-dimensional projective space $\mathbb{F}P^n$, the **flag metric** d is a metric on $\mathbb{F}P^n$ defined by a flag, i.e., an absolute consisting of a collection of

m-planes α_m , m = 0, ..., n-1, with α_{i-1} belonging to α_i for all $i \in \{1, ..., n-1\}$. The metric space $(\mathbb{F}P^n, d)$ is abbreviated by F^n and is called a *flag space*.

If one chooses an affine coordinate system $(x_i)_i$ in a space F^n , so that the vectors of the lines passing through the (n-m-1)-plane α_{n-m-1} are defined by the condition $x_1 = \dots x_m = 0$, then the flag metric d(x,y) between the points $x = (x_1, \dots, x_n)$ and $y = (y_1, \dots, y_n)$ is defined by

$$d(x,y) = |x_1 - y_1|$$
, if $x_1 \neq y_1$, $d(x,y) = |x_2 - y_2|$, if $x_1 = y_1, x_2 \neq y_2, \dots$
 $\dots, d(x,y) = |x_k - y_k|$, if $x_1 = y_1, \dots, x_{k-1} = y_{k-1}, x_k \neq y_k, \dots$

• Projective determination of a metric

The **projective determination of a metric** is an introduction, in subsets of a projective space, of a metric such that these subsets become isomorphic to a Euclidean, *hyperbolic*, or *elliptic space*.

To obtain a Euclidean determination of a metric in $\mathbb{R}P^n$, one should distinguish in this space an (n-1)-dimensional hyperplane π , called the hyperplane at infinity, and define \mathbb{E}^n as the subset of the projective space obtained by removing from it this hyperplane π . In terms of homogeneous coordinates, π consists of all points $(x_1 : \ldots : x_n : 0)$, and \mathbb{E}^n consists of all points $(x_1 : \ldots : x_n : x_{n+1})$ with $x_{n+1} \neq 0$. Hence, it can be written as $\mathbb{E}^n = \{x \in \mathbb{R}P^n : x = (x_1 : \ldots : x_n : 1)\}$. The Euclidean metric d_E on \mathbb{E}^n is defined by

$$\sqrt{\langle x-y, x-y \rangle}$$
,

where, for any $x = (x_1 : ... : x_n : 1), y = (y_1 : ... : y_n : 1) \in \mathbb{E}^n$, one has $\langle x, y \rangle = \sum_{i=1}^n x_i y_i$.

To obtain a hyperbolic determination of a metric in $\mathbb{R}P^n$, a set D of interior points of a real oval hypersurface Ω of order two in $\mathbb{R}P^n$ is considered. The hyperbolic metric d_{hyp} on D is defined by

$$\frac{r}{2}|\ln(x,y,z,t)|,$$

where z and t are the points of intersection of the straight line $l_{x,y}$ through the points x and y with Ω , (x,y,z,t) is the *cross-ratio* of the points x,y,z,t, and r is a fixed positive constant. If, for any $x=(x_1:\ldots:x_{n+1}),y=(y_1:\ldots:y_{n+1})\in\mathbb{R}P^n$, the *scalar product* $\langle x,y\rangle=-x_1y_1+\sum_{i=2}^{n+1}x_iy_i$ is defined, the hyperbolic metric on the set $D=\{x\in\mathbb{R}P^n:\langle x,x\rangle<0\}$ can be written as

$$r \operatorname{arccosh} \frac{|\langle x, y \rangle|}{\sqrt{\langle x, x \rangle \langle y, y \rangle}},$$

where r is a fixed positive constant, and arccosh denotes the non-negative values of the inverse hyperbolic cosine.

To obtain an elliptic determination of a metric in $\mathbb{R}P^n$, one should consider, for any $x=(x_1:\ldots:x_{n+1}),y=(y_1:\ldots:y_{n+1})\in\mathbb{R}P^n$, the inner product $\langle x,y\rangle=\sum_{i=1}^{n+1}x_iy_i$. The elliptic metric d_{ell} on $\mathbb{R}P^n$ is defined now by

$$r \arccos \frac{|\langle x, y \rangle|}{\sqrt{\langle x, x \rangle \langle y, y \rangle}},$$

where r is a fixed positive constant, and arccos is the inverse cosine in $[0, \pi]$. In all the considered cases, some hypersurfaces of the second order remain invariant under given **motions**, i.e., projective transformations preserving a given metric. These hypersurfaces are called *absolutes*. In the case of a Euclidean determination of a metric, the absolute is an imaginary (n-2)-dimensional oval surface of order two, in fact, the degenerate absolute $x_1^2 + \cdots + x_n^2 = 0$, $x_{n+1} = 0$. In the case of a hyperbolic determination of a metric, the absolute is a real (n-1)-dimensional oval hypersurface of order two, in the simplest case, the absolute $-x_1^2 + x_2^2 + \cdots + x_{n+1}^2 = 0$. In the case of an elliptic determination of a metric, the absolute is an imaginary (n-1)-dimensional oval hypersurface of order two, in fact, the absolute $x_1^2 + \cdots + x_{n+1}^2 = 0$.

6.3 Affine Geometry

An *n*-dimensional *affine space* over a field \mathbb{F} is a set A^n (the elements of which are called *points* of the affine space) to which corresponds an *n*-dimensional vector space V over \mathbb{F} (called the *space associated to* A^n) such that, for any $a \in A^n$, $A = a + V = \{a + v : v \in V\}$. In the other words, if $a = (a_1, \ldots, a_n)$ and $b = (b_1, \ldots, b_n) \in A^n$, then the vector $\overrightarrow{ab} = (b_1 - a_1, \ldots, b_n - a_n)$ belongs to V. In an affine space, one can add a vector to a point to get another point, and subtract points to get vectors, but one cannot add points, since there is no origin. Given points $a, b, c, d \in A^n$ such that $c \neq d$, and the vectors \overrightarrow{ab} and \overrightarrow{cd} are collinear, the scalar λ , defined by $\overrightarrow{ab} = \lambda \overrightarrow{cd}$, is called the *affine ratio* of ab and cd, and is denoted by $\frac{ab}{cd}$.

An affine transformation (or affinity) is a bijection of A^n onto itself which preserves collinearity (i.e., all points lying on a line initially, still lie on a line after transformation) and ratios of distances (for example, the midpoint of a line segment remains the midpoint after transformation). In this sense, affine indicates a special class of projective transformations that do not move any objects from the affine space to the plane at infinity or conversely. Any affine transformation is a composition of rotations, translations, dilations, and shears. The set of all affine transformations of A^n forms a group $Aff(A^n)$,

called the *general affine group* of A^n . Each element $f \in Aff(A)$ can be given by a formula f(a) = b, $b_i = \sum_{j=1}^n p_{ij}a_j + c_j$, where $((p_{ij}))$ is an invertible matrix.

The subgroup of $Aff(A^n)$, consisting of affine transformations with $det((p_{ij}))=1$, is called the equi-affine group of A^n . An equi-affine space is an affine space with the equi-affine group of transformations. The fundamental invariants of an equi-affine space are volumes of parallelepipeds. In an equi-affine plane A^2 , any two vectors v_1, v_2 have an invariant $|v_1 \times v_2|$ (the modulus of their cross product) – the surface area of the parallelogram constructed on v_1 and v_2 . Given a non-rectilinear curve $\gamma = \gamma(t)$, its affine parameter (or equi-affine arc length) is an invariant parameter defined by $s = \int_{t_0}^t |\gamma' \times \gamma''|^{1/3} dt$. The invariant $k = \frac{d^2 \gamma}{ds^2} \times \frac{d^3 \gamma}{ds^3}$ is called the equi-affine curvature of γ . Passing to the general affine group, two more invariants of the curve are considered: the affine arc length $\sigma = \int k^{1/2} ds$, and the affine curvature $k = \frac{1}{13/2} \frac{dk}{ds}$.

curvature $k = \frac{1}{k^{3/2}} \frac{dk}{ds}$. For A^n , n > 2, the affine parameter (or equi-affine arc length) of a curve $\gamma = \gamma(t)$ is defined by $s = \int_{t_0}^t |(\gamma', \gamma'', \dots, \gamma^{(n)})|^{\frac{2}{n(n+1)}} dt$, where the invariant (v_1, \dots, v_n) is the (oriented) volume spanned by the vectors v_1, \dots, v_n , which is equal to the determinant of the $n \times n$ matrix whose *i*-th column is the vector v_i .

• Affine distance

Given an affine plane A^2 , a line element (a, l_a) of A^2 consists of a point $a \in A^2$ together with a straight line $l_a \subset A^2$ passing through a.

The **affine distance** is a distance on the set of all line elements of A^2 defined by

$$2f^{1/3}$$
,

where, for a given line elements (a, l_a) and (b, l_b) , f is the surface area of the triangle abc if c is the point of intersection of the straight lines l_a and l_b . The affine distance between (a, l_a) and (b, l_b) can be interpreted as the affine length of the arc ab of a parabola such that l_a and l_b are tangent to the parabola at a and b, respectively.

• Affine pseudo-distance

Let A^2 be an equi-affine plane, and let $\gamma = \gamma(s)$ be a curve in A^2 defined as a function of the affine parameter s. The **affine pseudo-distance** dp_{aff} for A^2 is defined by

$$dp_{aff}(a,b) = \left| \overrightarrow{ab} \times \frac{d\gamma}{ds} \right|,$$

i.e., is equal to the surface area of the parallelogram constructed on the vectors \overrightarrow{ab} and $\frac{d\gamma}{ds}$, where b is an arbitrary point in A^2 , a is a point on γ , and $\frac{d\gamma}{ds}$ is the tangent vector to the curve γ at the point a.

The **affine pseudo-distance** for an *equi-affine space* A^3 can be defined in a similar manner as

$$\left| (\overrightarrow{ab}, \frac{d\gamma}{ds}, \frac{d^2\gamma}{ds^2}) \right|,$$

where $\gamma = \gamma(s)$ is a curve in A^3 , defined as a function of the affine parameter $s, b \in A^3$, a is a point of γ , and the vectors $\frac{d\gamma}{ds}, \frac{d^2\gamma}{ds^2}$ are obtained at the point a.

For A^n , n > 3, we have $dp_{aff}(a,b) = |(\overrightarrow{ab}, \frac{d\gamma}{ds}, \dots, \frac{d^{n-1}\gamma}{ds^{n-1}})|$. For an arbitrary parametrization $\gamma = \gamma(t)$, one obtains $dp_{aff}(a,b) = |(\overrightarrow{ab}, \gamma', \dots, \gamma^{(n-1)})||(\gamma', \dots, \gamma^{(n-1)})||\frac{1-n}{1+n}|$.

• Affine metric

The **affine metric** is a metric on a non-developable surface $r = r(u_1, u_2)$ in an equi-affine space A^3 , given by its **metric tensor** $((g_{ij}))$:

$$g_{ij} = \frac{a_{ij}}{|det((a_{ij}))|^{1/4}},$$

where $a_{ij} = (\partial_1 r, \partial_2 r, \partial_{ij} r), i, j \in \{1, 2\}.$

6.4 Non-Euclidean Geometry

The term non-Euclidean Geometry describes both Hyperbolic Geometry (or Lobachevsky Geometry, Lobachevsky–Bolyai–Gauss Geometry) and Elliptic Geometry (sometimes called also Riemannian Geometry) which are contrasted with Euclidean Geometry (or Parabolic Geometry). The essential difference between Euclidean and non-Euclidean Geometry is the nature of parallel lines. In Euclidean Geometry, if we start with a line l and a point a, which is not on l, then there is only one line through a that is parallel to l. In Hyperbolic Geometry there are infinitely many lines through a parallel to l. In Elliptic Geometry, parallel lines do not exist.

The *Spherical Geometry* is also "non-Euclidean," but it fails the axiom that any two points determine exactly one line.

• Spherical metric

Let $S^n(0,r)=\{x\in\mathbb{R}^{n+1}:\sum_{i=1}^{n+1}x_i^2=r^2\}$ be the sphere in \mathbb{R}^{n+1} with the center 0 and the radius r>0.

The spherical metric (or great circle metric) d_{sph} is a metric on $S^{n}(0,r)$ defined by

$$r \arccos\left(\frac{\left|\sum_{i=1}^{n+1} x_i y_i\right|}{r^2}\right)$$

where arccos is the inverse cosine in $[0,\pi]$. It is the length of the *great circle* arc, passing through x and y. In terms of the standard *inner product* $\langle x,y\rangle = \sum_{i=1}^{n+1} x_i y_i$ on \mathbb{R}^{n+1} , the spherical metric can be written as $r \arccos \frac{|\langle x,y\rangle|}{\sqrt{\langle x,x\rangle\langle y,y\rangle}}$.

The metric space $(S^n(0,r), d_{sph})$ is called *n*-dimensional spherical space. It is a space of curvature $1/r^2$, and r is the radius of curvature. It is a model of *n*-dimensional Spherical Geometry. The great circles of the sphere are its geodesics and all geodesics are closed and of the same length. (See, for example, [Blum70].)

• Elliptic metric

Let $\mathbb{R}P^n$ be the real *n*-dimensional projective space. The **elliptic metric** d_{ell} is a metric on $\mathbb{R}P^n$ defined by

$$r \arccos \frac{|\langle x, y \rangle|}{\sqrt{\langle x, x \rangle \langle y, y \rangle}},$$

where, for any $x = (x_1 : ... : x_{n+1})$ and $y = (y_1 : ... : y_{n+1}) \in \mathbb{R}P^n$, one has $\langle x, y \rangle = \sum_{i=1}^{n+1} x_i y_i$, r is a fixed positive constant, and arccos is the inverse cosine in $[0, \pi]$.

The metric space $(\mathbb{R}P^n, d_{ell})$ is called *n-dimensional elliptic space*. It is a model of *n*-dimensional *Elliptic Geometry*. It is the space of curvature $1/r^2$, and r is the radius of curvature. As $r \to \infty$, the metric formulas of Elliptic Geometry yield formulas of Euclidean Geometry (or become meaningless).

If $\mathbb{R}P^n$ is viewed as the set $E^n(0,r)$, obtained from the sphere $S^n(0,r)=\{x\in\mathbb{R}^{n+1}:\sum_{i=1}^{n+1}x_i^2=r^2\}$ in \mathbb{R}^{n+1} with center 0 and radius r by identifying diametrically-opposite points, then the elliptic metric on $E^n(0,r)$ can be written as $d_{sph}(x,y)$ if $d_{sph}(x,y)\leq \frac{\pi}{2}r$, and as $\pi r-d_{sph}(x,y)$ if $d_{sph}(x,y)>\frac{\pi}{2}r$, where d_{sph} is the **spherical metric** on $S^n(0,r)$. Thus, no two points of $E^n(0,r)$ have distance exceeding $\frac{\pi}{2}r$. The elliptic space $(E^2(0,r),d_{ell})$ is called the *Poincaré sphere*.

If $\mathbb{R}P^n$ is viewed as the set E^n of lines through the zero element 0 in \mathbb{R}^{n+1} , then the elliptic metric on E^n is defined as the angle between the corresponding subspaces.

An *n*-dimensional elliptic space is a *Riemannian space* of constant positive curvature. It is the only such space which is topologically equivalent to a projective space. (See, for example, [Blum70], [Buse55].)

• Hermitian elliptic metric

Let $\mathbb{C}P^n$ be the *n*-dimensional complex projective space. The **Hermitian** elliptic metric d_{ell}^H (see, for example, [Buse55]) is a metric on $\mathbb{C}P^n$ defined by

$$r \arccos \frac{|\langle x, y \rangle|}{\sqrt{\langle x, x \rangle \langle y, y \rangle}},$$

where, for any $x = (x_1 : ... : x_{n+1})$ and $y = (y_1 : ... : y_{n+1}) \in \mathbb{C}P^n$, one has $\langle x, y \rangle = \sum_{i=1}^{n+1} \overline{x}_i y_i$, r is a fixed positive constant, and arccos is the inverse cosine in $[0, \pi]$.

The metric space $(\mathbb{C}P^n, d_{ell}^H)$ is called *n*-dimensional *Hermitian elliptic space* (cf. **Fubini–Study metric** in Chap. 7).

• Elliptic plane metric

The elliptic plane metric is the elliptic metric on the elliptic plane $\mathbb{R}P^2$. If $\mathbb{R}P^2$ is viewed as the *Poincaré sphere* (i.e., a sphere in \mathbb{R}^3 with identified diametrically-opposite points) of diameter 1 tangent to the extended complex plane $\overline{\mathbb{C}} = \mathbb{C} \cup \{\infty\}$ at the point z = 0, then, under the stereographic projection from the "north pole" (0,0,1), $\overline{\mathbb{C}}$ with identified points z and $-\frac{1}{z}$ is a model of the elliptic plane, and the elliptic plane metric d_{ell} on it is defined by its line element $ds^2 = \frac{|dz|^2}{(1+|z|^2)^2}$.

• Pseudo-elliptic distance

The **pseudo-elliptic distance** (or *elliptic pseudo-distance*) dp_{ell} is a distance on the extended complex plane $\overline{\mathbb{C}} = \mathbb{C} \cup \{\infty\}$, with identified points z and $-\frac{1}{z}$ defined by

$$\left| \frac{z-u}{1+\overline{z}u} \right|$$
.

In fact, $dp_{ell}(z, u) = \tan d_{ell}(z, u)$, where d_{ell} is the **elliptic plane metric**.

• Hyperbolic metric

Let $\mathbb{R}P^n$ be the *n*-dimensional real projective space. Let, for any $x = (x_1 : \ldots : x_{n+1})$ and $y = (y_1 : \ldots : y_{n+1}) \in \mathbb{R}P^n$, the scalar product $\langle x, y \rangle = -x_1y_1 + \sum_{i=2}^{n+1} x_iy_i$ be considered.

The **hyperbolic metric** d_{hyp} is a metric on the set $H^n = \{x \in \mathbb{R}P^n : \langle x, x \rangle < 0\}$ defined by

$$r$$
arccosh $\frac{|\langle x, y \rangle|}{\sqrt{\langle x, x \rangle \langle y, y \rangle}}$,

where r is a fixed positive constant, and arccosh denotes the non-negative values of the inverse hyperbolic cosine. In this construction, the points of H^n can be viewed as the one-spaces of the *pseudo-Euclidean space* $\mathbb{R}^{n,1}$ inside the cone $C = \{x \in \mathbb{R}^{n,1} : \langle x, x \rangle = 0\}$.

The metric space (H^n, d_{hyp}) is called *n-dimensional hyperbolic space*. It is a model of *n*-dimensional *Hyperbolic Geometry*. It is the space of curvature $-1/r^2$, and r is the radius of curvature. Replacement of r by ir transforms all metric formulas of Hyperbolic Geometry into the corresponding formulas of Elliptic Geometry. As $r \to \infty$, both systems yield formulas of Euclidean Geometry (or become meaningless).

If H^n is viewed as the set $\{x \in \mathbb{R}^n : \sum_{i=1}^n x_i^2 < K\}$, where K > 1 is an arbitrary fixed constant, the hyperbolic metric can be written as

$$\frac{r}{2}\ln\frac{1+\sqrt{1-\gamma(x,y)}}{1-\sqrt{1-\gamma(x,y)}},$$

where $\gamma(x,y) = \frac{(K-\sum_{i=1}^n x_i^2)(K-\sum_{i=1}^n y_i^2)}{(K-\sum_{i=1}^n x_i y_i)^2}$, and r is a positive number with $\tanh \frac{1}{r} = \frac{1}{\sqrt{K}}$.

If H^n is viewed as a submanifold of the (n+1)-dimensional pseudo-Euclidean space $\mathbb{R}^{n,1}$ with the scalar product $\langle x,y\rangle = -x_1y_1 + \sum_{i=2}^{n+1} x_iy_i$ (in fact, as the top sheet $\{x \in \mathbb{R}^{n,1} : \langle x, x \rangle = -1, x_1 > 0\}$ of the twosheeted hyperboloid of revolution), then the hyperbolic metric on H^n is induced from the **pseudo-Riemannian metric** on $\mathbb{R}^{n,1}$ (cf. **Lorentz** metric in Chap. 26).

An n-dimensional hyperbolic space is a Riemannian space of constant negative curvature. It is the only such space which is **complete** and topologically equivalent to an Euclidean space. (See, for example, [Blum70], [Buse55].)

• Hermitian hyperbolic metric

Let $\mathbb{C}P^n$ be the *n*-dimensional complex projective space. Let, for any x= $(x_1:\ldots:x_{n+1})$ and $y=(y_1:\ldots:y_{n+1})\in\mathbb{C}P^n$, the scalar product $\langle x,y\rangle = -\overline{x}_1y_1 + \sum_{i=2}^{n+1} \overline{x}_iy_i$ be considered. The **Hermitian hyperbolic metric** d_{hyp}^H (see, for example, [Buse55])

is a metric on the set $\mathbb{C}H^n = \{x \in \mathbb{C}P^n : \langle x, x \rangle < 0\}$ defined by

$$r$$
arccosh $\frac{|\langle x, y \rangle|}{\sqrt{\langle x, x \rangle \langle y, y \rangle}}$,

where r is a fixed positive constant, and arccosh denotes the non-negative values of the inverse hyperbolic cosine.

The metric space $(\mathbb{C}H^n, d_{hyp}^H)$ is called *n*-dimensional Hermitian hyperbolic space.

• Poincaré metric

The **Poincaré metric** d_P is the **hyperbolic metric** for the *Poincaré disk* model (or conformal disk model) of Hyperbolic Geometry. In this model every point of the unit disk $\Delta = \{z \in \mathbb{C} : |z| < 1\}$ is called a hyperbolic point, the disk Δ itself is called the hyperbolic plane, circular arcs (and diameters) in Δ which are orthogonal to the absolute $\Omega = \{z \in \mathbb{C} : |z| = 1\}$ are called hyperbolic straight lines. Every point of Ω is called an ideal point. The angular measurements in this model are the same as in Hyperbolic Geometry. The Poincaré metric on Δ is defined by its line element

$$ds^{2} = \frac{|dz|^{2}}{(1 - |z|^{2})^{2}} = \frac{dz_{1}^{2} + dz_{2}^{2}}{(1 - z_{1}^{2} - z_{2}^{2})^{2}}.$$

The distance between two points z and u of Δ can be written as

$$\frac{1}{2}\ln\frac{|1-z\overline{u}|+|z-u|}{|1-z\overline{u}|-|z-u|}=\operatorname{arctanh}\frac{|z-u|}{|1-z\overline{u}|}.$$

In terms of *cross-ratio*, it is equal to

$$\frac{1}{2}\ln(z, u, z^*, u^*) = \frac{1}{2}\ln\frac{(z^* - z)(u^* - u)}{(z^* - u)(u^* - z)},$$

where z^* and u^* are the points of intersection of the hyperbolic straight line passing through z and u with Ω , z^* on the side of u, and u^* on the side of z.

In the Poincaré half-plane model of Hyperbolic Geometry the hyperbolic plane is the upper half-plane $H^2 = \{z \in \mathbb{C} : z_2 > 0\}$, and the hyperbolic lines are semi-circles and half-lines which are orthogonal to the real axis. The absolute (i.e., the set of ideal points) is the real axis together with the point at infinity. The angular measurements in the model are the same as in Hyperbolic Geometry.

The line element of the **Poincaré metric** on H^2 is given by

$$ds^{2} = \frac{|dz|^{2}}{(\Im z)^{2}} = \frac{dz_{1}^{2} + dz_{2}^{2}}{z_{2}^{2}}.$$

The distance between two points z, u can be written as

$$\frac{1}{2}\ln\frac{|z-\overline{u}|+|z-u|}{|z-\overline{u}|-|z-u|}=\operatorname{arctanh}\frac{|z-u|}{|z-\overline{u}|}.$$

In terms of cross-ratio, it is equal to

$$\frac{1}{2}\ln(z, u, z^*, u^*) = \frac{1}{2}\ln\frac{(z^* - z)(u^* - u)}{(z^* - u)(u^* - z)},$$

where z^* is the ideal point of the half-line emanating from z and passing through u, and u^* is the ideal point of the half-line emanating from u and passing through z.

In general, the **hyperbolic metric** in any domain $D \subset \mathbb{C}$ with at least three boundary points is defined as the preimage of the Poincaré metric in Δ under a conformal mapping $f: D \to \Delta$. Its line element has the form

$$ds^{2} = \frac{|f'(z)|^{2}|dz|^{2}}{(1 - |f(z)|^{2})^{2}}.$$

The distance between two points z and u in D can be written as

$$\frac{1}{2} \ln \frac{|1 - f(z)\overline{f(u)}| + |f(z) - f(u)|}{|1 - f(z)\overline{f(u)}| - |f(z) - f(u)|}.$$

Pseudo-hyperbolic distance

The pseudo-hyperbolic distance (or Gleason distance, hyperbolic pseudo-distance) dp_{hyp} is a metric on the unit disk $\Delta = \{z \in \mathbb{C} : |z| < 1\}$, defined by

 $\left| \frac{z-u}{1-\overline{z}u} \right|$.

In fact, $dp_{hyp}(z, u) = \tanh d_P(z, u)$, where d_P is the **Poincaré metric** on Δ .

• Cayley-Klein-Hilbert metric

The Cayley-Klein-Hilbert metric d_{CKH} is the hyperbolic metric for the Klein model (or projective disk model, Beltrami-Klein model) for Hyperbolic Geometry. In this model the hyperbolic plane is realized as the unit disk $\Delta = \{z \in \mathbb{C} : |z| < 1\}$, and the hyperbolic straight lines are realized as the chords of Δ . Every point of the absolute $\Omega = \{z \in \mathbb{C} : |z| = 1\}$ is called an ideal point. The angular measurements in this model are distorted. The Cayley-Klein-Hilbert metric on Δ is given by its metric tensor $((g_{ij}))$, i, j = 1, 2:

$$g_{11} = \frac{r^2(1-z_2^2)}{(1-z_1^2-z_2^2)^2}, \ g_{12} = \frac{r^2z_1z_2}{(1-z_1^2-z_2^2)^2}, \ g_{22} = \frac{r^2(1-z_1^2)}{(1-z_1^2-z_2^2)^2},$$

where r is an arbitrary positive constant. The distance between points z and u in Δ can be written as

$$r \operatorname{arccosh} \left(\frac{1 - z_1 u_1 - z_2 u_2}{\sqrt{1 - z_1^2 - z_2^2} \sqrt{1 - u_1^2 - u_2^2}} \right),$$

where arccosh denotes the non-negative values of the inverse hyperbolic cosine.

• Weierstrass metric

Given a real *n*-dimensional **inner product space** $(\mathbb{R}^n, \langle , \rangle)$, $n \geq 2$, the **Weierstrass metric** d_W is a metric on \mathbb{R}^n defined by

$$\operatorname{arccosh}(\sqrt{1+\langle x,x\rangle}\sqrt{1+\langle y,y\rangle}-\langle x,y\rangle),$$

where arccosh denotes the non-negative values of the inverse hyperbolic cosine.

Here, $(x, \sqrt{1 + \langle x, x \rangle}) \in \mathbb{R}^n \oplus \mathbb{R}$ are the Weierstrass coordinates of $x \in \mathbb{R}^n$, and the metric space (\mathbb{R}^n, d_W) can be identified with the Weierstrass model of Hyperbolic Geometry.

The Cayley–Klein–Hilbert metric $d_{CKH}(x,y) = \operatorname{arccosh} \frac{1-\langle x,y\rangle}{\sqrt{1-\langle x,x\rangle}\sqrt{1-\langle y,y\rangle}}$ on the open ball $B^n = \{x \in \mathbb{R}^n : \langle x,x\rangle < 1\}$ can be obtained from d_W by $d_{GKH}(x,y) = d_W(\mu(x),\mu(y))$, where $\mu: \mathbb{R}^n \to B^n$ is the Weierstrass mapping: $\mu(x) = \frac{x}{\sqrt{1-\langle x,x\rangle}}$.

• Harnack metric

Given a domain $D \subset \mathbb{R}^n$, $n \geq 2$, the **Harnack metric** is a metric on D defined by

$$\sup_{f} |\log \frac{f(x)}{f(y)}|,$$

where the supremum is taken over all positive functions which are harmonic on D.

• Quasi-hyperbolic metric

Given a domain $D \subset \mathbb{R}^n$, $n \geq 2$, the **quasi-hyperbolic metric** is a metric on D defined by

$$\inf_{\gamma \in \Gamma} \int_{\gamma} \frac{|dz|}{\rho(z)},$$

where the infimum is taken over the set Γ of all rectifiable curves connecting x and y in D, $\rho(z) = \inf_{u \in \partial D} ||z - u||_2$ is the distance between z and the boundary ∂D of D, and $||.||_2$ is the Euclidean norm on \mathbb{R}^n .

This metric is **Gromov hyperbolic** if the domain D is uniform, i.e., there exist constants C, C' such that each pair of points $x, y \in D$ can be joined by a rectifiable curve $\gamma = \gamma(x, y) \in D$ of length $l(\gamma)$ at most C|x-y|, and $\min\{l(\gamma(x,z)), l(\gamma(z,y))\} \leq C'd(z,\partial D)$ holds for all $z \in \gamma$. For n=2, one can define the **hyperbolic metric** on D by

$$\inf_{\gamma \in \Gamma} \int_{\gamma} \frac{2|f'(z)|}{1 - |f(z)|^2} |dz|,$$

where $f:D\to \Delta$ is any conformal mapping of D onto the $unit\ disk\ \Delta=\{z\in\mathbb{C}:|z|<1\}$. For $n\geq 3$, this metric is defined only for the half-hyperplane H^n and for the $open\ unit\ ball\ B^n$ as the infimum over all $\gamma\in\Gamma$ of the integrals $\int_{\gamma}\frac{|dz|}{z_n}$ and $\int_{\gamma}\frac{2|dz|}{1-||z||_2^2}$, respectively.

The quasi-hyperbolic metric is the **inner metric** (cf. Chap. 4) of the **Vuorinen metric**.

• Apollonian metric

Let $D \subset \mathbb{R}^n$, $D \neq \mathbb{R}^n$, be a *domain* such that the complement of D is not contained in a hyperplane or a sphere.

The **Apollonian metric** (or **Barbilian metric**, [Barb35]) is a metric on D defined by the cross-ratio in the following way:

$$\sup_{a,b \in \partial D} \ln \frac{||a-x||_2 ||b-y||_2}{||a-y||_2 ||b-x||_2},$$

where ∂D is the boundary of D, and $||.||_2$ is the Euclidean norm on \mathbb{R}^n . This metric is **Gromov hyperbolic**.

• Half-Apollonian metric

Given a domain $D \subset \mathbb{R}^n$, $D \neq \mathbb{R}^n$, the half-Apollonian metric η_D (Hästo and Lindén 2004) is a metric on D, defined by

$$\sup_{a \in \partial D} \left| \ln \frac{||a - y||_2}{||a - x||_2} \right|,$$

where ∂D is the boundary of D, and $||.||_2$ is the Euclidean norm on \mathbb{R}^n .

This metric is **Gromov hyperbolic** only if the domain is $\mathbb{R}^n \setminus \{x\}$, i.e., D has only one boundary point.

• Gehring metric

Given a domain $D \subset \mathbb{R}^n$, $D \neq \mathbb{R}^n$, the **Gehring metric** $\tilde{j_D}$ (Gehring 1982) is a metric on D, defined by

$$\frac{1}{2}\ln\left(\left(1+\frac{||x-y||_2}{\rho(x)}\right)\left(1+\frac{||x-y||_2}{\rho(y)}\right)\right),$$

where $||.||_2$ is the Euclidean norm on \mathbb{R}^n , and $\rho(x) = \inf_{u \in \partial D} ||x - u||_2$ is the distance between x and the boundary ∂D of D.

This metric is **Gromov hyperbolic**.

• Vuorinen metric

Given a domain $D \subset \mathbb{R}^n$, $D \neq \mathbb{R}^n$, the **Vuorinen metric** j_D (Vuorinen 1988) is a metric on D defined by

$$\ln\left(1+\frac{||x-y||_2}{\min\{\rho(x),\rho(y)\}}\right),\,$$

where $||.||_2$ is the Euclidean norm on \mathbb{R}^n , and $\rho(x) = \inf_{u \in \partial D} ||x - u||_2$ is the distance between x and the boundary ∂D of D.

This metric is **Gromov hyperbolic** only if the domain is $\mathbb{R}^n \setminus \{x\}$, i.e., D has only one boundary point.

• Ferrand metric

Given a domain $D \subset \mathbb{R}^n$, $D \neq \mathbb{R}^n$, the **Ferrand metric** σ_D (Ferrand 1987) is a metric on D defined by

$$\inf_{\gamma \in \Gamma} \int_{\gamma} \sup_{a,b \in \partial D} \frac{||a-b||_2}{||z-a||_2||z-b||_2} |dz|,$$

where the infimum is taken over the set Γ of all rectifiable curves connecting x and y in D, ∂D is the boundary of D, and $||.||_2$ is the Euclidean norm on \mathbb{R}^n .

This metric is **Gromov hyperbolic** if D is uniform, i.e., there exist constants C, C' such that each pair of points $x, y \in D$ can be joined by a rectifiable curve $\gamma \in D$ of length $l(\gamma)$ at most C|x-y|, and $\min\{l(\gamma(x,z)), l(\gamma(z,y))\} \leq C'd(z,\partial D)$ holds for all $z \in \gamma$.

The Ferrand metric is the inner metric (cf. Chap. 4) of the Seittenranta metric.

• Seittenranta metric

Given a domain $D \subset \mathbb{R}^n$, $D \neq \mathbb{R}^n$, the **Siettenranta metric** δ_D (Siettenranta 1999) is a metric on D defined by

$$\sup_{a,b\in\partial D} \ln \left(1 + \frac{||a-x||_2||b-y||_2}{||a-b||_2||x-y||_2} \right),$$

where ∂D is the boundary of D, and $||.||_2$ is the Euclidean norm on \mathbb{R}^n . This metric is **Gromov hyperbolic**.

• Modulus metric

Let $D \subset \mathbb{R}^n$, $D \neq \mathbb{R}^n$, be a *domain*, whose boundary ∂D has positive capacity.

The **modulus metric** (Gal 1960) μ_D (Gál 1960) is a metric on D, defined by

$$\inf_{C_{xy}} M(\Delta(C_{xy}, \partial D, D)),$$

where $M(\Gamma)$ is the conformal modulus of the curve family Γ , and C_{xy} is a continuum such that for some $\gamma:[0,1]\to D$ we have the following properties: $C_{xy}=\gamma([0,1]), \gamma(0)=x$, and $\gamma(1)=y$ (cf. **extremal metric** in Chap. 8).

This metric is **Gromov hyperbolic** if D is the open ball $B^n = \{x \in \mathbb{R}^n : \langle x, x \rangle < 1\}$ or a simply connected domain in \mathbb{R}^2 .

• Ferrand second metric

Let $D \subset \mathbb{R}^n$, $D \neq \mathbb{R}^n$, be a *domain* such that $|\mathbb{R}^n \setminus \{D\}| \geq 2$. The **Ferrand second metric** λ_D^* (Ferrand 1997) is a metric on D defined by

$$\left(\inf_{C_x,C_y} M(\Delta(C_x,C_y,D))\right)^{\frac{1}{1-n}},$$

where $M(\Gamma)$ is the conformal modulus of the curve family Γ , and C_z , z=x,y, is a continuum such that, for some $\gamma_z:[0,1]\to D$, $C_z=\gamma([0,1])$, $z\in |\gamma_z|$, and $\gamma_z(t)\to \partial D$ as $t\to 1$ (cf. **extremal metric** in Chap. 8).

This metric is **Gromov hyperbolic** if D is the open ball $B^n = \{x \in \mathbb{R}^n : \langle x, x \rangle < 1\}$ or a simply connected domain in \mathbb{R}^2 .

• Parabolic distance

The **parabolic distance** is a metric on \mathbb{R}^{n+1} , considered as $\mathbb{R}^n \times \mathbb{R}$ defined by

$$\sqrt{(x_1-y_1)^2+\cdots+(x_n-y_n)^2}+|t_x-t_y|^{1/m}, m \in \mathbb{N},$$

for any $x = (x_1, \dots, x_n, t_x), y = (y_1, \dots, y_n, t_y) \in \mathbb{R}^n \times \mathbb{R}$.

The space $\mathbb{R}^n \times \mathbb{R}$ can be interpreted as multidimensional *space-time*. Usually, the value m=2 is applied. There exist some variants of the parabolic distance, for example, the parabolic distance

$$\sup\{|x_1 - y_1|, |x_2 - y_2|^{1/2}\}\$$

on \mathbb{R}^2 (cf. also **Rickman's rug metric** in Chap. 19), or the **half-space parabolic distance** on $\mathbb{R}^3_+ = \{x \in \mathbb{R}^3 : x_1 \geq 0\}$ defined by

$$\frac{|x_1 - y_1| + |x_2 - y_2|}{\sqrt{x_1} + \sqrt{x_2} + \sqrt{|x_2 - y_2|}} + \sqrt{|x_3 - y_3|}.$$

Chapter 7 Riemannian and Hermitian Metrics

Riemannian Geometry is a multidimensional generalization of the intrinsic geometry of two-dimensional surfaces in the Euclidean space \mathbb{E}^3 . It studies real smooth manifolds equipped with Riemannian metrics, i.e., collections of positive-definite symmetric bilinear forms $((g_{ij}))$ on their tangent spaces which vary smoothly from point to point. The geometry of such (Riemannian) manifolds is based on the line element $ds^2 = \sum_{i,j} g_{ij} dx_i dx_j$. This gives, in particular, local notions of angle, length of curve, and volume. From these notions some other global quantities can be derived, by integrating local contributions. Thus, the value ds is interpreted as the length of the vector (dx_1, \ldots, dx_n) , and it is called the infinitesimal distance. The arc length of a curve γ is expressed by $\int_{\gamma} \sqrt{\sum_{i,j} g_{ij} dx_i dx_j}$, and then the intrinsic metric on a Riemannian manifold is defined as the infimum of lengths of curves joining two given points of the manifold.

Therefore, a Riemannian metric is not an ordinary metric, but it induces an ordinary metric, in fact, the intrinsic metric, sometimes called **Riemannian distance**, on any connected Riemannian manifold. A Riemannian metric is an infinitesimal form of the corresponding Riemannian distance.

As particular special cases of Riemannian Geometry, there occur *Euclidean Geometry* as well as the two standard types, *Elliptic Geometry* and *Hyperbolic Geometry*, of *Non-Euclidean Geometry*.

If the bilinear forms $((g_{ij}))$ are non-degenerate but indefinite, one obtains $Pseudo-Riemannian\ Geometry$. In the case of dimension four (and $signature\ (1,3)$) it is the main object of the General Theory of Relativity. If $ds = F(x_1, \ldots, x_n, dx_1, \ldots, dx_n)$, where F is a real positive-definite convex function which can not be given as the square root of a symmetric bilinear form (as in the Riemannian case), one obtains the $Finsler\ Geometry$ generalizing Riemannian Geometry.

Hermitian Geometry studies complex manifolds equipped with Hermitian metrics, i.e., collections of positive-definite symmetric sesquilinear forms (or $\frac{3}{2}$ -linear forms) since they are linear in one argument and antilinear in the other) on their tangent spaces, which vary smoothly from point to point. It is a complex analog of Riemannian Geometry. A special class of Hermitian

metrics form **Kähler metrics** which have a closed fundamental form w. A generalization of Hermitian metrics give **complex Finsler metrics** which can not be written in terms of a bilinear symmetric positive-definite sesqulinear form.

7.1 Riemannian metrics and generalizations

A real n-dimensional manifold M^n with boundary is a **Hausdorff space** in which every point has an open neighborhood homeomorphic to either an open subset of \mathbb{E}^n , or an open subset of the closed half of \mathbb{E}^n . The set of points which have an open neighborhood homeomorphic to \mathbb{E}^n is called the *interior* (of the manifold); it is always non-empty. The complement of the interior is called the boundary (of the manifold); it is an (n-1)-dimensional manifold. If the boundary of M^n is empty, one obtains a real n-dimensional manifold without boundary.

A manifold without boundary is called *closed* if it is compact, and *open* otherwise.

An open set of M^n together with a homeomorphism between the open set and an open set of \mathbb{E}^n is called a *coordinate chart*. A collection of charts which cover M^n is called an *atlas* on M^n . The homeomorphisms of two overlapping charts provide a transition mapping from a subset of \mathbb{E}^n to some other subset of \mathbb{E}^n . If all these mappings are continuously differentiable, then M^n is called a *differentiable manifold*. If all the connecting mappings are k times continuously differentiable, then the manifold is called a C^k manifold; if they are infinitely often differentiable, then the manifold is called a *smooth manifold* (or C^{∞} manifold).

An atlas of a manifold is called *oriented* if the coordinate transformations between charts are all positive, i.e., the Jacobians of the coordinate transformations between any two charts are positive at every point. An *orientable manifold* is a manifold admitting an oriented atlas.

Manifolds inherit many local properties of the Euclidean space. In particular, they are locally path-connected, locally compact, and locally metrizable. Every smooth Riemannian manifold embedds isometrically (Nash 1956) in some finite-dimensional Euclidean space.

Associated with every point on a differentiable manifold is a tangent space and its dual, a cotangent space. Formally, let M^n be a C^k manifold, $k \geq 1$, and p a point of M^n . Fix a chart $\varphi: U \to \mathbb{E}^n$, where U is an open subset of M^n containing p. Suppose that two curves $\gamma^1: (-1,1) \to M^n$ and $\gamma^2: (-1,1) \to M^n$ with $\gamma^1(0) = \gamma^2(0) = p$ are given such that $\varphi \cdot \gamma^1$ and $\varphi \cdot \gamma^2$ are both differentiable at 0. Then γ^1 and γ^2 are called tangent at 0 if the ordinary derivatives of $\varphi \cdot \gamma^1$ and $\varphi \cdot \gamma^2$ coincide at 0: $(\varphi \cdot \gamma^1)'(0) = (\varphi \cdot \gamma^2)'(0)$. If the functions $\varphi \cdot \gamma^i: (-1,1) \to \mathbb{E}^n$, i=1,2, are given by n real-valued component functions $(\varphi \cdot \gamma^i)_1(t), \ldots, (\varphi \cdot \gamma^i)_n(t)$, the condition above means that their

Jacobians $\left(\frac{d(\varphi \cdot \gamma^i)_1(t)}{dt}, \dots, \frac{d(\varphi \cdot \gamma^i)_n(t)}{dt}\right)$ coincide at 0. This is an equivalence relation, and the equivalence class $\gamma'(0)$ of the curve γ is called a *tangent vector* of M^n at p. The *tangent space* $T_p(M^n)$ of M^n at p is defined as the set of all tangent vectors at p. The function $(d\varphi)_p: T_p(M^n) \to \mathbb{E}^n$ defined by $(d\varphi)_p(\gamma'(0)) = (\varphi \cdot \gamma)'(0)$, is bijective and can be used to transfer the vector space operations from \mathbb{E}^n over to $T_p(M^n)$.

All the tangent spaces $T_p(M^n)$, $p \in M^n$, when "glued together," form the tangent bundle $T(M^n)$ of M^n . Any element of $T(M^n)$ is a pair (p,v), where $v \in T_p(M^n)$. If for an open neighborhood U of p the function $\varphi : U \to \mathbb{R}^n$ is a coordinate chart, then the preimage V of U in $T(M^n)$ admits a mapping $\psi : V \to \mathbb{R}^n \times \mathbb{R}^n$ defined by $\psi(p,v) = (\varphi(p), d\varphi(p))$. It defines the structure of a smooth 2n-dimensional manifold on $T(M^n)$. The cotangent bundle $T^*(M^n)$ of M^n is obtained in similar manner using cotangent spaces $T_p^*(M^n)$, $p \in M^n$.

A vector field on a manifold M^n is a section of its tangent bundle $T(M^n)$, i.e., a smooth function $f: M^n \to T(M^n)$ which assigns to every point $p \in M^n$ a vector $v \in T_p(M^n)$.

A connection (or covariant derivative) is a way of specifying a derivative of a vector field along another vector field on a manifold. Formally, the covariant derivative ∇ of a vector u (defined at a point $p \in M^n$) in the direction of the vector v (defined at the same point p) is a rule that defines a third vector at p, called $\nabla_v u$, which has the properties of a derivative. A Riemannian metric uniquely defines a special covariant derivative called the Levi-Civita connection. It is the torsion-free connection ∇ of the tangent bundle, preserving the given Riemannian metric.

The Riemann curvature tensor R is the standard way to express the curvature of Riemannian manifolds. The Riemann curvature tensor can be given in terms of the Levi–Civita connection ∇ by the following formula:

$$R(u, v)w = \nabla_u \nabla_v w - \nabla_v \nabla_u w - \nabla_{[u, v]} w,$$

where R(u,v) is a linear transformation of the tangent space of the manifold M^n ; it is linear in each argument. If $u=\frac{\partial}{\partial x_i}$ and $v=\frac{\partial}{\partial x_j}$ are coordinate vector fields, then [u,v]=0, and the formula simplifies to $R(u,v)w=\nabla_u\nabla_v w-\nabla_v\nabla_u w$, i.e., the curvature tensor measures anti-commutativity of the covariant derivative. The linear transformation $w\to R(u,v)w$ is also called the *curvature transformation*.

The Ricci curvature tensor (or Ricci curvature) Ric is obtained as the trace of the full curvature tensor R. It can be thought of as a Laplacian of the Riemannian metric tensor in the case of Riemannian manifolds. Ricci curvature is a linear operator on the tangent space at a point. Given an orthonormal basis $(e_i)_i$ in the tangent space $T_p(M^n)$, we have

$$Ric(u) = \sum_{i} R(u, e_i)e_i.$$

The value of Ric(u) does not depend on the choice of an orthonormal basis. Starting with dimension four, the Ricci curvature does not describe the curvature tensor completely.

The Ricci scalar (or scalar curvature) Sc of a Riemannian manifold M^n is the full trace of the curvature tensor; given an orthonormal basis $(e_i)_i$ at $p \in M^n$, we have

$$Sc = \sum_{i,j} \langle R(e_i, e_j)e_j, e_i \rangle = \sum_i \langle Ric(e_i), e_i \rangle.$$

The sectional curvature $K(\sigma)$ of a Riemannian manifold M^n is defined as the Gauss curvature of an σ -section at a point $p \in M^n$. Here, given a 2-plane σ in the tangent space $T_p(M^n)$, a σ -section is a locally-defined piece of surface which has the plane σ as a tangent plane at p, obtained from geodesics which start at p in the directions of the image of σ under the exponential mapping.

• Metric tensor

The **metric tensor** (or basic tensor, fundamental tensor) is a symmetric tensor of rank 2, that is used to measure distances and angles in a real n-dimensional differentiable manifold M^n . Once a local coordinate system $(x_i)_i$ is chosen, the metric tensor appears as a real symmetric $n \times n$ matrix $((g_{ij}))$.

The assignment of a metric tensor on an n-dimensional differentiable manifold M^n introduces a scalar product (i.e., symmetric bilinear, but in general not positive-definite, form) \langle,\rangle_p on the tangent space $T_p(M^n)$ at any point $p \in M^n$ defined by

$$\langle x, y \rangle_p = g_p(x, y) = \sum_{i,j} g_{ij}(p) x_i y_j,$$

where $g_{ij}(p)$ is a value of the metric tensor at the point $p \in M^n$, and $x = (x_1, \ldots, x_n)$, $y = (y_1, \ldots, y_n) \in T_p(M^n)$. The collection of all these scalar products is called the **metric** g with the metric tensor $((g_{ij}))$. The length ds of the vector (dx_1, \ldots, dx_n) is expressed by the quadratic differential form

$$ds^2 = \sum_{i,j} g_{ij} dx_i dx_j,$$

which is called the *line element* (or *first fundamental form*) of the metric g. The *length* of a curve γ is expressed by the formula $\int_{\gamma} \sqrt{\sum_{i,j} g_{ij} dx_i dx_j}$. In general it may be real, purely imaginary, or zero (an *isotropic curve*).

The signature of a metric tensor is the pair (p,q) of positive (p) and negative (q) eigenvalues of the matrix $((g_{ij}))$. The signature is said to be indefinite if both p and q are non-zero, and positive-definite if q = 0.

A Riemannian metric is a metric g with a positive-definite signature (p,0), and a pseudo-Riemanian metric is a metric g with an indefinite signature (p,q).

• Non-degenerate metric

A non-degenerate metric is a metric g with the metric tensor $((g_{ij}))$, for which the *metric discriminant* $det((g_{ij})) \neq 0$. All Riemannian and pseudo-Riemannian metrics are non-degenerate.

A degenerate metric is a metric g with the metric tensor $((g_{ij}))$ for which the metric discriminant $\det((g_{ij})) = 0$ (cf. semi-Riemannian metric and semi-pseudo-Riemannian metric). A manifold with a degenerate metric is called an *isotropic manifold*.

• Diagonal metric

A diagonal metric is a metric g with a metric tensor $((g_{ij}))$ which is zero for $i \neq j$. The Euclidean metric is a diagonal metric, as its metric tensor has the form $g_{ii} = 1, g_{ij} = 0$ for $i \neq j$.

• Riemannian metric

Consider a real n-dimensional differentiable manifold M^n in which each tangent space is equipped with an $inner\ product$ (i.e., a symmetric positive-definite bilinear form) which varies smoothly from point to point.

A **Riemannian metric** on M^n is a collection of inner products \langle , \rangle_p on the tangent spaces $T_p(M^n)$, one for each $p \in M^n$.

Every inner product \langle , \rangle_p is completely defined by inner products $\langle e_i, e_j \rangle_p = g_{ij}(p)$ of elements e_1, \ldots, e_n of a standard basis in \mathbb{E}^n , i.e., by the real symmetric and positive-definite $n \times n$ matrix $((g_{ij})) = ((g_{ij}(p)))$, called a **metric tensor**. In fact, $\langle x, y \rangle_p = \sum_{i,j} g_{ij}(p) x_i y_j$, where $x = (x_1, \ldots, x_n)$ and $y = (y_1, \ldots, y_n) \in T_p(M^n)$. The smooth function g completely determines the Riemannian metric.

A Riemannian metric on M^n is not an ordinary metric on M^n . However, for a connected manifold M^n , every Riemannian metric on M^n induces an ordinary metric on M^n , in fact, the **intrinsic metric** of M^n ; for any points $p, q \in M^n$ the **Riemannian distance** between them is defined as

$$\inf_{\gamma} \int_{0}^{1} \langle \frac{d\gamma}{dt}, \frac{d\gamma}{dt} \rangle^{\frac{1}{2}} dt = \inf_{\gamma} \int_{0}^{1} \sqrt{\sum_{i,j} g_{ij} \frac{dx_{i}}{dt} \frac{dx_{j}}{dt}} dt,$$

where the infimum is taken over all rectifiable curves $\gamma:[0,1]\to M^n,$ connecting p and q.

A Riemannian manifold (or Riemannian space) is a real n-dimensional differentiable manifold M^n equipped with a Riemannian metric. The theory of Riemannian spaces is called Riemannian Geometry. The simplest examples of Riemannian spaces are Euclidean spaces, hyperbolic spaces, and elliptic spaces. A Riemannian space is called complete if it is a complete metric space.

• Conformal metric

A conformal structure on a vector space V is a class of pairwise-homothetic Euclidean metrics on V. Any Euclidean metric d_E on V defines a conformal structure $\{\lambda d_E : \lambda > 0\}$.

A conformal structure on a manifold is a field of conformal structures on the tangent spaces or, equivalently, a class of conformally equivalent Riemannian metrics. Two Riemannian metrics g and h on a smooth manifold M^n are called conformally equivalent if $g = f \cdot h$ for some positive function f on M^n , called a conformal factor.

A **conformal metric** is a Riemannian metric that represents the conformal structure (cf. **conformally invariant metric** in Chap. 8).

Conformal space

The **conformal space** (or *inversive space*) is the Euclidean space \mathbb{E}^n extended by an ideal point (at infinity). Under *conformal* transformations, i.e., continuous transformations preserving local angles, the ideal point can be taken to be an ordinary point. Therefore, in a conformal space a sphere is indistinguishable from a plane: a plane is a sphere passing through the ideal point.

Conformal spaces are considered in Conformal Geometry (or Angle-Preserving Geometry, Möbius geometry, Inversive Geometry) in which properties of figures are studied that are invariant under conformal transformations. It is the set of transformations that map spheres into spheres, i.e., generated by the Euclidean transformations together with inversions which in coordinate form are conjugate to $x_i \to \frac{r^2 x_i}{\sum_j x_j^2}$, where r is the radius of the inversion. An inversion in a sphere becomes an everywhere well-defined automorphism of period two. Any angle inverts into an equal angle.

The two-dimensional conformal space is the *Riemann sphere*, on which the conformal transformations are given by the *Möbius transformations* $z \to \frac{az+b}{cz+d}$, $ad-bc \neq 0$.

In general, a *conformal mapping* between two Riemannian manifolds is a diffeomorphism between them such that the pulled back metric is *conformally equivalent* to the original one. A **conformal Euclidean space** is a *Riemannian space* admitting a conformal mapping onto an Euclidean space.

In the General Theory of Relativity, conformal transformations are considered on the $Minkowski\ space\ \mathbb{R}^{1,3}$ extended by two ideal points.

• Space of constant curvature

A space of constant curvature is a Riemannian space M^n for which the sectional curvature $K(\sigma)$ is constant in all two-dimensional directions σ .

A **space form** is a connected complete space of constant curvature. A **flat space** is a space form of zero curvature.

The Euclidean space and the flat torus are space forms of zero curvature (i.e., flat spaces), the sphere is a space form of positive curvature, the *hyperbolic space* is a space form of negative curvature.

• Generalized Riemannian spaces

A generalized Riemannian space is a metric space with the intrinsic metric, subject to certain restrictions on the curvature. Such spaces include spaces of bounded curvature, Riemannian spaces, etc. Generalized Riemannian spaces differ from Riemannian spaces not only by greater generality, but also by the fact that they are defined and investigated on the basis of their metric alone, without coordinates.

A space of bounded curvature ($\leq k$ and $\geq k'$) is a generalized Riemannian space defined by the condition: for any sequence of geodesic triangles T_n contracting to a point we have

$$k \ge \overline{\lim} \frac{\overline{\delta}(T_n)}{\sigma(T_n^0)} \ge \underline{\lim} \frac{\overline{\delta}(T_n)}{\sigma(T_n^0)} \ge k',$$

where a geodesic triangle T=xyz is the triplet of geodesic segments [x,y], [y,z], [z,x] (the sides of T) connecting in pairs three different points x,y,z, $\overline{\delta}(T)=\alpha+\beta+\gamma-\pi$ is the excess of the geodesic triangle T, and $\sigma(T^0)$ is the area of a Euclidean triangle T^0 with the sides of the same lengths. The intrinsic metric on the space of bounded curvature is called a metric of bounded curvature.

Such a space turns out to be Riemannian under two additional conditions: local compactness of the space (this ensures the condition of local existence of geodesics), and local extendibility of geodesics. If in this case k = k', it is a Riemannian space of constant curvature k (cf. space of geodesics in Chap. 6).

A space of curvature $\leq k$ is defined by the condition $\overline{\lim} \frac{\overline{\delta}(T_n)}{\sigma(T_n^0)} \leq k$. In such a space any point has a neighborhood in which the sum $\alpha + \beta + \gamma$ of the angles of a geodesic triangle T does not exceed the sum $\alpha_k + \beta_k + \gamma_k$ of the angles of a triangle T^k with sides of the same lengths in a space of constant curvature k. The intrinsic metric of such space is called a k-concave metric.

A space of curvature \geq k is defined by the condition $\varliminf \frac{\overline{\delta}(T_n)}{\sigma(T_n^0)} \geq k$. In such a space any point has a neighborhood in which $\alpha + \beta + \gamma \geq \alpha_k + \beta_k + \gamma_k$ for triangles T and T^k . The intrinsic metric of such space is called a K-concave metric.

An Alexandrov space is a generalized Riemannian space with upper, lower or integral curvature bounds.

• Complete Riemannian metric

A Riemannian metric g on a manifold M^n is called **complete** if M^n forms a complete metric space with respect to g. Any Riemannian metric on a compact manifold is complete.

• Ricci-flat metric

A **Ricci-flat metric** is a Riemannian metric with vanished Ricci curvature tensor.

A Ricci-flat manifold is a Riemannian manifold equipped with a Ricci-flat metric. Ricci-flat manifolds represent vacuum solutions to the Einstein field equation, and are special cases of $K\ddot{a}hler$ -Einstein manifolds. Important Ricci-flat manifolds are Calabi-Yau manifolds, and hyper- $K\ddot{a}hler$ manifolds.

• Osserman metric

An **Osserman metric** is a Riemannian metric for which the Riemannian curvature tensor R is Osserman. It means that the eigenvalues of the Jacobi operator $\mathcal{J}(x): y \to R(y,x)x$ are constant on the unit sphere S^{n-1} in \mathbb{E}^n , i.e., they are independent of the unit vectors x.

• G-invariant metric

A G-invariant metric is a Riemannian metric g on a differentiable manifold M^n , that does not change under any of the transformations of a given $Lie\ group\ (G,\cdot,id)$ of transformations. The group (G,\cdot,id) is called the $group\ of\ motions$ (or $group\ of\ isometries$) of the Riemannian space (M^n,g) .

• Ivanov-Petrova metric

Let R be the Riemannian curvature tensor of a Riemannian manifold M^n , and let $\{x,y\}$ be an orthogonal basis for an oriented 2-plane π in the tangent space $T_p(M^n)$ at a point p of M^n .

The **Ivanov–Petrova metric** is a Riemannian metric on M^n for which the eigenvalues of the antisymmetric curvature operator $\mathcal{R}(\pi) = R(x,y)$ [IvSt95] depend only on the point p of a Riemannian manifold M^n , but not upon the plane π .

• Zoll metric

A **Zoll metric** is a Riemannian metric on a smooth manifold M^n whose geodesics are all simple closed curves of an equal length. A two-dimensional sphere S^2 admits many such metrics, besides the obvious metrics of constant curvature. In terms of cylindrical coordinates (z,θ) $(z \in [-1,1], \theta \in [0,2\pi])$, the *line element*

$$ds^{2} = \frac{(1+f(z))^{2}}{1-z^{2}}dz^{2} + (1-z^{2})d\theta^{2}$$

defines a Zoll metric on S^2 for any smooth odd function $f:[-1,1] \to (-1,1)$ which vanishes at the end points of the interval.

• Cycloidal metric

The **cycloidal metric** is a Riemannian metric on the half-plane $\mathbb{R}^2_+ = \{x \in \mathbb{R}^2 : x_1 \geq 0\}$ defined by the *line element*

$$ds^2 = \frac{dx_1^2 + dx_2^2}{2x_1}.$$

It is called *cycloidal* because its geodesics are cycloid curves. The corresponding distance d(x,y) between two points $x,y\in\mathbb{R}^2_+$ is equivalent to the distance

$$\rho(x,y) = \frac{|x_1 - y_1| + |x_2 - y_2|}{\sqrt{x_1} + \sqrt{x_2} + \sqrt{|x_2 - y_2|}}$$

in the sense that $d \leq C\rho$, and $\rho \leq Cd$ for some positive constant C.

• Berger metric

The **Berger metric** is a Riemannian metric on the *Berger sphere* (i.e., the three-sphere S^3 squashed in one direction) defined by the *line element*

$$ds^{2} = d\theta^{2} + \sin^{2}\theta d\phi^{2} + \cos^{2}\alpha (d\psi + \cos\theta d\phi)^{2},$$

where α is a constant, and θ , ϕ , ψ are Euler angles.

• Carnot-Carathéodory metric

A distribution (or polarization) on a manifold M^n is a subbundle of the tangent bundle $T(M^n)$ of M^n . Given a distribution $H(M^n)$, a vector field in $H(M^n)$ is called horizontal. A curve γ on M^n is called horizontal (or distinguished, admissible) with respect to $H(M^n)$ if $\gamma'(t) \in H_{\gamma(t)}(M^n)$ for any t.

A distribution $H(M^n)$ is called *completely non-integrable* if the Lie brackets of $H(M^n)$, i.e., $[\cdots, [H(M^n), H(M^n)]]$, span the tangent bundle $T(M^n)$, i.e., for all $p \in M^n$ any tangent vector v from $T_p(M^n)$ can be presented as a linear combination of vectors of the following types: $u, [u, w], [u, [w, t]], [u, [w, t, s]], \cdots \in T_p(M^n)$, where all vector fields u, w, t, s, \ldots are horizontal

The Carnot-Carathéodory metric (or CC metric, sub-Riemannian metric, control metric) is a metric on a manifold M^n with a completely non-integrable horizontal distribution $H(M^n)$ defined as the section g_C of positive-definite scalar products on $H(M^n)$. The distance $d_C(p,q)$ between any points $p, q \in M^n$ is defined as the infimum of the g_C -lengths of the horizontal curves joining p and q.

A sub-Riemannian manifold (or polarized manifold) is a manifold M^n equipped with a Carnot–Carathéodory metric. It is a generalization of a Riemannian manifold. Roughly, in order to measure distances in a sub-Riemannian manifold, one is allowed to go only along curves tangent to horizontal spaces.

• Pseudo-Riemannian metric

Consider a real *n*-dimensional differentiable manifold M^n in which every tangent space $T_p(M^n)$, $p \in M^n$, is equipped with a *scalar product* which varies smoothly from point to point and is non-degenerate, but indefinite.

A **pseudo-Riemannian metric** on M^n is a collection of scalar products \langle , \rangle_p on the tangent spaces $T_p(M^n)$, $p \in M^n$, one for each $p \in M^n$.

Every scalar product \langle , \rangle_p is completely defined by scalar products $\langle e_i, e_j \rangle_p = g_{ij}(p)$ of elements e_1, \ldots, e_n of a standard basis in \mathbb{E}^n , i.e., by the real symmetric indefinite $n \times n$ matrix $((g_{ij})) = ((g_{ij}(p)))$, called a **metric tensor** (cf. **Riemannian metric** in which case this tensor is not only non-degenerate but, moreover, positive-definite). In fact,

 $\langle x,y\rangle_p = \sum_{i,j} g_{ij}(p)x_iy_j$, where $x=(x_1,\ldots,x_n)$ and $y=(y_1,\ldots,y_n)\in T_p(M^n)$. The smooth function g determines the pseudo-Riemannian metric.

The length ds of the vector (dx_1, \ldots, dx_n) is given by the quadratic differential form

$$ds^2 = \sum_{i,j} g_{ij} dx_i dx_j.$$

The length of a curve $\gamma:[0,1]\to M^n$ is expressed by the formula $\int_{\gamma}\sqrt{\sum_{i,j}g_{ij}dx_idx_j}=\int_0^1\sqrt{\sum_{i,j}g_{ij}\frac{dx_i}{dt}\frac{dx_j}{dt}}dt$. In general it may be real, purely imaginary or zero (an *isotropic curve*).

A pseudo-Riemannian metric on M^n is a metric with a fixed, but indefinite signature (p,q), p+q=n. A pseudo-Riemannian metric is non-degenerate, i.e., its metric discriminant $\det((g_{ij})) \neq 0$. Therefore, it is a non-degenerate indefinite metric.

A pseudo-Riemannian manifold (or pseudo-Riemannian space) is a real n-dimensional differentiable manifold M^n equipped with a pseudo-Riemannian metric. The theory of pseudo-Riemannian spaces is called $Pseudo-Riemannian\ Geometry$.

• Pseudo-Euclidean distance

The model space of a **pseudo-Riemannian space** of signature (p,q) is the *pseudo-Euclidean space* $\mathbb{R}^{p,q}$, p+q=n, which is a real *n*-dimensional vector space \mathbb{R}^n equipped with the metric tensor $((g_{ij}))$ of signature (p,q) defined by $g_{11} = \cdots = g_{pp} = 1$, $g_{p+1,p+1} = \cdots = g_{nn} = -1$, $g_{ij} = 0$ for $i \neq j$. The *line element* of the corresponding metric is given by

$$ds^{2} = dx_{1}^{2} + \dots + dx_{p}^{2} - dx_{p+1}^{2} - \dots - dx_{n}^{2}.$$

The **pseudo-Euclidean distance** of signature (p, q = n - p) on \mathbb{R}^n is defined, for $x, y \in \mathbb{R}^n$, by

$$d_{pE}^{2}(x,y) = \sum_{i=1}^{p} (x_{i} - y_{i})^{2} - \sum_{i=p+1}^{n} (x_{i} - y_{i})^{2}.$$

Such a pseudo-Euclidean space can be seen as $\mathbb{R}^p \times i\mathbb{R}^q$, where $i = \sqrt{-1}$. The pseudo-Euclidean space with (p,q) = (1,3) is used as space-time model of Special Relativity; cf. **Minkowsky metric** in Chap. 26. The points correspond to *events*; the line spanned by events x and y is *space-like* if d(x,y) > 0 and time-like if d(x,y) < 0. If d(x,y) > 0, then $\sqrt{d(x,y)}$ is Euclidean distance and if d(x,y) < 0, then $\sqrt{|d(x,y)|}$ is the life time of a particle (from x to y).

The general quadratic-form distance for two points $x, y \in \mathbb{R}^n$, is defined by $\sqrt{(x-y)^T A(x-y)}$, where A is a real non-singular symmetric $n \times n$ matrix; cf. Mahalonobis distance in Chap. 17. The pseudo-Euclidean distance of signature (p, q = n - p) is the case $A = diag(a_i)$ with $a_i = 1$ for $1 \le i \le p$ and $a_i = -1$ for $p + 1 \le i \le n$.

• Lorentzian metric

A Lorentzian metric (or Lorentz metric) is a pseudo-Riemannian metric of signature (1, p).

A Lorentzian manifold is a manifold equipped with a Lorentzian metric. The Minkowski space $\mathbb{R}^{1,p}$ with the flat Minkowski metric is a model of it, in the same way as Riemannian manifolds can be modeled on Euclidean space.

• Osserman Lorentzian metric

An Osserman Lorentzian metric is a Lorentzian metric for which the Riemannian curvature tensor R is Osserman, i.e., the eigenvalues of the Jacobi operator $\mathcal{J}(x): y \to R(y,x)x$ are independent of the unit vectors x.

A Lorentzian manifold is Osserman if and only if it is of constant curvature.

• Blaschke metric

The **Blaschke metric** on a non-degenerate hypersurface is a pseudo-Riemannian metric, associated to the affine normal of the immersion $\phi: M^n \to \mathbb{R}^{n+1}$, where M^n is an *n*-dimensional manifold, and \mathbb{R}^{n+1} is considered as an affine space.

• Semi-Riemannian metric

A semi-Riemannian metric on a real n-dimensional differentiable manifold M^n is a degenerate Riemannian metric, i.e., a collection of positive-semi-definite scalar products $\langle x,y\rangle_p=\sum_{i,j}g_{ij}(p)x_iy_j$ on the tangent spaces $T_p(M^n),\ p\in M^n$; the metric discriminant $det((g_{ij}))=0$.

A semi-Riemannian manifold (or semi-Riemannian space) is a real n-dimensional differentiable manifold M^n equipped with a semi-Riemannian metric.

The model space of a semi-Riemannian manifold is the semi-Euclidean space R_d^n , $d \geq 1$ (sometimes denoted also by \mathbb{R}_{n-d}^n), i.e., a real n-dimensional vector space \mathbb{R}^n equipped with a semi-Riemannian metric. It means that there exists a scalar product of vectors such that, relative to a suitably chosen basis, the scalar product $\langle x, x \rangle$ of any vector with itself has the form $\langle x, x \rangle = \sum_{i=1}^{n-d} x_i^2$. The number $d \geq 1$ is called the defect (or deficiency) of the space.

• Grushin metric

The **Grushin metric** is a semi-Riemannian metric on \mathbb{R}^2 defined by the *line element*

$$ds^2 = dx_1^2 + \frac{dx_2^2}{x_1^2}.$$

• Agmon distance

Given a Schrödinger operator $H(h) = -h^2 \Delta + V(x)$ on $L_2(\mathbb{R}^d)$, where V is a potential and h is the Planck constant, consider a semi-Riemannian metric on \mathbb{R}^d with respect to the energy $E_0(h) = h^{-\alpha}e_0$ defined by the line element

$$ds^{2} = \max\{0, V(x) - E_{0}(h)\}dx^{2}.$$

Then the **Agmon distance** on \mathbb{R}^d is the corresponding Riemannian distance defined, for any $x, y \in \mathbb{R}^d$, by

$$\inf_{\gamma}\{\int_{0}^{1}\sqrt{\max\{V(\gamma(s))-E_{0}(h),0\}}\cdot|\gamma^{'}(s)|ds:\gamma(0)=x,\gamma(1)=y,\gamma\in C^{1}\}.$$

• Semi-pseudo-Riemannian metric

A semi-pseudo-Riemannian metric on a real n-dimensional differentiable manifold M^n is a degenerate pseudo-Riemannian metric, i.e., a collection of degenerate indefinite scalar products $\langle x,y\rangle_p = \sum_{i,j} g_{ij}(p)x_iy_j$ on the tangent spaces $T_p(M^n)$, $p \in M^n$; the metric discriminant $det((g_{ij})) = 0$. In fact, a semi-pseudo-Riemannian metric is a **degenerate indefinite** metric.

A semi-pseudo-Riemannian manifold (or semi-pseudo-Riemannian space) is a real n-dimensional differentiable manifold M^n equipped with a semi-pseudo-Riemannian metric.

The model space of a semi-pseudo-Riemannian manifold is the semi-pseudo-Euclidean space $\mathbb{R}^n_{l_1,\ldots,l_r}$, i.e., a real n-dimensional vector space \mathbb{R}^n equipped with a semi-pseudo-Riemannian metric. It means that there exist r scalar products $\langle x,y\rangle_a=\sum \epsilon_{i_a}x_{i_a}y_{i_a}$, where $a=1,\ldots r, \ 0=m_0< m_1<\cdots< m_r=n,\ i_a=m_{a-1}+1,\ldots m_a,\ \epsilon_{i_a}=\pm 1,\ \text{and}\ -1\ \text{occurs}\ l_a$ times among the numbers ϵ_{i_a} . The product $\langle x,y\rangle_a$ is defined for those vectors for which all coordinates $x_i,i\leq m_{a-1}$ or $i>m_a+1$ are zero. The first scalar square of an arbitrary vector x is a degenerate quadratic form $\langle x,x\rangle_1=-\sum_{i=1}^{l_1}x_i^2+\sum_{j=l_1+1}^{n-d}x_j^2$. The number $l_1\geq 0$ is called the index, and the number $d=n-m_1$ is called the defect of the space. If $l_1=\cdots=l_r=0$, we obtain a semi-Euclidean space. The spaces \mathbb{R}^n_{a} and $\mathbb{R}^n_{k,l}$ are called quasi-Euclidean spaces.

The semi-pseudo-non-Euclidean space $\mathbb{S}^n_{l_1,\ldots,l_r}$ can be defined as a hypersphere in $\mathbb{R}^{n+1}_{l_1,\ldots,l_r}$ with identified antipodal points. If $l_1=\cdots=l_r=0$, the space $\mathbb{S}^n_{\substack{m_1,\ldots,m_{r-1}\\m_1,\ldots,m_{r-1}}}$ is called a semi-elliptic space (or semi-non-Euclidean space). If there exist $l_i\neq 0$, the space $\mathbb{S}^n_{\substack{l_1,\ldots,l_r\\m_1,\ldots,m_{r-1}}}$ is called a semi-hyperbolic space.

• Finsler metric

Consider a real *n*-dimensional differentiable manifold M^n in which every tangent space $T_p(M^n)$, $p \in M^n$, is equipped with a Banach norm ||.|| such that the Banach norm as a function of position is smooth, and the matrix $((g_{ij}))$,

$$g_{ij} = g_{ij}(p, x) = \frac{1}{2} \frac{\partial^2 ||x||^2}{\partial x_i \partial x_j},$$

is positive-definite for any $p \in M^n$ and any $x \in T_p(M^n)$.

A **Finsler metric** on M^n is a collection of Banach norms ||.|| on the tangent spaces $T_p(M^n)$, one for each $p \in M^n$. The *line element* of this metric has the form

$$ds^2 = \sum_{i,j} g_{ij} dx_i dx_j.$$

The Finsler metric can be given by a real positive-definite convex function F(p,x) of coordinates of $p \in M^n$ and components of vectors $x \in T_p(M^n)$ acting at the point p. F(p,x) is positively homogeneous of degree one in x: $F(p,\lambda x) = \lambda F(p,x)$ for every $\lambda > 0$. The value of F(p,x) is interpreted as the length of the vector x. The Finsler metric tensor has the form $((g_{ij})) = ((\frac{1}{2} \frac{\partial^2 F^2(p,x)}{\partial x_i \partial x_j}))$. The length of a curve $\gamma : [0,1] \to M^n$ is given by $\int_0^1 F(p,\frac{dp}{dt})dt$. For each fixed p the Finsler metric tensor is Riemannian in the variables x.

The Finsler metric is a generalization of the Riemannian metric, where the general definition of the length ||x|| of a vector $x \in T_p(M^n)$ is not necessarily given in the form of the square root of a symmetric bilinear form as in the Riemannian case.

A Finsler manifold (or Finsler space) is a real n-dimensional differentiable manifold M^n equipped with a Finsler metric. The theory of Finsler spaces is called Finsler Geometry. The difference between a Riemannian space and a Finsler space is that the former behaves locally like a Euclidean space, and the latter locally like a $Minkowskian\ space$ or, analytically, the difference is that to an ellipsoid in the Riemannian case there corresponds an arbitrary convex surface which has the origin as the center.

A generalized Finsler space is a space with the intrinsic metric, subject to certain restrictions on the behavior of shortest curves, i.e., the curves with length equal to the distance between their ends. Such spaces include spaces of geodesics, Finsler spaces, etc. Generalized Finsler spaces differ from Finsler spaces not only in their greater generality, but also in the fact that they are defined and investigated starting from a metric, without coordinates.

• Kropina metric

The **Kropina metric** is a Finsler metric F_{Kr} on a real *n*-dimensional manifold M^n defined by

$$\frac{\sum_{i,j} g_{ij} x_i x_j}{\sum_i b_i(p) x_i}$$

for any $p \in M^n$ and $x \in T_p(M^n)$, where $((g_{ij}))$ is a Riemannian metric tensor, and $b(p) = (b_i(p))$ is a vector field.

• Randers metric

The **Randers metric** is a Finsler metric F_{Ra} on a real *n*-dimensional manifold M^n defined by

$$\sqrt{\sum_{i,j} g_{ij} x_i x_j} + \sum_i b_i(p) x_i$$

for any $p \in M^n$ and $x \in T_p(M^n)$, where $((g_{ij}))$ is a Riemannian metric tensor, and $b(p) = (b_i(p))$ is a vector field.

• Klein metric

The **Klein metric** is a Riemannian metric on the *open unit ball* $B^n = \{x \in \mathbb{R}^n : ||x||_2 < 1\}$ in \mathbb{R}^n defined by

$$\frac{\sqrt{||y||_2^2-\left(||x||_2^2||y||_2^2-\langle x,y\rangle^2\right)}}{1-||x||_2^2}$$

for any $x \in B^n$ and $y \in T_x(B^n)$, where $||.||_2$ is the Euclidean norm on \mathbb{R}^n , and \langle , \rangle is the ordinary *inner product* on \mathbb{R}^n .

• Funk metric

The **Funk metric** is a Finsler metric F_{Fu} on the open unit ball $B^n = \{x \in \mathbb{R}^n : ||x||_2 < 1\}$ in \mathbb{R}^n defined by

$$\frac{\sqrt{||y||_2^2-(||x||_2^2||y||_2^2-\langle x,y\rangle^2)}+\langle x,y\rangle}{1-||x||_2^2}$$

for any $x \in B^n$ and $y \in T_x(B^n)$, where $||.||_2$ is the Euclidean norm on \mathbb{R}^n , and \langle,\rangle is the ordinary inner product on \mathbb{R}^n . It is a **projective metric**.

• Shen metric

Given a vector $a \in \mathbb{R}^n$, $||a||_2 < 1$, the **Shen metric** is a Finsler metric F_{Sh} on the open unit ball $B^n = \{x \in \mathbb{R}^n : ||x||_2 < 1\}$ in \mathbb{R}^n defined by

$$\frac{\sqrt{||y||_2^2 - (||x||_2^2||y||_2^2 - \langle x, y \rangle^2)} + \langle x, y \rangle}{1 - ||x||_2^2} + \frac{\langle a, y \rangle}{1 + \langle a, x \rangle}$$

for any $x \in B^n$ and $y \in T_x(B^n)$, where $||.||_2$ is the Euclidean norm on \mathbb{R}^n , and \langle , \rangle is the ordinary *inner product* on \mathbb{R}^n . It is a **projective metric**. For a = 0 it becomes the **Funk metric**.

• Berwald metric

The **Berwald metric** is a Finsler metric F_{Be} on the open unit ball $B^n = \{x \in \mathbb{R}^n : ||x||_2 < 1\}$ in \mathbb{R}^n defined by

$$\frac{\left(\sqrt{||y||_2^2-(||x||_2^2||y||_2^2-\langle x,y\rangle^2)}+\langle x,y\rangle\right)^2}{(1-||x||_2^2)^2\sqrt{||y||_2^2-(||x||_2^2||y||_2^2-\langle x,y\rangle^2)}}$$

for any $x \in B^n$ and $y \in T_x(B^n)$, where $||.||_2$ is the Euclidean norm on \mathbb{R}^n , and \langle , \rangle is the ordinary inner product on \mathbb{R}^n . It is a **projective metric**.

In general, every Finsler metric on a manifold M^n induces a spray (second-order homogeneous ordinary differential equation) $y_i \frac{\partial}{\partial x_i} - 2G^i \frac{\partial}{\partial y_i}$ which determines the geodesics. A Finsler metric is called a **Berwald**

metric if the spray coefficients $G^i = G^i(x, y)$ are quadratic in $y \in T_x(M^n)$ at any point $x \in M^n$, i.e., $G^i = \frac{1}{2}\Gamma^i_{jk}(x)y^iy^k$. Every Berwald metric is affinely equivalent to a Riemannian metric.

• Douglas metric

A **Douglas metric** a Finsler metric for which the *spray coefficients* $G^i = G^i(x, y)$ have the following form:

$$G^{i} = \frac{1}{2} \Gamma^{i}_{jk}(x) y_i y_k + P(x, y) y_i.$$

Every Finsler metric which is projectively equivalent to a **Berwald metric** is a Douglas metric. Every known Douglas metric is (locally) projectively equivalent to a Berwald metric.

• Bryant metric

Let α be an angle with $|\alpha| < \frac{\pi}{2}$. Let, for any $x, y \in \mathbb{R}^n$, $A = ||y||_2^4 \sin^2 2\alpha + (||y||_2^2 \cos 2\alpha + ||x||_2^2||y||_2^2 - \langle x, y \rangle^2)^2$, $B = ||y||_2^2 \cos 2\alpha + ||x||_2^2||y||_2^2 - \langle x, y \rangle^2$, $C = \langle x, y \rangle \sin 2\alpha$, $D = ||x||_2^4 + 2||x||_2^2 \cos 2\alpha + 1$. Then one obtains a (**projective**) Finsler metric F by

$$\sqrt{\frac{\sqrt{A}+B}{2D} + \left(\frac{C}{D}\right)^2} + \frac{C}{D}.$$

On the two-dimensional unit sphere S^2 , it is the **Bryant metric**.

• Kawaguchi metric

The **Kawaguchi metric** is a metric on a smooth n-dimensional manifold M^n , given by the arc element ds of a regular curve x = x(t), $t \in [t_0, t_1]$, expressed by the formula

$$ds = F(x, \frac{dx}{dt}, \dots, \frac{d^kx}{dt^k})dt,$$

where the metric function F satisfies Zermelo's conditions: $\sum_{s=1}^k sx^{(s)}$ $F_{(s)i} = F$, $\sum_{s=r}^k \binom{s}{k} x^{(s-r+1)i} F_{(s)i} = 0$, $x^{(s)i} = \frac{d^s x^i}{dt^s}$, $F_{(s)i} = \frac{\partial F}{\partial x^{(s)i}}$, and $r = 2, \ldots, k$. These conditions ensure that the arc element ds is independent of the parametrization of the curve x = x(t).

A Kawaguchi manifold (or Kawaguchi space) is a smooth manifold equipped with a Kawaguchi metric. It is a generalization of a Finsler manifold.

• DeWitt supermetric

The **DeWitt supermetric** (or Wheeler–DeWitt supermetric) $G = ((G_{ijkl}))$ calculates distances between metrics on a given manifold, and it is a generalization of a Riemannian (or pseudo-Riemannian) metric $g = ((g_{ij}))$.

More exactly, for a given connected smooth three-dimensional manifold M^3 , consider the space $\mathcal{M}(M^3)$ of all Riemannian (or pseudo-Riemannian) metrics on M^3 . Identifying points of $\mathcal{M}(M^3)$ that are related by a diffeomorphism of M^3 , one obtains the space $Geom(M^3)$ of 3-geometries (of fixed topology), points of which are the classes of diffeomorphically equivalent metrics. The space $Geom(M^3)$ is called a superspace. It plays an important role in several formulations of Quantum Gravity.

A supermetric, i.e., a "metric on metrics," is a metric on $\mathcal{M}(M^3)$ (or on $Geom(M^3)$) which is used for measuring distances between metrics on M^3 (or between their equivalence classes). Given a metric $g = ((g_{ij})) \in \mathcal{M}(M^3)$, we obtain

$$||\delta g||^2 = \int_{M^3} d^3x G^{ijkl}(x) \delta g_{ij}(x) \delta g_{kl}(x),$$

where G^{ijkl} is the inverse of the **DeWitt supermetric**

$$G_{ijkl} = \frac{1}{2\sqrt{\det((g_{ij}))}} (g_{ik}g_{jl} + g_{il}g_{jk} - \lambda g_{ij}g_{kl}).$$

The value λ parameterizes the distance between metrics in $\mathcal{M}(M^3)$, and may take any real value except $\lambda = \frac{2}{3}$, for which the supermetric is *singular*.

• Lund-Regge supermetric

The Lund-Regge supermetric (or simplicial supermetric) is an analog of the **DeWitt supermetric**, used to measure the distances between simplicial 3-geometries in a simplicial configuration space.

More exactly, given a closed simplicial three-dimensional manifold M^3 consisting of several tetrahedra (i.e., 3-simplices), a simplicial geometry on M^3 is fixed by an assignment of values to the squared edge lengths of M^3 , and a flat Riemannian Geometry to the interior of each tetrahedron consistent with those values. The squared edge lengths should be positive and constrained by the triangle inequalities and their analogs for the tetrahedra, i.e., all squared measures (lengths, areas, volumes) must be non-negative (cf. tetrahedron inequality in Chap. 3). The set $\mathcal{T}(M^3)$ of all simplicial geometries on M^3 is called a simplicial configuration space.

The Lund–Regge supermetric $((G_{mn}))$ on $\mathcal{T}(M^3)$ is induced from the DeWitt supermetric on $\mathcal{M}(M^3)$, using for representations of points in $\mathcal{T}(M^3)$ such metrics in $\mathcal{M}(M^3)$ which are piecewise flat in the tetrahedra.

• Space of Lorentz metrics

Let M^n be an n-dimensional compact manifold, and $\mathcal{L}(M^n)$ the set of all **Lorentz metrics** (i.e., the quasi-Riemannian metrics of signature (n-1,1)) on M^n .

Given a Riemannian metric g on M^n , one can identify the vector space $S^2(M^n)$ of all symmetric 2-tensors with the vector space of endomorphisms

of the tangent to M^n , which are symmetric with respect to g. In fact, if \tilde{h} is the endomorphism associated to a tensor h, then the distance on $S^2(M^n)$ is given by

$$d_g(h,t) = \sup_{x \in M^n} \sqrt{tr(\tilde{h}_x - \tilde{t}_x)^2}.$$

The set $\mathcal{L}(M^n)$ equipped with the distance d_g is an open subset of $S^2(M^n)$ called the **space of Lorentz metrics**. Cf. **manifold triangulation metric** in Chap. 9.

• Perelman supermetric proof

The Thurston's Geometrization Conjecture is that, after two well-known splittings, any three-dimensional manifold admits, as remaining components, only one of 8 Thurston model geometries. If true, this conjecture implies the validity of the famous Poincaré Conjecture of 1904, that any 3-manifold, in which every simple closed curve can be deformed continuously to a point, is homeomorphic to the 3-sphere.

In 2003, Perelman gave a sketch of a proof of Thurston's conjecture using a kind of supermetric approach to the space of all Riemannian metrics on a given smooth 3-manifold. In a *Ricci flow* the distances decrease in directions of positive curvature since the metric is time-dependent. Perelman's modification of the standard Ricci flow permitted systematic elimination of arising singularities.

7.2 Riemannian metrics in Information Theory

Some special Riemannian metrics are commonly used in Information Theory. A list of such metrics is given below.

• Fisher information metric

In Statistics, Probability, and Information Geometry, the **Fisher information metric** (or **Fisher metric**, **Rao metric**) is a Riemannian metric for a statistical differential manifold (see, for example, [Amar85], [Frie98]). It addresses the differential geometry properties of families of classical probability densities.

Formally, let $p_{\theta} = p(x, \theta)$ be a family of densities, indexed by n parameters $\theta = (\theta_1, \dots, \theta_n)$ which form the parameter manifold P. The **Fisher information metric** $g = g_{\theta}$ on P is a Riemannian metric, defined by the Fisher information matrix $(I(\theta)_{ij})$, where

$$I(\theta)_{ij} = \mathbb{E}_{\theta} \left[\frac{\partial \ln p_{\theta}}{\partial \theta_i} \cdot \frac{\partial \ln p_{\theta}}{\partial \theta_j} \right] = \int \frac{\partial \ln p(x,\theta)}{\partial \theta_i} \frac{\partial \ln p(x,\theta)}{\partial \theta_j} p(x,\theta) dx.$$

It is a symmetric bilinear form which gives a classical measure (Rao measure) for the statistical distinguishability of distribution parameters. Putting $i(x, \theta) = -\ln p(x, \theta)$, one obtains an equivalent formula

$$I(\theta)_{ij} = \mathbb{E}_{\theta} \left[\frac{\partial^2 i(x,\theta)}{\partial \theta_i \partial \theta_j} \right] = \int \frac{\partial^2 i(x,\theta)}{\partial \theta_i \partial \theta_j} p(x,\theta) dx.$$

In a coordinate-free language, we get

$$I(\theta)(u, v) = \mathbb{E}_{\theta} \left[u(\ln p_{\theta}) \cdot v(\ln p_{\theta}) \right],$$

where u and v are vectors tangent to the parameter manifold P, and $u(\ln p_{\theta}) = \frac{d}{dt} \ln p_{\theta+tu|t=0}$ is the derivative of $\ln p_{\theta}$ along the direction u.

A manifold of densities M is the image of the parameter manifold P under the mapping $\theta \to p_{\theta}$ with certain regularity conditions. A vector u tangent to this manifold is of the form $u = \frac{d}{dt} p_{\theta + tu|t=0}$, and the Fisher metric $g = g_p$ on M, obtained from the metric g_{θ} on P, can be written as

$$g_p(u,v) = \mathbb{E}_p \left[\frac{u}{p} \cdot \frac{v}{p} \right].$$

• Fisher-Rao metric

Let $\mathcal{P}_n = \{p \in \mathbb{R}^n : \sum_{i=1}^n p_i = 1, p_i > 0\}$ be the simplex of strictly positive probability vectors. An element $p \in \mathcal{P}_n$ is a density of the *n*-point set $\{1, \ldots, n\}$ with $p(i) = p_i$. An element u of the tangent space $T_p(\mathcal{P}_n) = \{u \in \mathbb{R}^n : \sum_{i=1}^n u_i = 0\}$ at a point $p \in \mathcal{P}_n$ is a function on $\{1, \ldots, n\}$ with $u(i) = u_i$.

The **Fisher–Rao metric** g_p on \mathcal{P}_n is a Riemannian metric defined by

$$g_p(u,v) = \sum_{i=1}^n \frac{u_i v_i}{p_i}$$

for any $u, v \in T_p(\mathcal{P}_n)$, i.e., it is the **Fisher information metric** on \mathcal{P}_n . The Fisher–Rao metric is the unique (up to a constant factor) Riemannian metric on \mathcal{P}_n , contracting under stochastic maps [Chen72].

The Fisher–Rao metric is isometric, by $p \to 2(\sqrt{p_1}, \dots, \sqrt{p_n})$, with the standard metric on an open subset of the sphere of radius two in \mathbb{R}^n . This identification of \mathcal{P}_n allows one to obtain on \mathcal{P}_n the **geodesic distance**, called the **Fisher distance** (or **Bhattacharya distance** 1), by

$$2\arccos(\sum_i p_i^{1/2}q_i^{1/2}).$$

The Fisher–Rao metric can be extended to the set $\mathcal{M}_n = \{p \in \mathbb{R}^n, p_i > 0\}$ of all finite strictly positive measures on the set $\{1, \ldots, n\}$. In this case, the geodesic distance on \mathcal{M}_n can be written as

$$2(\sum_{i}(\sqrt{p_i}-\sqrt{q_i})^2)^{1/2}$$

for any $p, q \in \mathcal{M}_n$ (cf. **Hellinger metric** in Chap. 14).

• Monotone metric

Let M_n be the set of all complex $n \times n$ matrices. Let $\mathcal{M} \subset M_n$ be the manifold of all complex positive-definite $n \times n$ matrices. Let $\mathcal{D} \subset \mathcal{M}$, $\mathcal{D} = \{\rho \in \mathcal{M} : Tr\rho = 1\}$, be the manifold of all density matrices. The tangent space of \mathcal{M} at $\rho \in \mathcal{M}$ is $T_{\rho}(\mathcal{M}) = \{x \in M_n : x = x^*\}$, i.e., the set of all $n \times n$ Hermitian matrices. The tangent space $T_{\rho}(\mathcal{D})$ at $\rho \in \mathcal{D}$ is the subspace of traceless (i.e., with trace 0) matrices in $T_{\rho}(\mathcal{M})$.

A Riemannian metric λ on \mathcal{M} is called **monotone metric** if the inequality

$$\lambda_{h(\rho)}(h(u), h(u)) \le \lambda_{\rho}(u, u)$$

holds for any $\rho \in \mathcal{M}$, any $u \in T_{\rho}(\mathcal{M})$, and any completely positive trace preserving mapping h, called *stochastic mapping*. In fact [Petz96], λ is monotone if and only if it can be written as

$$\lambda_{\rho}(u, v) = Tr \ u J_{\rho}(v),$$

where J_{ρ} is an operator of the form $J_{\rho} = \frac{1}{f(L_{\rho}/R_{\rho})R_{\rho}}$. Here L_{ρ} and R_{ρ} are the left and the right multiplication operators, and $f:(0,\infty)\to\mathbb{R}$ is an operator monotone function which is symmetric, i.e., $f(t)=tf(t^{-1})$, and normalized, i.e., f(1)=1. Then $J_{\rho}(v)=\rho^{-1}v$ if v and ρ are commute, i.e., any monotone metric is equal to the **Fisher information metric** on commutative submanifolds. Therefore, monotone metrics generalize the Fisher information metric on the class of probability densities (classical or commutative case) to the class of density matrices (quantum or noncommutative case) which are used in Quantum Statistics and Information Theory. In fact, \mathcal{D} is the space of faithful states of an n-level quantum system.

A monotone metric $\lambda_{\rho}(u,v) = Tr \ u \frac{1}{f(L_{\rho}/R_{\rho})R_{\rho}}(v)$ can be rewritten as $\lambda_{\rho}(u,v) = Tr \ u \ c(L_{\rho},R_{\rho})(v)$, where the function $c(x,y) = \frac{1}{f(x/y)y}$ is the Morozova-Chentsov function related to λ .

The **Bures metric** is the smallest monotone metric, obtained for $f(t) = \frac{1+t}{2}$ (for $c(x,y) = \frac{2}{x+y}$). In this case $J_{\rho}(v) = g$, $\rho g + g\rho = 2v$, is the symmetric logarithmic derivative.

The **right logarithmic derivative metric** is the greatest monotone metric, corresponding to the function $f(t) = \frac{2t}{1+t}$ (for $c(x,y) = \frac{x+y}{2xy}$). In this case $J_{\rho}(v) = \frac{1}{2}(\rho^{-1}v + v\rho^{-1})$ is the *right logarithmic derivative*.

The **Bogolubov–Kubo–Mori metric** is obtained for $f(x) = \frac{x-1}{\ln x}$ (for $c(x,y) = \frac{\ln x - \ln x}{x-y}$). It can be written as $\lambda_{\rho}(u,v) = \frac{\partial^2}{\partial s \partial t} Tr(\rho + su) \ln(\rho + tv)|_{s,t=0}$.

The Wigner-Yanase-Dyson metrics λ_{ρ}^{α} are monotone for $\alpha \in [-3,3]$. For $\alpha = \pm 1$, we obtain the Bogolubov-Kubo-Mori metric; for $\alpha = \pm 3$ we obtain the right logarithmic derivative metric. The smallest in the family is the Wigner-Yanase metric, obtained for $\alpha = 0$.

• Bures metric

The Bures metric (or statistical metric) is a monotone metric on the manifold \mathcal{M} of all complex positive-definite $n \times n$ matrices defined by

$$\lambda_{\rho}(u, v) = Tr \ u J_{\rho}(v),$$

where $J_{\rho}(v) = g$, $\rho g + g\rho = 2v$, is the symmetric logarithmic derivative. It is the smallest monotone metric.

For any $\rho_1, \rho_2 \in \mathcal{M}$ the Bures distance, i.e., the geodesic distance defined by the Bures metric, can be written as

$$2\sqrt{Tr\rho_1 + Tr\rho_2 - 2Tr(\rho_1^{1/2}\rho_2\rho_1^{1/2})^{1/2}}.$$

On the submanifold $\mathcal{D} = \{ \rho \in \mathcal{M} : Tr\rho = 1 \}$ of density matrices it has the form

$$2\arccos Tr(\rho_1^{1/2}\rho_2\rho_1^{1/2})^{1/2}$$
.

• Right logarithmic derivative metric

The **right logarithmic derivative metric** (or *RLD-metric*) is a **monotone metric** on the manifold \mathcal{M} of all complex positive-definite $n \times n$ matrices defined by

$$\lambda_{\rho}(u,v) = Tr \ u J_{\rho}(v),$$

where $J_{\rho}(v) = \frac{1}{2}(\rho^{-1}v + v\rho^{-1})$ is the right logarithmic derivative. It is the greatest monotone metric.

• Bogolubov-Kubo-Mori metric

The **Bogolubov–Kubo–Mori metric** (or BKM-metric) is a **monotone metric** on the manifold \mathcal{M} of all complex positive-definite $n \times n$ matrices defined by

$$\lambda_{\rho}(u,v) = \frac{\partial^2}{\partial s \partial t} Tr(\rho + su) \ln(\rho + tv)|_{s,t=0}.$$

• Wigner-Yanase-Dyson metrics

The Wigner-Yanase-Dyson metrics (or WYD-metrics) form a family of metrics on the manifold \mathcal{M} of all complex positive-definite $n \times n$ matrices defined by

$$\lambda_{\rho}^{\alpha}(u,v) = \frac{\partial^2}{\partial t \partial s} Tr f_{\alpha}(\rho + tu) f_{-\alpha}(\rho + sv)|_{s,t=0},$$

where $f_{\alpha}(x)=\frac{2}{1-\alpha}x^{\frac{1-\alpha}{2}}$, if $\alpha\neq 1$, and is $\ln x$, if $\alpha=1$. These metrics are monotone for $\alpha\in[-3,3]$. For $\alpha=\pm 1$ one obtains the

Bogolubov–Kubo–Mori metric; for $\alpha=\pm 3$ one obtains the right logarithmic derivative metric.

The Wigner–Yanase metric (or WY-metric) λ_{ρ} is the Wigner–Yanase–Dyson metric λ_{ρ}^{0} , obtained for $\alpha=0$. It can be written as

$$\lambda_{\rho}(u,v) = 4Tr \ u(\sqrt{L_{\rho}} + \sqrt{R_{\rho}})^{2}(v),$$

and is the smallest metric in the family. For any $\rho_1, \rho_2 \in \mathcal{M}$ the **geodesic** distance defined by the WY-metric, has the form

$$2\sqrt{Tr\rho_1 + Tr\rho_2 - 2Tr(\rho_1^{1/2}\rho_2^{1/2})}.$$

On the submanifold $\mathcal{D}=\{\rho\in\mathcal{M}:Tr\rho=1\}$ of density matrices it is equal to

$$2\arccos Tr(\rho_1^{1/2}\rho_2^{1/2}).$$

• Connes metric

Roughly, the **Connes metric** is a generalization (from the space of all probability measures of a set X, to the *state space* of any *unital* C^* -algebra) of the **Kantorovich–Mallows–Monge–Wasserstein metric** defined as the **Lipschitz distance between measures**.

Let M^n be a smooth n-dimensional manifold. Let $A = C^{\infty}(M^n)$ be the (commutative) algebra of smooth complex-valued functions on M^n , represented as multiplication operators on the Hilbert space $H = L^2(M^n, S)$ of square integrable sections of the spinor bundle on M^n by $(f\xi)(p) = f(p)\xi(p)$ for all $f \in A$ and for all $\xi \in H$. Let D be the Dirac operator. Let the commutator [D, f] for $f \in A$ be the Clifford multiplication by the gradient ∇f so that its operator norm ||.|| in H is given by $||[D, f]|| = \sup_{p \in M^n} ||\nabla f||$.

The Connes metric is the intrinsic metric on M^n , defined by

$$\sup_{f \in A, ||[D, f]|| \le 1} |f(p) - f(q)|.$$

This definition can also be applied to discrete spaces, and even generalized to "non-commutative spaces" (unital C^* -algebras). In particular, for a labeled connected locally finite graph G = (V, E) with the vertex-set $V = \{v_1, \ldots, v_n, \ldots\}$, the Connes metric on V is defined by

$$\sup_{||[D,f]||=||df|| \le 1} |f_{v_i} - f_{v_j}|$$

for any $v_i, v_j \in V$, where $\{f = \sum f_{v_i} v_i : \sum |f_{v_i}|^2 < \infty\}$ is the set of formal sums f, forms a Hilbert space, and ||[D, f]|| can be obtained by $||[D, f]|| = \sup_i (\sum_{k=1}^{deg(v_i)} (f_{v_k} - f_{v_i})^2)^{\frac{1}{2}}$.

7.3 Hermitian metrics and generalizations

A vector bundle is a geometrical construct where to every point of a topological space M we attach a vector space so that all those vector spaces "glued together" form another topological space E. A continuous mapping $\pi: E \to M$ is called a projection E on M. For every $p \in M$, the vector space $\pi^{-1}(p)$ is called a fiber of the vector bundle. A real (complex) vector bundle is a vector bundle $\pi: E \to M$ whose fibers $\pi^{-1}(p)$, $p \in M$, are real (complex) vector spaces.

In a real vector bundle, for every $p \in M$, the fiber $\pi^{-1}(p)$ locally looks like the vector space \mathbb{R}^n , i.e., there is an open neighborhood U of p, a natural number n, and a homeomorphism $\varphi: U \times \mathbb{R}^n \to \pi^{-1}(U)$ such that, for all $x \in U$ and $v \in \mathbb{R}^n$, one has $\pi(\varphi(x,v)) = v$, and the mapping $v \to \varphi(x,v)$ yields an isomorphism between \mathbb{R}^n and $\pi^{-1}(x)$. The set U, together with φ , is called a local trivialization of the bundle. If there exists a "global trivialization," then a real vector bundle $\pi: M \times \mathbb{R}^n \to M$ is called trivial. Similarly, in a complex vector bundle, for every $p \in M$, the fiber $\pi^{-1}(p)$ locally looks like the vector space \mathbb{C}^n . The basic example of a complex vector bundle is the trivial bundle $\pi: U \times \mathbb{C}^n \to U$, where U is an open subset of \mathbb{R}^k .

Important special cases of a real vector bundle are the tangent bundle $T(M^n)$ and the cotangent bundle $T^*(M^n)$ of a real n-dimensional manifold $M_{\mathbb{R}}^n = M^n$. Important special cases of a complex vector bundle are the tangent bundle and the cotangent bundle of a complex n-dimensional manifold.

Namely, a complex n-dimensional manifold $M^n_{\mathbb{C}}$ is a topological space in which every point has an open neighborhood homeomorphic to an open set of the n-dimensional complex vector space \mathbb{C}^n , and there is an atlas of charts such that the change of coordinates between charts is analytic. The (complex) tangent bundle $T_{\mathbb{C}}(M^n_{\mathbb{C}})$ of a complex manifold $M^n_{\mathbb{C}}$ is a vector bundle of all (complex) tangent spaces of $M^n_{\mathbb{C}}$ at every point $p \in M^n_{\mathbb{C}}$. It can be obtained as a complexification $T_{\mathbb{R}}(M^n_{\mathbb{R}}) \otimes \mathbb{C} = T(M^n) \otimes \mathbb{C}$ of the corresponding real tangent bundle, and is called the complexified tangent bundle of $M^n_{\mathbb{C}}$.

The complexified cotangent bundle of $M^n_{\mathbb{C}}$ is obtained similarly as $T^*(M^n)\otimes \mathbb{C}$. Any complex n-dimensional manifold $M^n_{\mathbb{C}}=M^n$ can be regarded as a special case of a real 2n-dimensional manifold equipped with a complex structure on each tangent space. A complex structure on a real vector space V is the structure of a complex vector space on V that is compatible with the original real structure. It is completely determined by the operator of multiplication by the number i, the role of which can be taken by an arbitrary linear transformation $J: V \to V$, $J^2 = -id$, where id is the identity mapping.

A connection (or covariant derivative) is a way of specifying a derivative of a vector field along another vector field in a vector bundle. A **metric** connection is a linear connection in a vector bundle $\pi: E \to M$, equipped with a bilinear form in the fibers, for which parallel displacement along an arbitrary piecewise-smooth curve in M preserves the form, that is, the scalar

product of two vectors remains constant under parallel displacement. In the case of a non-degenerate symmetric bilinear form, the metric connection is called the *Euclidean connection*. In the case of non-degenerate antisymmetric bilinear form, the metric connection is called the *symplectic connection*.

• Bundle metric

A **bundle metric** is a metric on a vector bundle.

• Hermitian metric

A Hermitian metric on a complex vector bundle $\pi: E \to M$ is a collection of Hermitian inner products (i.e., positive-definite symmetric sesquilinear forms) on every fiber $E_p = \pi^{-1}(p), \ p \in M$, that varies smoothly with the point p in M. Any complex vector bundle has a Hermitian metric.

The basic example of a vector bundle is the trivial bundle $\pi: U \times \mathbb{C}^n \to U$, where U is an open set in \mathbb{R}^k . In this case a Hermitian inner product on \mathbb{C}^n , and hence, a Hermitian metric on the bundle $\pi: U \times \mathbb{C}^n \to U$, is defined by

$$\langle u, v \rangle = u^T H \overline{v},$$

where H is a positive-definite Hermitian matrix, i.e., a complex $n \times n$ matrix such that $H^* = \overline{H}^T = H$, and $\overline{v}^T H v > 0$ for all $v \in \mathbb{C}^n \setminus \{0\}$. In the simplest case, one has $\langle u, v \rangle = \sum_{i=1}^n u_i \overline{v}_i$.

An important special case is a Hermitian metric h on a complex manifold M^n , i.e., on the complexified tangent bundle $T(M^n) \otimes \mathbb{C}$ of M^n . This is the Hermitian analog of a Riemannian metric. In this case h = g + iw, and its real part g is a Riemannian metric, while its imaginary part w is a non-degenerate antisymmetric bilinear form, called a fundamental form. Here g(J(x), J(y)) = g(x, y), w(J(x), J(y)) = w(x, y), and w(x, y) = g(x, J(y)), where the operator J is an operator of complex structure on M^n ; as a rule, J(x) = ix. Any of the forms g, w determines h uniquely. The term Hermitian metric can also refer to the corresponding Riemannian metric g, which gives M^n a Hermitian structure.

On a complex manifold a Hermitian metric h can be expressed in local coordinates by a Hermitian symmetric tensor $((h_{ij}))$:

$$h = \sum_{i,j} h_{ij} dz_i \otimes d\overline{z}_j,$$

where $((h_{ij}))$ is a positive-definite Hermitian matrix. The associated fundamental form w is then written as $w = \frac{i}{2} \sum_{i,j} h_{ij} dz_i \wedge d\overline{z}_j$.

A Hermitian manifold (or Hermitian space) is a complex manifold equipped with a Hermitian metric.

• Kähler metric

A Kähler metric (or Kählerian metric) is a Hermitian metric h = g + iw on a complex manifold M^n whose fundamental form w is closed, i.e., satisfies the condition dw = 0. A Kähler manifold is a complex manifold equipped with a Kähler metric.

If h is expressed in local coordinates, i.e., $h = \sum_{i,j} h_{ij} dz_i \otimes d\overline{z}_j$, then the associated fundamental form w can be written as $w = \frac{i}{2} \sum_{i,j} h_{ij} dz_i \wedge d\overline{z}_j$, where \wedge is the wedge product which is antisymmetric, i.e., $dx \wedge dy = -dy \wedge dx$ (hence, $dx \wedge dx = 0$). In fact, w is a differential 2-form on M^n , i.e., a tensor of rank 2 that is antisymmetric under exchange of any pair of indices: $w = \sum_{i,j} f_{ij} dx^i \wedge dx^j$, where f_{ij} is a function on M^n . The exterior derivative dw of w is defined by $dw = \sum_{i,j} \sum_k \frac{\partial f_{ij}}{\partial x_k} dx_k \wedge dx_i \wedge dx_k$. If dw = 0, then w is a symplectic (i.e., closed non-degenerate) differential 2-form. Such differential 2-forms are called Kähler forms.

The metric on a Kähler manifold locally satisfies

$$h_{ij} = \frac{\partial^2 K}{\partial z_i \partial \overline{z}_j}$$

for some function K, called the Kähler potential.

The term $K\ddot{a}hler\ metric$ can also refer to the corresponding Riemannian metric g, which gives M^n a Kähler structure. Then a Kähler manifold is defined as a complex manifold which carries a Riemannian metric and a Kähler form on the underlying real manifold.

• Hessian metric

Given a smooth f on an open subset of a real vector space, the associated **Hessian metric** is defined by

$$g_{ij} = \frac{\partial^2 f}{\partial x_i \partial x_j}.$$

A Hessian metric is also called an **affine Kähler metric** since a Kähler metric on a complex manifold has an analogous description as $\frac{\partial^2 f}{\partial z_i \partial \overline{z_i}}$.

• Calabi-Yau metric

The Calabi-Yau metric is a Kähler metric which is Ricci-flat.

A Calabi-Yau manifold (or Calabi-Yau space) is a simply-connected complex manifold equipped with a Calabi-Yau metric. It can be considered as a 2n-dimensional (six-dimensional being particularly interesting) smooth manifold with holonomy group (i.e., the set of linear transformations of tangent vectors arising from parallel transport along closed loops) in the special unitary group.

• Kähler-Einstein metric

A Kähler–Einstein metric (or Einstein metric) is a Kähler metric on a complex manifold M^n whose $Ricci\ curvature\ tensor$ is proportional to the metric tensor. This proportionality is an analog of the $Einstein\ field\ equation$ in the General Theory of Relativity.

A Kähler–Einstein manifold (or Einstein manifold) is a complex manifold equipped with a Kähler–Einstein metric. In this case the Ricci curvature tensor, considered as an operator on the tangent space, is just multiplication by a constant.

Such a metric exists on any domain $D \subset \mathbb{C}^n$ that is bounded and pseudo-convex. It can be given by the line element

$$ds^{2} = \sum_{i,j} \frac{\partial^{2} u(z)}{\partial z_{i} \partial \overline{z}_{j}} dz_{i} d\overline{z}_{j},$$

where u is a solution to the boundary value problem: $det(\frac{\partial^2 u}{\partial z_i \partial \overline{z}_j}) = e^{2u}$ on D, and $u = \infty$ on ∂D .

The Kähler–Einstein metric is a complete metric. On the unit disk $\Delta = \{z \in \mathbb{C} : |z| < 1\}$ it is coincides with the **Poincaré metric**.

• Hodge metric

The **Hodge metric** is a **Kähler metric** whose fundamental form w defines an integral cohomology class or, equivalently, has integral periods.

A Hodge manifold (or Hodge variety) is a complex manifold equipped with a Hodge metric. A compact complex manifold is a Hodge manifold if and only if it is isomorphic to a smooth algebraic subvariety of some complex projective space.

• Fubini-Study metric

The Fubini-Study metric (or Cayley-Fubini-Study metric) is a Kähler metric on a complex projective space $\mathbb{C}P^n$ defined by a Hermitian inner product \langle , \rangle in \mathbb{C}^{n+1} . It is given by the line element

$$ds^2 = \frac{\langle x, x \rangle \langle dx, dx \rangle - \langle x, d\overline{x} \rangle \langle \overline{x}, dx \rangle}{\langle x, x \rangle^2}.$$

The distance between two points $(x_1 : ... : x_{n+1})$ and $(y_1 : ... : y_{n+1}) \in \mathbb{C}P^n$, where $x = (x_1, ..., x_{n+1})$ and $y = (y_1, ..., y_{n+1}) \in \mathbb{C}^{n+1} \setminus \{0\}$, is equal to

$$\arccos \frac{|\langle x, y \rangle|}{\sqrt{\langle x, x \rangle \langle y, y \rangle}}.$$

The Fubini–Study metric is a **Hodge metric**. The space $\mathbb{C}P^n$ endowed with the Fubini–Study metric is called a *Hermitian elliptic space* (cf. **Hermitian elliptic metric**).

• Bergman metric

The **Bergman metric** is a **Kähler metric** on a bounded *domain* $D \subset \mathbb{C}^n$ defined by the *line element*

$$ds^{2} = \sum_{i,j} \frac{\partial^{2} \ln K(z,z)}{\partial z_{i} \partial \overline{z}_{j}} dz_{i} d\overline{z}_{j},$$

where K(z, u) is the Bergman kernel function. The Bergman metric is invariant under all automorphisms of D; it is complete if D is homogeneous.

For the unit disk $\Delta = \{z \in \mathbb{C} : |z| < 1\}$ the Bergman metric coincides with the **Poincaré metric** (cf. also **Bergman** p-metric in Chap. 13).

The Bergman kernel function is defined as follows. Consider a domain $D \subset \mathbb{C}^n$ in which there exists analytic functions $f \neq 0$ of class $L_2(D)$ with respect to the Lebesgue measure. The set of these functions forms the **Hilbert space** $L_{2,a}(D) \subset L_2(D)$ with an orthonormal basis $(\phi_i)_i$. The Bergman kernel function in the domain $D \times D \subset \mathbb{C}^{2n}$ is defined by $K_D(z,u) = K(z,u) = \sum_{i=1}^{\infty} \phi_i(z) \overline{\phi_i(u)}$.

• Hyper-Kähler metric

A hyper-Kähler metric is a Riemannian metric g on a 4n-dimensional $Riemannian\ manifold$ which is compatible with a quaternionic structure on the tangent bundle of the manifold.

Thus, the metric g is Kählerian with respect to the three Kähler structures (I, w_I, g) , (J, w_J, g) , and (K, w_K, g) , corresponding to the complex structures, as endomorphisms of the tangent bundle which satisfy the quaternionic relationship

$$I^2 = J^2 = K^2 = IJK = -JIK = -1.$$

A hyper-Kähler manifold is a Riemannian manifold equipped with a hyper-Kähler metric. It is a special case of a Kähler manifold. All hyper-Kähler manifolds are Ricci-flat. Compact four-dimensional hyper-Kähler manifolds are called K_3 -surfaces; they are studied in Algebraic Geometry.

• Calabi metric

The Calabi metric is a hyper-Kähler metric on the cotangent bundle $T^*(\mathbb{C}P^{n+1})$ of a complex projective space $\mathbb{C}P^{n+1}$. For n=4k+4, this metric can be given by the line element

$$ds^{2} = \frac{dr^{2}}{1 - r^{-4}} + \frac{1}{4}r^{2}(1 - r^{-4})\lambda^{2} + r^{2}(\nu_{1}^{2} + \nu_{2}^{2}) +$$

$$+\frac{1}{2}(r^2-1)(\sigma_{1\alpha}^2+\sigma_{2\alpha}^2)+\frac{1}{2}(r^2+1)(\varSigma_{1\alpha}^2+\varSigma_{2\alpha}^2),$$

where $(\lambda, \nu_1, \nu_2, \sigma_{1\alpha}, \sigma_{2\alpha}, \Sigma_{1\alpha}, \Sigma_{2\alpha})$, with α running over k values, are left-invariant one-forms (i.e., linear real-valued functions) on the coset SU(k+2)/U(k). Here U(k) is the unitary group consisting of complex $k \times k$ unitary matrices, and SU(k) is the special unitary group of complex $k \times k$ unitary matrices with determinant 1.

For k=0, the Calabi metric coincides with the **Eguchi–Hanson** metric.

• Stenzel metric

The **Stenzel metric** is a **hyper-Kähler metric** on the cotangent bundle $T^*(S^{n+1})$ of a sphere S^{n+1} .

• SO(3)-invariant metric

An SO(3)-invariant metric is a four-dimensional hyper-Kähler metric with the *line element* given, in the Bianchi-IX formalism, by

$$ds^2 = f^2(t)dt^2 + a^2(t)\sigma_1^2 + b^2(t)\sigma_2^2 + c^2(t)\sigma_3^2,$$

where the invariant one-forms σ_1 , σ_2 , σ_3 of SO(3) are expressed in terms of Euler angles θ , ψ , ϕ as $\sigma_1 = \frac{1}{2}(\sin\psi d\theta - \sin\theta\cos\psi d\phi)$, $\sigma_2 = -\frac{1}{2}(\cos\psi d\theta + \sin\theta\sin\psi d\phi)$, $\sigma_3 = \frac{1}{2}(d\psi + \cos\theta d\phi)$, and the normalization has been chosen so that $\sigma_i \wedge \sigma_j = \frac{1}{2}\epsilon_{ijk}d\sigma_k$. The coordinate t of the metric can always be chosen so that $f(t) = \frac{1}{2}abc$, using a suitable reparametrization.

• Atiyah-Hitchin metric

The Atiyah–Hitchin metric is a complete regular SO(3)-invariant metric with the line element

$$ds^2 = \frac{1}{4}a^2b^2c^2\left(\frac{dk}{k(1-k^2)K^2}\right)^2 + a^2(k)\sigma_1^2 + b^2(k)\sigma_2^2 + c^2(k)\sigma_3^2,$$

where a,b,c are functions of k, ab=-K(k)(E(k)-K(k)), $bc=-K(k)(E(k)-(1-k^2)K(k))$, ac=-K(k)E(k), and K(k), E(k) are the complete elliptic integrals, respectively, of the first and second kind, with 0 < k < 1. The coordinate t is given by the change of variables $t=-\frac{2K(1-k^2)}{\pi K(k)}$ up to an additive constant.

• Taub-NUT metric

The Taub-NUT metric is a complete regular SO(3)-invariant metric with the $line\ element$

$$ds^{2} = \frac{1}{4} \frac{r+m}{r-m} dr^{2} + (r^{2}-m^{2})(\sigma_{1}^{2}+\sigma_{2}^{2}) + 4m^{2} \frac{r-m}{r+m} \sigma_{3}^{2},$$

where m is the relevant moduli parameter, and the coordinate r is related to t by $r = m + \frac{1}{2mt}$.

• Eguchi-Hanson metric

The Eguchi–Hanson metric is a complete regular SO(3)-invariant metric with the line element

$$ds^{2} = \frac{dr^{2}}{1 - \left(\frac{a}{r}\right)^{4}} + r^{2} \left(\sigma_{1}^{2} + \sigma_{2}^{2} + \left(1 - \left(\frac{a}{r}\right)^{4}\right)\sigma_{3}^{2}\right),$$

where a is the moduli parameter, and the coordinate r is related to t by $r^2 = a^2 \coth(a^2 t)$.

The Eguchi–Hanson metric coincides with the four-dimensional Calabi metric.

• Complex Finsler metric

A **complex Finsler metric** is an upper semi-continuous function $F: T(M^n) \to \mathbb{R}_+$ on a complex manifold M^n with the analytic tangent bundle $T(M^n)$ satisfying the following conditions:

1. F^2 is smooth on \check{M}^n , where \check{M}^n is the complement in $T(M^n)$ of the zero section.

- 2. F(p,x) > 0 for all $p \in M^n$ and $x \in \check{M}_p^n$.
- 3. $F(p, \lambda x) = |\lambda| F(p, x)$ for all $p \in M^n$, $x \in T_p(M^n)$, and $\lambda \in \mathbb{C}$.

The function $G = F^2$ can be locally expressed in terms of the coordinates $(p_1, \ldots, p_n, x_1, \ldots, x_n)$; the Finsler metric tensor of the complex Finsler metric is given by the matrix $((G_{ij})) = ((\frac{1}{2} \frac{\partial^2 F^2}{\partial x_i \partial \overline{x}_j}))$, called the Levi matrix. If the matrix $((G_{ij}))$ is positive-definite, the complex Finsler metric F is called strongly pseudo-convex.

• Distance-decreasing semi-metric

Let d be a semi-metric which can be defined on some class \mathcal{M} of complex manifolds containing the unit disk $\Delta = \{z \in \mathbb{C} : |z| < 1\}$. It is called **distance-decreasing** for all analytic mappings if, for any analytic mapping $f: M_1 \to M_2$ with $M_1, M_2 \in \mathcal{M}$, the inequality $d(f(p), f(q)) \leq d(p, q)$ holds for all $p, q \in M_1$.

Cf. Kobayashi metric, Carathéodory metric and Wu metric.

• Kobayashi metric

Let D be a *domain* in \mathbb{C}^n . Let $\mathcal{O}(\Delta, D)$ be the set of all analytic mappings $f: \Delta \to D$, where $\Delta = \{z \in \mathbb{C} : |z| < 1\}$ is the *unit disk*.

The Kobayashi metric (or Kobayashi–Royden metric) F_K is a complex Finsler metric defined by

$$F_K(z, u) = \inf\{\alpha > 0 : \exists f \in \mathcal{O}(\Delta, D), f(0) = z, \alpha f'(0) = u\}$$

for all $z \in D$ and $u \in \mathbb{C}^n$. It is a generalization of the **Poincaré metric** to higher-dimensional domains. Then $F_K(z,u) \geq F_C(z,u)$, where F_C is the **Carathéodory metric**. If D is convex, and $d(z,u) = \inf\{\lambda : z + \frac{u}{\alpha} \in D \text{ if } |\alpha| > \lambda\}$, then $\frac{d(z,u)}{2} \leq F_K(z,u) = F_C(z,u) \leq d(z,u)$.

Given a complex manifold M^n , the **Kobayashi semi-metric** F_K is defined by

$$F_K(p, u) = \inf\{\alpha > 0 : \exists f \in \mathcal{O}(\Delta, M^n), f(0) = p, \alpha f'(0) = u\}$$

for all $p \in M^n$ and $u \in T_p(M^n)$. $F_K(p, u)$ is a semi-norm of the tangent vector u, called the *Kobayashi semi-norm*. F_K is a metric if M^n is taut, i.e., $\mathcal{O}(\Delta, M^n)$ is a normal family.

The Kobayashi semi-metric is an infinitesimal form of the so-called **Kobayashi semi-distance** (or *Kobayashi pseudo-distance*) K_{M^n} on M^n , defined as follows. Given $p, q \in M^n$, a chain of disks α from p to q is a collection of points $p = p^0, p^1, \ldots, p^k = q$ of M^n , pairs of points $a^1, b^1, \ldots; a^k, b^k$ of the unit disk Δ , and analytic mappings $f_1, \ldots f_k$ from Δ into M^n , such that $f_j(a^j) = p^{j-1}$ and $f_j(b^j) = p^j$ for all j. The length $l(\alpha)$ of a chain α is the sum $d_P(a^1, b^1) + \cdots + d_P(a^k, b^k)$, where d_P is the Poincaré metric. The Kobayashi semi-metric K_{M^n} on M^n is defined by

$$K_{M^n}(p,q) = \inf_{\alpha} l(\alpha),$$

where the infimum is taken over all lengths $l(\alpha)$ of chains of disks α from p to q.

The Kobayashi semi-distance is **distance-decreasing** for all analytic mappings. It is the greatest semi-metric among all semi-metrics on M^n , that are distance-decreasing for all analytic mappings from Δ into M^n , where distances on Δ are measured in the Poincaré metric. K_{Δ} is the Poincaré metric, and $K_{\mathbb{C}^n} \equiv 0$.

A manifold is called *Kobayashi hyperbolic* if the Kobayashi semi-distance is a metric on it. In fact, a manifold is Kobayashi hyperbolic if and only if it is biholomorphic to a bounded homogeneous domain.

• Kobayashi-Busemann metric

Given a complex manifold M^n , the Kobayashi-Busemann semimetric on M^n is the double dual of the Kobayashi semi-metric. It is a metric if M^n is taut.

• Carathéodory metric

Let D be a *domain* in \mathbb{C}^n . Let $\mathcal{O}(D, \Delta)$ be the set of all analytic mappings $f: D \to \Delta$, where $\Delta = \{z \in \mathbb{C} : |z| < 1\}$ is the *unit disk*.

The Carathéodory metric F_C is a complex Finsler metric defined by

$$F_{C}(z,u) = \sup\{|f'(z)u| : f \in \mathcal{O}(D,\Delta)\}$$

for any $z \in D$ and $u \in \mathbb{C}^n$. It is a generalization of the **Poincaré metric** to higher-dimensional domains. Then $F_C(z,u) \leq F_K(z,u)$, where F_K is the Kobayashi metric. If D is convex and $d(z,u) = \inf\{\lambda : z + \frac{u}{\alpha} \in D \text{ if } |\alpha| > \lambda\}$, then $\frac{d(z,u)}{2} \leq F_C(z,u) = F_K(z,u) \leq d(z,u)$.

Given a complex manifold M^n , the Carathéodory semi-metric F_C is defined by

$$F_C(p, u) = \sup\{|f'(p)u| : f \in \mathcal{O}(M^n, \Delta)\}$$

for all $p \in M^n$ and $u \in T_p(M^n)$. F_C is a metric if M^n is taut.

The Carathéodory semi-distance (or Carathéodory pseudo-distance) C_{M^n} is a semi-metric on a complex manifold M^n , defined by

$$C_{M^n}(p,q) = \sup\{d_P(f(p), f(q)) : f \in \mathcal{O}(M^n, \Delta)\},\$$

where d_P is the Poincaré metric. In general, the integrated semimetric of the infinitesimal Carathéodory semi-metric is **internal** for the Carathéodory semi-distance, but does not coincides with it.

The Carathéodory semi-distance is **distance-decreasing** for all analytic mappings. It is the smallest distance-decreasing semi-metric. C_{Δ} coincides with the Poincaré metric, and $C_{\mathbb{C}^n} \equiv 0$.

• Azukawa metric

Let D be a domain in \mathbb{C}^n . Let $g_D(z,u) = \sup\{f(u) : f \in K_D(z)\}$, where $K_D(z)$ is the set of all logarithmically plurisubharmonic functions

 $f: D \to [0,1)$ such that there exist M, r > 0 with $f(u) \leq M||u-z||_2$ for all $u \in B(z,r) \subset D$; here $||.||_2$ is the l_2 -norm on \mathbb{C}^n , and $B(z,r) = \{x \in \mathbb{C}^n: ||z-x||_2 < r\}$.

The Azukawa metric (in general, a semi-metric) F_A is a complex Finsler metric defined by

$$F_A(z, u) = \lim_{\lambda \to 0} \sup \frac{1}{|\lambda|} g_D(z, z + \lambda u)$$

for all $z \in D$ and $u \in \mathbb{C}^n$. It "lies between" the **Carathéodory metric** F_C and the **Kobayashi metric** F_K : $F_C(z, u) \leq F_A(z, u) \leq F_K(z, u)$ for all $z \in D$ and $u \in \mathbb{C}^n$. If D is convex, then all these metrics coincide.

The Azukawa metric is an infinitesimal form of the so-called **Azukawa** semi-distance.

• Sibony semi-metric

Let D be a domain in \mathbb{C}^n . Let $K_D(z)$ be the set of all logarithmically plurisubharmonic functions $f:D\to [0,1)$ such that there exist M,r>0 with $f(u)\leq M||u-z||_2$ for all $u\in B(z,r)\subset D$; here $||.||_2$ is the l_2 -norm on \mathbb{C}^n , and $B(z,r)=\{x\in\mathbb{C}^n:||z-x||_2< r\}$. Let $C^2_{loc}(z)$ be the set of all functions of class C^2 on some open neighborhood of z.

The Sibony semi-metric F_S is a complex Finsler semi-metric defined by

$$F_S(z, u) = \sup_{f \in K_D(z) \cap C_{loc}^2(z)} \sqrt{\sum_{i,j} \frac{\partial^2 f}{\partial z_i \partial \overline{z}_j}(z) u_i \overline{u}_j}$$

for all $z \in D$ and $u \in \mathbb{C}^n$. When it is a metric, it "lies between" the **Carathéodory metric** F_C and the **Kobayashi metric** F_K : $F_C(z, u) \le F_S(z, u) \le F_A(z, u) \le F_K(z, u)$ for all $z \in D$ and $u \in \mathbb{C}^n$, where F_A is the **Azukawa metric**. If D is convex, then all these metrics coincide.

The Sibony semi-metric is an infinitesimal form of the so-called **Sibony** semi-distance.

• Wu metric

The **Wu metric** W_{M^n} is an upper-semi-continuous Hermitian metric on a complex manifold M^n , that is **distance-decreasing** for all analytic mappings. In fact, for two n-dimensional complex manifolds M_1^n and M_2^n , the inequality $W_{M_2^n}(f(p), f(q)) \leq \sqrt{n}W_{M_1^n}(p, q)$ holds for all $p, q \in M_1^n$.

Invariant metrics, including the Carathéodory, Kobayashi, Bergman, and Kähler–Einstein metrics, play an important role in Complex Function Theory and Convex Geometry. The Carathéodory and Kobayashi metrics are used mostly because of the distance-decreasing property. But they are almost never Hermitian. On the other hand, the Bergman metric and the Kähler–Einstein metric are Hermitian (in fact, Kählerian), but the distance-decreasing property, in general, fails for them.

• Teichmüller metric

A Riemann surface R is a one-dimensional complex manifold. Two Riemann surfaces R_1 and R_2 are called conformally equivalent if there exists a bijective analytic function (i.e., a conformal homeomorphism) from R_1 into R_2 . More precisely, consider a fixed closed Riemann surface R_0 of a given genus $g \geq 2$. For a closed Riemann surface R of genus g, construct a pair (R, f), where $f: R_0 \to R$ is a homeomorphism. Two pairs (R, f) and (R_1, f_1) are called conformally equivalent if there exists a conformal homeomorphism $h: R \to R_1$ such that the mapping $(f_1)^{-1} \cdot h \cdot f: R_0 \to R_0$ is homotopic to the identity.

An abstract Riemann surface $R^* = (R, f)^*$ is the equivalence class of all Riemann surfaces, conformally equivalent to R. The set of all equivalence classes is called the Teichmüller space $T(R_0)$ of the surface R_0 . For closed surfaces R_0 of given genus g, the spaces $T(R_0)$ are isometrically isomorphic, and one can speak of the Teichmüller space T_g of surfaces of genus g. T_g is a complex manifold. If R_0 is obtained from a compact surface of genus $g \geq 2$ by removing n points, then the complex dimension of T_g is 3g-3+n.

The **Teichmüller metric** is a metric on T_q defined by

$$\frac{1}{2}\inf_{h}\ln K(h)$$

for any $R_1^*, R_2^* \in T_g$, where $h: R_1 \to R_2$ is a quasi-conformal homeomorphism, homotopic to the identity, and K(h) is the maximal dilation of h. In fact, there exists a unique extremal mapping, called the Teichmüller mapping, which minimizes the maximal dilation of all such h, and the distance between R_1^* and R_2^* is equal to $\frac{1}{2} \ln K$, where the constant K is the dilation of the Teichmüller mapping.

In terms of the extremal length $ext_{R^*}(\gamma)$, the distance between R_1^* and R_2^* is

$$\frac{1}{2}\ln\sup_{\gamma}\frac{ext_{R_1^*}(\gamma)}{ext_{R_2^*}(\gamma)},$$

where the supremum is taken over all simple closed curves on R_0 .

The Teichmüller space T_g , with the Teichmüller metric on it, is a **geodesic** metric space (moreover, a **straight** G-**space**) but it is neither **Gromov hyperbolic**, nor a **globally non-positively Busemann curved** metric space.

The **Thurston quasi-metric** on the *Teichmüller space* T_g is defined by

$$\frac{1}{2}\inf_{h}\ln||h||_{Lip}$$

for any $R_1^*, R_2^* \in T_g$, where $h: R_1 \to R_2$ is a quasi-conformal homeomorphism, homotopic to the identity, and $||.||_{Lip}$ is the Lipschitz norm on the set of all injective functions $f: X \to Y$ defined by $||f||_{Lip} = \sup_{x,y \in X, x \neq y} \frac{d_Y(f(x),f(y))}{d_X(x,y)}$.

The moduli space R_g of conformal classes of Riemann surfaces of genus

The moduli space R_g of conformal classes of Riemann surfaces of genus g is obtained by factorization of T_g by some countable group of automorphisms of it, called the modular group. Examples of metrics related to moduli and Teichmüller spaces are, besides the Teichmüller metric, the Weil-Petersson metric, Quillen metric, Carathéodory metric, Kobayashi metric, Bergman metric, Cheng-Yau-Mok metric, McMullen metric, asymptotic Poincaré metric, Ricci metric, perturbed Ricci metric, VHS-metric.

• Weil-Petersson metric

The Weil-Petersson metric is a Kähler metric on the Teichmüller space $T_{g,n}$ of abstract Riemann surfaces of genus g with n punctures and negative Euler characteristic.

The Weil-Peterson metric is Gromov hyperbolic if and only if (Brock and Farb 2006) the complex dimension 3g - 3 + n of $T_{g,n}$ is at most 2.

• Gibbons-Manton metric

The **Gibbons–Manton metric** is a 4n-dimensional **hyper-Kähler metric** on the moduli space of n-monopoles, admitted an isometric action of the n-dimensional torus T^n . It can be described also as a hyper-Kähler quotient of a flat quaternionic vector space.

Zamolodchikov metric

The **Zamolodchikov metric** is a metric on the moduli space of twodimensional conformal field theories.

• Metrics on determinant lines

Let M^n be an n-dimensional compact smooth manifold, and let F be a flat vector bundle over M^n . Let $H^{\bullet}(M^n,F) = \bigoplus_{i=0}^n H^i(M^n,F)$ be the de Rham cohomology of M^n with coefficients in F. Given an n-dimensional vector space V, the determinant line $\det V$ of V is defined as the top exterior power of V, i.e., $\det V = \wedge^n V$. Given a finite-dimensional graded vector space $V = \bigoplus_{i=0}^n V_i$, the determinant line of V is defined as the tensor product $\det V = \bigotimes_{i=0}^n (\det V_i)^{(-1)^i}$. Thus, the determinant line $\det H^{\bullet}(M^n,F)$ of the cohomology $H^{\bullet}(M^n,F)$ can be written as $\det H^{\bullet}(M^n,F) = \bigotimes_{i=0}^n (\det H^i(M^n,F))^{(-1)^i}$.

The **Reidemeister metric** is a metric on $detH^{\bullet}(M^n, F)$, defined by a given smooth triangulation of M^n , and the classical *Reidemeister-Franz* torsion.

Let g^F and $g^{T(M^n)}$ be smooth metrics on the vector bundle F and tangent bundle $T(M^n)$, respectively. These metrics induce a canonical L_2 -metric $h^{H^{\bullet}(M^n,F)}$ on $H^{\bullet}(M^n,F)$. The **Ray–Singler metric** on $det H^{\bullet}(M^n,F)$ is defined as the product of the metric induced on

 $det H^{\bullet}(M^n, F)$ by $h^{H^{\bullet}(M^n, F)}$ with the Ray-Singler analytic torsion. The **Milnor metric** on $det H^{\bullet}(M^n, F)$ can be defined in a similar manner using the Milnor analytic torsion. If g^F is flat, the above two metrics coincide with the Reidemeister metric. Using a co-Euler structure, one can define a modified Ray-Singler metric on $det H^{\bullet}(M^n, F)$.

The **Poincaré**–**Reidemeister metric** is a metric on the cohomological determinant line $detH^{\bullet}(M^n, F)$ of a closed connected oriented odd-dimensional manifold M^n . It can be constructed using a combination of the Reidemeister torsion with the Poincaré duality. Equivalently, one can define the Poincaré–Reidemeister scalar product on $detH^{\bullet}(M^n, F)$ which completely determines the Poincaré–Reidemeister metric but contains an additional sign or phase information.

The **Quillen metric** is a metric on the inverse of the cohomological determinant line of a compact Hermitian one-dimensional complex manifold. It can be defined as the product of the L_2 -metric with the Ray–Singler analytic torsion.

• Kähler supermetric

The **Kähler supermetric** is a generalization of the **Kähler metric** for the case of a *supermanifold*. A *supermanifold* is a generalization of the usual manifold with *fermonic* as well as *bosonic* coordinates. The bosonic coordinates are ordinary numbers, whereas the fermonic coordinates are *Grassmann numbers*.

• Hofer metric

A symplectic manifold (M^n, w) , n = 2k, is a smooth even-dimensional manifold M^n equipped with a symplectic form, i.e., a closed non-degenerate 2-form, w.

A Lagrangian manifold is a k-dimensional smooth submanifold L^k of a symplectic manifold (M^n, w) , n = 2k, such that the form w vanishes identically on L^k , i.e., for any $p \in L^k$ and any $x, y \in T_p(L^k)$, one has w(x,y) = 0.

Let $L(M^n, \Delta)$ be the set of all Lagrangian submanifolds of a closed symplectic manifold (M^n, w) , diffeomorphic to a given Lagrangian submanifold Δ . A smooth family $\alpha = \{L_t\}_t$, $t \in [0, 1]$, of Lagrangian submanifolds $L_t \in L(M^n, \Delta)$ is called an *exact path* connecting L_0 and L_1 , if there exists a smooth mapping $\Psi : \Delta \times [0, 1] \to M^n$ such that, for every $t \in [0, 1]$, one has $\Psi(\Delta \times \{t\}) = L_t$, and $\Psi * w = dH_t \wedge dt$ for some smooth function $H : \Delta \times [0, 1] \to \mathbb{R}$. The Hofer length $l(\alpha)$ of an exact path α is defined by $l(\alpha) = \int_0^1 \{ \max_{p \in \Delta} H(p, t) - \min_{p \in \Delta} H(p, t) \} dt$.

The **Hofer metric** on the set $L(M^n, \Delta)$ is defined by

$$\inf_{\alpha} l(\alpha)$$

for any $L_0, L_1 \in L(M^n, \Delta)$, where the infimum is taken over all exact paths on $L(M^n, \Delta)$, that connect L_0 and L_1 .

The Hofer metric can be defined in similar way on the group $Ham(M^n, w)$ of Hamiltonian diffeomorphisms of a closed symplectic manifold (M^n, w) , whose elements are time-one mappings of Hamiltonian flows ϕ_t^H : it is $\inf_{\alpha} l(\alpha)$, where the infimum is taken over all smooth paths $\alpha = \{\phi_t^H\}$, $t \in [0, 1]$, connecting ϕ and ψ .

• Sasakian metric

A **Sasakian metric** is a metric of positive scalar curvature on a *contact* manifold, naturally adapted to the *contact structure*.

A contact manifold equipped with a Sasakian metric is called a *Sasakian* space, and is an odd-dimensional analog of *Kähler manifolds*.

• Cartan metric

A Killing form (or Cartan–Killing form) on a finite-dimensional Lie algebra Ω over a field \mathbb{F} is a symmetric bilinear form

$$B(x,y) = Tr(ad_x \cdot ad_y),$$

where Tr denotes the trace of a linear operator, and ad_x is the image of x under the adjoint representation of Ω , i.e., the linear operator on the vector space Ω defined by the rule $z \to [x, z]$, where [,] is the Lie bracket.

Let $e_1, \ldots e_n$ be a basis for the Lie algebra Ω , and $[e_i, e_j] = \sum_{k=1}^n \gamma_{ij}^k e_k$, where γ_{ij}^k are corresponding *structure constants*. Then the Killing form is given by

$$B(x_i, x_j) = g_{ij} = \sum_{k,l=1}^{n} \gamma_{il}^k \gamma_{ik}^l.$$

In Theoretical Physics, the **metric tensor** $((g_{ij}))$ is called a **Cartan metric**.

Chapter 8 Distances on Surfaces and Knots

8.1 General surface metrics

A *surface* is a real two-dimensional *manifold* M^2 , i.e., a **Hausdorff space**, each point of which has a neighborhood which is homeomorphic to a plane \mathbb{E}^2 , or a closed half-plane (cf. Chap. 7).

A compact orientable surface is called *closed* if it has no boundary, and it is called a *surface with boundary*, otherwise. There are compact non-orientable surfaces (closed or with boundary); the simplest such surface is the *Möbius strip*. Non-compact surfaces without boundary are called *open*.

Any closed connected surface is homeomorphic to either a sphere with, say, g (cylindric) handles, or a sphere with, say, g cross-caps (i.e., caps with a twist like Möbius strip in them). In both cases the number g is called the genus of the surface. In the case of handles, the surface is orientable; it is called a torus (doughnut), double torus, and triple torus for g=1,2 and 3, respectively. In the case of cross-caps, the surface is non-orientable; it is called the real projective plane, Klein bottle, and Dyck's surface for g=1,2 and 3, respectively. The genus is the maximal number of disjoint simple closed curves which can be cut from a surface without disconnecting it (the Jordan curve theorem for surfaces).

The Euler-Poincaré characteristic of a surface is (the same for all polyhedral decompositions of a given surface) the number $\chi = v - e + f$, where v, e and f are, respectively, the number of vertices, edges and faces of the decomposition. Then $\chi = 2 - 2g$ if the surface is orientable, and $\chi = 2 - g$ if not. Every surface with boundary is homeomorphic to a sphere with an appropriate number of (disjoint) holes (i.e., what remains if an open disk is removed) and handles or cross-caps. If h is the number of holes, then $\chi = 2 - 2g - h$ holds if the surface is orientable, and $\chi = 2 - g - h$ if not.

The connectivity number of a surface is the largest number of closed cuts that can be made on the surface without separating it into two or more parts. This number is equal to $3-\chi$ for closed surfaces, and $2-\chi$ for surfaces with boundaries. A surface with connectivity number 1, 2 and 3 is called, respectively, simply, doubly and triply connected. A sphere is simply connected, while a torus is triply connected.

A surface can be considered as a metric space with its own **intrinsic metric**, or as a figure in space. A surface in \mathbb{E}^3 is called *complete* if it is a **complete** metric space with respect to its intrinsic metric.

A surface is called *differentiable*, *regular*, or *analytic*, respectively, if in a neighborhood of each of its points it can be given by an expression

$$r = r(u, v) = r(x_1(u, v), x_2(u, v), x_3(u, v)),$$

where the position vector r = r(u, v) is a differentiable, regular (i.e., a sufficient number of times differentiable), or real analytic, respectively, vector function satisfying the condition $r_u \times r_v \neq 0$.

Any regular surface has the intrinsic metric with the *line element* (or *first fundamental form*)

$$ds^{2} = dr^{2} = E(u, v)du^{2} + 2F(u, v)dudv + G(u, v)dv^{2},$$

where $E(u, v) = \langle r_u, r_u \rangle$, $F(u, v) = \langle r_u, r_v \rangle$, $G(u, v) = \langle r_v, r_v \rangle$. The length of a curve, defined on the surface by the equations u = u(t), v = v(t), $t \in [0, 1]$, is computed by

$$\int_0^1 \sqrt{Eu'^2 + 2Fu'v' + Gv'^2} dt,$$

and the distance between any points $p, q \in M^2$ is defined as the infimum of the lengths of all curves on M^2 , connecting p and q. A **Riemannian metric** is a generalization of the first fundamental form of a surface.

For surfaces, two kinds of curvature are considered: Gaussian curvature, and mean curvature. To compute these curvatures at a given point of the surface, consider the intersection of the surface with a plane, containing a fixed normal vector, i.e., a vector which is perpendicular to the surface at this point. This intersection is a plane curve. The curvature k of this plane curve is called the normal curvature of the surface at the given point. If we vary the plane, the normal curvature k will change, and there are two extremal values, the maximal curvature k_1 , and the minimal curvature k_2 , called the principal curvatures of the surface. A curvature is taken to be positive if the curve turns in the same direction as the surface's chosen normal, otherwise it is taken to be negative.

The Gaussian curvature is $K = k_1 k_2$ (it can be given entirely in terms of the first fundamental form). The mean curvature is $H = \frac{1}{2}(k_1 + k_2)$.

A minimal surface is a surface with mean curvature zero or, equivalently, a surface of minimum area subject to constraints on the location of its boundary.

A Riemann surface is a one-dimensional complex manifold, or a twodimensional real manifold with a complex structure, i.e., in which the local coordinates in neighborhoods of points are related by complex analytic functions. It can be thought of as a deformed version of the complex plane. All Riemann surfaces are orientable. Closed Riemann surfaces are geometrical models of complex algebraic curves. Every connected Riemann surface can be turned into a complete two-dimensional Riemannian manifold with constant curvature -1, 0, or 1. The Riemann surfaces with curvature -1 are called hyperbolic, and the unit disk $\Delta = \{z \in \mathbb{C} : |z| < 1\}$ is the canonical example. The Riemann surfaces with curvature 0 are called parabolic, and \mathbb{C} is a typical example. The Riemann surfaces with curvature 1 are called elliptic, and the Riemann sphere $\mathbb{C} \cup \{\infty\}$ is a typical example.

• Regular metric

The intrinsic metric of a surface is **regular** if it can be specified by the *line element*

$$ds^2 = Edu^2 + 2Fdudv + Gdv^2,$$

where the coefficients of the form ds^2 are regular functions.

Any regular surface, given by an expression r = r(u, v), has a regular metric with the *line element* ds^2 , where $E(u, v) = \langle r_u, r_u \rangle$, $F(u, v) = \langle r_u, r_v \rangle$, $G(u, v) = \langle r_v, r_v \rangle$.

• Analytic metric

The intrinsic metric on a surface is **analytic** if it can be specified by the *line element*

$$ds^2 = Edu^2 + 2Fdudv + Gdv^2,$$

where the coefficients of the form ds^2 are real analytic functions.

Any analytic surface, given by an expression r = r(u, v), has an analytic metric with the line element ds^2 , where $E(u, v) = \langle r_u, r_u \rangle$, $F(u, v) = \langle r_u, r_v \rangle$, $G(u, v) = \langle r_v, r_v \rangle$.

• Metric of positive curvature

A metric of positive curvature is the intrinsic metric on a *surface of* positive curvature, i.e., a surface in \mathbb{E}^3 that has positive Gaussian curvature at every point.

Metric of negative curvature

A metric of negative curvature is the intrinsic metric on a *surface* of negative curvature, i.e., a surface in \mathbb{E}^3 that has negative Gaussian curvature at every point.

A surface of negative curvature locally has a saddle-like structure. The intrinsic geometry of a surface of constant negative curvature (in particular, of a *pseudo-sphere*) locally coincides with the geometry of the *Lobachevsky plane*. There exists no surface in \mathbb{E}^3 whose intrinsic geometry coincides completely with the geometry of the Lobachevsky plane (i.e., a complete regular surface of constant negative curvature).

• Metric of non-positive curvature

A metric of non-positive curvature is the intrinsic metric on a *saddle-like surface*. A *saddle-like surface* is a generalization of a surface of negative curvature: a twice continuously-differentiable surface is a saddle-like surface if and only if at each point of the surface its Gaussian curvature is non-positive.

These surfaces can be seen as antipodes of *convex surfaces*, but they do not form such a natural class of surfaces as do convex surfaces.

• Metric of non-negative curvature

A metric of non-negative curvature is the intrinsic metric on a *convex* surface.

A convex surface is a domain (i.e., a connected open set) on the boundary of a convex body in \mathbb{E}^3 (in some sense, it is an antipode of a saddle-like surface).

The entire boundary of a convex body is called a *complete convex surface*. If the body is finite (bounded), the complete convex surface is called *closed*. Otherwise, it is called *infinite* (an infinite convex surface is homeomorphic to a plane or to a circular cylinder).

Any convex surface M^2 in \mathbb{E}^3 is a surface of bounded curvature. The total Gaussian curvature $w(A) = \int \int_A K(x) d\sigma(x)$ of a set $A \subset M^2$ is always non-negative (here $\sigma(.)$ is the area, and K(x) is the Gaussian curvature of M^2 at a point x), i.e., a convex surface can be seen as a surface of non-negative curvature.

The intrinsic metric of a convex surface is a **convex metric** (not to be confused with **metric convexity** from Chap. 1) in the sense of Surface Theory, i.e., it displays the *convexity condition*: the sum of the angles of any triangle whose sides are shortest curves is not less that π .

• Metric with alternating curvature

A metric with alternating curvature is the intrinsic metric on a surface with alternating (positive or negative) Gaussian curvature.

• Flat metric

A flat metric is the intrinsic metric on a *developable surface*, i.e., a surface, on which the Gaussian curvature is everywhere zero.

• Metric of bounded curvature

A metric of bounded curvature is the intrinsic metric ρ on a surface of bounded curvature.

A surface M^2 with an intrinsic metric ρ is called a *surface of bounded* curvature if there exists a sequence of **Riemannian metrics** ρ_n , defined on M^2 , such that for any compact set $A \subset M^2$ one has $\rho_n \to \rho$ uniformly, and the sequence $|w_n|(A)$ is bounded, where $|w|_n(A) = \int \int_A |K(x)| d\sigma(x)$ is the *total absolute curvature* of the metric ρ_n (here K(x) is the Gaussian curvature of M^2 at a point x, and $\sigma(.)$ is the *area*).

• Λ-metric

A Λ -metric (or *metric of type* Λ) is a **complete** metric on a surface with curvature bounded from above by a negative constant.

A Λ -metric does not have embeddings into \mathbb{E}^3 . It is a generalization of the classical result of Hilbert (1901): no complete regular surface of constant negative curvature (i.e., a surface whose intrinsic geometry coincides completely with the geometry of the Lobachevsky plane) exists in \mathbb{E}^3 .

• (h, Δ) -metric

A (h, Δ) -metric is a metric on a surface with a slowly-changing negative curvature.

A complete (h, Δ) -metric does not permit a regular isometric embedding in three-dimensional Euclidean space (cf. Λ -metric).

• G-distance

A connected set G of points on a surface M^2 is called a *geodesic region* if, for each point $x \in G$, there exists a $disk\ B(x,r)$ with center at x, such that $B_G = G \cap B(x,r)$ has one of the following forms: $B_G = B(x,r)$ (x is a regular interior point of G); B_G is a semi-disk of B(x,r) (x is a regular boundary point of G); B_G is a sector of B(x,r) other than a semi-disk (x is an angular point of G); B_G consists of a finite number of sectors of B(x,r) with no common points except x (a nodal point of G).

The G-distance between any x and $y \in G$ is the greatest lower bound of the lengths of all rectifiable curves connecting x and $y \in G$ and completely contained in G.

• Conformally invariant metric

Let R be a Riemann surface. A local parameter (or local uniformizing parameter, local uniformizer) is a complex variable z considered as a continuous function $z_{p_0} = \phi_{p_0}(p)$ of a point $p \in R$ which is defined everywhere in some neighborhood (parametric neighborhood) $V(p_0)$ of a point $p_0 \in R$ and which realizes a homeomorphic mapping (parametric mapping) of $V(p_0)$ onto the disk (parametric disk) $\Delta(p_0) = \{z \in \mathbb{C} : |z| < r(p_0)\}$, where $\phi_{p_0}(p_0) = 0$. Under a parametric mapping, any point function g(p), defined in the parametric neighborhood $V(p_0)$, goes into a function of the local parameter z: $g(p) = g(\phi_{p_0}^{-1}(z)) = G(z)$.

A conformally invariant metric is a differential $\rho(z)|dz|$ on the Riemann surface R which is invariant with respect to the choice of the local parameter z. Thus, to each local parameter z ($z : U \to \overline{\mathbb{C}}$) a function $\rho_z : z(U) \to [0, \infty]$ is associated such that, for any local parameters z_1 and z_2 , we have

$$\frac{\rho_{z_2}(z_2(p))}{\rho_{z_1}(z_1(p))} = \left|\frac{dz_1(p)}{dz_2(p)}\right| \text{ for any } p \in U_1 \cap U_2.$$

Every linear differential $\lambda(z)dz$ and every quadratic differential $Q(z)dz^2$ induce conformally invariant metrics $|\lambda(z)||dz|$ and $|Q(z)|^{1/2}||dz|$, respectively (cf. Q-metric).

• Q-metric

An Q-metric is a conformally invariant metric $\rho(z)|dz| = |Q(z)|^{1/2}$ |dz| on a Riemann surface R, defined by a quadratic differential $Q(z)dz^2$.

A quadratic differential $Q(z)dz^2$ is a non-linear differential on a Riemann surface R which is invariant with respect to the choice of the local parameter z. Thus, to each local parameter z ($z: U \to \overline{\mathbb{C}}$) a function $Q_z: z(U) \to \overline{\mathbb{C}}$ is associated such that, for any local parameters z_1 and z_2 , we have

$$\frac{Q_{z_2}(z_2(p))}{Q_{z_1}(z_1(p))} = \left(\frac{dz_1(p)}{dz_2(p)}\right)^2 \text{ for any } p \in U_1 \cap U_2.$$

• Extremal metric

An extremal metric is a conformally invariant metric in the modulus problem for a family Γ of locally rectifiable curves on a Riemann surface R which realizes the infimum in the definition of the modulus $M(\Gamma)$.

Formally, let Γ be a family of locally rectifiable curves on a Riemann surface R, let P be a non-empty class of conformally invariant metrics $\rho(z)|dz|$ on R such that $\rho(z)$ is square-integrable in the z-plane for every local parameter z, and the integrals

$$A_{\rho}(R) = \int \int_{R} \rho^{2}(z) dx dy \ \text{ and } \ L_{\rho}(\Gamma) = \inf_{\gamma \in \Gamma} \int_{y} \rho(z) |dz|$$

are not simultaneously equal to 0 or ∞ (each of the above integrals is understood as a Lebesgue integral). The modulus of the family of curves Γ is defined by

$$M(\Gamma) = \inf_{\rho \in P} \frac{A_{\rho}(R)}{(L_{\rho}(\Gamma))^2}.$$

The extremal length of the family of curves Γ is equal to $\sup_{\rho \in P} \frac{(L_{\rho}(\Gamma))^2}{A_{\rho}(R)}$, i.e., is the reciprocal of $M(\Gamma)$.

The modulus problem for Γ is defined as follows: let P_L be the subclass of P such that, for any $\rho(z)|dz| \in P_L$ and any $\gamma \in \Gamma$, one has $\int_{\gamma} \rho(z)|dz| \geq 1$. If $P_L \neq \emptyset$, then the modulus $M(\Gamma)$ of the family Γ can be written as $M(\Gamma) = \inf_{\rho \in P_L} A_{\rho}(R)$. Every metric from P_L is called an **admissible** metric for the modulus problem on Γ . If there exists ρ^* for which

$$M(\Gamma) = \inf_{\rho \in P_L} A_{\rho}(R) = A_{\rho^*}(R),$$

the metric $\rho^*|dz|$ is called an **extremal metric** for the modulus problem on Γ .

• Fréchet surface metric

Let (X,d) be a metric space, M^2 a compact two-dimensional manifold, f a continuous mapping $f: M^2 \to X$, called a parameterized surface, and $\sigma: M^2 \to M^2$ a homeomorphism of M^2 onto itself. Two parameterized surfaces f_1 and f_2 are called equivalent if $\inf_{\sigma} \max_{p \in M^2} d(f_1(p), f_2(\sigma(p))) = 0$, where the infimum is taken over all possible homeomorphisms σ . A class f^* of parameterized surfaces, equivalent to f, is called a Fréchet surface. It is a generalization of the notion of a surface in Euclidean space to the case of an arbitrary metric space (X,d).

The **Fréchet surface metric** on the set of all Fréchet surfaces is defined by

$$\inf_{\sigma} \max_{p \in M^2} d(f_1(p), f_2(\sigma(p)))$$

for any Fréchet surfaces f_1^* and f_2^* , where the infimum is taken over all possible homeomorphisms σ . Cf. the **Fréchet metric** in Chap. 1.

8.2 Intrinsic metrics on surfaces

In this section we list intrinsic metrics, given by their *line elements* (which, in fact, are two-dimensional **Riemannian metrics**), for some selected surfaces.

• Quadric metric

A quadric (or quadratic surface, surface of second order) is a set of points in \mathbb{E}^3 , whose coordinates in a Cartesian coordinate system satisfy an algebraic equation of degree two. There are 17 classes of such surfaces. Among them are: ellipsoids, one-sheet and two-sheet hyperbolicids, elliptic paraboloids, hyperbolic paraboloids, elliptic, hyperbolic and parabolic cylinders, and conical surfaces.

For example, a *cylinder* can be given by the following parametric equations:

$$x_1(u, v) = a \cos v, \ x_2(u, v) = a \sin v, \ x_3(u, v) = u.$$

The intrinsic metric on it is given by the line element

$$ds^2 = du^2 + a^2 dv^2.$$

An *elliptic cone* (i.e., a cone with elliptical cross-section) has the following equations:

$$x_1(u,v) = a\frac{h-u}{h}\cos v, \ x_2(u,v) = b\frac{h-u}{h}\sin v, \ x_3(u,v) = u,$$

where h is the *height*, a is the *semi-major axis*, and b is the *semi-minor axis* of the cone. The intrinsic metric on it is given by the *line element*

$$ds^{2} = \frac{h^{2} + a^{2}\cos^{2}v + b^{2}\sin^{2}v}{h^{2}}du^{2} + 2\frac{(a^{2} - b^{2})(h - u)\cos v \sin v}{h^{2}}dudv +$$

$$+\frac{(h-u)^2(a^2\sin^2v+b^2\cos^2v)}{b^2}dv^2.$$

• Sphere metric

A sphere is a quadric, given by the Cartesian equation $(x_1 - a)^2 + (x_2 - b)^2 + (x_3 - c)^2 = r^2$, where the point (a, b, c) is the center of the sphere,

and r > 0 is the radius of the sphere. The sphere of radius r, centered at the origin, can be given by the following parametric equations:

$$x_1(\theta,\phi) = r\sin\theta\cos\phi, \ x_2(\theta,\phi) = r\sin\theta\sin\phi, \ x_3(\theta,\phi) = r\cos\theta,$$

where the azimuthal angle $\phi \in [0, 2\pi)$, and the polar angle $\theta \in [0, \pi]$.

The intrinsic metric on it (in fact, the two-dimensional **spherical metric**) is given by the *line element*

$$ds^2 = r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2.$$

A sphere of radius r has constant positive Gaussian curvature equal to r.

• Ellipsoid metric

An *ellipsoid* is a *quadric* given by the Cartesian equation $\frac{x_1^2}{a^2} + \frac{x_2^2}{b^2} + \frac{x_3^2}{c^2} = 1$, or by the following parametric equations:

$$x_1(\theta,\phi) = a\cos\phi\sin\theta, \ x_2(\theta,\phi) = b\sin\phi\sin\theta, \ x_3(\theta,\phi) = c\cos\theta,$$

where the azimuthal angle $\phi \in [0, 2\pi)$, and the polar angle $\theta \in [0, \pi]$. The intrinsic metric on it is given by the line element

$$ds^{2} = (b^{2}\cos^{2}\phi + a^{2}\sin^{2}\phi)\sin^{2}\theta d\phi^{2} + (b^{2} - a^{2})\cos\phi\sin\phi\cos\theta\sin\theta d\theta d\phi +$$
$$+((a^{2}\cos^{2}\phi + b^{2}\sin^{2}\phi)\cos^{2}\theta + c^{2}\sin^{2}\theta)d\theta^{2}.$$

• Spheroid metric

A spheroid is an ellipsoid having two axes of equal length. It is also a rotation surface, given by the following parametric equations:

$$x_1(u, v) = a \sin v \cos u, \ x_2(u, v) = a \sin v \sin u, \ x_3(u, v) = c \cos v,$$

where $0 \le u < 2\pi$, and $0 \le v \le \pi$.

The intrinsic metric on it is given by the line element

$$ds^2 = a^2 \sin^2 v du^2 + \frac{1}{2}(a^2 + c^2 + (a^2 - c^2)\cos(2v))dv^2.$$

• Hyperboloid metric

A hyperboloid is a quadric which may be one- or two-sheeted. The one-sheeted hyperboloid is a surface of revolution obtained by rotating a hyperbola about the perpendicular bisector to the line between the foci, while the two-sheeted hyperboloid is a surface of revolution obtained by rotating a hyperbola about the line joining the foci. The one-sheeted circular hyperboloid, oriented along the x_3 -axis, is given by the Cartesian equation $\frac{x_1^2}{a^2} + \frac{x_2^2}{a^2} - \frac{x_3^2}{c^2} = 1$, or by the following parametric equations:

$$x_1(u,v) = a\sqrt{1+u^2}\cos v, \ x_2(u,v) = a\sqrt{1+u^2}\sin v, \ x_3(u,v) = cu,$$

where $v \in [0, 2\pi)$. The intrinsic metric on it is given by the line element

$$ds^{2} = \left(c^{2} + \frac{a^{2}u^{2}}{u^{2} + 1}\right)du^{2} + a^{2}(u^{2} + 1)dv^{2}.$$

• Rotation surface metric

A rotation surface (or surface of revolution) is a surface generated by rotating a two-dimensional curve about an axis. It is given by the following parametric equations:

$$x_1(u, v) = \phi(v)\cos u, \ x_2(u, v) = \phi(v)\sin u, \ x_3(u, v) = \psi(v).$$

The intrinsic metric on it is given by the line element

$$ds^{2} = \phi^{2}du^{2} + (\phi^{'2} + \psi^{'2})dv^{2}.$$

• Pseudo-sphere metric

A pseudo-sphere is a half of the rotation surface generated by rotating a tractrix about its asymptote. It is given by the following parametric equations:

$$x_1(u,v) = \operatorname{sech} u \cos v, \ x_2(u,v) = \operatorname{sech} u \sin v, \ x_3(u,v) = u - \tanh u,$$

where $u \ge 0$, and $0 \le v < 2\pi$. The intrinsic metric on it is given by the line element

$$ds^2 = \tanh^2 u du^2 + \operatorname{sech}^2 u dv^2.$$

The pseudo-sphere has constant negative Gaussian curvature equal to -1, and in this sense is an analog of the sphere which has constant positive Gaussian curvature.

• Torus metric

A torus is a surface having genus one. A torus azimuthally symmetric about the x_3 -axis is given by the Cartesian equation $(c - \sqrt{x_1^2 + x_2^2})^2 + x_3^2 = a^2$, or by the following parametric equations:

$$x_1(u, v) = (c + a\cos v)\cos u, \ x_2(u, v) = (c + a\cos v)\sin u, \ x_3(u, v) = a\sin v,$$

where c > a, and $u, v \in [0, 2\pi)$.

The intrinsic metric on it is given by the line element

$$ds^{2} = (c + a\cos v)^{2}du^{2} + a^{2}dv^{2}.$$

• Helical surface metric

A helical surface (or surface of screw motion) is a surface described by a plane curve γ which, while rotating around an axis at a uniform rate, also

advances along that axis at a uniform rate. If γ is located in the plane of the axis of rotation x_3 and is defined by the equation $x_3 = f(u)$, the position vector of the helical surface is

$$r = (u\cos v, u\sin v, f(u) = hv), h = \text{const},$$

and the intrinsic metric on it is given by the line element

$$ds^{2} = (1 + f'^{2})du^{2} + 2hf'dudv + (u^{2} + h^{2})dv^{2}.$$

If f = const, one has a helicoid; if h = 0, one has a rotation surface.

• Catalan surface metric

The Catalan surface is a minimal surface, given by the following equations:

$$x_1(u,v) = u - \sin u \cosh v, \ x_2(u,v) = 1 - \cos u \cosh v,$$
$$x_3(u,v) = 4 \sin\left(\frac{u}{2}\right) \sinh\left(\frac{v}{2}\right).$$

The intrinsic metric on it is given by the line element

$$ds^{2} = 2\cosh^{2}\left(\frac{v}{2}\right)\left(\cosh v - \cos u\right)du^{2} + 2\cosh^{2}\left(\frac{v}{2}\right)\left(\cosh v - \cos u\right)dv^{2}.$$

• Monkey saddle metric

The monkey saddle is a surface, given by the Cartesian equation $x_3 = x_1(x_1^2 - 3x_2^2)$, or by the following parametric equations:

$$x_1(u,v) = u$$
, $x_2(u,v) = v$, $x_3(u,v) = u^3 - 3uv^2$.

This is a surface which a monkey can straddle with both legs and his tail. The intrinsic metric on it is given by the *line element*

$$ds^2 = (1 + (su^2 - 3v^2)^2)du^2 - 2(18uv(u^2 - v^2))dudv + (1 + 36u^2v^2)dv^2).$$

8.3 Distances on knots

A knot is a closed, non-self-intersecting curve that is embedded in S^3 . The $trivial\ knot$ (or unknot) O is a closed loop that is not knotted. A knot can be generalized to a link which is a set of disjoint knots. Every link has its $Seifert\ surface$, i.e., a compact oriented surface with the given link as boundary.

Two knots (links) are called *equivalent* if one can be smoothly deformed into another. Formally, a link is defined as a smooth one-dimensional *submanifold* of the 3-sphere S^3 ; a knot is a link consisting of one component; two

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links L_1 and L_2 are called *equivalent* if there exists an orientation-preserving homeomorphism $f: S^3 \to S^3$ such that $f(L_1) = L_2$.

All the information about a knot can be described using a *knot diagram*. It is a projection of a knot onto a plane such that no more than two points of the knot are projected to the same point on the plane, and at each such point it is indicated which strand is closest to the plane, usually by erasing part of the lower strand. Two different knot diagrams may both represent the same knot. Much of Knot Theory is devoted to telling when two knot diagrams represent the same knot.

An unknotting operation is an operation which changes the overcrossing and the undercrossing at a double point of a given knot diagram. The unknotting number of a knot K is the minimum number of unknotting operations needed to deform a diagram of K into that of the trivial knot, where the minimum is taken over all diagrams of K. Roughly, the unknotting number is the smallest number of times a knot K must be passed through itself to until it.

An \sharp -unknotting operation in a diagram of a knot K is an analog of the unknotting operation for a \sharp -part of the diagram consisting of two pairs of parallel strands with one of the pair overcrossing another. Thus, an \sharp -unknotting operation changes the overcrossing and the undercrossing at each vertex of obtained quadrangle.

• Gordian distance

The **Gordian distance** is a metric on the set of all knots defined, for given knots K and K', as the minimum number of unknotting operations needed to deform a diagram of K into that of K', where the minimum is taken over all diagrams of K from which one can obtain diagrams of K'. The unknotting number of K is equal to the Gordian distance between K and the trivial knot O.

Let rK be the knot obtained from K by taking its mirror image, and let -K be the knot with the reversed orientation. The **positive reflection distance** $Ref_+(K)$ is the Gordian distance between K and rK. The **negative reflection distance** $Ref_-(K)$ is the Gordian distance between K and -rK. The **inversion distance** Inv(K) is the Gordian distance between K and -K.

The Gordian distance is the case k=1 of the C_k -distance which is the minimum number of C_k -moves needed to transform K into K'; Habiro (1994) and Goussarov (1995), independently proved that, for k>1, it is finite if and only if both knots have the same Vassiliev invariants of order less than k. A C_1 -move is a single crossing change, a C_2 -move (or delta-move) is a simultaneous crossing change for three arcs forming a triangle. C_2 - and C_3 -distances are called delta distance and clasp-pass distance, respectively.

• #-Gordian distance

The \sharp -Gordian distance (see, for example, [Mura85]) is a metric on the set of all knots defined, for given knots K and K', as the minimum number

of \sharp -unknotting operations needed to deform a diagram of K into that of K', where the minimum is taken over all diagrams of K from which one can obtain diagrams of K'.

Let rK be the knot obtained from K by taking its mirror image, and let -K be the knot with the reversed orientation. The **positive** \sharp -reflection distance $Ref_+^{\sharp}(K)$ is the \sharp -Gordian distance between K and rK. The negative \sharp -reflection distance $Ref_-^{\sharp}(K)$ is the \sharp -Gordian distance between K and -rK. The \sharp -inversion distance $Inv^{\sharp}(K)$ is the \sharp -Gordian distance between K and -K.

• Knot complement hyperbolic metric

The *complement* of a knot K (or a link L) is $S^3 \setminus K$ (or $S^3 \setminus L$, respectively). A knot (or, in general, a link) is called *hyperbolic* if its complement supports a complete Riemannian metric of constant curvature -1. In this case, the metric is called a **knot** (or **link**) **complement hyperbolic metric**, and it is unique.

A knot is hyperbolic if and only if (Thurston 1978) it is not a *satellite* knot (then it supports a complete locally homogeneous Riemannian metric) and not a torus knot (does not lie on a trivially embedded torus in S^3). The complement of any non-trivial knot supports a complete non-positively curved Riemannian metric.

Chapter 9

Distances on Convex Bodies, Cones, and Simplicial Complexes

9.1 Distances on convex bodies

A convex body in the n-dimensional Euclidean space \mathbb{E}^n is a compact convex subset of \mathbb{E}^n . It is called proper if it has non-empty interior. Let K denote the space of all convex bodies in \mathbb{E}^n , and let K_p be the subspace of all proper convex bodies.

Any metric space (K, d) on K is called a *metric space of convex bodies*. Metric spaces of convex bodies, in particular the metrization by the **Hausdorff metric**, or by the **symmetric difference metric**, play a basic role in the foundations of analysis in Convex Geometry (see, for example, [Grub93]).

For $C, D \in K \setminus \{\emptyset\}$ the Minkowski addition and the Minkowski non-negative scalar multiplication are defined by $C+D=\{x+y:x\in C,y\in D\}$, and $\alpha C=\{\alpha x:x\in C\},\ \alpha\geq 0$, respectively. The Abelian semi-group (K,+) equipped with non-negative scalar multiplication operators can be considered as a convex cone.

The support function $h_C: S^{n-1} \to \mathbb{R}$ of $C \in K$ is defined by $h_C(u) = \sup\{\langle u, x \rangle : x \in C\}$ for any $u \in S^{n-1}$, where S^{n-1} is the (n-1)-dimensional unit sphere in \mathbb{E}^n , and \langle , \rangle is the inner product in \mathbb{E}^n .

Given a set $X \subset \mathbb{E}^n$, its $convex\ hull\ conv(X)$ is the minimal convex set containing X.

• Area deviation

The area deviation (or template metric) is a metric on the set K_p in \mathbb{E}^2 (i.e., on the set of plane convex disks), defined by

$$A(C\triangle D)$$
,

where A(.) is the area, and \triangle is the symmetric difference. If $C \subset D$, then it is equal to A(D) - A(C).

• Perimeter deviation

The **perimeter deviation** is a metric on K_p in \mathbb{E}^2 , defined by

$$2p(conv(C \cup D)) - p(C) - p(D),$$

where p(.) is the *perimeter*. In the case $C \subset D$, it is equal to p(D) - p(C).

• Mean width metric

The **mean width metric** is a metric on K_p in \mathbb{E}^2 , defined by

$$2W(conv(C \cup D)) - W(C) - W(D),$$

where W(.) is the mean width: $W(C) = p(C)/\pi$, and p(.) is the perimeter.

• Pompeiu-Hausdorff-Blaschke metric

The **Pompeiu–Hausdorff–Blaschke metric** is a metric on K, defined by

$$\max\{\sup_{x \in C} \inf_{y \in D} ||x - y||_2, \sup_{y \in D} \inf_{x \in C} ||x - y||_2\},$$

where $||.||_2$ is the Euclidean norm on \mathbb{E}^n .

In terms of support functions, using Minkowski addition, the metric is

$$\sup_{u \in S^{n-1}} |h_C(u) - h_D(u)| = ||h_C - h_D||_{\infty} =$$

$$= \inf\{\lambda \ge 0 : C \subset D + \lambda \overline{B}^n, D \subset C + \lambda \overline{B}^n\},\$$

where \overline{B}^n is the *unit ball* of \mathbb{E}^n .

This metric can be defined using any norm on \mathbb{R}^n instead of the Euclidean norm. It can be defined for the space of bounded closed subsets of any metric space.

• Pompeiu-Eggleston metric

The **Pompeiu–Eggleston metric** is a metric on K, defined by

$$\sup_{x \in C} \inf_{y \in D} ||x - y||_2 + \sup_{y \in D} \inf_{x \in C} ||x - y||_2,$$

where $||.||_2$ is the Euclidean norm on \mathbb{E}^n .

In terms of support functions, using Minkowski addition, the metric is

$$\max\{0, \sup_{u \in S^{n-1}} (h_C(u) - h_D(u))\} + \max\{0, \sup_{u \in S^{n-1}} (h_D(u) - h_C(u))\} =$$

$$=\inf\{\lambda\geq 0: C\subset D+\lambda\overline{B}^n\}+\inf\{\lambda\geq 0: D\subset C+\lambda\overline{B}^n\},$$

where \overline{B}^n is the unit ball of \mathbb{E}^n .

This metric can be defined using any norm on \mathbb{R}^n instead of the Euclidean norm. It can be defined for the space of bounded closed subsets of any metric space.

• McClure-Vitale metric

Given $1 \le p \le \infty$, the **McClure–Vitale metric** is a metric on K, defined by

$$\left(\int_{S^{n-1}} |h_C(u) - h_D(u)|^p d\sigma(u)\right)^{\frac{1}{p}} = ||h_C - h_D||_p.$$

• Florian metric

The **Florian metric** is a metric on K, defined by

$$\int_{S^{n-1}} |h_C(u) - h_D(u)| d\sigma(u) = ||h_C - h_D||_1.$$

It can be expressed in the form $2S(conv(C \cup D)) - S(C) - S(D)$ for n = 2 (cf. **perimeter deviation**); it can be expressed also in the form $nk_n(2W(conv(C \cup D)) - W(C) - W(D))$ for $n \ge 2$ (cf. **mean width metric**).

Here S(.) is the surface area, k_n is the volume of the unit ball \overline{B}^n of \mathbb{E}^n , and W(.) is the mean width: $W(C) = \frac{1}{nk_n} \int_{S^{n-1}} (h_C(u) + h_C(-u)) d\sigma(u)$.

• Sobolev distance

The **Sobolev distance** is a metric on K, defined by

$$||h_C - h_D||_w$$

where $||.||_w$ is the *Sobolev 1-norm* on the set $G_{S^{n-1}}$ of all real continuous functions on the *unit sphere* S^{n-1} of \mathbb{E}^n .

The Sobolev 1-norm is defined by $||f||_w = \langle f, f \rangle_w^{1/2}$, where \langle , \rangle_w is an inner product on $G_{S^{n-1}}$, given by

$$\langle f, g \rangle_w = \int_{S^{n-1}} (fg + \nabla_s(f, g)) dw_0, \ w_0 = \frac{1}{n \cdot k_n} w,$$

 $\nabla_s(f,g) = \langle grad_s f, grad_s g \rangle, \langle, \rangle$ is the inner product in \mathbb{E}^n , and $grad_s$ is the gradient on S^{n-1} (see [ArWe92]).

• Shephard metric

The **Shephard metric** is a metric on K_p , defined by

$$\ln(1+2\inf\{\lambda\geq 0: C\subset D+\lambda(D-D), D\subset C+\lambda(C-C)\}).$$

Nikodym metric

The **Nikodym metric** is a metric on K_p , defined by

$$V(C\triangle D)$$
,

where V(.) is the *volume* (i.e., the Lebesgue *n*-dimensional measure), and \triangle is the *symmetric difference*. For n=2, one obtains the **area deviation**.

• Steinhaus metric

The Steinhaus metric (or homogeneous symmetric difference metric, Steinhaus distance) is a metric on K_p , defined by

$$\frac{V(C\triangle D)}{V(C\cup D)},$$

where V(.) is the *volume*. So, it is $\frac{d_{\triangle}(C,D)}{V(C\cup D)}$, where d_{\triangle} is the **Nikodym** metric.

This metric is **bounded**; it is affine invariant, while the Nikodym metric is invariant only under volume-preserving affine transformations.

• Eggleston distance

The Eggleston distance (or symmetric surface area deviation) is a distance on K_p , defined by

$$S(C \cup D) - S(C \cap D),$$

where S(.) is the *surface area*. It is not a metric.

• Asplund metric

The **Asplund metric** is a metric on the space K_p/\approx of affine-equivalence classes in K_p , defined by

$$\ln\inf\{\lambda \geq 1: \exists T: \mathbb{E}^n \to \mathbb{E}^n \text{ affine, } x \in \mathbb{E}^n, C \subset T(D) \subset \lambda C + x\}$$

for any equivalence classes C^* and D^* with the representatives C and D, respectively.

• Macbeath metric

The Macbeath metric is a metric on the space K_p/\approx of affine-equivalence classes in K_p , defined by

$$\ln\inf\{|\det T\cdot P|:\exists T,P:\mathbb{E}^n\to\mathbb{E}^n \text{ regular affine}, C\subset T(D), D\subset P(C)\}$$

for any equivalence classes C^* and D^* with the representatives C and D, respectively.

Equivalently, it can be written as

$$\ln \delta_1(C, D) + \ln \delta_1(D, C),$$

where $\delta_1(C,D) = \inf_T \{ \frac{V(T(D))}{V(C)}; C \subset T(D) \}$, and T is a regular affine mapping of \mathbb{E}^n onto itself.

• Banach-Mazur metric

The **Banach–Mazur metric** is a metric on the space K_{po}/\sim of the equivalence classes of proper 0-symmetric convex bodies with respect to linear transformations, defined by

$$\ln\inf\{\lambda\geq 1: \exists T: \mathbb{E}^n \to \mathbb{E}^n \text{ linear, } C\subset T(D)\subset \lambda C\}$$

for any equivalence classes C^* and D^* with the representatives C and D, respectively.

It is a special case of the **Banach–Mazur distance** between *n*-dimensional *normed spaces*.

• Separation distance

The **separation distance** between two disjoint convex bodies C and D in \mathbb{E}^n is (Buckley 1985) the minimum Euclidean distance (in general, the **set-set distance** between any two disjoint subsets of \mathbb{E}^n): $\inf\{||x-y||_2 : x \in C, y \in D\}$, while $\sup\{||x-y||_2 : x \in C, y \in D\}$ is called the **spanning distance**.

• Penetration depth distance

The **penetration depth distance** between two inter-penetrating convex bodies C and D in \mathbb{E}^n (in general, between any two inter-penetrating subsets of \mathbb{E}^n) is (Cameron and Culley 1986) defined as the minimum translation distance that one body undergoes to make the interiors of C and D disjoint:

$$\min\{||t||_2 : interior(C+t) \cap D = \emptyset\}.$$

Keerthi and Sridharan (1991) considered $||t||_{1}$ - and $||t||_{\infty}$ -analogues of the above definition.

Cf. penetration distance in Chap. 23 and penetration depth in Chap. 24.

• Growth distances

Let $C, D \in K_p$ be two compact convex bodies with non-empty interior. Fix their seed points $p_C \in int C$ and $p_D \in int D$; usually, they are the centroids of C and D. The growth function g(C, D) is the minimal number $\lambda > 0$, such that

$$(\{p_C\} + \lambda(C \setminus \{p_C\})) \cap (\{p_D\} + \lambda(D \setminus \{p_D\})) \neq \emptyset.$$

It is the amount objects must be grown if g(C, D) > 1 (i.e., $C \cap D = \emptyset$), or contracted if g(C, D) > 1 (i.e., $int C \cap int D \neq \emptyset$) from their internal seed points until their surfaces just touch. The **growth separation distance** $d_S(C, D)$ and the **growth penetration distance** $d_P(C, D)$ [OnGi96] are defined as

$$d_S(C,D) = \max\{0, r_{CD}(g(C,D)-1)\} \text{ and }$$

$$d_S(C, D) = \max\{0, r_{CD}(1 - g(C, D))\},\$$

where r_{CD} is the scaling coefficient (usually, the sum of radii of circumscribing spheres for the sets $C \setminus \{p_C\}$ and $D \setminus \{p_D\}$).

The one-sided growth distance between disjoint C and D (Leven and Sharir 1987) is $-1 + \min \lambda > 0 : (\{p_C\} + \lambda\{(C \setminus \{p_C\})\}) \cap D \neq \emptyset\}$.

• Minkowski difference

The **Minkowski difference** on the set of all compact subsets, in particular, on the set of all *sculptured objects* (or *free form objects*), of \mathbb{R}^3 is defined by

$$A - B = \{x - y : x \in A, y \in B\}.$$

If we consider object B to be free to move with fixed orientation, the Minkowski difference is a set containing all the translations that bring B to intersect with A. The closest point from the Minkowski difference boundary, $\partial(A-B)$, to the origin gives the **separation distance** between A and B.

If both objects intersect, the origin is inside of their Minkowski difference, and the obtained distance can be interpreted as a **penetration** depth distance.

• Demyanov distance

Given $C \in K_p$ and $u \in S^{n-1}$, denote, if $|\{c \in C : \langle u, c \rangle = h_C(u)\}| = 1$, this unique point by y(u, C) (exposed point of C in direction u).

The Demyanov difference $A \ominus B$ of two subsets $A, B \in K_p$ is the closure of

$$conv(\cup_{T(A)\cap T(B)}\{y(u,A)-y(u,B)\}),$$

where
$$T(C) = \{u \in S^{n-1} : |\{c \in C : \langle u, c \rangle = h_C(u)\}| = 1\}.$$

The **Demyanov distance** between two subsets $A, B \in K_p$ is defined by

$$||A \ominus B|| = \max_{c \in A \ominus B} ||c||_2.$$

It is shown in [BaFa07] that $||A \ominus B|| = \sup_{\alpha} ||St_{\alpha}(A) - St_{\alpha}(M)||_2$, where $St_{\alpha}(C)$ is a generalized Steiner point and the supremum is over all "sufficiently smooth" probabilistic measures α .

• Maximum polygon distance

The **maximum polygon distance** is a distance between two convex polygons $P = (p_1, \ldots, p_n)$ and $Q = (q_1, \ldots, q_m)$, defined by

$$\max_{i,j} ||p_i - q_j||_2, \ i \in \{1, \dots, n\}, \ j \in \{1, \dots, m\},$$

where $||.||_2$ is the Euclidean norm.

• Grenander distance

Let $P = (p_1, \ldots, p_n)$ and $Q = (q_1, \ldots, q_m)$ be two disjoint convex polygons, and let $L(p_i, q_j), L(p_l, q_m)$ be two intersecting *critical support lines* for P and Q. Then the **Grenander distance** between P and Q is defined by

$$||p_i - q_j||_2 + ||p_l - q_m||_2 - \Sigma(p_i, p_l) - \Sigma(g_j, q_m),$$

where $||.||_2$ is the Euclidean norm, and $\Sigma(p_i, p_l)$ is the sum of the edges lengths of the polynomial chain p_i, \ldots, p_l .

Here $P = (p_1, \ldots, p_n)$ is a convex polygon with the vertices in standard form, i.e., the vertices are specified according to cartesian coordinates in a clockwise order, and no three consecutive vertices are collinear. A line L is a line of support of P if the interior of P lies completely to one side

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of L. Given two disjoint polygons P and Q, the line $L(p_i, q_j)$ is a *critical* support line if it is a line of support for P at p_i , a line of support for Q at q_j , and P and Q lie on opposite sides of $L(p_i, q_j)$.

9.2 Distances on cones

A convex cone C in a real vector space V is a subset C of V such that $C+C\subset C$, $\lambda C\subset C$ for any $\lambda\geq 0$, and $C\cap (-C)=\{0\}$. A cone C induces a partial order on V by

$$x \leq y$$
 if and only if $y - x \in C$.

The order \leq respects the vector structure of V, i.e., if $x \leq y$ and $z \leq u$, then $x+z \leq y+u$, and if $x \leq y$, then $\lambda x \leq \lambda y$, $\lambda \in \mathbb{R}$, $\lambda \geq 0$. Elements $x,y \in V$ are called *comparable* and denoted by $x \sim y$ if there exist positive real numbers α and β such that $\alpha y \leq x \leq \beta y$. Comparability is an equivalence relation; its equivalence classes (which belong to C or to C) are called *parts* (or *components*, *constituents*).

Given a convex cone C, a subset $S = \{x \in C : T(x) = 1\}$, where $T : V \to \mathbb{R}$ is some positive linear functional, is called a *cross-section* of C.

A convex cone C is called *almost Archimedean* if the closure of its restriction to any two-dimensional subspace is also a cone.

• Thompson part metric

Given a convex cone C in a real vector space V, the **Thompson part** metric on a part $K \subset C \setminus \{0\}$ is defined by

$$\ln \max\{m(x,y),m(y,x)\}$$

for any $x, y \in K$, where $m(x, y) = \inf\{\lambda \in \mathbb{R} : y \leq \lambda x\}$.

If C is almost Archimedean, then K equipped with the Thompson part metric is a **complete** metric space. If C is finite-dimensional, then one obtains a chord space, i.e., a metric space in which there is a distinguished set of geodesics, satisfying certain axioms. The positive cone $\mathbb{R}^n_+ = \{(x_1, \ldots, x_n) : x_i \geq 0 \text{ for } 1 \leq i \leq n\}$ equipped with the Thompson part metric is isometric to a normed space which one may think of as being flat.

If C is a closed cone in \mathbb{R}^n with non-empty interior, then $int \ C$ can be considered as an n-dimensional manifold M^n . If for any tangent vector $v \in T_p(M^n)$, $p \in M^n$, we define a norm $||v||_p^T = \inf\{\alpha > 0 : -\alpha p \le v \le \alpha p\}$, then the length of any piecewise differentiable curve $\gamma : [0,1] \to M^n$ can be written as $l(\gamma) = \int_0^1 ||\gamma'(t)||_{\gamma(t)}^T dt$, and the distance between x and y is equal to $\inf_{\gamma} l(\gamma)$, where the infimum is taken over all such curves γ with $\gamma(0) = x$ and $\gamma(1) = y$.

Hilbert projective semi-metric

Given a convex cone C in a real vector space V, the **Hilbert projective semi-metric** is a semi-metric on $C\setminus\{0\}$, defined, for any $x,y\in C\setminus\{0\}$, by

$$\ln(m(x,y)\cdot m(y,x)),$$

where $m(x,y) = \inf\{\lambda \in \mathbb{R} : y \leq \lambda x\}$. It is equal to 0 if and only if $x = \lambda y$ for some $\lambda > 0$, and it becomes a metric on the space of rays of the cone. If C is finite-dimensional, and S is a cross-section of C (in particular, $S = \{x \in \mathbb{C} : ||x|| = 1\}$, where ||.|| is a norm on V), then, for any distinct

points $x, y \in S$, their distance is equal to $|\ln(x, y, z, t)|$, where z, t are the points of the intersection of the line $l_{x,y}$ with the boundary of S, and (x, y, z, t) is the cross-ratio of x, y, z, t.

If C is almost Archimedean and finite-dimensional, then each part of C is a chord space under the Hilbert projective metric. The Lorentz cone $\{(t, x_1, \dots, x_n) \in \mathbb{R}^{n+1} : t^2 > x_1^2 + \dots + x_n^2\}$ equipped with the Hilbert projective metric is isometric to the *n*-dimensional hyperbolic space. The positive cone $\mathbb{R}^n_+ = \{(x_1, \dots, x_n) : x_i \geq 0 \text{ for } 1 \leq i \leq n\}$ with the Hilbert projective metric is isometric to a normed space which can be seen as being flat.

If C is a closed cone in \mathbb{R}^n with non-empty interior, then int C can be considered as an n-dimensional manifold M^n . If for any tangent vector $v \in T_p(M^n), p \in M^n$, we define a semi-norm $||v||_p^H = m(p,v) - m(v,p)$, then the length of any piecewise differentiable curve $\gamma:[0,1]\to M^n$ can be written as $l(\gamma)=\int_0^1||\gamma'(t)||_{\gamma(t)}^Hdt$, and the distance between x and y is equal to $\inf_{\gamma} l(\gamma)$, where the infimum is taken over all such curves γ with $\gamma(0) = x$ and $\gamma(1) = y$.

• Bushell metric

Given a convex cone C in a real vector space V, the **Bushell metric** on the set $S = \{x \in C : \sum_{i=1}^{n} |x_i| = 1\}$ (in general, on any *cross-section* of C) is defined by

$$\frac{1-m(x,y)\cdot m(y,x)}{1+m(x,y)\cdot m(y,x)}$$

for any $x, y \in S$, where $m(x, y) = \inf\{\lambda \in \mathbb{R} : y \leq \lambda x\}$. In fact, it is equal to $tanh(\frac{1}{2}h(x,y))$, where h is the **Hilbert projective semi-metric**.

k-oriented distance

A simplicial cone C in \mathbb{R}^n is defined as the intersection of n (open or closed) half-spaces, each of whose supporting planes contain the origin 0. For any set M of n points on the unit sphere, there is a unique simplicial cone Cthat contains these points. The axes of the cone C can be constructed as the set of the n rays, where each ray originates at the origin, and contains one of the points from M.

Given a partition $\{C_1, \ldots, C_k\}$ of \mathbb{R}^n into a set of simplicial cones C_1, \ldots, C_k , the k-oriented distance is a metric on \mathbb{R}^n , defined by

$$d_k(x-y)$$

for all $x, y \in \mathbb{R}^n$, where, for any $x \in C_i$, the value $d_k(x)$ is the length of the shortest path from the origin 0 to x traveling only in directions parallel to the axes of C_i .

• Cones over metric space

A cone over a metric space (X,d) is the quotient space $Con(X,d) = (X \times [0,1])/(X \times \{0\})$ obtained from the product $X \times \mathbb{R}_{\geq 0}$ by collapsing the *fiber* (subspace $X \times \{0\}$) to a point (the apex of the cone). Cf. metric cone structure, tangent metric cone in Chap. 1.

Let a metric on Con(X) be defined, for any $(x,t), (y,s) \in Con(X,d)$, by

$$\sqrt{t^2 + s^2 - 2ts\cos(\min\{d(x,y),\pi\})}.$$

The cone Con(X, d) with this metric is called the *Euclidean cone over the metric space* (X, d).

If (X, d) is a compact metric space with diameter <2, the **Krakus metric** is a metric on Con(X, d) defined, for any $(x, t), (y, s) \in Con(X, d)$, by

$$\min\{s, t\}d(x, y) + |t - s|.$$

The cone Con(X, d) with the Krakus metric admits a unique *midpoint* for each pair of its points if (X, d) has this property.

If M^n is a manifold with a (pseudo) Riemannian metric g, one can consider a metric $dr^2 + r^2g$ (more generally, a metric $\frac{1}{k}dr^2 + r^2g$, $k \neq 0$) on the cone $Con(M^n) = M^n \times \mathbb{R}_{>0}$.

• Suspension metric

A spherical cone (or suspension) $\Sigma(X)$ over a metric space (X,d) is the quotient of the product $X \times [0,a]$ obtained by identifying all points in the fibers $X \times \{0\}$ and $X \times \{a\}$.

If (X, d) is a **length space** (cf. Chap. 6) with diameter $diam(X) \leq \pi$, and $a = \pi$, the **suspension metric** is a metric on $\Sigma(X)$, defined, for any $(x, t), (y, s) \in \Sigma(X)$, by

 $\arccos(\cos t \cos s + \sin t \sin s \cos d(x, y)).$

9.3 Distances on simplicial complexes

An r-dimensional simplex (or geometrical simplex, hypertetrahedron) is the convex hull of r+1 points of \mathbb{E}^n which do not lie in any (r-1)-plane. The boundary of an r-simplex has r+1 0-faces (polytope vertices), $\frac{r(r+1)}{2}$ 1-faces

(polytope edges), and $\binom{r+1}{i+1}$ *i-faces*, where $\binom{r}{i}$ is the binomial coefficient. The *content* (i.e., the *hypervolume*) of a simplex can be computed using the *Cayley–Menger determinant*. The regular simplex of dimension r is denoted by α_r .

Roughly, a geometrical simplicial complex is a space with a triangulation, i.e., a decomposition of it into closed simplices such that any two simplices either do not intersect or intersect only along a common face.

An abstract simplicial complex S is a set, whose elements are called vertices, in which a family of finite non-empty subsets, called simplices, is distinguished, such that every non-empty subset of a simplex s is a simplex, called a face of s, and every one-element subset is a simplex. A simplex is called i-dimensional if it consists of i+1 vertices. The dimension of S is the maximal dimension of its simplices. For every simplicial complex S there exists a triangulation of a polyhedron whose simplicial complex is S. This geometric simplicial complex, denoted by S, is called the geometric realization of S.

• Simplicial metric

Let S be an abstract simplicial complex, and GS a geometric simplicial complex which is a geometric realization of S. The points of GS can be identified with the functions $\alpha: S \to [0,1]$ for which the set $\{x \in S : \alpha(x) \neq 0\}$ is a simplex in S, and $\sum_{x \in S} \alpha(x) = 1$. The number $\alpha(x)$ is called the x-th barycentric coordinate of α .

The **simplicial metric** on GS (Lefschetz 1939) is the Euclidean metric on it:

$$\sqrt{\sum_{x \in S} (\alpha(x) - \beta(x))^2}.$$

Tukey (1939) found another metric on GS, topologically equivalent to a simplicial one. His **polyhedral metric** is the **intrinsic metric** on GS, defined as the infimum of the lengths of the polygonal lines joining the points α and β such that each link is within one of the simplices. An example of a polyhedral metric is the intrinsic metric on the surface of a convex polyhedron in \mathbb{E}^3 .

A polyhedral metric can be considered on a complex of simplices in a *space of constant curvature* and, in general, on complexes which are *manifolds*.

Manifold triangulation metric

Let M^n be a compact PL (piecewise-linear) n-dimensional manifold. A triangulation of M^n is a simplicial complex such that its corresponding polyhedron is PL-homeomorphic to M^n . Let T_{M^n} be the set of all $combinatorial \ types$ of triangulations, where two triangulations are equivalent if they are simplicially isomorphic.

Every such triangulation can be seen as a metric on the smooth manifold M if one assigns the unit length for any of its one-dimensional simplices; so, T_{M^n} can be seen as a discrete analogue of the space of Riemannian structures, i.e., isometry classes of Riemannian metrics on M^n .

A manifold triangulation metric between two triangulations x and y is (Nabutovsky and Ben-Av 1993) an editing metric on T_{M^n} , i.e., the minimal number of elementary moves, from a given finite list of operations, needed to obtain y from x.

For example, the *bistellar move* consists of replacing a subcomplex of a given triangulation, which is simplicially isomorphic to a subcomplex of the boundary of the standard (n+1)-simplex, by the complementary subcomplex of the boundary of an (n+1)-simplex, containing all remaining n-simplices and their faces. Every triangulation can be obtained from any other triangulation by a finite sequence of bistellar moves (Pachner 1986).

• Polyhedral chain metric

An r-dimensional polyhedral chain A in \mathbb{E}^n is a linear expression $\sum_{i=1}^m d_i t_i^r$, where, for any i, the value t_i^r is an r-dimensional simplex of \mathbb{E}^n . The boundary of a chain is the linear combination of boundaries of the simplices in the chain. The boundary of an r-dimensional chain is an (r-1)-dimensional chain.

A polyhedral chain metric is a norm metric

$$||A - B||$$

on the set $C_r(\mathbb{E}^n)$ of all r-dimensional polyhedral chains. As a norm ||.|| on $C_r(\mathbb{E}^n)$ one can take:

- 1. The mass of a polyhedral chain, i.e., $|A| = \sum_{i=1}^{m} |d_i| |t_i^r|$, where $|t^r|$ is the volume of the cell t_i^r .
- 2. The flat norm of a polyhedral chain, i.e., $|A|^{\flat} = \inf_{D} \{|A \partial D| + |D|\}$, where |D| is the mass of D, ∂D is the boundary of D, and the infimum is taken over all (r+1)-dimensional polyhedral chains; the completion of the metric space $(C_r(\mathbb{E}^n), |.|^{\flat})$ by the flat norm is a **separable Banach** space, denoted by $C_r^{\flat}(\mathbb{E}^n)$, its elements are known as r-dimensional flat chains.
- 3. The sharp norm of a polyhedral chain, i.e.,

$$|A|^{\sharp} = \inf\left(\frac{\sum_{i=1}^{m} |d_i||t_i^r||v_i|}{r+1} + |\sum_{i=1}^{m} d_i T_{v_i} t_i^r|^{\flat}\right),$$

where $|A|^{\flat}$ is the flat norm of A, and the infimum is taken over all *shifts* v (here $T_v t^r$ is the cell obtained by shifting t^r by a vector v of length |v|); the completion of the metric space $(C_r(\mathbb{E}^n), |.|^{\sharp})$ by the sharp norm is a separable Banach space, denoted by $C_r^{\sharp}(\mathbb{E}^n)$, and its elements are called r-dimensional sharp chains. A flat chain of finite mass is a sharp chain. If r = 0, than $|A|^{\flat} = |A|^{\sharp}$.

The metric space of *polyhedral co-chains* (i.e., linear functions of polyhedral chains) can be defined in similar way. As a norm of a polyhedral co-chain X one can take:

- 1. The co-mass of a polyhedral co-chain, i.e., $|X|=\sup_{|A|=1}|X(A)|,$ where X(A) is the value of the co-chain X on a chain A
- 2. The flat co-norm of a polyhedral co-chain, i.e., $|X|^{\flat} = \sup_{|A|^{\flat}=1} |X|^{\flat}$
- 3. The sharp co-norm of a polyhedral co-chain, i.e., $|X|^{\sharp} = \sup_{|A|^{\sharp}=1} |X(A)|$

Part III Distances in Classical Mathematics

Chapter 10 Distances in Algebra

10.1 Group metrics

A group (G,\cdot,e) is a set G of elements with a binary operation \cdot , called the group operation, that together satisfy the four fundamental properties of closure $(x\cdot y\in G$ for any $x,y\in G)$, associativity $(x\cdot (y\cdot z)=(x\cdot y)\cdot z$ for any $x,y,z\in G)$, the identity property $(x\cdot e=e\cdot x=x$ for any $x\in G)$, and the inverse property (for any $x\in G$, there exists an element $x^{-1}\in G$ such that $x\cdot x^{-1}=x^{-1}\cdot x=e)$. In additive notation, a group (G,+,0) is a set G with a binary operation + such that the following properties hold: $x+y\in G$ for any $x,y\in G$, x+(y+z)=(x+y)+z for any $x,y,z\in G$, x+0=0+x=x for any $x\in G$, and, for any $x\in G$, there exists an element $-x\in G$ such that x+(-x)=(-x)+x=0. A group (G,\cdot,e) is called finite if the set G is finite. A group (G,\cdot,e) is called Abelian if it is commutative, i.e., $x\cdot y=y\cdot x$ for any $x,y\in G$.

Most metrics considered in this section are **group norm metrics** on a group (G, \cdot, e) , defined by

$$||x \cdot y^{-1}||$$

(or, sometimes, by $||y^{-1} \cdot x||$), where ||.|| is a *group norm*, i.e., a function $||.||:G \to \mathbb{R}$ such that, for any $x,y \in G$, we have the following properties:

- 1. $||x|| \ge 0$, with ||x|| = 0 if and only if x = e.
- 2. $||x|| = ||x^{-1}||$.
- 3. $||x \cdot y|| \le ||x|| + ||y||$ (triangle inequality).

In additive notation, a group norm metric on a group (G, +, 0) is defined by ||x + (-y)|| = ||x - y||, or, sometimes, by ||(-y) + x||.

The simplest example of a group norm metric is the **bi-invariant ultrametric** (sometimes called the *Hamming metric*) $||x \cdot y^{-1}||_H$, where $||x||_H = 1$ for $x \neq e$, and $||e||_H = 0$.

• Bi-invariant metric

A metric (in general, a semi-metric) d on a group (G,\cdot,e) is called **bi-invariant** if

$$d(x,y) = d(x \cdot z, y \cdot z) = d(z \cdot x, z \cdot y)$$

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for any $x, y, z \in G$ (cf. translation invariant metric in Chap. 5). Any group norm metric on an Abelian group is bi-invariant.

A metric (in general, a semi-metric) d on a group (G, \cdot, e) is called a **right-invariant metric** if $d(x,y) = d(x \cdot z, y \cdot z)$ for any $x,y,z \in G$, i.e., the operation of right multiplication by an element z is a **motion** of the metric space (G,d). Any group norm metric, defined by $||x \cdot y^{-1}||$, is right-invariant.

A metric (in general, a semi-metric) d on a group (G, \cdot, e) is called a **left-invariant metric** if $d(x,y) = d(z \cdot x, z \cdot y)$ holds for any $x,y,z \in G$, i.e., the operation of left multiplication by an element z is a motion of the metric space (G,d). Any group norm metric, defined by $||y^{-1} \cdot x||$, is left-invariant.

Any right-invariant or left-invariant (in particular, bi-invariant) metric d on G is a group norm metric, since one can define a group norm on G by ||x|| = d(x, 0).

• Positively homogeneous metric

A metric (in general, a distance) d on an Abelian group (G, +, 0) is called **positively homogeneous** if

$$d(mx, my) = md(x, y)$$

for all $x, y \in G$ and all $m \in \mathbb{N}$, where mx is the sum of m terms all equal to x.

• Translation discrete metric

A group norm metric (in general, a group norm semi-metric) on a group (G, \cdot, e) is called **translation discrete** if the *translation distances* (or *translation numbers*)

$$\tau_G(x) = \lim_{n \to \infty} \frac{||x^n||}{n}$$

of the non-torsion elements x (i.e., such that $x^n \neq e$ for any $n \in \mathbb{N}$) of the group with respect to that metric are bounded away from zero.

If the numbers $\tau_G(x)$ are just non-zero, such a group norm metric is called a **translation proper metric**.

• Word metric

Let (G, \cdot, e) be a finitely-generated group with a set A of generators (i.e., A is finite, and every element of G can be expressed as a product of finitely many elements A and their inverses). The word length $w_W^A(x)$ of an element $x \in G \setminus \{e\}$ is defined by

$$w_W^A(x) = \inf\{r : x = a_1^{\epsilon_1} \dots a_r^{\epsilon_r}, a_i \in A, \epsilon_i \in \{\pm 1\}\},\$$

and $w_W^A(e) = 0$.

The word metric d_W^A associated with A is a group norm metric on G, defined by

$$w_W^A(x \cdot y^{-1}).$$

As the word length w_W^A is a group norm on G, d_W^A is **right-invariant**. Sometimes it is defined as $w_W^A(y^{-1} \cdot x)$, and then it is **left-invariant**. In fact, d_W^A is the maximal metric on G that is right-invariant, and such that the distance from any element of A or A^{-1} to the identity element e is equal to one.

If A and B are two finite sets of generators of the group (G, \cdot, e) , then the identity mapping between the metric spaces (G, d_W^A) and (G, d_W^B) is a **quasi-isometry**, i.e., the word metric is unique up to quasi-isometry.

The word metric is the **path metric** of the Cayley graph Γ of (G, \cdot, e) , constructed with respect to A. Namely, Γ is a graph with the vertex-set G in which two vertices x and $y \in G$ are connected by an edge if and only if $y = a^{\epsilon}x$, $\epsilon = \pm 1$, $a \in A$.

• Weighted word metric

Let (G, \cdot, e) be a finitely-generated group with a set A of generators. Given a bounded weight function $w: A \to (0, \infty)$, the weighted word length $w_{WW}^A(x)$ of an element $x \in G \setminus \{e\}$ is defined by

$$w_{WW}^{A}(x) = \inf \left\{ \sum_{i=1}^{t} w(a_i), t \in \mathbb{N} : x = a_1^{\epsilon_1} \dots a_t^{\epsilon_t}, a_i \in A, \epsilon_i \in \{\pm 1\} \right\},$$

and $w_{WW}^A(e) = 0$.

The weighted word metric d_{WW}^A associated with A is a group norm metric on G, defined by

$$w_{WW}^A(x \cdot y^{-1}).$$

As the weighted word length w_{WW}^A is a group norm on G, d_{WW}^A is **right-invariant**. Sometimes it is defined as $w_{WW}^A(y^{-1} \cdot x)$, and then it is **left-invariant**.

The metric d_{WW}^A is the supremum of semi-metrics d on G with the property that $d(e,a) \leq w(a)$ for any $a \in A$.

The metric d_{WW}^A is a **coarse-path metric**, and every right-invariant coarse path metric is a weighted word metric up to **coarse isometry**.

The metric d_{WW}^A is the **path metric** of the weighted Cayley graph Γ_W of (G, \cdot, e) constructed with respect to A. Namely, Γ_W is a weighted graph with the vertex-set G in which two vertices x and $y \in G$ are connected by an edge with the weight w(a) if and only if $y = a^{\epsilon}x$, $\epsilon = \pm 1$, $a \in A$.

• Interval norm metric

An interval norm metric is a group norm metric on a finite group (G, \cdot, e) , defined by

 $||x \cdot y^{-1}||_{int},$

where $||.||_{int}$ is an interval norm on G, i.e., a group norm such that the values of $||.||_{int}$ form a set of consecutive integers starting with 0.

To each interval norm $||.||_{int}$ corresponds an ordered partition $\{B_0, \ldots, B_m\}$ of G with $B_i = \{x \in G : ||x||_{int} = i\}$ (cf. **Sharma–Kaushik distance** in Chap. 16). The Hamming norm and the Lee norm are special cases of interval norms. A generalized Lee norm is an interval norm for which each class has a form $B_i = \{a, a^{-1}\}$.

• C-metric

A C-metric d is a metric on a group (G, \cdot, e) satisfying the following conditions:

- 1. The values of d form a set of consecutive integers starting with 0.
- 2. The cardinality of the sphere $B(x,r) = \{y \in G : d(x,y) = r\}$ is independent of the particular choice of $x \in G$.

The word metric, the Hamming metric, and the Lee metric are C-metrics. Any interval norm metric is a C-metric.

• Order norm metric

Let (G, \cdot, e) be a finite Abelian group. Let ord(x) be the order of an element $x \in G$, i.e., the smallest positive integer n such that $x^n = e$. Then the function $||.||_{ord} : G \to \mathbb{R}$, defined by $||x||_{ord} = \ln ord(x)$, is a group norm on G, called the $order\ norm$.

The order norm metric is a group norm metric on G, defined by

$$||x \cdot y^{-1}||_{ord}.$$

• Monomorphism norm metric

Let (G, +, 0) be a group. Let (H, \cdot, e) be a group with a group norm $||.||_H$. Let $f: G \to H$ be a monomorphism of groups G and H, i.e., an injective function such that $f(x + y) = f(x) \cdot f(y)$ for any $x, y \in G$. Then the function $||.||_G^f: G \to \mathbb{R}$, defined by $||x||_G^f = ||f(x)||_H$, is a group norm on G, called the monomorphism norm.

The monomorphism norm metric is a group norm metric on G, defined by

$$||x-y||_G^f$$
.

• Product norm metric

Let (G, +, 0) be a group with a group norm $||.||_G$. Let (H, \cdot, e) be a group with a group norm $||.||_H$. Let $G \times H = \{\alpha = (x,y) : x \in G, y \in H\}$ be the Cartesian product of G and H, and $(x,y) \cdot (z,t) = (x+z,y\cdot t)$. Then the function $||.||_{G\times H} : G \times H \to \mathbb{R}$, defined by $||\alpha||_{G\times H} = ||(x,y)||_{G\times H} = ||x||_G + ||y||_H$, is a group norm on $G \times H$, called the product norm.

The **product norm metric** is a **group norm metric** on $G \times H$, defined by

$$||\alpha \cdot \beta^{-1}||_{G \times F}.$$

10.1 Group metrics

On the Cartesian product $G \times H$ of two finite groups with the *interval norms* $||.||_{G}^{int}$ and $||.||_{H}^{int}$, an interval norm $||.||_{G\times H}^{int}$ can be defined. In fact, $||\alpha||_{G\times H}^{int} = ||(x,y)||_{G\times H}^{int} = ||x||_{G} + (m+1)||y||_{H}$, where $m = \max_{a\in G} ||a||_{G}^{int}$.

• Quotient norm metric

Let (G, \cdot, e) be a group with a group norm $||.||_G$. Let (N, \cdot, e) be a normal subgroup of (G, \cdot, e) , i.e., xN = Nx for any $x \in G$. Let $(G/N, \cdot, eN)$ be the quotient group of G, i.e., $G/N = \{xN : x \in G\}$ with $xN = \{x \cdot a : a \in N\}$, and $xN \cdot yN = xyN$. Then the function $||.||_{G/N} : G/N \to \mathbb{R}$, defined by $||xN||_{G/N} = \min_{a \in N} ||xa||_X$, is a group norm on G/N, called the quotient norm.

A quotient norm metric is a group norm metric on G/N, defined by

 $||xN \cdot (yN)^{-1}||_{G/N} = ||xy^{-1}N||_{G/N}.$

If $G = \mathbb{Z}$ with the norm being the absolute value, and $N = m\mathbb{Z}$, $m \in \mathbb{N}$, then the quotient norm on $\mathbb{Z}/m\mathbb{Z} = \mathbb{Z}_m$ coincides with the *Lee norm*.

If a metric d on a group (G, \cdot, e) is **right-invariant**, then for any normal subgroup (N, \cdot, e) of (G, \cdot, e) the metric d induces a right-invariant metric (in fact, the **Hausdorff metric**) d^* on G/N by

$$d^*(xN, yN) = \max\{\max_{b \in yN} \min_{a \in xN} d(a, b), \max_{a \in xN} \min_{b \in yN} d(a, b)\}.$$

• Commutation distance

Let (G,\cdot,e) be a finite non-Abelian group. Let $Z(G)=\{c\in G:\ x\cdot c=c\cdot x\ \text{ for any }x\in G\}$ be the center of G. The commutation graph of G is defined as a graph with the vertex-set G in which distinct elements $x,y\in G$ are connected by an edge whenever they commute, i.e., $x\cdot y=y\cdot x$. (Darafsheh, 2009, consider non-commuting graph on $G\setminus Z(G)$.) Obviously, any two distinct elements $x,y\in G$ that do not commute, are connected in this graph by the path x,c,y, where c is any element of Z(G) (for example, e). A path $x=x^1,x^2,\ldots,x^k=y$ in the commutation graph is called an (x-y) N-path if $x^i\notin Z(G)$ for any $i\in\{1,\ldots,k\}$. In this case the elements $x,y\in G\setminus Z(G)$ are called N-connected.

The **commutation distance** (see [DeHu98]) d is an extended distance on G, defined by the following conditions:

- 1. d(x,x) = 0.
- 2. d(x,y) = 1 if $x \neq y$, and $x \cdot y = y \cdot x$.
- 3. d(x,y) is the minimum length of an (x-y) N-path for any N-connected elements x and $y \in G \setminus Z(G)$.
- 4. $d(x,y) = \infty$ if $x, y \in G \setminus Z(G)$ are not connected by any N-path.

Given a group G and a G-conjugacy class X in it, Bates, Bundy, Perkins and Rowley in 2003, 2004, 2007, 2008 considered *commuting graph* (X, E) whose vertex set is X and distinct vertices $x, y \in X$ are joined by an edge $e \in E$ whenever they commute.

• Modular distance

Let $(\mathbb{Z}_m, +, 0)$, $m \geq 2$, be a finite cyclic group. Let $r \in \mathbb{N}$, $r \geq 2$. The modular r-weight $w_r(x)$ of an element $x \in \mathbb{Z}_m = \{0, 1, \ldots, m\}$ is defined as $w_r(x) = \min\{w_r(x), w_r(m-x)\}$, where $w_r(x)$ is the arithmetic r-weight of the integer x. The value $w_r(x)$ can be obtained as the number of nonzero coefficients in the generalized non-adjacent form $x = e_n r^n + \ldots e_1 r + e_0$ with $e_i \in \mathbb{Z}$, $|e_i| < r$, $|e_i + e_{i+1}| < r$, and $|e_i| < |e_{i+1}|$ if $e_i e_{i+1} < 0$ (cf. arithmetic r-norm metric in Chap. 12).

The **modular distance** is a distance on \mathbb{Z}_m , defined by

$$w_r(x-y)$$
.

The modular distance is a metric for $w_r(m) = 1$, $w_r(m) = 2$, and for several special cases with $w_r(m) = 3$ or 4. In particular, it is a metric for $m = r^n$ or $m = r^n - 1$; if r = 2, it is a metric also for $m = 2^n + 1$ (see, for example, [Ernv85]).

The most popular metric on \mathbb{Z}_m is the **Lee metric**, defined by $||x-y||_{Lee}$, where $||x||_{Lee} = \min\{x, m-x\}$ is the *Lee norm* of an element $x \in \mathbb{Z}_m$.

• G-norm metric

Consider a finite field \mathbb{F}_{p^n} for a prime p and a natural number n. Given a compact convex centrally-symmetric body G in \mathbb{R}^n , define the G-norm of an element $x \in \mathbb{F}_{p^n}$ by $||x||_G = \inf\{\mu \geq 0 : x \in p\mathbb{Z}^n + \mu G\}$.

The G-norm metric is a group norm metric on \mathbb{F}_{p^n} , defined by

$$||x \cdot y^{-1}||_G.$$

• Permutation norm metric

Given a finite metric space (X, d), the **permutation norm metric** is a **group norm metric** on the group (Sym_X, \cdot, id) of all permutations of X (id is the *identity mapping*), defined by

$$||f \cdot g^{-1}||_{Sym},$$

where the group norm $||.||_{Sym}$ on Sym_X is given by $||f||_{Sym} = \max_{x \in X} d(x, f(x))$.

• Metric of motions

Let (X, d) be a metric space, and let $p \in X$ be a fixed element of X.

The **metric of motions** (see [Buse55]) is a metric on the group (Ω, \cdot, id) of all **motions** of (X, d) (id is the $identity\ mapping$), defined by

$$\sup_{x \in X} d(f(x), g(x)) \cdot e^{-d(p,x)}$$

for any $f, g \in \Omega$ (cf. Busemann metric of sets in Chap. 3). If the space (X, d) is bounded, a similar metric on Ω can be defined as

$$\sup_{x \in X} d(f(x), g(x)).$$

10.1 Group metrics

Given a semi-metric space (X, d), the **semi-metric of motions** on (Ω, \cdot, id) is defined by

• General linear group semi-metric

Let \mathbb{F} be a locally compact non-discrete topological field. Let $(\mathbb{F}^n, ||.||_{\mathbb{F}^n})$, $n \geq 2$, be a normed vector space over \mathbb{F} . Let ||.|| be the operator norm associated with the normed vector space $(\mathbb{F}^n, ||.||_{\mathbb{F}^n})$. Let $GL(n, \mathbb{F})$ be the general linear group over \mathbb{F} . Then the function $|.|_{op} : GL(n, \mathbb{F}) \to \mathbb{R}$, defined by $|g|_{op} = \sup\{|\ln ||g||, |\ln ||g^{-1}||\}$, is a semi-norm on $GL(n, \mathbb{F})$.

The **general linear group semi-metric** is a semi-metric on the group $GL(n, \mathbb{F})$, defined by

$$|g \cdot h^{-1}|_{op}$$
.

It is a **right-invariant** semi-metric which is unique, up to **coarse isometry**, since any two norms on \mathbb{F}^n are **bi-Lipschitz equivalent**.

• Generalized torus semi-metric

Let (T,\cdot,e) be a generalized torus, i.e., a topological group which is isomorphic to a direct product of n multiplicative groups \mathbb{F}_i^* of locally compact non-discrete topological fields \mathbb{F}_i ; then there is a proper continuous homomorphism $v:T\to\mathbb{R}^n$, namely, $v(x_1,\ldots,x_n)=(v_1(x_1),\ldots,v_n(x_n))$, where $v_i:\mathbb{F}_i^*\to\mathbb{R}$ are proper continuous homomorphisms from the \mathbb{F}_i^* to the additive group \mathbb{R} , given by the logarithm of the valuation. Every other proper continuous homomorphism $v':T\to\mathbb{R}^n$ is of the form $v'=\alpha\cdot v$ with $\alpha\in GL(n,\mathbb{R})$. If ||.|| is a norm on \mathbb{R}^n , one obtains the corresponding semi-norm $||x||_T=||v(x)||$ on T.

The **generalized torus semi-metric** is defined on the group (T, \cdot, e) by

$$||xy^{-1}||_T = ||v(xy^{-1})|| = ||v(x) - v(y)||.$$

• Stable norm metric

Given a Riemannian manifold (M, g), the **stable norm metric** is a **group norm metric** on its real homology group $H_k(M, \mathbb{R})$, defined by the following stable norm $||h||_s$: the infimum of the Riemannian k-volumes of real cycles representing h.

The Riemannian manifold (\mathbb{R}^n, g) is within finite **Gromov–Hausdoff distance** (cf. Chap. 1) from an *n*-dimensional normed vector space $(\mathbb{R}^n, ||.||_s)$.

If (M, g) is a compact connected oriented Riemannian manifold, then the manifold $H_1(M, \mathbb{R})/H_1(M, \mathbb{R})$ with metric induced by $||.||_s$ is called the Albanese torus (or Jacobi torus) of (M, g). This Albanese metric is a flat metric (cf. Chap. 8).

• Heisenberg metric

Let (H, \cdot, e) be the (real) Heisenberg group \mathcal{H}^n , i.e., a group on the set $H = \mathbb{R}^n \times \mathbb{R}^n \times \mathbb{R}$ with the group law $h \cdot h' = (x, y, t) \cdot (x', y', t') = (x+x', y+y', t+t'+2\sum_{i=1}^n (x_i'y_i-x_iy_i')$, and the identity e = (0,0,0). Let

 $|.|_{Heis}$ be the $Heisenberg\ gauge\ ({
m Cygan}\ 1978)$ on ${\mathcal H}^n$, defined by $|h|_{Heis}=|(x,y,t)|_{Heis}=((\sum_{i=1}^n(x_i^2+y_i^2))^2+t^2)^{1/4}$.

The Heisenberg metric (or Korányi metric, Cygan metric, gauge metric) d_{Heis} is a group norm metric on \mathcal{H}^n , defined by

$$|x^{-1} \cdot y|_{Heis}$$
.

One can identify the Heisenberg group $\mathcal{H}^{n-1} = \mathbb{C}^{n-1} \times \mathbb{R}$ with $\partial \mathbb{H}^n_{\mathbb{C}} \setminus \{\infty\}$, where $\mathbb{H}^n_{\mathbb{C}}$ is the Hermitian (i.e., complex) hyperbolic *n*-space, and ∞ is any point of its boundary $\partial \mathbb{H}^n_{\mathbb{C}}$. So, the usual hyperbolic metric of $\mathbb{H}^{n+1}_{\mathbb{C}}$ induces naturally a metric on \mathcal{H}^n . The **Hamenstädt distance** on $\partial \mathbb{H}^n_{\mathbb{C}} \setminus \{\infty\}$ (Hersonsky and Paulin 2004) is $\frac{1}{\sqrt{2}}d_{Heis}$.

Sometimes, the term $Cygan\ metric$ is reserved for the extension of the metric d_{Heis} on whole $\mathbb{H}^n_{\mathbb{C}}$ and (Apanasov 2004) for its generalization (via the $Carnot\ group\ \mathbb{F}^{n-1}\times Im\mathbb{F}$) on \mathbb{F} -hyperbolic spaces $\mathbb{H}^n_{\mathbb{F}}$ over numbers \mathbb{F} that can be complex numbers, or quaternions or, for n=2, octonions. Also, the generalization of d_{Heis} on Carnot groups of $Heisenberg\ type$ is called the $Cygan\ metric$.

The second natural metric on \mathcal{H}^n is the Carnot-Carathéodory metric (or CC metric, sub-Riemannian metric; cf. Chap. 7) d_C , defined as the length metric (cf. Chap. 6) using horizontal vector fields on \mathcal{H}^n . This metric is the internal metric (cf. Chap. 4) corresponding to d_{Heis} . The metric d_{Heis} is bi-Lipschitz equivalent with d_C but not with any Riemannian distance and, in particular, not with any Euclidean metric. For both metrics, the Heisenberg group \mathcal{H}^n is a fractal since its Hausdorff dimension, 2n+2, is strictly greater than its topological dimension, 2n+1.

• Metric between intervals

Let G be the set of all intervals [a,b] of \mathbb{R} . The set G forms semi-groups (G,+) and (G,\cdot) under addition $I+J=\{x+y:x\in I,y\in J\}$ and under multiplication $I\cdot J=\{x\cdot y:x\in I,y\in J\}$, respectively.

The **metric between intervals** is a metric on G, defined by

$$\max\{|I|,|J|\}$$

for all $I, J \in G$, where, for I = [a, b], one has |I| = |a - b|.

• Metric between games

Consider positional games, i.e., two-player nonrandom games of perfect information with real-valued outcomes. Play is alternating with a nonterminated game having move options for both players. Real-world examples include Chess, Go and Tic-Tac-Toe. Formally, let $F_{\mathbb{R}}$ be the universe of games defined inductively as follows:

- 1. Every real number $r \in \mathbb{R}$ belongs to $F_{\mathbb{R}}$ and is called an *atomic game*.
- 2. If $A, B \subset F_{\mathbb{R}}$ with $1 \leq |A|, |B| < \infty$, then $\{A|B\} \in F_{\mathbb{R}}$ (non-atomic game).

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Write any game $G = \{A|B\}$ as $\{G^L|G^R\}$, where $G^L = A$ and $G^R = B$ are the set of left and right moves of G, respectively.

 $F_{\mathbb{R}}$ becomes a commutative semi-group under the following addition operation:

- 1. If p and q are atomic games, then p+q is the usual addition in \mathbb{R} .
- 2. $p + \{g_{l_1}, \ldots | g_{r_1}, \ldots \} = \{g_{l_1} + p, \ldots | g_{r_1} + p, \ldots \}.$
- 3. If G and H are both non-atomic, then $\{G^L|G^R\} + \{H^L|H^R\} = \{I^L|I^R\}$, where $I^L = \{g_l + H, G + h_l : g_l \in G^L, h_l \in H^L\}$ and $I^R = \{g_r + H, G + h_r : g_r \in G^R, h_r \in H^R\}$.

For any game $G \in F_{\mathbb{R}}$, define the optimal outcomes $\overline{L}(G)$ and $\overline{R}(G)$ (if both players play optimally with Left and Right starting, respectively) as follows:

$$\overline{L}(p)=\overline{R}(p)=p$$
 and $\overline{L}(G)=\max\{\overline{R}(g_l):g_l\in G^L\},\ \overline{R}(G)=\max\{\overline{L}(g_r):g_r\in G^R\}.$

The **metric between games** G and H defined by Ettinger (2000) is the following **extended metric** on $F_{\mathbb{R}}$:

$$\sup_{X} |\overline{L}(G+X) - \overline{L}(H+X)| = \sup_{X} |\overline{R}(G+X) - \overline{R}(H+X)|.$$

• Helly semi-metric

Consider a game (A, \mathcal{B}, H) between player A and B. Here A and B are the *strategy sets* for players A and B respectively, and $H = H(\cdot, \cdot)$ is the *payoff function*, i.e., if player A plays $a \in A$ and player B plays $b \in B$, then A pays H(a,b) to B. A player's *strategy set* is the set of available to him *pure strategies*, i.e., complete algorithms for playing the game, indicating the move for every possible situation throughout it.

The **Helly semi-metric** between strategies $a_1 \in \mathcal{A}$ and $a_2 \in \mathcal{A}$ of A is defined by

$$\sup_{b\in\mathcal{B}}|H(a_1,b)-H(a_2,b)|.$$

• Factorial ring semi-metric

Let $(A, +, \cdot)$ be a factorial ring, i.e., a ring with unique factorization.

The **factorial ring semi-metric** is a semi-metric on the set $A \setminus \{0\}$, defined by

$$\ln \frac{l.c.m.(x,y)}{g.c.d.(x,y)},$$

where l.c.m.(x, y) is the least common multiple, and g.c.d.(x, y) is the greatest common divisor of elements $x, y \in A \setminus \{0\}$.

• Frankild-Sather-Wagstaff metric

Let $\mathcal{G}(R)$ be the set of isomorphism classes, up to a shift, of semidualizing complexes over a local Noetherian commutative ring R. An R-complex is a particular sequence of R-module homomorphisms; see [FrWa]) for exact Commutative Algebra definitions.

The **Frankild–Sather-Wagstaff metric** [FrWa] is a metric on $\mathcal{G}(R)$, defined, for any classes $[K], [L] \in \mathcal{G}(R)$, as the infimum of the *lengths* of chains of pairwise comparable elements starting with [K] and ending with [L].

10.2 Metrics on binary relations

A binary relation R on a set X is a subset of $X \times X$; it is the arc-set of the directed graph (X, R) with the vertex-set X.

A binary relation R which is $symmetric\ ((x,y)\in R \text{ implies } (y,x)\in R),$ $reflexive\ (\text{all } (x,x)\in R),$ and $transitive\ ((x,y),(y,z)\in R \text{ imply } (x,z)\in R)$ is called an $equivalence\ relation$ or a $partition\ (\text{of } X \text{ into equivalence classes}).$ Any q-ary $sequence\ x=(x_1,...,x_n),\ q\geq 2\ (\text{i.e., with } 0\leq x_i\leq q-1\ \text{for } 1\leq i\leq n),$ corresponds to the partition $\{B_0,\ldots,B_{q-1}\}$ of $V_n=\{1,\ldots,n\},$ where $B_i=\{1\leq i\leq n:x_i=j\}$ are the equivalence classes.

A binary relation R which is $antisymmetric\ ((x,y),(y,x)\in R\ imply\ x=y),$ reflexive, and transitive is called a $partial\ order,$ and the pair (X,R) is called a poset (partially ordered set). A partial order R on X is denoted also by \preceq with $x\preceq y$ if and only if $(x,y)\in R$. The order \preceq is called linear if any elements $x,y\in X$ are compatible, i.e., $x\preceq y$ or $y\preceq x.$

A poset (L, \preceq) is called a *lattice* if every two elements $x, y \in L$ have the $join \ x \lor y$ and the $meet \ x \land y$. All partitions of X form a lattice by refinement; it is a sublattice of the lattice (by set-inclusion) of all binary relations.

• Kemeny distance

The **Kemeny distance** between binary relations R_1 and R_2 on a set X is the **Hamming metric** $|R_1 \triangle R_2|$. It is twice the minimal number of inversions of pairs of adjacent elements of X which is necessary to obtain R_2 from R_1 .

If R_1, R_2 are partitions, then the Kemeny distance coincides with the Mirkin-Tcherny distance, and $1 - \frac{|R_1 \triangle R_2|}{n(n-1)}$ is the Rand index.

If binary relations R_1, R_2 are linear orders (or rankings, permutations) on the set X, then the Kemeny distance coincides with the **inversion** metric on permutations.

The **Drapal–Kepka distance** between distinct *quasigroups* (differing from groups in that they need not be associative) (X, +) and (X, \cdot) is defined by $|\{(x, y) : x + y \neq x \cdot y\}|$.

• Metrics between partitions

Let X be a finite set of cardinality n = |X|, and let A, B be non-empty subsets of X. Let P_X be the set of partitions of X, and $P,Q \in P_X$. Let B_1, \ldots, B_q be blocks in the partition P, i.e., the pairwise disjoint sets such that $X = B_1 \cup \cdots \cup B_q$, $q \geq 2$. Let $P \vee Q$ be the join of P and Q, and $P \vee Q$ the meet of P and Q in the lattice \mathbb{P}_X of partitions of X.

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Consider the following *editing operations* on partitions:

- An augmentation transforms a partition P of $A \setminus \{B\}$ into a partition of A by either including the objects of B in a block, or including B itself as a new block.

- An removal transforms a partition P of A into a partition of $A \setminus \{B\}$ by deleting the objects in B from each block that contains them.
- A division transforms one partition P into another by the simultaneous removal of B from B_i (where $B \subset B_i$, $B \neq B_i$), and augmentation of B as a new block.
- A merging transforms one partition P into another by the simultaneous removal of B from B_i (where $B = B_i$), and augmentation of B to B_j (where $j \neq i$).
- A transfer transforms one partition P into another by the simultaneous removal of B from B_i (where $B \subset B_i$), and augmentation of B to B_j (where $j \neq i$).

Define (see, for example, [Day81]), in terms of above operations, the following editing metrics on P_X :

- 1. The minimum number of augmentations and removals of single objects needed to transform P into Q.
- 2. The minimum number of divisions, mergings, and transfers of single objects needed to transform P into Q.
- 3. The minimum number of divisions, mergings, and transfers needed to transform P into Q.
- 4. The minimum number of divisions and mergings needed to transform P into Q; in fact, it is equal to $|P| + |Q| 2|P \vee Q|$.
- 5. $\sigma(P) + \sigma(Q) 2\sigma(P \wedge Q)$, where $\sigma(P) = \sum_{P_i \in P} |P_i|(|P_i| 1)$.
- 6. $e(P) + e(Q) 2e(P \wedge Q)$, where $e(P) = \log_2 n + \sum_{P_i \in P} \frac{|P_i|}{n} \log_2 \frac{|P_i|}{n}$.

The **Reignier distance** is the minimum number of elements that must be moved between the blocks of partition P in order to transform it into Q. (Cf. **Earth Mover distance** in Chap. 21 and the above metric 2.)

10.3 Metrics on lattices

Consider a poset (L, \preceq) . The *meet* (or *infimum*) $x \wedge y$ (if it exists) of two elements x and y is the unique element satisfying $x \wedge y \preceq x, y$, and $z \preceq x \wedge y$ if $z \preceq x, y$; similarly, the *join* (or *supremum*) $x \vee y$ (if it exists) is the unique element such that $x, y \preceq x \vee y$, and $x \vee y \preceq z$ if $x, y \preceq z$.

A poset (L, \preceq) is called a *lattice* if every two elements $x, y \in L$ have the join $x \vee y$ and the meet $x \wedge y$. A poset (L, \preceq) is called a *meet semi-lattice* (or

lower semi-lattice) if only the meet-operation is defined. A poset (L, \preceq) is called a join semi-lattice (or upper semi-lattice) if only the join-operation is defined.

A lattice $\mathbb{L} = (L, \leq, \vee, \wedge)$ is called a *semi-modular lattice* (or *semi-Dedekind lattice*) if the *modularity relation* xMy is symmetric: xMy implies yMx for any $x, y \in L$. The *modularity relation* here is defined as follows: two elements x and y are said to constitute a *modular pair*, in symbols xMy, if $x \wedge (y \vee z) = (x \wedge y) \vee z$ for any $z \leq x$.

A lattice \mathbb{L} in which every pair of elements is modular, is called a *modular lattice* (or *Dedekind lattice*). A lattice is modular if and only if the *modular law* is valid: if $z \leq x$, then $x \wedge (y \vee z) = (x \wedge y) \vee z$ for any y. A lattice is called *distributive* if $x \wedge (y \vee z) = (x \wedge y) \vee (x \wedge z)$ for any $x, y, z \in L$.

Given a lattice \mathbb{L} , a function $v: L \to \mathbb{R}_{\geq 0}$, satisfying $v(x \vee y) + v(x \wedge y) \leq v(x) + v(y)$ for all $x, y \in L$, is called a *subvaluation* on \mathbb{L} . A subvaluation v is called *isotone* if $v(x) \leq v(y)$ whenever $x \leq y$, and it is called *positive* if v(x) < v(y) whenever $x \leq y$, $x \neq y$.

A subvaluation v is called a *valuation* if it is isotone and $v(x \lor y) + v(x \land y) = v(x) + v(y)$ for all $x, y \in L$. An integer-valued valuation is called the *height* (or *length*) of \mathbb{L} .

• Lattice valuation metric

Let $\mathbb{L} = (L, \preceq, \vee, \wedge)$ be a lattice, and let v be an isotone subvaluation on \mathbb{L} . The lattice subvaluation semi-metric d_v on L is defined by

$$2v(x \vee y) - v(x) - v(y).$$

(It can be defined also on some semi-lattices.) If v is a positive subvaluation on \mathbb{L} , one obtains a metric, called the **lattice subvaluation metric**. If v is a valuation, d_v can be written as

$$v(x \vee y) - v(x \wedge y) = v(x) + v(y) - 2v(x \wedge y),$$

and is called the *valuation semi-metric*. If v is a positive valuation on \mathbb{L} , one obtains a metric, called the **lattice valuation metric**.

If $L = \mathbb{N}$ (the set of positive integers), $x \vee y = l.c.m.(x,y)$ (least common multiple), $x \wedge y = g.c.d.(x,y)$ (greatest common divisor), and the positive valuation $v(x) = \ln x$, then $d_v(x,y) = \ln \frac{l.c.m.(x,y)}{g.c.d(x,y)}$. This metric can be generalized on any factorial ring (i.e., a ring with unique factorization) equipped with a positive valuation v such that $v(x) \geq 0$ with equality only for the multiplicative unit of the ring, and v(xy) = v(x) + v(y). (Cf. ring semi-metric).

• Finite subgroup metric

Let (G, \cdot, e) be a group. Let $\mathbb{L} = (L, \subset, \cap)$ be the meet semi-lattice of all finite subgroups of the group (G, \cdot, e) with the meet $X \cap Y$ and the valuation $v(X) = \ln |X|$.

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The finite subgroup metric is a valuation metric on L, defined by

$$v(X) + v(Y) - 2v(X \wedge Y) = \ln \frac{|X||Y|}{(|X \cap Y|)^2}.$$

• Scalar and vectorial metrics

Let $\mathbb{L}=(L,\leq,\max,\min)$ be a lattice with the join $\max\{x,y\}$, and the meet $\min\{x,y\}$ on a set $L\subset[0,\infty)$ which has a fixed number a as the greatest element and is closed under *negation*, i.e., for any $x\in L$, one has $\overline{x}=a-x\in L$.

The scalar metric d on L is defined, for $x \neq y$, by

$$d(x,y) = \max\{\min\{x, \overline{y}\}, \min\{\overline{x}, y\}\}.$$

The scalar metric d^* on $L^* = L \cup \{*\}, * \notin L$, is defined, for $x \neq y$, by

$$d^*(x,y) = \begin{cases} d(x,y), & \text{if} \quad x,y \in L, \\ \max\{x,\overline{x}\}, & \text{if} \quad y = *, x \neq *, \\ \max\{y,\overline{y}\}, & \text{if} \quad x = *, y \neq *. \end{cases}$$

Given a norm ||.|| on \mathbb{R}^n , $n \geq 2$, the **vectorial metric** on L^n is defined by

$$||(d(x_1,y_1),\ldots,d(x_n,y_n))||,$$

and the **vectorial metric** on $(L^*)^n$ is defined by

$$||(d^*(x_1,y_1),\ldots,d^*(x_n,y_n))||.$$

The vectorial metric on $L_2^n = \{0,1\}^n$ with l_1 -norm on \mathbb{R}^n is the **Fréchet-Nikodym-Aronszyan distance**. The vectorial metric on $L_m^n = \{0, \frac{1}{m-1}, \dots, \frac{m-2}{m-1}, 1\}^n$ with l_1 -norm on \mathbb{R}^n is the **Sgarro** m-valued metric. The vectorial metric on $[0,1]^n$ with l_1 -norm on \mathbb{R}^n is the **Sgarro** fuzzy metric.

If L is L_m or [0,1], and $x = (x_1, \ldots, x_n, x_{n+1}, \ldots, x_{n+r})$, $y = (y_1, \ldots, y_n, *, \ldots, *)$, where * stands in r places, then the vectorial metric between x and y is the **Sgarro metric** (see, for example, [CSY01]).

• Metrics on Riesz space

A Riesz space (or vector lattice) is a partially ordered vector space (V_{Ri}, \preceq) in which the following conditions hold:

- 1. The vector space structure and the partial order structure are compatible: from $x \leq y$ it follows that $x + z \leq y + z$, and from x > 0, $\lambda \in \mathbb{R}, \lambda > 0$ it follows that $\lambda x > 0$.
- 2. For any two elements $x, y \in V_{Ri}$ there exists the join $x \vee y \in V_{Ri}$ (in particular, the join and the meet of any finite set of elements from V_{Ri} exist).

The **Riesz norm metric** is a **norm metric** on V_{Ri} , defined by

$$||x-y||_{Ri}$$

where $||.||_{Ri}$ is a $Riesz\ norm$, i.e., a norm on V_{Ri} such that, for any $x, y \in V_{Ri}$, the inequality $|x| \leq |y|$, where $|x| = (-x) \vee (x)$, implies $||x||_{Ri} \leq ||y||_{Ri}$. The space $(V_{Ri}, ||.||_{Ri})$ is called a $normed\ Riesz\ space$. In the case of completeness it is called a $Banach\ lattice$. All Riesz norms on a Banach lattice are equivalent.

An element $e \in V_{Ri}^+ = \{x \in V_{Ri} : x \succ 0\}$ is called a *strong unit* of V_{Ri} if for each $x \in V_{Ri}$ there exists $\lambda \in \mathbb{R}$ such that $|x| \leq \lambda e$. If a Riesz space V_{Ri} has a strong unit e, then $||x|| = \inf\{\lambda \in \mathbb{R} : |x| \leq \lambda e\}$ is a Riesz norm, and one obtains on V_{Ri} a Riesz norm metric

$$\inf\{\lambda \in \mathbb{R} : |x - y| \leq \lambda e\}.$$

A weak unit of V_{Ri} is an element e of V_{Ri}^+ such that $e \wedge |x| = 0$ implies x = 0. A Riesz space V_{Ri} is called Archimedean if, for any two $x, y \in V_{Ri}^+$, there exists a natural number n, such that $nx \leq y$. The **uniform metric** on an Archimedean Riesz space with a weak unit e is defined by

$$\inf\{\lambda \in \mathbb{R} : |x - y| \land e \leq \lambda e\}.$$

• Gallery distance of flags

Let \mathbb{L} be a lattice. A *chain* C in \mathbb{L} is a subset of L which is *linearly ordered*, i.e., any two elements of C are compatible. A *flag* is a chain in \mathbb{L} which is maximal with respect to inclusion. If \mathbb{L} is a semi-modular lattice, containing a finite flag, then \mathbb{L} has a unique minimal and a unique maximal element, and any two flags C, D in \mathbb{L} have the same cardinality, n+1. Then n is the height of the lattice \mathbb{L} . Two flags C, D in \mathbb{L} are called *adjacent* if either they are equal or D contains exactly one element not in C. A *gallery* from C to D of length m is a sequence of flags $C = C_0, C_1, \ldots, C_m = D$ such that C_{i-1} and C_i are adjacent for $i = 1, \ldots, m$.

A gallery distance of flags (see [Abel91]) is a distance on the set of all flags of a semi-modular lattice \mathbb{L} with finite height, defined as the minimum of lengths of galleries from C to D. It can be written as

$$|C\vee D|-|C|=|C\vee D|-|D|,$$

where $C \vee D = \{c \vee d : c \in C, d \in D\}$ is the upper sub-semi-lattice generated by C and D.

The gallery distance of flags is a special case of the **gallery metric** (of the *chamber system* consisting of flags).

Chapter 11 Distances on Strings and Permutations

An alphabet is a finite set \mathcal{A} , $|\mathcal{A}| \geq 2$, elements of which are called characters (or symbols). A string (or word) is a sequence of characters over a given finite alphabet \mathcal{A} . The set of all finite strings over the alphabet \mathcal{A} is denoted by $W(\mathcal{A})$. Examples of real world applications, using distances and similarities of string pairs, are Speech Recognition, Bioinformatics, Information Retrieval, Machine Translation, Lexicography, Dialectology.

A substring (or factor, chain, block) of the string $x = x_1 \dots x_n$ is any contiguous subsequence $x_i x_{i+1} \dots x_k$ with $1 \le i \le k \le n$. A prefix of a string $x_1 \dots x_n$ is any substring of it starting with x_1 ; a suffix is any substring of it finishing with x_n . If a string is a part of a text, then the delimiters (a space, a dot, a comma, etc.) are added to the alphabet A.

A vector is any finite sequence consisting of real numbers, i.e., a finite string over the infinite alphabet \mathbb{R} . A frequency vector (or discrete probability distribution) is any string $x_1 \dots x_n$ with all $x_i \geq 0$ and $\sum_{i=1}^n x_i = 1$. A permutation (or ranking) is any string $x_1 \dots x_n$ with all x_i being different numbers from $\{1, \dots, n\}$.

An editing operation is an operation on strings, i.e., a symmetric binary relation on the set of all considered strings. Given a set of editing operations $\mathcal{O} = \{O_1, \ldots, O_m\}$, the corresponding editing metric (or unit cost edit distance) between strings x and y is the minimum number of editing operations from \mathcal{O} needed to obtain y from x. It is the path metric of a graph with the vertex-set $W(\mathcal{A})$ and xy being an edge if y can be obtained from x by one of the operations from \mathcal{O} . In some applications, a cost function is assigned to each type of editing operation; then the editing distance is the minimal total cost of transforming x into y. Given a set of editing operations \mathcal{O} on strings, the corresponding necklace editing metric between cyclic strings x and y is the minimum number of editing operations from \mathcal{O} needed to obtain y from x, minimized over all rotations of x.

The main editing operations on strings are:

- Character indel, i.e., insertion or deletion of a character
- Character replacement
- Character swap, i.e., an interchange of adjacent characters

- Substring move, i.e., transforming, say, the string $x = x_1 \dots x_n$ into the string $x_1 \dots x_{i-1} \mathbf{x_i} \dots \mathbf{x_{k-1}} x_i \dots x_{j-1} x_k \dots x_n$
- Substring copy, i.e., transforming, say, $x = x_1 \dots x_n$ into $x_1 \dots x_{i-1} \mathbf{x_j} \dots \mathbf{x_{k-1}} x_i \dots x_n$
- Substring uncopy, i.e., the removal of a substring provided that a copy of it remains in the string

We list below the main distances on strings. However, some string distances will appear in Chaps. 15, 21 and 23, where they fit better, with respect to the needed level of generalization or specification.

11.1 Distances on general strings

• Levenstein metric

The Levenstein metric (or edit distance, shuffle-Hamming distance, Hamming+Gap metric) is (Levenstein 1965) an editing metric on W(A), obtained for \mathcal{O} consisting of only character replacements and indels.

The Levenstein metric $d_L(x,y)$ between strings $x = x_1 \dots x_m$ and $y = y_1 \dots y_n$ is equal to

$$\min\{d_H(x^*, y^*)\},\,$$

where x^* , y^* are strings of length k, $k \ge \max\{m, n\}$, over the alphabet $\mathcal{A}^* = \mathcal{A} \cup \{*\}$ so that, after deleting all new characters *, strings x^* and y^* shrink to x and y, respectively. Here, the gap is the new symbol *, and x^* , y^* are shuffles of strings x and y with strings consisting of only *.

The Levenstein similarity is $1 - \frac{d_L(x,y)}{\max\{m,n\}}$.

The **Damerau–Levenstein metric** (Damerau 1964) is an editing metric on W(A), obtained for \mathcal{O} consisting only of character replacements, indels and transpositions. In the Levenstein metric, a transposition corresponds to two editing operations: one insertion and one deletion.

The **constrained edit distance** (Oomen 1986) is the Levenstein metric, but the ranges for the number of replacements, insertions and deletions are specified.

• Editing metric with moves

The editing metric with moves is an editing metric on W(A) [Corm03], obtained for \mathcal{O} consisting of only substring moves and indels.

• Editing compression metric

The editing compression metric is an editing metric on W(A) [Corm03], obtained for \mathcal{O} consisting of only indels, copy and uncopy operations.

• Indel metric

The **indel metric** is an editing metric on W(A), obtained for \mathcal{O} consisting of only indels.

It is an analog of the **Hamming metric** $|X\Delta Y|$ between sets X and Y. For strings $x = x_1 \dots x_m$ and $y = y_1 \dots y_n$ it is equal to m + n - 2LCS(x,y), where the similarity LCS(x,y) is the length of the longest common subsequence of x and y.

The factor distance on W(A) is m+n-2LCF(x,y), where the similarity LCF(x,y) is the length of the longest common substring (factor) of x and y.

The *LCS ratio* and the *LCF ratio* are the similarities on W(A) defined by $\frac{LCS(x,y)}{\min\{m,n\}}$ and $\frac{LCF(x,y)}{\min\{m,n\}}$, respectively; sometimes, the denominator is $\max\{m,n\}$ or $\frac{m+n}{2}$.

• Swap metric

The **swap metric** is an editing metric on W(A), obtained for \mathcal{O} consisting only of character swaps.

• Edit distance with costs

Given a set of editing operations $\mathcal{O} = \{O_1, \ldots, O_m\}$ and a weight (or cost function) $w_i \geq 0$, assigned to each type O_i of operation, the **edit distance** with costs between strings x and y is the minimal total cost of an editing path between them, i.e., the minimal sum of weights for a sequence of operations transforming x into y.

The **normalized edit distance** between strings x and y (Marzal and Vidal 1993) is the minimum, over all editing paths P between them, of $\frac{W(P)}{L(P)}$, where W(P) and L(P) are the total cost and the length of the editing path P.

• Transduction edit distances

The **Levenstein metric** with costs between strings x and y is modeled in [RiYi98] as a memoryless stochastic transduction between x and y.

Each step of transduction generates either a character replacement pair (a,b), a deletion pair (a,\emptyset) , an insertion pair (\emptyset,b) , or the specific termination symbol t according to a probability function $\delta: E \cup \{t\} \to [0,1]$, where E is the set of all possible above pairs. Such a transducer induces a probability function on the set of all sequences of operations.

The **transduction edit distances** between strings x and y are [RiYi98] $\ln p$ of the following probabilities p:

for the **Viterbi edit distance**, the probability p of the most likely sequence of editing operations transforming x into y;

for the **stochastic edit distance**, the probability p of the string pair (x, y).

Those distances are never zero unless they are infinite for all other string pairs.

This model allows one to learn (in order to reduce error rate) the edit costs for the Levenstein metric from a corpus of examples (training set of string pairs). This learning is automatic; it reduces to estimating the parameters of above transducer.

• Bag distance

The **bag distance** (or multiset metric, counting filter) is a metric on W(A), defined (Navarro 1997) by

$$\max\{|X\backslash Y|, |Y\backslash X|\}$$

for any strings x and y, where X and Y are the bags of symbols (multisets of characters) in strings x and y, respectively, and, say, $|X \setminus Y|$ counts the number of elements in the multiset $X \setminus Y$. Cf. **metrics between multisets** in Chap. 1.

The bag distance is a (computationally) cheap approximation of the **Levenstein metric**.

• Marking metric

The marking metric is a metric on W(A) [EhHa88], defined by

$$\log_2((diff(x,y)+1)(diff(y,x)+1))$$

for any strings $x = x_1 \dots x_m$ and $y = y_1 \dots y_n$, where diff(x, y) is the minimal size |M| of a subset $M \subset \{1, \dots, m\}$ such that any substring of x, not containing any x_i with $i \in M$, is a substring of y.

Another metric, defined in [EhHa88], is $\log_2(diff(x,y) + diff(y,x) + 1)$.

• Transformation distance

The transformation distance is an editing distance with costs on $W(\mathcal{A})$ (Varre, Delahaye and Rivals 1999) obtained for \mathcal{O} consisting only of substring copy, uncopy and substring indels. The distance between strings x and y is the minimal cost of transformation x into y using these operations, where the cost of each operation is the length of its description. For example, the description of the copy requires a binary code specifying the type of operation, an offset between the substring locations in x and in y, and the length of the substring. A code for insertion specifies the type of operation, the length of the substring and the sequence of the substring.

• L_1 rearrangement distance

The L_1 rearrangement distance (Amir, Aumann, Indyk, Levy and Porat 2007) between strings $x = x_1 \dots x_m$ and $y = y_1 \dots y_m$ is equal to

$$\min_{\pi} \sum_{i=1}^{m} |i - \pi(i)|,$$

where $\pi: \{1, \ldots, m\} \to \{1, \ldots, m\}$ is a permutation transforming x into y; if there are no such permutations, the distance is equal to ∞ .

The L_{∞} rearrangement distance (Amir, Aumann, Indyk, Levy and Porat 2007) between strings $x=x_1\ldots x_m$ and $y=y_1\ldots y_m$ is $\min_{\pi}\max_{1\leq i\leq m}|i-\pi(i)|$ and, again, it is ∞ if such a permutation does not exist.

Cf. genome rearrangement distances in Chap. 23.

• Normalized information distance

The **normalized information distance** d between two binary strings x and y is a symmetric function on $W(\{0,1\})$ [LCLM04], defined by

$$\frac{\max\{K(x|y^*), K(y|x^*)\}}{\max\{K(x), K(y)\}}$$

Here, for binary strings u and v, u^* is a shortest binary program to compute u on an appropriate (i.e., using a Turing-complete language) universal computer, the Kolmogorov complexity (or algorithmic entropy) K(u) is the length of u^* (the ultimate compressed version of u), and K(u|v) is the length of the shortest program to compute u if v is provided as an auxiliary input.

The function d(x,y) is a metric up to small error term: $d(x,x) = O((K(x))^{-1})$, and $d(x,z) - d(x,y) - d(y,z) = O((\max\{K(x),K(y),K(z)\})^{-1})$. (Cf. d(x,y) the **information metric** (or *entropy metric*) H(X|Y) + H(Y|X) between stochastic sources X and Y.)

The Kolmogorov complexity is uncomputable and depends on the chosen computer language; so, instead of K(u), were proposed the *minimum message length* (shortest overall message) by Wallace (1968) and the *minimum description length* (largest compression of data) by Rissanen (1978).

The normalized compression distance is a distance on $W(\{0,1\})$ [LCLM04], [BGLVZ98], defined by

$$\frac{C(xy) - \min\{C(x), C(y)\}}{\max\{C(x), C(y)\}}$$

for any binary strings x and y, where C(x), C(y), and C(xy) denote the size of the compression (by fixed compressor C, such as gzip, bzip2, or PPMZ) of strings x, y, and their concatenation xy. This distance is not a metric. It is an approximation of the normalized information distance. A similar distance is defined by $\frac{C(xy)}{C(x)+C(y)} - \frac{1}{2}$.

• Lempel–Ziv distance

The **Lempel–Ziv distance** between two binary strings x and y of length n is

$$\max\{\frac{LZ(x|y)}{LZ(x)}, \frac{LZ(y|x)}{LZ(y)}\},$$

where $LZ(x) = \frac{|P(x)|\log|P(x)|}{n}$ is the Lempel-Ziv complexity of x, approximating its Kolmogorov complexity K(x). Here P(x) is the set of non-overlapping substrings into which x is parsed sequentially, so that the new substring is not yet contained in the set of substrings generated so far. For example, such a Lempel-Ziv parsing for x=00110010101010111 is 0|01|1|00|101|001|11. Now, $LZ(x|y)=\frac{|P(x)\setminus P(y)|\log|P(x)\setminus P(y)|}{n}$.

• Anthony-Hammer similarity

The **Anthony–Hammer similarity** between a binary string $x = x_1 \dots x_n$ and the set Y of binary strings $y = y_1 \dots y_n$ is the maximal number m such that, for every m-subset $M \subset \{1, \dots, n\}$, the substring of x, containing only x_i with $i \in M$, is a substring of some $y \in Y$ containing only y_i with $i \in M$.

Jaro similarity

Given strings $x=x_1\ldots x_m$ and $y=y_1\ldots y_n$, call a character x_i common with y if $x_i=y_j$, where $|i-j|\leq \frac{\min\{m,n\}}{2}$. Let $x^{'}=x_1^{'}\ldots x_{m^{'}}^{'}$ be all the characters of x which are common with y (in the same order as they appear in x), and let $y^{'}=y_1^{'}\ldots y_{n^{'}}^{'}$ be the analogous string for y.

The **Jaro** similarity $Jaro(x, y)^n$ between strings x and y is defined by

$$\frac{1}{3} \left(\frac{m^{'}}{m} + \frac{n^{'}}{n} + \frac{|\{1 \leq i \leq \min\{m^{'}, n^{'}\} : x_{i}^{'} = y_{i}^{'}\}|}{\min\{m^{'}, n^{'}\}} \right).$$

This and following two similarities are used in Record Linkage.

• Jaro-Winkler similarity

The **Jaro–Winkler similarity** between strings x and y is defined by

$$Jaro(x,y) + \frac{\max\{4, LCP(x,y)\}}{10}(1 - Jaro(x,y)),$$

where Jaro(x, y) is the **Jaro similarity**, and LCP(x, y) is the length of the longest common prefix of x and y.

• q-gram similarity

Given an integer $q \ge 1$ (usually, q is 2 or 3), the q-gram similarity between strings x and y is defined by

$$\frac{2q(x,y)}{q(x)+q(y)},$$

Sometimes, the strings x and y are padded before computing their q-gram similarity, i.e., q-1 special characters are added to their beginnings and ends. Padding increases the matching quality since q-grams at the beginning and end of strings are q-grams not matched to other q-grams.

The q-gram similarity is an example of **token-based similarities**, i.e., ones defined in terms of *tokens* (selected substrings or words). Here tokens

are q-grams. A generic **dictionary-based metric** between strings x and y is $|D(x)\Delta D(y)|$, where D(z) denotes the full *dictionary* of z, i.e., the set of all of its substrings.

• Prefix-Hamming metric

The **prefix-Hamming metric** between strings $x = x_1 \dots x_m$ and $y = y_1 \dots y_n$ is defined by

$$(\max\{m, n\} - \min\{m, n\}) + |\{1 \le i \le \min\{m, n\} : x_i \ne y_i\}|.$$

• Weighted Hamming metric

If (A, d) is a metric space, then the **weighted Hamming metric** between strings $x = x_1 \dots x_m$ and $y = y_1 \dots y_m$ is defined by

$$\sum_{i=1}^{m} d(x_i, y_i).$$

The term weighted Hamming metric (or weighted Hamming distance) is also used for $\sum_{1 \leq i \leq m, x_i \neq y_i} w_i$, where, for any $1 \leq i \leq m$, w(i) > 0 is its weight.

• Fuzzy Hamming distance

If (\mathcal{A}, d) is a metric space, the **fuzzy Hamming distance** between strings $x = x_1 \dots x_m$ and $y = y_1 \dots y_m$ is an **editing distance with costs** on $W(\mathcal{A})$ obtained for \mathcal{O} consisting of only indels, each of fixed cost q > 0, and character shifts (i.e., moves of 1-character substrings), where the cost of replacement of i by j is a function f(|i-j|). This distance is the minimal total cost of transforming x into y by these operations. Bookstein, Klein, Raita (2001) introduced this distance for Information Retrieval and proved that it is a metric if f is a monotonically increasing concave function on integers vanishing only at 0. The case f(|i-j|) = C|i-j|, where C > 0 is a constant and |i-j| is a time shift, corresponds to the Victor-Purpura spike train distance in Chap. 23.

Ralescu (2003) introduced, for Image Retrieval, another fuzzy Hamming distance on \mathbb{R}^m . The Ralescu distance between two strings $x = x_1 \dots x_m$ and $y = y_1 \dots y_m$ is the fuzzy cardinality of the difference fuzzy set $D_{\alpha}(x, y)$ (where α is a parameter) with membership function

$$\mu_i = 1 - e^{-\alpha(x_i - y_i)^2}, 1 \le i \le m.$$

The non-fuzzy cardinality of the fuzzy set $D_{\alpha}(x,y)$ approximating its fuzzy cardinality is $|\{1 \leq i \leq m : \mu_i > \frac{1}{2}\}|$.

• Needleman-Wunsch-Sellers metric

If (A, d) is a metric space, the **Needleman–Wunsch–Sellers metric** (or *global alignment metric*) is an **editing distance with costs** on W(A) [NeWu70], obtained for \mathcal{O} consisting of only indels, each of fixed

cost q > 0, and character replacements, where the cost of replacement of i by j is d(i, j). This metric is the minimal total cost of transforming x into y by these operations. Equivalently, it is

$$\min\{d_{wH}(x^*, y^*)\},\$$

where x^* , y^* are strings of length k, $k \ge \max\{m,n\}$, over the alphabet $\mathcal{A}^* = \mathcal{A} \cup \{*\}$, so that, after deleting all new characters *, strings x^* and y^* shrink to x and y, respectively. Here $d_{wH}(x^*, y^*)$ is the **weighted Hamming metric** between x^* and y^* with weight $d(x_i^*, y_i^*) = q$ (i.e., the editing operation is an indel) if one of x_i^* , y_i^* is *, and $d(x_i^*, y_i^*) = d(i, j)$, otherwise.

The Gotoh–Smith–Waterman distance (or string distance with affine gaps) is a more specialized editing metric with costs (see [Goto82]). It discounts mismatching parts at the beginning and end of the strings x, y, and introduces two indel costs: one for starting an affine gap (contiguous block of indels), and another one (lower) for extending a gap.

• Duncan metric

Consider the set X of all strictly increasing infinite sequences $x = \{x_n\}_n$ of positive integers. Define N(n,x) as the number of elements in $x = \{x_n\}_n$ which are less than n, and $\delta(x)$ as the density of x, i.e., $\delta(x) = \lim_{n \to \infty} \frac{N(n,x)}{n}$. Let Y be the subset of X consisting of all sequences $x = \{x_n\}_n$ for which $\delta(x) < \infty$.

The **Duncan metric** is a metric on Y, defined, for $x \neq y$, by

$$\frac{1}{1 + LCP(x, y)} + |\delta(x) - \delta(y)|,$$

where LCP(x,y) is the length of the longest common prefix of x and y.

• Martin metric

The Martin metric d^a between strings $x = x_1 \dots x_m$ and $y = y_1 \dots y_n$ is defined by

$$|2^{-m} - 2^{-n}| + \sum_{t=1}^{\max\{m,n\}} \frac{a_t}{|\mathcal{A}|^t} \sup_{z} |k(z,x) - k(z,y)|,$$

where z is any string of length t, k(z,x) is the Martin kernel of a Markov chain $M = \{M_t\}_{t=0}^{\infty}$, and the sequence $a \in \{a = \{a_t\}_{t=0}^{\infty} : a_t > 0, \sum_{t=1}^{\infty} a_t < \infty\}$ is a parameter.

• Baire metric

The **Baire metric** is an ultrametric between finite or infinite strings x and y, defined, for $x \neq y$, by

$$\frac{1}{1 + LCP(x, y)},$$

where LCP(x, y) is the length of the longest common prefix of x and y. Cf. **Baire space** in Chap. 2.

Given an infinite cardinal number κ and a set A of cardinality κ , the Cartesian product of countably many copies of A endowed with above ultametric $\frac{1}{1+LCP(x,y)}$ is called the **Baire space of weight** κ and denoted by $B(\kappa)$. In particular, $B(\aleph_0)$ (called the Baire zero-dimensional space) is homeomorphic to the space Irr of irrationals with **continued fraction metric** (cf. Chap. 12).

• Generalized Cantor metric

The **generalized Cantor metric** (or, sometimes, *Baire distance*) is an ultrametric between infinite strings x and y, defined, for $x \neq y$, by

$$a^{1+LCP(x,y)}$$
.

where a is a fixed number from the interval (0,1), and LCP(x,y) is the length of the longest common prefix of x and y.

This ultrametric space is **compact**. In the case $a = \frac{1}{2}$, the metric $\frac{1}{2^{1+LCP(x,y)}}$ was considered on a remarkable **fractal** (cf. Chap. 1) from [0,1], the *Cantor set*; cf. **Cantor metric** in Chap. 18.

Comyn and Dauchet (1985) and Kwiatkowska (1990) introduced some analogues of generalized Cantor metric for *traces*, i.e., equivalence classes of strings with respect to a congruence relation identifying strings x, y that are identical up to permutation of concurrent actions (xy = yx).

• Parentheses string metrics

Let P_n be the set of all strings on the alphabet $\{(,)\}$ generated by a grammar and having n open and n closed parentheses. A **parentheses** string metric is an editing metric on P_n (or on its subset) corresponding to a given set of editing operations.

For example, the **Monjardet metric** (Monjardet 1981) between two parentheses strings $x, y \in P_n$ is the minimum number of adjacent parentheses interchanges ["()" to ")(" or ")(" to "()"] needed to obtain y from x. It is the **Manhattan** metric between their representations p_x and p_x , where $p_z = (p_z(1), \ldots, p_z(n))$ and $p_z(i)$ is the number of open parentheses written before the i-th closed parenthese of $z \in P_n$.

There is a bijection between parentheses strings and binary trees; cf. the **tree rotation distance** in Chap. 15.

Similarly, Autord- $Dehornoy\ distance$ between shortest expressions x and y of a permutation as a product of transpositions, is the minimal number of $braid\ relations$ needed to get x from y.

• Schellenkens complexity quasi-metric

The Schellenkens complexity quasi-metric is a quasi-metric between infinite strings $x=x_0,x_1,\ldots,x_m,\ldots$ and $y=y_0,y_1,\ldots,y_n,\ldots$ over $\mathbb{R}_{\geq 0}$ with $\sum_{i=0}^{\infty} 2^{-i} \frac{1}{x_i} < \infty$ (seen as complexity functions), defined (Schellenkens 1995) by

$$\sum_{i=0}^{\infty} 2^{-i} \max\{0, \frac{1}{x_i} - \frac{1}{y_i}\}.$$

• Graev metrics

Let (X, d) be a metric space. Let $\overline{X} = X \cup X' \cup \{e\}$, where $X' = \{x' : x \in X\}$ is a disjoint copy of X, and $e \notin X \cup X'$. We use the notation (e')' = e and (x')' = x for any $x \in X$; also, the letters x, y, x_i, y_i will denote elements of \overline{X} . Let (\overline{X}, D) be a metric space such that D(x, y) = D(x', y') = d(x, y), D(x, e) = D(x', e) and D(x, y') = D(x', y) for all $x, y \in X$.

Denote by W(X) the set of all words over \overline{X} and, for each word $w \in W(X)$, denote by l(w) its length. A word $w \in W(X)$ is called *irreducible* if w = e or $w = x_0 \dots x_n$, where $x_i \neq e$ and $x_{i+1} \neq x_i'$ for $0 \leq i < n$.

For each word w over \overline{X} , denote by \widehat{w} the unique irreducible word obtained from w by successively replacing any occurrence of xx' in w by e and eliminating e from any occurrence of the form w_1ew_2 , where at least one of the words w_1 and w_2 is non-empty.

Denote by F(X) the set of all irreducible words over \overline{X} and, for $u, v \in F(X)$, define $u \cdot v = w'$, where w is the concatenation of words u and v. Then F(X) becomes a group; its identity element is the (non-empty) word e.

For any two words $v = x_0 \dots x_n$ and $u = y_0 \dots y_n$ over \overline{X} of the same length, let $\rho(v, u) = \sum_{i=0}^n D(x_i, y_i)$. The **Graev metric** between two irreducible words $u = u, v \in F(X)$ is defined [DiGa07] by

$$\inf\{\rho(u^*,v^*): u^*,v^* \in W(X), \ l(u^*) = l(v^*), \ \widehat{u^*} = u, \widehat{v^*} = v\}.$$

Graev proved that this metric is a **bi-invariant metric** on F(X), extending the metric d on X, and that F(X) is a topological group in the topology induced by it.

11.2 Distances on permutations

A permutation (or ranking) is any string $x_1
ldots x_n$ with all x_i being different numbers from $\{1, ..., n\}$; a signed permutation is any string $x_1
ldots x_n$ with all $|x_i|$ being different numbers from $\{1, ..., n\}$. Denote by (Sym_n, \cdot, id) the group of all permutations of the set $\{1, ..., n\}$, where id is the identity mapping.

The restriction, on the set Sym_n of all n-permutation vectors, of any metric on \mathbb{R}^n is a metric on Sym_n ; the main example is the l_p -metric $(\sum_{i=1}^n |x_i - y_i|^p)^{\frac{1}{p}}, \ p > 1$.

The main editing operations on permutations are:

- Block transposition, i.e., a substring move
- Character move, i.e., a transposition of a block consisting of only one character
- Character swap, i.e., interchanging of any two adjacent characters

- Character exchange, i.e., interchanging of any two characters (in Group Theory, it is called transposition)
- One-level character exchange, i.e., exchange of characters x_i and x_j , i < j, such that, for any k with i < k < j, either $\min\{x_i, x_j\} > x_k$, or $x_k > \max\{x_i, x_j\}$
- Block reversal, i.e., transforming, say, the permutation $x = x_1 \dots x_n$ into the permutation $x_1 \dots x_{i-1} \mathbf{x_j} \mathbf{x_{j-1}} \dots \mathbf{x_{i+1}} \mathbf{x_i} x_{j+1} \dots x_n$ (so, a swap is a reversal of a block consisting only of two characters)
- Signed reversal, i.e., a reversal in signed permutation, followed by multiplication on -1 of all characters of the reversed block

Below we list the most used editing and other metrics on Sym_n .

• Hamming metric on permutations

The Hamming metric on permutations d_H is an editing metric on Sym_n , obtained for \mathcal{O} consisting of only character replacements. It is a **bi-invariant** metric. Also, $n - d_H(x, y)$ is the number of fixed points of xy^{-1} .

• Spearman ρ distance

The **Spearman** ρ **distance** is the Euclidean metric on Sym_n :

$$\sqrt{\sum_{i=1}^{n} (x_i - y_i)^2}.$$

(Cf. Spearman ρ rank correlation in Chap. 17.)

• Spearman footrule distance

The **Spearman footrule distance** is the l_1 -metric on Sym_n :

$$\sum_{i=1}^{n} |x_i - y_i|.$$

(Cf. Spearman footrule similarity in Chap. 17.)

Both Spearman distances are **bi-invariant**.

• Kendall τ distance

The Kendall τ distance (or inversion metric, permutation swap metric) I is an editing metric on Sym_n , obtained for \mathcal{O} consisting only of character swaps.

In terms of Group Theory, I(x,y) is the number of adjacent transpositions needed to obtain x from y. Also, I(x,y) is the number of relative inversions of x and y, i.e., pairs (i,j), $1 \le i < j \le n$, with $(x_i - x_j)(y_i - y_j) < 0$. (Cf. Kendall τ rank correlation similarity in Chap. 17.)

In [BCFS97] the following metrics were also given, associated with the metric I(x, y):

- 1. $\min_{z \in Sym_n} (I(x, z) + I(z^{-1}, y^{-1}))$
- 2. $\max_{z \in Sym_n} I(zx, zy)$

- 3. $\min_{z \in Sym_n} I(zx, zy) = T(x, y)$, where T is the Cayley metric
- 4. Editing metric, obtained for \mathcal{O} consisting only of one-level character exchanges

• Daniels-Guilbaud semi-metric

The **Daniels–Guilbaud semi-metric** is a semi-metric on Sym_n , defined, for any $x, y \in Sym_n$, as the number of triples (i, j, k), $1 \le i < j < k \le n$, such that (x_i, x_j, x_k) is not a cyclic shift of (y_i, y_j, y_k) ; so, it is 0 if and only if x is a cyclical shift of y (see [Monj98]).

Cayley metric

The Cayley metric T is an editing metric on Sym_n , obtained for \mathcal{O} consisting only of character exchanges.

In terms of Group Theory, T(x, y) is the minimum number of transpositions needed to obtain x from y. Also, n - T(x, y) is the number of cycles in xy^{-1} . The metric T is **bi-invariant**.

• Ulam metric

The Ulam metric (or permutation editing metric) U is an editing metric on Sym_n , obtained for \mathcal{O} consisting only of character moves.

Equivalently, it is an editing metric, obtained for \mathcal{O} consisting only of indels. Also, $n-U(x,y)=LCS(x,y)=LIS(xy^{-1})$, where LCS(x,y) is the length of the longest common subsequence (not necessarily a substring) of x and y, while LIS(z) is the length of the longest increasing subsequence of $z \in Sym_n$.

This and the preceding six metrics are **right-invariant**.

• Reversal metric

The **reversal metric** is an editing metric on Sym_n , obtained for \mathcal{O} consisting only of block reversals.

• Signed reversal metric

The **signed reversal metric** (Sankoff 1989) is an editing metric on the set of all $2^n n!$ signed permutations of the set $\{1, \ldots, n\}$, obtained for \mathcal{O} consisting only of signed reversals.

This metric is used in Biology, where a signed permutation represents a single-chromosome genome, seen as a permutation of genes (along the chromosome) each having a direction (so, a sign + or -).

• Chain metric

The **chain metric** (or rearrangement metric) is a metric on Sym_n [Page65], defined, for any $x, y \in Sym_n$, as the minimum number, minus 1, of chains (substrings) y'_1, \ldots, y'_t of y, such that x can be parsed (concatenated) into, i.e., $x = y'_1 \ldots y'_t$.

• Lexicographic metric

The **lexicographic metric** (Golenko and Ginzburg 1973) is a metric on Sym_n , defined by

$$|N(x) - N(y)|,$$

where N(x) is the ordinal number of the position (among 1, ..., n!) occupied by the permutation x in the *lexicographic ordering* of the set Sym_n .

In the lexicographic ordering of Sym_n , $x = x_1 \dots x_n \prec y = y_1 \dots y_n$ if there exists $1 \leq i \leq n$ such that $x_1 = x_1, \dots, x_{i-1} = y_{i-1}$, but $x_i < y_i$.

• Fréchet permutation metric

The Fréchet permutation metric is the Fréchet product metric on the set Sym_{∞} of permutations of positive integers, defined by

$$\sum_{i=1}^{\infty} \frac{1}{2^i} \frac{|x_i - y_i|}{1 + |x_i - y_i|}.$$

Chapter 12 Distances on Numbers, Polynomials, and Matrices

12.1Metrics on numbers

Here we consider some of the most important metrics on the classical number systems: the semi-ring \mathbb{N} of natural numbers, the ring \mathbb{Z} of integers, and the fields \mathbb{Q} , \mathbb{R} , and \mathbb{C} of rational, real, and complex numbers, respectively. We consider also the algebra \mathcal{Q} of quaternions.

• Metrics on natural numbers

There are several well-known metrics on the set N of natural numbers:

- 1. |n-m|; the restriction of the **natural metric** (from \mathbb{R}) on \mathbb{N}
- 2. $p^{-\alpha}$, where α is the highest power of a given prime number p dividing m-n, for $m \neq n$ (and equal to 0 for m=n); the restriction of the p-adic metric (from \mathbb{Q}) on \mathbb{N}
- 3. $\ln \frac{l.c.m.(m,n)}{g.c.d.(m,n)}$; an example of the lattice valuation metric
- 4. $w_r(n-m)$, where $w_r(n)$ is the arithmetic r-weight of n; the restriction of the arithmetic r-norm metric (from \mathbb{Z}) on \mathbb{N}
- 5. $\frac{|n-m|}{mn}$ (cf. *M*-relative metric in Chap. 19)
 6. $1 + \frac{1}{m+n}$ for $m \neq n$ (and equal to 0 for m = n); the Sierpinski metric

Most of these metrics on \mathbb{N} can be extended on \mathbb{Z} . Moreover, any one of the above metrics can be used in the case of an arbitrary countable set X. For example, the **Sierpinski metric** is defined, in general, on a countable set $X = \{x_n : n \in \mathbb{N}\}$ by $1 + \frac{1}{m+n}$ for all $x_m, x_n \in X$ with $m \neq n$ (and is equal to 0, otherwise).

• Arithmetic r-norm metric

Let $r \in \mathbb{N}$, r > 2. The modified r-ary form of an integer x is a representation

$$x = e_n r^n + \dots + e_1 r + e_0,$$

where $e_i \in \mathbb{Z}$, and $|e_i| < r$ for all i = 0, ..., n. An r-ary form is called minimal if the number of non-zero coefficients is minimal. The minimal form is not unique, in general. But if the coefficients e_i , $0 \le i \le n-1$,

satisfy the conditions $|e_i + e_{i+1}| < r$, and $|e_i| < |e_{i+1}|$ if $e_i e_{i+1} < 0$, then the above form is unique and minimal; it is called the *generalized* non-adjacent form.

The arithmetic r-weight $w_r(x)$ of an integer x is the number of non-zero coefficients in a minimal r-ary form of x, in particular, in the generalized non-adjacent form. The **arithmetic** r-**norm metric** on \mathbb{Z} (see, for example, [Ernv85]) is defined by

$$w_r(x-y)$$
.

• p-adic metric

Let p be a prime number. Any non-zero rational number x can be represented as $x = p^{\alpha} \frac{c}{d}$, where c and d are integers not divisible by p, and α is a unique integer. The p-adic norm of x is defined by $|x|_p = p^{-\alpha}$. Moreover, $|0|_p = 0$ is defined.

The *p*-adic metric is a norm metric on the set \mathbb{Q} of rational numbers, defined by

$$|x-y|_p$$
.

This metric forms the basis for the algebra of p-adic numbers. In fact, the **Cauchy completion** of the metric space $(\mathbb{Q}, |x-y|_p)$ gives the field \mathbb{Q}_p of p-adic numbers; also the Cauchy completion of the metric space $(\mathbb{Q}, |x-y|)$ with the **natural metric** |x-y| gives the field \mathbb{R} of real numbers.

The **Gajić metric** is an **ultrametric** on the set \mathbb{Q} of rational numbers defined, for $x \neq y$ (via the integer part |z| of a real number z), by

$$\inf\{2^{-n} : n \in \mathbb{Z}, |2^n(x-e)| = |2^n(y-e)|\},\$$

where e is any fixed irrational number. This metric is **equivalent** to the **natural metric** |x - y| on \mathbb{Q} .

• Continued fraction metric on irrationals

The **continued fraction metric on irrationals** is a complete metric on the set Irr of irrational numbers defined, for $x \neq y$, by

$$\frac{1}{n}$$

where n is the first index for which the continued fraction expansions of x and y differ. This metric is **equivalent** to the **natural metric** |x-y| on Irr, which is non-complete and disconnected. Also, the $Baire\ zero-dimensional\ space\ B(\aleph_0)$ (cf. **Baire metric** in Chap. 11) is homeomorphic to Irr endowed with this metric.

• Natural metric

The natural metric (or absolute value metric, or the distance between numbers) is a metric on \mathbb{R} , defined by

$$|x - y| =$$

$$\begin{cases} y - x, & \text{if } x - y < 0, \\ x - y, & \text{if } x - y \ge 0. \end{cases}$$

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On \mathbb{R} all l_p -metrics coincide with the natural metric. The metric space $(\mathbb{R}, |x-y|)$ is called the *real line* (or *Euclidean line*).

There exist many other metrics on \mathbb{R} coming from |x-y| by some **metric transform** (cf. Chap. 4). For example: $\min\{1, |x-y|\}$, $\frac{|x-y|}{1+|x-y|}$, |x|+|x-y|+|y| (for $x\neq y$) and, for a given $0<\alpha<1$, the **generalized** absolute value metric $|x-y|^{\alpha}$.

• Zero bias metric

The **zero bias metric** is a metric on \mathbb{R} , defined by

$$1 + |x - y|$$

if one and only one of x and y is strictly positive, and by

$$|x-y|$$
,

otherwise, where |x-y| is the **natural metric** (see, for example, [Gile87]).

• Sorgenfrey quasi-metric

The **Sorgenfrey quasi-metric** is a quasi-metric d on \mathbb{R} , defined by

$$y-x$$

if $y \ge x$, and equal to 1 otherwise.

Some examples of similar quasi-metrics on \mathbb{R} are:

- 1. $d_1(x, y) = \max\{y x, 0\}.$
- 2. $d_2(x,y) = \min\{y-x,1\}$ if $y \ge x$, and equal to 1 otherwise.
- 3. $d_3(x,y) = y x$ if $y \ge x$, and equal to a(x-y) (for fixed a > 0) otherwise.
- 4. $d_4(x,y) = e^y e^x$ if $y \ge x$, and equal to $e^{-y} e^{-x}$ otherwise.

• Real half-line quasi-semi-metric

The **real half-line quasi-semi-metric** is defined on the half-line $\mathbb{R}_{>0}$ by

$$\max\{0, \ln\frac{y}{x}\}.$$

• Janous-Hametner metric

The **Janous–Hametner metric** is defined on the half-line $\mathbb{R}_{>0}$ by

$$\frac{|x-y|}{(x+y)^t},$$

where t = -1 or $0 \le t \le 1$, and |x - y| is the **natural metric**.

• Extended real line metric

An **extended real line metric** is a metric on $\mathbb{R} \cup \{+\infty\} \cup \{-\infty\}$. The main example (see, for example, [Cops68]) of such metric is given by

$$|f(x) - f(y)|,$$

where $f(x) = \frac{x}{1+|x|}$ for $x \in \mathbb{R}$, $f(+\infty) = 1$, and $f(-\infty) = -1$. Another metric, commonly used on $\mathbb{R} \cup \{+\infty\} \cup \{-\infty\}$, is defined by

$$|\arctan x - \arctan y|,$$

where $-\frac{1}{2}\pi < \arctan x < \frac{1}{2}\pi$ for $-\infty < x < \infty$, and $\arctan(\pm \infty) = \pm \frac{1}{2}\pi$.

• Complex modulus metric

The **complex modulus metric** is a metric on the set \mathbb{C} of complex numbers, defined by

$$|z-u|$$
,

where, for any $z \in \mathbb{C}$, the real number $|z| = |z_1 + z_2i| = \sqrt{z_1^2 + z_2^2}$ is the complex modulus. The metric space $(\mathbb{C}, |z - u|)$ is called the complex plane (or Argand plane).

Examples of other useful metrics on $\mathbb C$ are: the **British Rail metric**, defined by

$$|z| + |u|$$

for $z \neq u$ (and is equal to 0 otherwise); the *p*-relative metric, $1 \leq p \leq \infty$ (cf. (p,q)-relative metric in Chap. 19), defined by

$$\frac{|z-u|}{(|z|^p+|u|^p)^{\frac{1}{p}}}$$

for $|z| + |u| \neq 0$ (and is equal to 0 otherwise); for $p = \infty$ one obtains the **relative metric**, written for $|z| + |u| \neq 0$ as

$$\frac{|z-u|}{\max\{|z|,|u|\}}.$$

• Chordal metric

The **chordal metric** d_{χ} is a metric on the set $\overline{\mathbb{C}}=\mathbb{C}\cup\{\infty\}$, defined by

$$d_{\chi}(z,u) = \frac{2|z-u|}{\sqrt{1+|z|^2}\sqrt{1+|u|^2}}$$

for all $z, u \in \mathbb{C}$, and by

$$d_{\chi}(z,\infty) = \frac{2}{\sqrt{1+|z|^2}}$$

for all $z \in \mathbb{C}$ (cf. *M*-relative metric in Chap. 19). The metric space $(\overline{\mathbb{C}}, d_{\chi})$ is called the *extended complex plane*. It is homeomorphic and conformally equivalent to the *Riemann sphere*.

In fact, a *Riemann sphere* is a sphere in the Euclidean space \mathbb{E}^3 , considered as a metric subspace of \mathbb{E}^3 , onto which the extended complex plane

is one-to-one mapped under stereographic projection. The unit sphere $S^2 = \{(x_1, x_2, x_3) \in \mathbb{E}^3 : x_1^2 + x_2^2 + x_3^2 = 1\}$ can be taken as the Riemann sphere, and the plane $\overline{\mathbb{C}}$ can be identified with the plane $x_3 = 0$ such that the real axis coincides with the x_1 -axis, and the imaginary axis with the x_2 -axis. Under stereographic projection, each point $z \in \mathbb{C}$ corresponds to the point $(x_1, x_2, x_3) \in S^2$ obtained as the point where the ray drawn from the "north pole" (0,0,1) of the sphere to the point z meets the sphere S^2 ; the "north pole" corresponds to the point at infinity ∞ . The chordal distance between two points $p, q \in S^2$ is taken to be the distance between their preimages $z, u \in \overline{\mathbb{C}}$.

The chordal metric can be defined equivalently on $\overline{\mathbb{R}}^n = \mathbb{R}^n \cup \{\infty\}$. Thus, for any $x, y \in \mathbb{R}^n$, one has

$$d_{\chi}(x,y) = \frac{2||x-y||_2}{\sqrt{1+||x||_2^2}\sqrt{1+||y||_2^2}},$$

and for any $x \in \mathbb{R}^n$, one has

$$d_{\chi}(x,\infty) = \frac{2}{\sqrt{1 + ||x||_{2}^{2}}},$$

where $||.||_2$ is the ordinary Euclidean norm on \mathbb{R}^n .

The metric space (\mathbb{R}^n, d_{χ}) is called the **Möbius space**. It is a *Ptolemaic* metric space (cf. **Ptolemaic metric** in Chap. 1).

Given $\alpha > 0$, $\beta \geq 0$, $p \geq 1$, the **generalized chordal metric** is a metric on \mathbb{C} (in general, on $(\mathbb{R}^n, ||.||_2)$ and even on any Ptolemaic space (V, ||.||)), defined by

$$\frac{|z-u|}{(\alpha+\beta|z|^p)^{\frac{1}{p}}\cdot(\alpha+\beta|u|^p)^{\frac{1}{p}}}.$$

It can be easily generalized to $\overline{\mathbb{C}}$ (or $\overline{\mathbb{R}}^n$).

• Quaternion metric

Quaternions are members of a non-commutative division algebra \mathcal{Q} over the field \mathbb{R} , geometrically realizable in a four-dimensional space [Hami66]. The quaternions can be written in the form $q=q_1+q_2i+q_3j+q_4k$, $q_i \in \mathbb{R}$, where the quaternions i, j, and k, called the basic units, satisfy the following identities, known as Hamilton's rules: $i^2=j^2=k^2=-1$, and ij=-ji=k.

The quaternion norm ||q|| of $q = q_1 + q_2i + q_3j + q_4k \in \mathcal{Q}$ is defined by

$$||q|| = \sqrt{q\overline{q}} = \sqrt{q_1^2 + q_2^2 + q_3^2 + q_4^2}, \ \overline{q} = q_1 - q_2 i - q_3 j - q_4 k.$$

The quaternion metric is a norm metric ||x-y|| on the set Q of all quaternions.

12.2 Metrics on polynomials

A polynomial is an expression involving a sum of powers in one or more variables multiplied by coefficients. A polynomial in one variable (or monic polynomial) with constant real (complex) coefficients is given by $P = P(z) = \sum_{k=0}^{n} a_k z^k$, $a_k \in \mathbb{R}$ $(a_k \in \mathbb{C})$.

The set \mathcal{P} of all real (complex) polynomials forms a ring $(\mathcal{P}, +, \cdot, 0)$. It is also a vector space over \mathbb{R} (over \mathbb{C}).

• Polynomial norm metric

A polynomial norm metric (or polynomial bar metric) is a norm metric on the set \mathcal{P} of all real (complex) polynomials, defined by

$$||P - Q||$$

where ||.|| is a polynomial norm, i.e., a function $||.|| : \mathcal{P} \to \mathbb{R}$ such that, for all $P, Q \in \mathcal{P}$ and for any scalar k, we have the following properties:

- 1. ||P|| > 0, with ||P|| = 0 if and only if $P \equiv 0$.
- 2. ||kP|| = |k|||P||.
- 3. $||P + Q|| \le ||P|| + ||Q||$ (triangle inequality).

For the set \mathcal{P} several classes of norms are commonly used. The l_p -norm, $1 \leq p \leq \infty$, of a polynomial $P(z) = \sum_{k=0}^n a_k z^k$ is defined by

$$||P||_p = (\sum_{k=0}^n |a_k|^p)^{1/p},$$

giving the special cases $||P||_1 = \sum_{k=0}^n |a_k|$, $||P||_2 = \sqrt{\sum_{k=0}^n |a_k|^2}$, and $||P||_{\infty} = \max_{0 \le k \le n} |a_k|$. The value $||P||_{\infty}$ is called the *polynomial height*. The L_p -norm, $1 \le p \le \infty$, of a polynomial $P(z) = \sum_{k=0}^n a_k z^k$ is defined by

$$||P||_{L_p} = (\int_0^{2\pi} |P(e^{i\theta})|^p \frac{d\theta}{2\pi})^{\frac{1}{p}},$$

giving the special cases $||P||_{L_1} = \int_0^{2\pi} |P(e^{i\theta})| \frac{d\theta}{2\pi}$, $||P||_{L_2} = \sqrt{\int_0^{2\pi} |P(e^{i\theta})|^2 \frac{d\theta}{2\pi}}$, and $||P||_{L_{\infty}} = \sup_{|z|=1} |P(z)|$.

• Bombieri metric

The Bombieri metric (or polynomial bracket metric) is a polynomial norm metric on the set \mathcal{P} of all real (complex) polynomials, defined by

$$[P-Q]_p,$$

where $[.]_p$, $0 \le p \le \infty$, is the *Bombieri p-norm*. For a polynomial $P(z) = \sum_{k=0}^n a_k z^k$ it is defined by

$$[P]_p = \left(\sum_{k=0}^n \binom{n}{k}^{1-p} |a_k|^p\right)^{\frac{1}{p}},$$

where $\binom{n}{k}$ is a binomial coefficient.

• Metric space of roots

The **metric space of roots** is (Ćurgus and Mascioni 2006) the space (X, d) where X is the family of all multisets of complex numbers with n elements and the distance between multisets $U = \{u_1, \ldots, u_n\}$ and $V = \{v_1, \ldots, v_n\}$ is defined by the following analog of the **Fréchet metric**:

$$\min_{\tau \in Sym_n} \max_{1 \le j \le n} |u_j - v_{\tau(j)}|,$$

where τ is any permutation of $\{1, \ldots, n\}$. Here the set of roots of some monic complex polynomial of degree n is considered as a multiset with n elements. Cf. **metrics between multisets** in Chap. 1.

The function assigning to each polynomial the multiset of its roots is (Ćurgus and Mascioni 2006) a **homeomorphism** between the metric space of all monic complex polynomials of degree n with the **polynomial** norm metric l_{∞} and the metric space of roots.

12.3 Metrics on matrices

An $m \times n$ matrix $A = ((a_{ij}))$ over a field \mathbb{F} is a table consisting of m rows and n columns with the entries a_{ij} from \mathbb{F} . The set of all $m \times n$ matrices with real (complex) entries is denoted by $M_{m,n}$. It forms a group $(M_{m,n}, +, 0_{m,n})$, where $((a_{ij})) + ((b_{ij})) = ((a_{ij} + b_{ij}))$, and the matrix $0_{m,n} \equiv 0$, i.e., all its entries are equal to 0. It is also an mn-dimensional vector space over \mathbb{R} (over \mathbb{C}). The transpose of a matrix $A = ((a_{ij})) \in M_{m,n}$ is the matrix $A^T = ((a_{ij})) \in M_{n,m}$. The conjugate transpose (or adjoint) of a matrix $A = ((a_{ij})) \in M_{m,n}$ is the matrix $A^* = ((a_{ij})) \in M_{n,m}$.

A matrix is called a square matrix if m = n. The set of all square $n \times n$ matrices with real (complex) entries is denoted by M_n . It forms a ring $(M_n, +, \cdot, 0_n)$, where + and 0_n are defined as above, and $((a_{ij})) \cdot ((b_{ij})) = ((\sum_{k=1}^n a_{ik}b_{kj}))$. It is also an n^2 -dimensional vector space over \mathbb{R} (over \mathbb{C}). The trace of a square $n \times n$ matrix $A = ((a_{ij}))$ is defined to be the sum of the elements on the main diagonal (the diagonal from the upper left to the lower right) of A, i.e., $TrA = \sum_{i=1}^n a_{ii}$. A matrix $A = ((a_{ij})) \in M_n$ is called symmetric if $a_{ij} = a_{ji}$ for all $i, j \in \{1, \ldots, n\}$, i.e., if $A = A^T$. The identity matrix is $1_n = ((c_{ij}))$ with $c_{ii} = 1$, and $c_{ij} = 0$, $i \neq j$. A unitary matrix

 $U = ((u_{ij}))$ is a square matrix, defined by $U^{-1} = U^*$, where U^{-1} is the inverse matrix of U, i.e., $U \cdot U^{-1} = 1_n$. An orthonormal matrix is a matrix $A \in M_{m,n}$ such that $A^*A = 1_n$.

If for a matrix $A \in M_n$ there is a vector x such that $Ax = \lambda x$ for some scalar λ , then λ is called an *eigenvalue* of A with corresponding *eigenvector* x. Given a complex matrix $A \in M_{m,n}$, its *singular values* $s_i(A)$ are defined as the square roots of the eigenvalues of the matrix A^*A , where A^* is the conjugate transpose of A. They are non-negative real numbers $s_1(A) \geq s_2(A) \geq \ldots$

• Matrix norm metric

A matrix norm metric is a norm metric on the set $M_{m,n}$ of all real (complex) $m \times n$ matrices, defined by

$$||A - B||$$
,

where ||.|| is a matrix norm, i.e., a function $||.||: M_{m,n} \to \mathbb{R}$ such that, for all $A, B \in M_{m,n}$, and for any scalar k, we have the following properties:

- 1. $||A|| \ge 0$, with ||A|| = 0 if and only if $A = 0_{m,n}$.
- 2. ||kA|| = |k|||A||.
- 3. $||A + B|| \le ||A|| + ||B||$ (triangle inequality).

All matrix norm metrics on $M_{m,n}$ are equivalent. A matrix norm ||.|| on the set M_n of all real (complex) square $n \times n$ matrices is called *sub-multiplicative* if it is *compatible* with matrix multiplication, i.e., $||AB|| \le ||A|| \cdot ||B||$ for all $A, B \in M_n$. The set M_n with a sub-multiplicative norm is a *Banach algebra*.

The simplest example of a matrix norm metric is the **Hamming metric** on $M_{m,n}$ (in general, on the set $M_{m,n}(\mathbb{F})$ of all $m \times n$ matrices with entries from a field \mathbb{F}), defined by $||A - B||_H$, where $||A||_H$ is the *Hamming norm* of $A \in M_{m,n}$, i.e., the number of non-zero entries in A.

• Natural norm metric

A natural norm metric (or induced norm metric, subordinate norm metric) is a matrix norm metric on the set M_n of all real (complex) square $n \times n$ matrices, defined by

$$||A - B||_{nat}$$

where $||.||_{nat}$ is a natural norm on M_n .

The natural norm $||.||_{nat}$ on M_n , induced by the vector norm ||x||, $x \in \mathbb{R}^n$ $(x \in \mathbb{C}^n)$, is a sub-multiplicative matrix norm, defined by

$$||A||_{nat} = \sup_{||x|| \neq 0} \frac{||Ax||}{||x||} = \sup_{||x|| = 1} ||Ax|| = \sup_{||x|| \leq 1} ||Ax||.$$

The natural norm metric can be defined in similar way on the set $M_{m,n}$ of all $m \times n$ real (complex) matrices: given vector norms $||.||_{\mathbb{R}^m}$ on \mathbb{R}^m and $||.||_{\mathbb{R}^n}$ on \mathbb{R}^n , the natural norm $||A||_{nat}$ of a matrix $A \in M_{m,n}$, induced by $||.||_{\mathbb{R}^n}$ and $||.||_{\mathbb{R}^m}$, is a matrix norm, defined by $||A||_{nat} = \sup_{||x||_{\mathbb{R}^n} = 1} ||Ax||_{\mathbb{R}^m}$.

• Matrix p-norm metric

A matrix p-norm metric is a natural norm metric on M_n , defined by

$$||A-B||_{nat}^p$$

where $||.||_{nat}^p$ is the matrix p-norm, i.e., a natural norm, induced by the vector l_p -norm, $1 \le p \le \infty$:

$$||A||_{nat}^p = \max_{||x||_p = 1} ||Ax||_p$$
, where $||x||_p = (\sum_{i=1}^n |x_i|^p)^{1/p}$.

The maximum absolute column metric (more exactly, maximum absolute column sum norm metric) is the matrix 1-norm metric $||A-B||_{nat}^1$ on M_n . The matrix 1-norm $||.||_{nat}^1$, induced by the vector l_1 -norm, is also called the maximum absolute column sum norm. For a matrix $A = ((a_{ij})) \in M_n$ it can be written as

$$||A||_{nat}^1 = \max_{1 \le j \le n} \sum_{i=1}^n |a_{ij}|.$$

The maximum absolute row metric (more exactly, maximum absolute row sum norm metric) is the matrix ∞ -norm metric $||A-B||_{nat}^{\infty}$ on M_n . The matrix ∞ -norm $||.||_{nat}^{\infty}$, induced by the vector l_{∞} -norm, is also called the maximum absolute row sum norm. For a matrix $A = ((a_{ij})) \in M_n$ it can be written as

$$||A||_{nat}^{\infty} = \max_{1 \le j \le n} \sum_{j=1}^{n} |a_{ij}|.$$

The spectral norm metric is the matrix 2-norm metric $||A - B||_{nat}^2$ on M_n . The matrix 2-norm $||.||_{nat}^2$, induced by the vector l_2 -norm, is also called the *spectral norm* and denoted by $||.||_{sp}$. For a matrix $A = ((a_{ij})) \in M_n$, it can be written as

$$||A||_{sp} = (\text{ maximum eigenvalue of } A^*A)^{\frac{1}{2}},$$

where $A^* = ((\overline{a}_{ji})) \in M_n$ is the conjugate transpose of A (cf. **Ky Fan norm metric**).

• Frobenius norm metric

The Frobenius norm metric is a matrix norm metric on $M_{m,n}$, defined by

$$||A-B||_{Fr},$$

where $||.||_{Fr}$ is the Frobenius norm. For a matrix $A = ((a_{ij})) \in M_{m,n}$, it is

$$||A||_{Fr} = \sqrt{\sum_{i=1}^{m} \sum_{j=1}^{n} |a_{ij}|^2}.$$

It is also equal to the square root of the matrix trace of A^*A , where $A^* = ((\overline{a}_{ji}))$ is the conjugate transpose of A, or, equivalently, to the square root of the sum of eigenvalues λ_i of A^*A : $||A||_{Fr} = \sqrt{Tr(A^*A)} =$

 $\sqrt{\sum_{i=1}^{\min\{m,n\}} \lambda_i}$ (cf. **Schatten norm metric** in Chap. 13). This norm comes from an *inner product* on the space $M_{m,n}$, but it is not *submultiplicative* for m=n.

• (c, p)-norm metric

Let $k \in \mathbb{N}$, $k \leq \min\{m, n\}$, $c \in \mathbb{R}^k$, $c_1 \geq c_2 \geq \cdots \geq c_k > 0$, and $1 \leq p < \infty$. The (c, p)-norm metric is a matrix norm metric on $M_{m,n}$, defined by

$$||A-B||_{(c,p)}^k$$

where $||.||_{(c,p)}^k$ is the (c,p)-norm on $M_{m,n}$. For a matrix $A \in M_{m,n}$, it is defined by

$$||A||_{(c,p)}^k = (\sum_{i=1}^k c_i s_i^p(A))^{\frac{1}{p}},$$

where $s_1(A) \geq s_2(A) \geq \cdots \geq s_k(A)$ are the first k singular values of A. If p = 1, one obtains the c-norm. If, moreover, $c_1 = \cdots = c_k = 1$, one obtains the Ky Fan k-norm.

• Ky Fan norm metric

Given $k \in \mathbb{N}$, $k \leq \min\{m, n\}$, the **Ky Fan norm metric** is a **matrix norm metric** on $M_{m,n}$, defined by

$$||A-B||_{KF}^k$$

where $||.||_{KF}^k$ is the $Ky\ Fan\ k$ -norm on $M_{m,n}$. For a matrix $A \in M_{m,n}$, it is defined as the sum of its first $k\ singular\ values$:

$$||A||_{KF}^k = \sum_{i=1}^k s_i(A).$$

For k=1, it is the *spectral norm*. For $k=\min\{m,n\}$, one obtains the *trace norm*.

• Schatten norm metric

Given $1 \le p < \infty$, the Schatten norm metric is a matrix norm metric on $M_{m,n}$, defined by

$$||A - B||_{Sch}^p,$$

where $||.||_{Sch}^p$ is the *Schatten p-norm* on $M_{m,n}$. For a matrix $A \in M_{m,n}$, it is defined as the *p*-th root of the sum of the *p*-th powers of all its *singular values*:

$$||A||_{Sch}^p = \left(\sum_{i=1}^{\min\{m,n\}} s_i^p(A)\right)^{\frac{1}{p}}.$$

For p = 2, one obtains the *Frobenius norm* and, for p = 1, one obtains the trace norm.

• Trace norm metric

The trace norm metric is a matrix norm metric on $M_{m,n}$, defined by

$$||A-B||_{tr},$$

where $||.||_{tr}$ is the trace norm on $M_{m,n}$. For a matrix $A \in M_{m,n}$, it is defined as the sum of all its singular values:

$$||A||_{tr} = \sum_{i=1}^{\min\{m,n\}} s_i(A).$$

• Cut norm metric

The cut norm metric is a matrix norm metric on $M_{m,n}$, defined by

$$||A - B||_{cut}$$

where $||.||_{cut}$ is the *cut norm* on $M_{m,n}$ defined, for a matrix $A = ((a_{ij})) \in M_{m,n}$, as:

$$||A||_{cut} = \max_{I \subset \{1, \dots, m\}, J \subset \{1, \dots, n\}} |\sum_{i \in I, j \in J} a_{ij}|.$$

Cf. in Chap. 15 the **rectangle distance on weighted graphs** and the **cut semi-metric**, but the **weighted cut metric** in Chap. 19 is not related.

• Sym⁺ metric

Let Sym^+ be the set of all $n \times n$ real positive definite matrices, i.e., matrices A such that $x^TAx > 0$ for any non-zero vector $x \in \mathbb{R}^n$.

The Sym^+ metric is defined, for any $A, B \in Sym^+$, as

$$(\sum_{i=1}^n \log^2 \lambda_i)^{\frac{1}{2}},$$

where $\lambda_1, ..., \lambda_n$ are the eigenvalues of the matrix AB^{-1} .

The Sym^+ metric is the **Riemannian distance**, arising from the Riemannian metric, called the *trace metric*: $ds^2 = Tr(A^{-1}dA)^2$.

• Distances between graphs of matrices

The graph G(A) of a complex $m \times n$ matrix A is the range (i.e., the span of columns) of the matrix $R(A) = ([IA^T])^T$. So, G(A) is a subspace of \mathbb{C}^{m+n} of all vectors v, for which the equation R(A)x = v has a solution.

A distance between graphs of matrices A and B is a distance between the subspaces G(A) and G(B). It can be an **angle distance** between subspaces or, for example, the following distance (cf. also the **Kadets distance** in Chap. 1 and the **gap metric** in Chap. 18).

The **spherical gap distance** between subspaces A and B is defined by

$$\max\{\max_{x \in S(A)} d_E(x, S(B)), \max_{y \in S(B)} d_E(y, S(A))\},\$$

where S(A), S(B) are the unit spheres of the subspaces A, B, d(z, C) is the **point-set distance** inf_{$u \in C$} d(z, y) and $d_E(z, y)$ is the Euclidean distance.

• Angle distances between subspaces

Consider the Grassmannian space G(m, n) of all n-dimensional subspaces of Euclidean space \mathbb{E}^m ; it is a compact Riemannian manifold of dimension n(m-n).

Given two subspaces $A, B \in G(m, n)$, the principal angles $\frac{\pi}{2} \ge \theta_1 \ge \dots \ge \theta_n \ge 0$ between them are defined, for $k = 1, \dots, n$, inductively by

$$\cos \theta_k = \max_{x \in A} \max_{y \in B} x^T y = (x^k)^T y^k$$

subject to the conditions $||x||_2 = ||y||_2 = 1$, $x^T x^i = 0$, $y^T y^i = 0$, for $1 \le i \le k-1$, where $||.||_2$ is the Euclidean norm.

The principal angles can also be defined in terms of orthonormal matrices Q_A and Q_B spanning subspaces A and B, respectively: in fact, n ordered *singular values* of the matrix $Q_AQ_B \in M_n$ can be expressed as cosines $\cos \theta_1, \ldots, \cos \theta_n$.

The **geodesic distance** between subspaces A and B is (Wong 1967) defined by

$$\sqrt{2\sum_{i=1}^n\theta_i^2}.$$

The Martin distance between subspaces A and B is defined by

$$\sqrt{\ln \prod_{i=1}^{n} \frac{1}{\cos^2 \theta_i}}.$$

In the case when the subspaces represent *autoregressive models*, the Martin distance can be expressed in terms of the *cepstrum* of the autocorrelation functions of the models (cf. **Martin cepstrum distance** in Chap. 21).

The **Asimov distance** between subspaces A and B is defined by

$$\theta_1$$
.

It can be expressed also in terms of the **Finsler metric** on the manifold G(m, n).

The gap distance between subspaces A and B is defined by

$$\sin \theta_1$$
.

It can be expressed also in terms of *orthogonal projectors* as the l_2 -norm of the difference of the projectors onto A and B, respectively. Many versions of this distance are used in Control Theory (cf. **gap metric** in Chap. 18).

The **Frobenius distance** between subspaces A and B is defined by

$$\sqrt{2\sum_{i=1}^{n}\sin^2\theta_i}.$$

It can be expressed also in terms of *orthogonal projectors* as the *Frobenius norm* of the difference of the projectors onto A and B, respectively. A similar distance $\sqrt{\sum_{i=1}^{n}\sin^{2}\theta_{i}}$ is called the **chordal distance**.

• Semi-metrics on resemblances

The following two semi-metrics are defined for any two resemblances d_1 and d_2 on a given finite set X (moreover, for any two real symmetric matrices).

The **Lerman semi-metric** (cf. **Kendall** τ **distance** on permutations in Chap. 11) is defined by

$$\frac{|\{(\{x,y\},\{u,v\}): (d_1(x,y)-d_1(u,v))(d_2(x,y)-d_2(u,v))<0\}|}{\binom{|X|+1}{2}^2},$$

where $(\{x,y\},\{u,v\})$ is any pair of unordered pairs of elements x,y,u,v from X.

The **Kaufman semi-metric** is defined by

$$\frac{|\{(\{x,y\},\{u,v\}): (d_1(x,y)-d_1(u,v))(d_2(x,y)-d_2(u,v))<0\}|}{|\{(\{x,y\},\{u,v\}): (d_1(x,y)-d_1(u,v))(d_2(x,y)-d_2(u,v))\neq0\}|}.$$

Chapter 13

Distances in Functional Analysis

Functional Analysis is the branch of Mathematics concerned with the study of spaces of functions. This usage of the word functional goes back to the calculus of variations which studies functions whose argument is a function. In the modern view, Functional Analysis is seen as the study of complete normed vector spaces, i.e., Banach spaces.

For any real number $p \geq 1$, an example of a Banach space is given by L_p -space of all Lebesgue-measurable functions whose absolute value's p-th power has finite integral.

A Hilbert space is a Banach space in which the norm arises from an *inner product*. Also, in Functional Analysis are considered *continuous linear operators* defined on Banach and Hilbert spaces.

13.1 Metrics on function spaces

Let $I \subset \mathbb{R}$ be an open interval (i.e., a non-empty connected open set) in \mathbb{R} . A real function $f: I \to \mathbb{R}$ is called real analytic on I if it agrees with its Taylor series in an open neighborhood U_{x_0} of every point $x_0 \in I$: $f(x) = \sum_{n=0}^{\infty} \frac{f^{(n)}(x_0)}{n!} (x - x_0)^n$ for any $x \in U_{x_0}$. Let $D \subset \mathbb{C}$ be a domain (i.e., a convex open set) in \mathbb{C} .

A complex function $f: D \to \mathbb{C}$ is called *complex analytic* (or, simply, analytic) on D if it agrees with its Taylor series in an open neighborhood of every point $z_0 \in D$. A complex function f is analytic on D if and only if it is holomorphic on D, i.e., if it has a complex derivative $f'(z_0) = \lim_{z \to z_0} \frac{f(z) - f(z_0)}{z - z_0}$ at every point $z_0 \in D$.

• Integral metric

The **integral metric** is the L_1 -metric on the set $C_{[a,b]}$ of all continuous real (complex) functions on a given segment [a,b], defined by

$$\int_{a}^{b} |f(x) - g(x)| dx.$$

The corresponding metric space is abbreviated by $C^1_{[a,b]}$. It is a Banach space.

In general, for any **compact** (or *countably compact*) topological space X the integral metric can be defined on the set of all continuous functions $f: X \to \mathbb{R}$ (\mathbb{C}) by $\int_{Y} |f(x) - g(x)| dx$.

• Uniform metric

The uniform metric (or sup metric) is the L_{∞} -metric on the set $C_{[a,b]}$ of all real (complex) continuous functions on a given segment [a,b], defined by

$$\sup_{x \in [a,b]} |f(x) - g(x)|.$$

The corresponding metric space is abbreviated by $C^{\infty}_{[a,b]}$. It is a Banach space.

A generalization of $C^{\infty}_{[a,b]}$ is the space of continuous functions C(X), i.e., a metric space on the set of all continuous (more generally, bounded) functions $f: X \to \mathbb{C}$ of a topological space X with the L_{∞} -metric $\sup_{x \in X} |f(x) - g(x)|$.

In the case of the metric space C(X,Y) of continuous (more generally, bounded) functions $f: X \to Y$ from one **metric compactum** (X, d_X) to another (Y, d_Y) , the sup metric between two functions $f, g \in C(X,Y)$ is defined by $\sup_{x \in X} d_Y(f(x), g(x))$.

The metric space $C^{\infty}_{[a,b]}$, as well as the metric space $C^1_{[a,b]}$, are two of the most important cases of the metric space $C^p_{[a,b]}$, $1 \leq p \leq \infty$, on the set $C_{[a,b]}$ with the L_p -metric $(\int_a^b |f(x)-g(x)|^p dx)^{\frac{1}{p}}$. The space $C^p_{[a,b]}$ is an example of an L_p -space.

• Dogkeeper distance

Given a metric space (X, d), the **dogkeeper distance** is a metric on the set of all functions $f : [0, 1] \to X$, defined by

$$\inf_{\sigma} \sup_{t \in [0,1]} d(f(t), g(\sigma(t))),$$

where $\sigma : [0,1] \to [0,1]$ is a continuous, monotone increasing function such that $\sigma(0) = 0$, $\sigma(1) = 1$. This metric is a special case of the **Fréchet metric**. It is used for measuring the distances between curves.

• Bohr metric

Let \mathbb{R} be a metric space with a metric ρ . A continuous function $f: \mathbb{R} \to \mathbb{R}$ is called almost periodic if, for every $\epsilon > 0$, there exists $l = l(\epsilon) > 0$ such that every interval $[t_0, t_0 + l(\epsilon)]$ contains at least one number τ for which $\rho(f(t), f(t+\tau)) < \epsilon$ for $-\infty < t < +\infty$.

The **Bohr metric** is the **norm metric** ||f - g|| on the set AP of all almost periodic functions, defined by the norm

$$||f|| = \sup_{-\infty < t < +\infty} |f(t)|.$$

It makes AP a Banach space. Some generalizations of almost periodic functions were obtained using other norms; cf. **Stepanov distance**, **Weyl distance**, **Besicovitch distance** and **Bochner metric**.

• Stepanov distance

The **Stepanov distance** is a distance on the set of all measurable functions $f: \mathbb{R} \to \mathbb{C}$ with summable p-th power on each bounded integral, defined by

$$\sup_{x \in \mathbb{R}} \left(\frac{1}{l} \int_{x}^{x+l} |f(x) - g(x)|^{p} dx \right)^{1/p}.$$

The Weyl distance is a distance on the same set, defined by

$$\lim_{l \to \infty} \sup_{x \in \mathbb{R}} \left(\frac{1}{l} \int_x^{x+l} |f(x) - g(x)|^p dx \right)^{1/p}.$$

Corresponding to these distances one has the *generalized Stepanov* and Weyl almost periodic functions.

• Besicovitch distance

The **Besicovitch distance** is a distance on the set of all measurable functions $f: \mathbb{R} \to \mathbb{C}$ with summable p-th power on each bounded integral, defined by

$$\left(\overline{\lim}_{T\to\infty}\frac{1}{2T}\int_{-T}^{T}|f(x)-g(x)|^pdx\right)^{1/p}.$$

The generalized Besicovitch almost periodic functions correspond to this distance.

• Bochner metric

Given a measure space $(\Omega, \mathcal{A}, \mu)$, a Banach space $(V, ||.||_V)$, and $1 \leq p \leq \infty$, the Bochner space (or Lebesgue-Bochner space) $L^p(\Omega, V)$ is the set of all measurable functions $f: \Omega \to V$ such that $||f||_{L^p(\Omega, V)} \leq \infty$.

Here the Bochner norm $||f||_{L^p(\Omega,V)}$ is defined by $(\int_{\Omega} ||f(\omega)||_V^p d\mu(\omega))^{\frac{1}{p}}$ for $1 \leq p < \infty$, and, for $p = \infty$, by $ess \sup_{\omega \in \Omega} ||f(\omega)||_V$.

• Bergman p-metric

Given $1 \leq p < \infty$, let $L_p(\Delta)$ be the L_p -space of Lebesgue measurable functions f on the unit disk $\Delta = \{z \in \mathbb{C} : |z| < 1\}$ with $||f||_p = \left(\int_{\Delta} |f(z)|^p \mu(dz)\right)^{\frac{1}{p}} < \infty$.

The Bergman space $L_p^a(\Delta)$ is the subspace of $L_p(\Delta)$ consisting of analytic functions, and the **Bergman** p-metric is the L_p -metric on $L_p^a(\Delta)$ (cf. Bergman metric in Chap. 7). Any Bergman space is a Banach space.

• Bloch metric

The Bloch space B on the unit disk $\Delta = \{z \in \mathbb{C} : |z| < 1\}$ is the set of all analytic functions f on Δ such that $||f||_B = \sup_{z \in \Delta} (1 - |z|^2) |f'(z)| < \infty$. Using the complete semi-norm $||.||_B$, a norm on B is defined by

$$||f|| = |f(0)| + ||f||_B.$$

The **Bloch metric** is the **norm metric** ||f - g|| on B. It makes B a Banach space.

• Besov metric

Given $1 , the Besov space <math>B_p$ on the unit disk $\Delta = \{z \in \mathbb{C} : |z| < 1\}$ is the set of all analytic functions f in Δ such that $||f||_{B_p} = \left(\int_{\Delta} (1-|z|^2)^p |f'(z)|^p d\lambda(z)\right)^{\frac{1}{p}} < \infty$, where $d\lambda(z) = \frac{\mu(dz)}{(1-|z|^2)^2}$ is the Möbius invariant measure on Δ . Using the complete semi-norm $||.||_{B_p}$, the Besov norm on B_p is defined by

$$||f|| = |f(0)| + ||f||_{B_n}$$

The **Besov metric** is the **norm metric** ||f - g|| on B_p . It makes B_p a Banach space.

The set B_2 is the classical *Dirichlet space* of functions analytic on Δ with square integrable derivative, equipped with the **Dirichlet metric**. The *Bloch space* B can be considered as B_{∞} .

• Hardy metric

Given $1 \leq p < \infty$, the Hardy space $H^p(\Delta)$ is the class of functions, analytic on the unit disk $\Delta = \{z \in \mathbb{C} : |z| < 1\}$, and satisfying the following growth condition for the Hardy norm $||.||_{H^p}$:

$$||f||_{H^p(\Delta)} = \sup_{0 < r < 1} \left(\frac{1}{2\pi} \int_0^{2\pi} |f(re^{i\theta})|^p d\theta \right)^{\frac{1}{p}} < \infty.$$

The **Hardy metric** is the **norm metric** $||f-g||_{H^p(\Delta)}$ on $H^p(\Delta)$. It makes $H^p(\Delta)$ a Banach space.

In Complex Analysis, the Hardy spaces are analogs of the L_p -spaces of Functional Analysis. Such spaces are applied in Mathematical Analysis itself, and also to Scattering Theory and Control Theory (cf. Chap. 18).

Part metric

The **part metric** is a metric on a *domain* D of \mathbb{R}^2 , defined for any $x, y \in \mathbb{R}^2$ by

$$\sup_{f \in H^+} \left| \ln \left(\frac{f(x)}{f(y)} \right) \right|,$$

where H^+ is the set of all positive harmonic functions on the domain D. A twice-differentiable real function $f:D\to\mathbb{R}$ is called harmonic on D if its Laplacian $\Delta f=\frac{\partial^2 f}{\partial x_1^2}+\frac{\partial^2 f}{\partial x_2^2}$ vanishes on D.

• Orlicz metric

Let M(u) be an even convex function of a real variable which is increasing for u positive, and $\lim_{u\to 0} u^{-1}M(u) = \lim_{u\to \infty} u(M(u))^{-1} = 0$. In this case the function p(v) = M'(v) does not decrease on $[0, \infty)$, p(0) = 0

 $\lim_{v\to 0} p(v) = 0$, and p(v) > 0 when v > 0. Writing $M(u) = \int_0^{|u|} p(v) dv$, and defining $N(u) = \int_0^{|u|} p^{-1}(v) dv$, one obtains a pair (M(u), N(u)) of complementary functions.

Let (M(u), N(u)) be a pair of complementary functions, and let G be a bounded closed set in \mathbb{R}^n . The $Orlicz\ space\ L_M^*(G)$ is the set of Lebesgue-measurable functions f on G satisfying the following growth condition for the $Orlicz\ norm\ ||f||_M$:

$$||f||_{M}=\sup\left\{\int_{G}f(t)g(t)dt:\int_{G}N(g(t))dt\leq1\right\}<\infty.$$

The **Orlicz metric** is the norm metric $||f - g||_M$ on $L_M^*(G)$. It makes $L_M^*(G)$ a Banach space [Orli32].

When $M(u) = u^p$, $1 , <math>L_M^*(G)$ coincides with the space $L_p(G)$, and, up to scalar factor, the L_p -norm $||f||_p$ coincides with $||f||_M$. The Orlicz norm is equivalent to the Luxemburg norm $||f||_{(M)} = \inf\{\lambda > 0: \int_G M(\lambda^{-1}f(t))dt \le 1\}$; in fact, $||f||_{(M)} \le ||f||_M \le 2||f||_{(M)}$.

• Orlicz-Lorentz metric

Let $w:(0,\infty)\to(0,\infty)$ be a non-increasing function. Let $M:[0,\infty)\to[0,\infty)$ be a non-decreasing and convex function with M(0)=0. Let G be a bounded closed set in \mathbb{R}^n .

The Orlicz-Lorentz space $L_{w,M}(G)$ is the set of all Lebesgue-measurable functions f on G satisfying the following growth condition for the Orlicz-Lorentz norm $||f||_{w,M}$:

$$||f||_{w,M} = \inf \left\{ \lambda > 0 : \int_0^\infty w(x) M\left(\frac{f^*(x)}{\lambda}\right) dx \le 1 \right\} < \infty,$$

where $f^*(x) = \sup\{t : \mu(|f| \ge t) \ge x\}$ is the non-increasing rearrangement of f.

The Orlicz-Lorentz metric is the norm metric $||f - g||_{w,M}$ on $L_{w,M}(G)$. It makes $L_{w,M}(G)$ a Banach space.

The Orlicz–Lorentz space is a generalization of the Orlicz space $L_M^*(G)$ (cf. Orlicz metric), and the Lorentz space $L_{w,q}(G)$, $1 \leq q < \infty$, of all Lebesgue-measurable functions f on G satisfying the following growth condition for the Lorentz norm:

$$||f||_{w,q} = \left(\int_0^\infty w(x)(f^*(x))^q\right)^{\frac{1}{q}} < \infty.$$

• Hölder metric

Let $L^{\alpha}(G)$ be the set of all bounded continuous functions f, defined on a subset G of \mathbb{R}^n , and satisfying the *Hölder condition* on G. Here, a function f satisfies the *Hölder condition* at a point $y \in G$ with index (or order) α , $0 < \alpha \le 1$, and with coefficient A(y), if $|f(x) - f(y)| \le A(y)|x - y|^{\alpha}$ for

all $x \in G$ sufficiently close to y. If $A = \sup_{y \in G} (A(y)) < \infty$, the Hölder condition is called uniform on G, and A is called the $H\"{o}lder$ coefficient of G. The quantity $|f|_{\alpha} = \sup_{x,y \in G} \frac{|f(x) - f(y)|}{|x - y|^{\alpha}}$, $0 \le \alpha \le 1$, is called the $H\"{o}lder$ α -semi-norm of f, and the $H\"{o}lder$ norm of f is defined by

$$||f||_{L^{\alpha}(G)} = \sup_{x \in G} |f(x)| + |f|_{\alpha}.$$

The **Hölder metric** is the **norm metric** $||f-g||_{L^{\alpha}(G)}$ on $L^{\alpha}(G)$. It makes $L^{\alpha}(G)$ a Banach space.

• Sobolev metric

The Sobolev space $W^{k,p}$ is a subset of an L_p -space such that f and its derivatives up to order k have a finite L_p -norm. Formally, given a subset G of \mathbb{R}^n , define

$$W^{k,p} = W^{k,p}(G) = \{ f \in L_p(G) : f^{(i)} \in L_p(G), 1 \le i \le k \},$$

where $f^{(i)} = \partial_{x_1}^{\alpha_1} \dots \partial_{x_n}^{\alpha_n} f$, $\alpha_1 + \dots + \alpha_n = i$, and the derivatives are taken in a weak sense. The *Sobolev norm* on $W^{k,p}$ is defined by

$$||f||_{k,p} = \sum_{i=0}^{k} ||f^{(i)}||_{p}.$$

In fact, it is enough to take only the first and last in the sequence, i.e., the norm defined by $||f||_{k,p} = ||f||_p + ||f^{(k)}||_p$ is equivalent to the norm above.

For $p=\infty$, the Sobolev norm is equal to the essential supremum of |f|: $||f||_{k,\infty}=ess\sup_{x\in G}|f(x)|$, i.e., it is the infimum of all numbers $a\in\mathbb{R}$ for which |f(x)|>a on a set of measure zero.

The Sobolev metric is the norm metric $||f-g||_{k,p}$ on $W^{k,p}$. It makes $W^{k,p}$ a Banach space.

The Sobolev space $W^{k,2}$ is denoted by H^k . It is a Hilbert space for the inner product $\langle f,g\rangle_k=\sum_{i=1}^k\langle f^{(i)},g^{(i)}\rangle_{L_2}=\sum_{i=1}^k\int_Gf^{(i)}\overline{g}^{(i)}\mu(d\omega)$. Sobolev spaces are the modern replacement for the space C^1 (of func-

Sobolev spaces are the modern replacement for the space C^1 (of functions having continuous derivatives) for solutions of partial differential equations.

• Variable exponent space metrics

Let G be a non-empty open subset of \mathbb{R}^n , and let $p:G \to [1,\infty)$ be a measurable bounded function, called a variable exponent. The variable exponent Lebesgue space $L_{p(.)}(G)$ is the set of all measurable functions $f:G\to\mathbb{R}$ for which the modular $\varrho_{p(.)}(f)=\int_G|f(x)|^{p(x)}dx$ is finite. The Luxemburg norm on this space is defined by

$$||f||_{p(.)} = \inf\{\lambda > 0 : \varrho_{p(.)}(f/\lambda) \le 1\}.$$

The variable exponent Lebesgue space metric is the norm metric $||f-g||_{p(.)}$ on $L_{p(.)}(G)$.

A variable exponent Sobolev space $W^{1,p(.)}(G)$ is a subspace of $L_{p(.)}(G)$ consisting of functions f whose distributional gradient exists almost everywhere and satisfies the condition $|\nabla f| \in L_{p(.)}(G)$. The norm

$$||f||_{1,p(.)} = ||f||_{p(.)} + ||\nabla f||_{p(.)}$$

makes $W^{1,p(.)}(G)$ a Banach space. The variable exponent Sobolev space metric is the norm metric $||f-g||_{1,p(.)}$ on $W^{1,p(.)}$.

• Schwartz metric

The Schwartz space (or space of rapidly decreasing functions) $S(\mathbb{R}^n)$ is the class of all Schwartz functions, i.e., infinitely-differentiable functions $f: \mathbb{R}^n \to \mathbb{C}$ that decrease at infinity, as do all their derivatives, faster than any inverse power of x. More precisely, f is a Schwartz function if we have the following growth condition:

$$||f||_{\alpha\beta} = \sup_{x \in \mathbb{R}^n} |x_1^{\beta_1} \dots x_n^{\beta_n} \frac{\partial^{\alpha_1 + \dots + \alpha_n} f(x_1, \dots, x_n)}{\partial x_1^{\alpha_1} \dots \partial x_n^{\alpha_n}}| < \infty$$

for any non-negative integer vectors α and β . The family of *semi-norms* $||.||_{\alpha\beta}$ defines a **locally convex** topology of $S(\mathbb{R}^n)$ which is **metrizable** and complete. The **Schwartz metric** is a metric on $S(\mathbb{R}^n)$ which can be obtained using this topology (cf. **countably normed space** in Chap. 2).

The corresponding metric space on $S(\mathbb{R}^n)$ is a Fréchet space in the sense of Functional Analysis, i.e., a locally convex F-space.

• Bregman quasi-distance

Let $G \subset \mathbb{R}^n$ be a closed set with the non-empty interior G^0 . Let f be a Bregman function with zone G.

The **Bregman quasi-distance** $D_f: G \times G^0 \to \mathbb{R}_{\geq 0}$ is defined by

$$D_f(x,y) = f(x) - f(y) - \langle \nabla f(y), x - y \rangle,$$

where $\nabla f = (\frac{\partial f}{\partial x_1}, \dots \frac{\partial f}{\partial x_p})$. $D_f(x,y) = 0$ if and only if x = y. Also $D_f(x,y) + D_f(y,z) - D_f(x,z) = \langle \nabla f(z) - \nabla f(y), x - y \rangle$ but, in general, D_f does not satisfy the triangle inequality, and is not symmetric.

A real-valued function f whose effective domain contains G is called a Bregman function with zone G if the following conditions hold:

- 1. f is continuously differentiable on G^0 .
- 2. f is strictly convex and continuous on G.
- 3. For all $\delta \in \mathbb{R}$ the partial level sets $\Gamma(x,\delta) = \{y \in G^0 : D_f(x,y) \leq \delta\}$ are bounded for all $x \in G$.
- 4. If $\{y_n\}_n \subset G^0$ converges to y^* , then $D_f(y^*, y_n)$ converges to 0.

5. If $\{x_n\}_n \subset G$ and $\{y_n\}_n \subset G^0$ are sequences such that $\{x_n\}_n$ is bounded, $\lim_{n\to\infty} y_n = y^*$, and $\lim_{n\to\infty} D_f(x_n,y_n) = 0$, then $\lim_{n\to\infty} x_n = y^*$.

When $G = \mathbb{R}^n$, a sufficient condition for a strictly convex function to be a Bregman function has the form: $\lim_{||x|| \to \infty} \frac{f(x)}{||x||} = \infty$.

13.2 Metrics on linear operators

A linear operator is a function $T: V \to W$ between two vector spaces V, W over a field \mathbb{F} , that is compatible with their linear structures, i.e., for any $x, y \in V$ and any scalar $k \in \mathbb{F}$, we have the following properties: T(x + y) = T(x) + T(y), and T(kx) = kT(x).

• Operator norm metric

Consider the set of all linear operators from a normed space $(V, ||.||_V)$ into a normed space $(W, ||.||_W)$. The operator norm ||T|| of a linear operator $T: V \to W$ is defined as the largest value by which T stretches an element of V, i.e.,

$$||T|| = \sup_{||v||_V \neq 0} \frac{||T(v)||_W}{||v||_V} = \sup_{||v||_V = 1} ||T(v)||_W = \sup_{||v||_V \leq 1} ||T(v)||_W.$$

A linear operator $T:V\to W$ from a normed space V into a normed space W is called *bounded* if its operator norm is finite. For normed spaces, a linear operator is bounded if and only if it is *continuous*.

The **operator norm metric** is a **norm metric** on the set B(V, W) of all bounded linear operators from V into W, defined by

$$||T-P||$$
.

The space (B(V, W), ||.||) is called the *space of bounded linear operators*. This metric space is **complete** if W is. If V = W is complete, the space B(V, V) is a *Banach algebra*, as the operator norm is a *sub-multiplicative norm*.

A linear operator $T:V\to W$ from a Banach space V into another Banach space W is called compact if the image of any bounded subset of V is a relatively compact subset of W. Any compact operator is bounded (and, hence, continuous). The space (K(V,W),||.||) on the set K(V,W) of all compact operators from V into W with the operator norm ||.|| is called the $space\ of\ compact\ operators$.

• Nuclear norm metric

Let B(V, W) be the space of all bounded linear operators mapping a Banach space $(V, ||.||_V)$ into another Banach space $(W, ||.||_W)$. Let the Banach

dual of V be denoted by V', and the value of a functional $x' \in V'$ at a vector $x \in V$ by $\langle x, x' \rangle$. A linear operator $T \in B(V, W)$ is called a nuclear operator if it can be represented in the form $x \mapsto T(x) = \sum_{i=1}^{\infty} \langle x, x_i' \rangle y_i$, where $\{x_i'\}_i$ and $\{y_i\}_i$ are sequences in V' and W, respectively, such that $\sum_{i=1}^{\infty} ||x_i'||_{V'}||y_i||_W < \infty$. This representation is called nuclear, and can be regarded as an expansion of T as a sum of operators of rank 1 (i.e., with one-dimensional range). The nuclear norm of T is defined as

$$||T||_{nuc} = \inf \sum_{i=1}^{\infty} ||x_{i}^{'}||_{V^{'}} ||y_{i}||_{W},$$

where the infimum is taken over all possible nuclear representations of T.

The **nuclear norm metric** is the **norm metric** $||T - P||_{nuc}$ on the set N(V, W) of all nuclear operators mapping V into W. The space $(N(V, W), ||...|_{nuc})$, called the *space of nuclear operators*, is a Banach space.

A nuclear space is defined as a **locally convex** space for which all continuous linear functions into an arbitrary Banach space are nuclear operators. A nuclear space is constructed as a projective limit of Hilbert spaces H_{α} with the property that, for each $\alpha \in I$, one can find $\beta \in I$ such that $H_{\beta} \subset H_{\alpha}$, and the embedding operator $H_{\beta} \ni x \to x \in H_{\alpha}$ is a Hilbert-Schmidt operator. A normed space is nuclear if and only if it is finite-dimensional.

• Finite nuclear norm metric

Let F(V,W) be the space of all linear operators of finite rank (i.e., with finite-dimensional range) mapping a Banach space $(V,||.||_V)$ into another Banach space $(W,||.||_W)$. A linear operator $T \in F(V,W)$ can be represented in the form $x \mapsto T(x) = \sum_{i=1}^n \langle x, x_i' \rangle y_i$, where $\{x_i'\}_i$ and $\{y_i\}_i$ are sequences in V' (Banach dual of V) and W, respectively, and $\langle x, x' \rangle$ is the value of a functional $x' \in V'$ at a vector $x \in V$. The finite nuclear norm of T is defined as

$$||T||_{fnuc} = \inf \sum_{i=1}^{n} ||x_{i}^{'}||_{V^{'}}||y_{i}||_{W},$$

where the infimum is taken over all possible finite representations of T.

The finite nuclear norm metric is the norm metric $||T-P||_{fnuc}$ on F(V,W). The space $(F(V,W),||.||_{fnuc})$ is called the space of operators of finite rank. It is a dense linear subspace of the space of nuclear operators N(V,W).

• Hilbert-Schmidt norm metric

Consider the set of all linear operators from a Hilbert space $(H_1, ||.||_{H_1})$ into a Hilbert space $(H_2, ||.||_{H_2})$. The Hilbert-Schmidt norm $||T||_{HS}$ of a linear operator $T: H_1 \to H_2$ is defined by

$$||T||_{HS} = (\sum_{\alpha \in I} ||T(e_{\alpha})||_{H_2}^2)^{1/2},$$

where $(e_{\alpha})_{\alpha \in I}$ is an orthonormal basis in H_1 . A linear operator $T: H_1 \to H_2$ is called a *Hilbert–Schmidt operator* if $||T||^2_{HS} < \infty$.

The Hilbert–Schmidt norm metric is the norm metric $||T-P||_{HS}$ on the set $S(H_1, H_2)$ of all Hilbert–Schmidt operators from H_1 into H_2 .

For $H_1=H_2=H$, the algebra S(H,H)=S(H) with the Hilbert–Schmidt norm is a Banach algebra. It contains operators of finite rank as a dense subset, and is contained in the space K(H) of compact operators. An inner product \langle,\rangle_{HS} on S(H) is defined by $\langle T,P\rangle_{HS}=\sum_{\alpha\in I}\langle T(e_\alpha),P(e_\alpha)\rangle$, and $||T||_{HS}=\langle T,T\rangle_{HS}^{1/2}$. Therefore, S(H) is a Hilbert space (independent of the chosen basis $(e_\alpha)_{\alpha\in I}$).

• Trace-class norm metric

Given a Hilbert space H, the trace-class norm of a linear operator $T: H \to H$ is defined by

$$||T||_{tc} = \sum_{\alpha \in I} \langle |T|(e_{\alpha}), e_{\alpha} \rangle,$$

where |T| is the absolute value of T in the Banach algebra B(H) of all bounded operators from H into itself, and $(e_{\alpha})_{\alpha \in I}$ is an orthonormal basis of H. An operator $T: H \to H$ is called a trace-class operator if $||T||_{tc} < \infty$. Any such operator is the product of two Hilbert-Schmidt operators.

The trace-class norm metric is the norm metric $||T - P||_{tc}$ on the set L(H) of all trace-class operators from H into itself. The set L(H) with the norm $||.||_{tc}$ forms a Banach algebra which is contained in the algebra K(H) (of all compact operators from H into itself), and contains the algebra S(H) (of all Hilbert-Schmidt operators from H into itself).

• Schatten p-class norm metric

Let $1 \le p < \infty$. Given a separable Hilbert space H, the Schatten p-class norm of a compact linear operator $T: H \to H$ is defined by

$$||T||_{Sch}^p = \left(\sum_n |s_n|^p\right)^{\frac{1}{p}},$$

where $\{s_n\}_n$ is the sequence of singular values of T. A compact operator $T: H \to H$ is called a Schatten p-class operator if $||T||_{Sch}^p < \infty$.

The Schatten p-class norm metric is the norm metric $||T - P||_{Sch}^p$ on the set $S_p(H)$ of all Schatten p-class operators from H onto itself. The set $S_p(H)$ with the norm $||.||_{Sch}^p$ forms a Banach space. $S_1(H)$ is the trace-class of H, and $S_2(H)$ is the Hilbert-Schmidt class of H (cf. also Schatten norm metric in Chap. 12).

• Continuous dual space

For any vector space V over some field, its *algebraic dual space* is the set of all linear functionals on V.

Let (V, ||.||) be a normed vector space. Let V' be the set of all continuous linear functionals T from V into the base field $(\mathbb{R} \text{ or } \mathbb{C})$. Let ||.||' be the operator norm on V', defined by

$$||T||' = \sup_{||x|| \le 1} |T(x)|.$$

The space (V', ||.||') is a Banach space which is called the **continuous** dual (or *Banach dual*) of (V, ||.||).

In fact, the continuous dual of the metric space l_p^n (l_p^∞) is l_q^n (l_q^∞) , respectively). The continuous dual of l_1^n (l_1^∞) is l_∞^n (l_∞^∞) , respectively). The continuous duals of the Banach spaces C (consisting of all convergent sequences, with the l_∞ -metric) and C_0 (consisting of all sequences converging to zero, with the l_∞ -metric) are both naturally identified with l_1^∞ .

• Distance constant of operator algebra

Let \mathcal{A} be an operator algebra contained in B(H), the set of all bounded operators on a Hilbert space H. For any operator $T \in B(H)$, let $\beta(T, \mathcal{A}) = \sup\{||P^{\perp}TP|| : P \text{ is a projection, and } P^{\perp}\mathcal{A}P = (0)\}$. Let $dist(T, \mathcal{A})$ be the distance of T from the algebra \mathcal{A} , i.e., the smallest norm of an operator T - A, where A runs over \mathcal{A} . The smallest positive constant C (if it exists) such that, for any operator $T \in B(H)$,

$$dist(T, \mathcal{A}) \leq C\beta(T, \mathcal{A})$$

is called the **distance constant** for the algebra A.

Chapter 14 Distances in Probability Theory

A probability space is a measurable space (Ω, \mathcal{A}, P) , where \mathcal{A} is the set of all measurable subsets of Ω , and P is a measure on \mathcal{A} with $P(\Omega) = 1$. The set Ω is called a sample space. An element $a \in \mathcal{A}$ is called an event. In particular, an elementary event is a subset of Ω that contains only one element. P(a) is called the probability of the event a. The measure P on \mathcal{A} is called a probability measure, or (probability) distribution law, or simply (probability) distribution.

A random variable X is a measurable function from a probability space (Ω, \mathcal{A}, P) into a measurable space, called a *state space* of possible values of the variable; it is usually taken to be the real numbers with the *Borel \sigma-algebra*, so $X: \Omega \to \mathbb{R}$. The range \mathcal{X} of the random variable X is called the *support* of the distribution P; an element $x \in \mathcal{X}$ is called a *state*.

A distribution law can be uniquely described via a cumulative distribution function (CDF, distribution function, cumulative density function) F(x)which describes the probability that a random value X takes on a value at most x: $F(x) = P(X \le x) = P(\omega \in \Omega : X(\omega) \le x)$.

So, any random variable X gives rise to a probability distribution which assigns to the interval [a,b] the probability $P(a \le X \le b) = P(\omega \in \Omega : a \le X(\omega) \le b)$, i.e., the probability that the variable X will take a value in the interval [a,b].

A distribution is called discrete if F(x) consists of a sequence of finite jumps at x_i ; a distribution is called continuous if F(x) is continuous. We consider (as in the majority of applications) only discrete or absolutely continuous distributions, i.e., the CDF function $F: \mathbb{R} \to \mathbb{R}$ is absolutely continuous. It means that, for every number $\epsilon > 0$, there is a number $\delta > 0$ such that, for any sequence of pairwise disjoint intervals $[x_k, y_k]$, $1 \le k \le n$, the inequality $\sum_{1 \le k \le n} (y_k - x_k) < \delta$ implies the inequality $\sum_{1 \le k \le n} |F(y_k) - F(x_k)| < \epsilon$.

A distribution law also can be uniquely defined via a probability density function (PDF, density function, probability function) p(x) of the underlying random variable. For an absolutely continuous distribution, the CDF is almost everywhere differentiable, and the PDF is defined as the derivative p(x) = F'(x) of the CDF; so, $F(x) = P(X \le x) = \int_{-\infty}^{x} p(t)dt$, and

 $\int_a^b p(t)dt = P(a \leq X \leq b)$. In the discrete case, the PDF (the density of the random variable X) is defined by its values $p(x_i) = P(X = x)$; so $F(x) = \sum_{x_i \leq x} p(x_i)$. In contrast, each elementary event has probability zero in any continuous case.

The random variable X is used to "push-forward" the measure P on Ω to a measure dF on \mathbb{R} . The underlying probability space is a technical device used to guarantee the existence of random variables and sometimes to construct them.

For simplicity, we usually present the discrete version of probability metrics, but many of them are defined on any measurable space; see [Bass89], [Cha08]. For a probability distance d on random quantities, the conditions P(X=Y)=1 or equality of distributions imply (and characterize) d(X,Y)=0; such distances are called [Rach91] compound or simple distances, respectively. In many cases, some ground distance d is given on the state space $\mathcal X$ and the presented distance is a lifting of it to a distance on distributions.

In Statistics, many of the distances below, between distributions P_1 and P_2 , are used as measures of goodness of fit between estimated, P_2 , and theoretical, P_1 , distributions. Also, in Statistics, a distance that not satisfy the triangle inequality, is often called a **distance statistic**; a statistic is a function of a sample which is independent of its distribution.

Below we use the notation $\mathbb{E}[X]$ for the expected value (or mean) of the random variable X: in the discrete case $\mathbb{E}[X] = \sum_x xp(x)$, in the continuous case $\mathbb{E}[X] = \int xp(x)dx$. The variance of X is $\mathbb{E}[(X - \mathbb{E}[X])^2]$. Also we denote $p_X = p(x) = P(X = x)$, $F_X = F(x) = P(X \le x)$, p(x,y) = P(X = x, Y = y).

14.1 Distances on random variables

All distances in this section are defined on the set **Z** of all random variables with the same support \mathcal{X} ; here $X, Y \in \mathbf{Z}$.

• p-average compound metric

Given $p \geq 1$, the *p*-average compound metric (or L_p -metric between variables) is a metric on **Z** with $\mathcal{X} \subset \mathbb{R}$ and $\mathbb{E}[|Z|^p] < \infty$ for all $Z \in \mathbf{Z}$, defined by

$$(\mathbb{E}[|X - Y|^p])^{1/p} = (\sum_{(x,y) \in \mathcal{X} \times \mathcal{X}} |x - y|^p p(x,y))^{1/p}.$$

For p = 2 and ∞ , it is called, respectively, the mean-square distance and essential supremum distance between variables.

• Absolute moment metric

Given $p \geq 1$, the **absolute moment metric** is a metric on **Z** with $\mathcal{X} \subset \mathbb{R}$ and $\mathbb{E}[|Z|^p] < \infty$ for all $Z \in \mathbf{Z}$, defined by

$$(|(\mathbb{E}[|X|^p])^{1/p} - (\mathbb{E}[|Y|^p])^{1/p}|.$$

For p = 1 it is called the *engineer metric*.

• Indicator metric

The **indicator metric** is a metric on **Z**, defined by

$$\mathbb{E}[1_{X \neq Y}] = \sum_{(x,y) \in \mathcal{X} \times \mathcal{X}} 1_{x \neq y} p(x,y) = \sum_{(x,y) \in \mathcal{X} \times \mathcal{X}, x \neq y} p(x,y).$$

(Cf. **Hamming metric** in Chap. 1.)

• Ky Fan metric K

The **Ky Fan metric** K is a metric K on **Z**, defined by

$$\inf\{\epsilon > 0 : P(|X - Y| > \epsilon) < \epsilon\}.$$

It is the case d(x,y) = |X - Y| of the **probability distance**.

• Ky Fan metric K^*

The **Ky Fan metric** K^* is a metric K^* on **Z**, defined by

$$\mathbb{E}\left[\frac{|X-Y|}{1+|X-Y|}\right] = \sum_{(x,y)\in\mathcal{X}\times\mathcal{X}} \frac{|x-y|}{1+|x-y|} p(x,y).$$

• Probability distance

Given a metric space (\mathcal{X}, d) , the **probability distance** on **Z** is defined by

$$\inf\{\epsilon > 0 : P(d(X,Y) > \epsilon) < \epsilon\}.$$

14.2 Distances on distribution laws

All distances in this section are defined on the set \mathcal{P} of all distribution laws such that corresponding random variables have the same range \mathcal{X} ; here $P_1, P_2 \in \mathcal{P}$.

• L_p -metric between densities

The L_p -metric between densities is a metric on \mathcal{P} (for a countable \mathcal{X}), defined, for any $p \geq 1$, by

$$(\sum_{x} |p_1(x) - p_2(x)|^p)^{\frac{1}{p}}.$$

For p = 1, one half of it is called the **total variation metric** (or *variational distance*, *trace-distance*). For p = 2, it is the **Patrick-Fisher distance**. The *point metric* $\sup_{x} |p_1(x) - p_2(x)|$ corresponds to $p = \infty$.

The **Lissak-Fu distance** with parameter $\alpha > 0$ is defined as $\sum_{x} |p_1(x) - p_2(x)|^{\alpha}$.

• Bayesian distance

The error probability in classification is the following error probability of the optimal Bayes rule for the classification into 2 classes with a priori probabilities ϕ , $1-\phi$ and corresponding densities p_1 , p_2 of the observations:

$$P_e = \sum_{x} \min(\phi p_1(x), (1 - \phi)p_2(x)).$$

The **Bayesian distance** on \mathcal{P} is defined by $1 - P_e$.

For the classification into m classes with a priori probabilities ϕ_i , $1 \le i \le m$, and corresponding densities p_i of the observations, the error probability becomes

$$P_e = 1 - \sum_{x} p(x) \max_{i} P(C_i|x),$$

where $P(C_i|x)$ is the a posteriori probability of the class C_i given the observation x and $p(x) = \sum_{i=1}^{m} \phi_i P(x|C_i)$. The general mean distance between m classes C_i (cf. m-hemi-metric in Chap. 3) is defined (Van der Lubbe 1979), for $\alpha > 0$ and $\beta > 1$, by

$$\sum_{x} p(x) (\sum_{i} P(C_{i}|x)^{\beta})^{\alpha}.$$

The case $\alpha=1,\beta=2$ corresponds to the *Bayesian distance* in Devijver (1974); the case $\beta=\frac{1}{\alpha}$ was considered in Trouborst, Baker, Boekee and Boxma (1974).

• Mahalanobis semi-metric

The **Mahalanobis semi-metric** (or *quadratic distance*) is a semi-metric on \mathcal{P} (for $\mathcal{X} \subset \mathbb{R}^n$), defined by

$$\sqrt{(\mathbb{E}_{P_1}[X] - \mathbb{E}_{P_2}[X])^T A^{-1}(\mathbb{E}_{P_1}[X] - \mathbb{E}_{P_2}[X])}$$

for a given positive-definite matrix A.

• Engineer semi-metric

The **engineer semi-metric** is a semi-metric on \mathcal{P} (for $\mathcal{X} \subset \mathbb{R}$), defined by

$$|\mathbb{E}_{P_1}[X] - \mathbb{E}_{P_2}[X]| = |\sum_x x(p_1(x) - p_2(x))|.$$

• Stop-loss metric of order m

The stop-loss metric of order m is a metric on \mathcal{P} (for $\mathcal{X} \subset \mathbb{R}$), defined by

$$\sup_{t \in \mathbb{R}} \sum_{x > t} \frac{(x - t)^m}{m!} (p_1(x) - p_2(x)).$$

• Kolmogorov-Smirnov metric

The Kolmogorov–Smirnov metric (or Kolmogorov metric, uniform metric) is a metric on \mathcal{P} (for $\mathcal{X} \subset \mathbb{R}$), defined by

$$\sup_{x \in \mathbb{R}} |P_1(X \le x) - P_2(X \le x)|.$$

The **Kuiper distance** on \mathcal{P} is defined by

$$\sup_{x \in \mathbb{R}} (P_1(X \le x) - P_2(X \le x)) + \sup_{x \in \mathbb{R}} (P_2(X \le x) - P_1(X \le x)).$$

(Cf. Pompeiu-Eggleston metric in Chap. 9.)

The Anderson–Darling distance on \mathcal{P} is defined by

$$\sup_{x \in \mathbb{R}} \frac{|(P_1(X \le x) - P_2(X \le x))|}{\ln \sqrt{(P_1(X \le x)(1 - P_1(X \le x))}}.$$

The Crnkovic-Drachma distance is defined by

$$\sup_{x \in \mathbb{R}} (P_1(X \le x) - P_2(X \le x)) \ln \frac{1}{\sqrt{(P_1(X \le x)(1 - P_1(X \le x))}} +$$

$$+ \sup_{x \in \mathbb{R}} (P_2(X \le x) - P_1(X \le x)) \ln \frac{1}{\sqrt{(P_1(X \le x)(1 - P_1(X \le x))}}.$$

The above three distances are used in Statistics as measures of *goodness* of fit, especially, for VaR (Value at Risk) measurements in Finance.

• Cramer-von Mises distance

The Cramer-von Mises distance is a distance on \mathcal{P} (for $\mathcal{X} \subset \mathbb{R}$), defined by

$$\int_{-\infty}^{+\infty} (P_1(X \le x) - P_2(X \le x))^2 dx.$$

This is the squared L_2 -metric between cumulative density functions.

• Levy–Sibley metric

The **Levy metric** is a metric on \mathcal{P} (for $\mathcal{X} \subset \mathbb{R}$ only), defined by

$$\inf\{\epsilon > 0: P_1(X \le x - \epsilon) - \epsilon \le P_2(X \le x) \le P_1(X \le x + \epsilon) + \epsilon \text{ for any } x \in \mathbb{R}\}.$$

It is a special case of the **Prokhorov metric** for $(\mathcal{X}, d) = (\mathbb{R}, |x - y|)$.

• Prokhorov metric

Given a metric space (\mathcal{X}, d) , the **Prokhorov metric** on \mathcal{P} is defined by

$$\inf\{\epsilon > 0: P_1(X \in B) \le P_2(X \in B^{\epsilon}) + \epsilon \text{ and } P_2(X \in B) \le P_1(X \in B^{\epsilon}) + \epsilon\},$$

where B is any Borel subset of \mathcal{X} , and $B^{\epsilon} = \{x : d(x,y) < \epsilon, y \in B\}.$

It is the smallest (over all joint distributions of pairs (X, Y) of random variables X, Y such that the marginal distributions of X and Y are P_1 and P_2 , respectively) **probability distance** between random variables X and Y.

• Dudley metric

Given a metric space (\mathcal{X}, d) , the **Dudley metric** on \mathcal{P} is defined by

$$\sup_{f \in F} |\mathbb{E}_{P_1}[f(X)] - \mathbb{E}_{P_2}[f(X)]| = \sup_{f \in F} |\sum_{x \in \mathcal{X}} f(x)(p_1(x) - p_2(x))|,$$

where $F = \{f : \mathcal{X} \to \mathbb{R}, ||f||_{\infty} + Lip_d(f) \leq 1\}$, and $Lip_d(f) = \sup_{x,y \in \mathcal{X}, x \neq y} \frac{|f(x) - f(y)|}{d(x,y)}$.

• Szulga metric

Given a metric space (\mathcal{X}, d) , the **Szulga metric** on \mathcal{P} is defined by

$$\sup_{f \in F} |(\sum_{x \in \mathcal{X}} |f(x)|^p p_1(x))^{1/p} - (\sum_{x \in \mathcal{X}} |f(x)|^p p_2(x))^{1/p}|,$$

where $F = \{f : X \to \mathbb{R}, Lip_d(f) \leq 1\}$, and $Lip_d(f) = \sup_{x,y \in \mathcal{X}, x \neq y} \frac{|f(x) - f(y)|}{d(x,y)}$.

• Zolotarev semi-metric

The **Zolotarev semi-metric** is a semi-metric on \mathcal{P} , defined by

$$\sup_{f \in F} |\sum_{x \in \mathcal{X}} f(x)(p_1(x) - p_2(x))|,$$

where F is any set of functions $f: \mathcal{X} \to \mathbb{R}$ (in the continuous case, F is any set of such bounded continuous functions); cf. **Szulga metric**, **Dudley metric**.

• Convolution metric

Let G be a separable locally compact abelian group, and let C(G) be the set of all real bounded continuous functions on G vanishing at infinity. Fix a function $g \in C(G)$ such that |g| is integrable with respect to the Haar measure on G, and $\{\beta \in G^* : \widehat{g}(\beta) = 0\}$ has empty interior; here G^* is the dual group of G, and \widehat{g} is the Fourier transform of g.

The **convolution metric** (or *smoothing metric*) is defined (Yukich 1985), for any two finite signed Baire measures P_1 and P_2 on G, by

$$\sup_{x \in G} | \int_{y \in G} g(xy^{-1})(dP_1 - dP_2)(y) |.$$

This metric can also be seen as the difference $T_{P_1}(g) - T_{P_2}(g)$ of convolution operators on C(G) where, for any $f \in C(G)$, the operator $T_P f(x)$ is $\int_{y \in G} \widehat{f}(xy^{-1}) dP(y).$ • Discrepancy metric

Given a metric space (\mathcal{X}, d) , the **discrepancy metric** on \mathcal{P} is defined by

$$\sup\{|P_1(X \in B) - P_2(X \in B)| : B \text{ is any closed ball}\}.$$

• Bi-discrepancy semi-metric

The **bi-discrepancy semi-metric** is a semi-metric evaluating the proximity of distributions P_1 , P_2 (over different collections A_1 , A_2 of measurable sets), defined in the following way:

$$D(P_1, P_2) + D(P_2, P_1),$$

where $D(P_1, P_2) = \sup \{\inf \{P_2(C) : B \subset C \in A_2\} - P_1(B) : B \in A_1 \}$ (discrepancy).

• Le Cam distance

The Le Cam distance is a semi-metric, evaluating the proximity of probability distributions P_1, P_2 (on different spaces $\mathcal{X}_1, \mathcal{X}_2$), defined in the following way:

$$\max\{\delta(P_1, P_2), \delta(P_2, P_1)\},\$$

where $\delta(P_1, P_2) = \inf_B \sum_{x_2 \in \mathcal{X}_2} |BP_1(X_2 = x_2) - BP_2(X_2 = x_2)|$ is the *Le Cam deficiency*. Here $BP_1(X_2 = x_2) = \sum_{x_1 \in \mathcal{X}_1} p_1(x_1)b(x_2|x_1)$, where Bis a probability distribution over $\mathcal{X}_1 \times \mathcal{X}_2$, and

$$b(x_2|x_1) = \frac{B(X_1 = x_1, X_2 = x_2)}{B(X_1 = x_1)} = \frac{B(X_1 = x_1, X_2 = x_2)}{\sum_{x \in \mathcal{X}_2} B(X_1 = x_1, X_2 = x)}.$$

So, $BP_2(X_2 = x_2)$ is a probability distribution over \mathcal{X}_2 , since $\sum_{x_2 \in \mathcal{X}_2}$ $b(x_2|x_1) = 1.$

Le Cam distance is not a probabilistic distance, since P_1 and P_2 are defined over different spaces; it is a distance between statistical experiments (models).

• Skorokhod–Billingsley metric

The Skorokhod-Billingsley metric is a metric on \mathcal{P} , defined by

$$\inf_{f} \max \left\{ \sup_{x} |P_1(X \leq x) - P_2(X \leq f(x))|, \sup_{x} |f(x) - x|, \sup_{x \neq y} \left| \ln \frac{f(y) - f(x)}{y - x} \right| \right\},$$

where $f: \mathbb{R} \to \mathbb{R}$ is any strictly increasing continuous function.

• Skorokhod metric

The **Skorokhod metric** is a metric on \mathcal{P} , defined by

$$\inf\{\epsilon > 0 : \max\{\sup_{x} |P_1(X < x) - P_2(X \le f(x))|, \sup_{x} |f(x) - x|\} < \epsilon\},\$$

where $f: \mathbb{R} \to \mathbb{R}$ is a strictly increasing continuous function.

• Birnbaum-Orlicz distance

The **Birnbaum–Orlicz distance** is a distance on \mathcal{P} , defined by

$$\sup_{x \in \mathbb{R}} f(|P_1(X \le x) - P_2(X \le x)|),$$

where $f: \mathbb{R}_{\geq 0} \to \mathbb{R}_{\geq 0}$ is any non-decreasing continuous function with f(0) = 0, and $f(2t) \leq Cf(t)$ for any t > 0 and some fixed $C \geq 1$. It is a **near-metric**, since the C-triangle inequality $d(P_1, P_2) \leq C(d(P_1, P_3) + d(P_3, P_2))$ holds.

Birnbaum–Orlicz distance is also used, in Functional Analysis, on the set of all integrable functions on the segment [0,1], where it is defined by $\int_0^1 H(|f(x)-g(x)|)dx$, where H is a non-decreasing continuous function from $[0,\infty)$ onto $[0,\infty)$ which vanishes at the origin and satisfies the *Orlicz condition*: $\sup_{t>0} \frac{H(2t)}{H(t)} < \infty$.

• Kruglov distance

The **Kruglov distance** is a distance on \mathcal{P} , defined by

$$\int f(P_1(X \le x) - P_2(X \le x)) dx,$$

where $f: \mathbb{R}_{\geq 0} \to \mathbb{R}_{\geq 0}$ is any even strictly increasing function with f(0) = 0, and $f(s+t) \leq C(f(s)+f(t))$ for any $s,t \geq 0$ and some fixed $C \geq 1$. It is a **near-metric**, since the C-triangle inequality $d(P_1, P_2) \leq C(d(P_1, P_3) + d(P_3, P_2))$ holds.

• Burbea-Rao distance

Consider a continuous convex function $\phi(t):(0,\infty)\to\mathbb{R}$ and put $\phi(0)=\lim_{t\to 0}\phi(t)\in(-\infty,\infty]$. The convexity of ϕ implies non-negativity of the function $\delta_{\phi}:[0,1]^2\to(-\infty,\infty]$, defined by $\delta_{\phi}(x,y)=\frac{\phi(x)+\phi(y)}{2}-\phi(\frac{x+y}{2})$ if $(x,y)\neq(0,0)$, and $\delta_{\phi}(0,0)=0$.

The corresponding Burbea-Rao distance on \mathcal{P} is defined by

$$\sum_{x} \delta_{\phi}(p_1(x), p_2(x)).$$

• Bregman distance

Consider a differentiable convex function $\phi(t):(0,\infty)\to\mathbb{R}$, and put $\phi(0)=\lim_{t\to 0}\phi(t)\in(-\infty,\infty]$. The convexity of ϕ implies that the

function $\delta_{\phi}:[0,1]^2 \to (-\infty,\infty]$ defined by continuous extension of $\delta_{\phi}(u,v)=\phi(u)-\phi(v)-\phi'(v)(u-v),\ 0< u,v\leq 1,\ \text{on}\ [0,1]^2$ is non-negative.

The corresponding **Bregman distance** on \mathcal{P} is defined by

$$\sum_{1}^{m} \delta_{\phi}(p_i, q_i).$$

(Cf. Bregman quasi-distance.)

• f-divergence of Csizar

The f-divergence of Csizar is a function on $\mathcal{P} \times \mathcal{P}$, defined by

$$\sum_{x} p_2(x) f\left(\frac{p_1(x)}{p_2(x)}\right),\,$$

where f is a continuous convex function $f: \mathbb{R}_{>0} \to \mathbb{R}$.

The cases $f(t) = t \ln t$ and $f(t) = (t-1)^2/2$ correspond to the **Kullback–Leibler distance** and to the χ^2 -distance below, respectively. The case f(t) = |t-1| corresponds to the L_1 -metric between densities, and the case $f(t) = 4(1 - \sqrt{t})$ (as well as $f(t) = 2(t+1) - 4\sqrt{t}$) corresponds to the squared **Hellinger metric**.

Semi-metrics can also be obtained, as the square root of the f-divergence of Csizar, in the cases $f(t)=(t-1)^2/(t+1)$ (the Vajda–Kus semi-metric), $f(t)=|t^a-1|^{1/a}$ with $0< a \le 1$ (the generalized Matusita distance), and $f(t)=\frac{(t^a+1)^{1/a}-2^{(1-a)/a}(t+1)}{1-1/\alpha}$ (the Osterreicher semi-metric).

• Fidelity similarity

The fidelity similarity (or Bhattacharya coefficient, Hellinger affinity) on \mathcal{P} is

$$\rho(P_1, P_2) = \sum_{x} \sqrt{p_1(x)p_2(x)}.$$

• Hellinger metric

In terms of the fidelity similarity ρ , the Hellinger metric (or Hellinger-Kakutani metric) on \mathcal{P} is defined by

$$(2\sum_{x}(\sqrt{p_1(x)}-\sqrt{p_2(x)})^2)^{\frac{1}{2}}=2(1-\rho(P_1,P_2))^{\frac{1}{2}}.$$

Sometimes, $(\sum_x (\sqrt{p_1(x)} - \sqrt{p_2(x)})^2)^{\frac{1}{2}}$ is called the **Matusita distance**, while $(\sum_x (\sqrt{p_1(x)} - \sqrt{p_2(x)})^2)$ is called the *squared-chord distance*.

• Harmonic mean similarity

The harmonic mean similarity is a similarity on \mathcal{P} , defined by

$$2\sum_{x} \frac{p_1(x)p_2(x)}{p_1(x) + p_2(x)}.$$

• Bhattacharya distance 1

In terms of the fidelity similarity ρ , the Bhattacharya distance 1 on \mathcal{P} is

$$(\arccos \rho(P_1, P_2))^2$$
.

Twice this distance is used also in Statistics and Machine Learning, where it is called the **Fisher distance**.

• Bhattacharya distance 2

In terms of the fidelity similarity ρ , the Bhattacharya distance 2 on \mathcal{P} is

$$-\ln \rho(P_1, P_2).$$

• χ^2 -distance

The χ^2 -distance (or Pearson χ^2 -distance) is a quasi-distance on \mathcal{P} , defined by

$$\sum_{x} \frac{(p_1(x) - p_2(x))^2}{p_2(x)}.$$

The **Neyman** χ^2 -distance is a quasi-distance on \mathcal{P} , defined by

$$\sum_{x} \frac{(p_1(x) - p_2(x))^2}{p_1(x)}.$$

The probabilistic **symmetric** χ^2 -measure is a distance on \mathcal{P} , defined by

$$2\sum_{x} \frac{(p_1(x) - p_2(x))^2}{p_1(x) + p_2(x)}.$$

The half of the probabilistic symmetric χ^2 -measure is called *squared* χ^2 .

• Separation quasi-distance

The **separation distance** is a quasi-distance on \mathcal{P} (for a countable \mathcal{X}) defined by

$$\max_{x} \left(1 - \frac{p_1(x)}{p_2(x)} \right).$$

(Not to be confused with separation distance in Chap. 9.)

• Kullback-Leibler distance

The Kullback-Leibler distance (or relative entropy, information deviation, information gain, KL-distance) is a quasi-distance on \mathcal{P} , defined by

$$KL(P_1, P_2) = \mathbb{E}_{P_1}[\ln L] = \sum_x p_1(x) \ln \frac{p_1(x)}{p_2(x)},$$

where $L = \frac{p_1(x)}{p_2(x)}$ is the *likelihood ratio*. Therefore,

$$KL(P_1, P_2) = -\sum_{x} (p_1(x) \ln p_2(x)) + \sum_{x} (p_1(x) \ln p_1(x)) = H(P_1, P_2) - H(P_1),$$

where $H(P_1)$ is the *entropy* of P_1 , and $H(P_1, P_2)$ is the *cross-entropy* of P_1 and P_2 .

If P_2 is the product of marginals of P_1 (say, $p_2(x,y) = p_1(x)p_1(y)$), the KL-distance $KL(P_1, P_2)$ is called the *Shannon information quantity* and (cf. **Shannon distance**) is equal to $\sum_{(x,y)\in\mathcal{X}\times\mathcal{X}} p_1(x,y) \ln \frac{p_1(x,y)}{p_1(x)p_1(y)}$.

• Skew divergence

The **skew divergence** is a quasi-distance on \mathcal{P} , defined by

$$KL(P_1, aP_2 + (1-a)P_1),$$

where $a \in [0,1]$ is a constant, and KL is the **Kullback–Leibler distance**.

The cases a = 1 and $a = \frac{1}{2}$ correspond to $KL(P_1, P_2)$ and K-divergence.

• Jeffrey divergence

The **Jeffrey divergence** (or J-divergence, divergence distance, KL2-distance) is a symmetric version of the **Kullback–Leibler distance**, defined by

$$KL(P_1, P_2) + KL(P_2, P_1) = \sum_{x} (p_1(x) - p_2(x)) \ln \frac{p_1(x)}{p_2(x)}.$$

For $P_1 \to P_2$, the Jeffrey divergence behaves like the χ^2 -distance.

• Jensen-Shannon divergence

The Jensen–Shannon divergence is defined by

$$aKL(P_1, P_3) + (1 - a)KL(P_2, P_3),$$

where $P_3 = aP_1 + (1-a)P_2$, and $a \in [0,1]$ is a constant (cf. clarity similarity).

In terms of entropy $H(P) = -\sum_{x} p(x) \ln p(x)$, the Jensen–Shannon divergence is equal to $H(aP_1 + (1-a)P_2) - aH(P_1) - (1-a)H(P_2)$.

• Topsøe distance

Let P_3 denote $\frac{1}{2}(P_1+P_2)$. The **Topsøe distance** (or *information statistics*) is a symmetric version of the **Kullback–Leibler distance** (or rather of the K-divergence $KL(P_1, P_3)$):

$$KL(P_1,P_3) + KL(P_2,P_3) = \sum_{x} \left(p_1(x) \ln \frac{p_1(x)}{p_3(x)} + p_2(x) \ln \frac{p_2(x)}{p_3(x)} \right).$$

The Topsøe distance is twice the **Jensen–Shannon divergence** with $a = \frac{1}{2}$. Some authors use the term *Jensen–Shannon divergence* only for this value of a. It is not a metric, but its square root is a metric.

The **Taneja distance** is defined by

$$\sum_{x} p_3(x) \ln \frac{p_3(x)}{\sqrt{p_1(x)p_2(x)}}.$$

• Resistor-average distance

The Johnson–Simanović's **resistor-average distance** is a symmetric version of the **Kullback–Leibler distance** on \mathcal{P} which is defined by the harmonic sum

$$\left(\frac{1}{KL(P_1, P_2)} + \frac{1}{KL(P_2, P_1)}\right)^{-1}$$
.

Cf. resistance metric for graphs in Chap. 15.

• Ali-Silvey distance

The Ali–Silvey distance is a quasi-distance on \mathcal{P} , defined by the functional

$$f(\mathbb{E}_{P_1}[g(L)]),$$

where $L = \frac{p_1(x)}{p_2(x)}$ is the *likelihood ratio*, f is a non-decreasing function on \mathbb{R} , and g is a continuous convex function on $\mathbb{R}_{\geq 0}$ (cf. f-divergence of Csizar).

The case f(x) = x, $g(x) = x \ln x$ corresponds to the **Kullback–Leibler** distance; the case $f(x) = -\ln x$, $g(x) = x^t$ corresponds to the **Chernoff** distance.

• Chernoff distance

The Chernoff distance (or $R\acute{e}nyi\ cross-entropy$) is a distance on \mathcal{P} , defined by

$$\max_{t \in [0,1]} D_t(P_1, P_2),$$

where $0 \le t \le 1$ and $D_t(P_1, P_2) = -\ln \sum_x (p_1(x))^t (p_2(x))^{1-t}$ (called the *Chernoff coefficient* or *Hellinger path*), which is proportional to the **Rényi** distance.

The case $t = \frac{1}{2}$ corresponds to the **Bhattacharya distance** 2.

Rényi distance

The **Rényi distance** (or order t Rényi entropy) is a quasi-distance on \mathcal{P} , defined, for any constant $0 \le t < 1$, by

$$\frac{1}{1-t}\ln\sum_{x}p_{2}(x)\left(\frac{p_{1}(x)}{p_{2}(x)}\right)^{t}.$$

The limit of the Rényi distance, for $t \to 1$, is the Kullback–Leibler distance. For $t = \frac{1}{2}$, one half of the Rényi distance is the Bhattacharya distance 2 (cf. f-divergence of Csizar and Chernoff distance).

• Clarity similarity

The clarity similarity is a similarity on \mathcal{P} , defined by

$$(KL(P_1, P_3) + KL(P_2, P_3)) - (KL(P_1, P_2) + KL(P_2, P_1)) =$$

$$= \sum_{x} \left(p_1(x) \ln \frac{p_2(x)}{p_3(x)} + p_2(x) \ln \frac{p_1(x)}{p_3(x)} \right),$$

where KL is the **Kullback–Leibler distance**, and P_3 is a fixed probability law. It was introduced in [CCL01] with P_3 being the probability distribution of English.

• Shannon distance

Given a measure space (Ω, \mathcal{A}, P) , where the set Ω is finite and P is a probability measure, the entropy (or Shannon information entropy) of a function $f: \Omega \to X$, where X is a finite set, is defined by

$$H(f) = -\sum_{x \in X} P(f = x) \log_a(P(f = x));$$

here a=2, e, or 10 and the unit of entropy is called a bit, nat, or dit (digit), respectively. The function f can be seen as a partition of the measure space. For any two such partitions $f:\Omega\to X$ and $g:\Omega\to Y$, denote by H(f,g) the entropy of the partition $(f,g):\Omega\to X\times Y$ (joint entropy), and by H(f|g) the conditional entropy (or equivocation); then the **Shannon distance** between f and g is a metric defined by

$$H(f|g) + H(g|f) = 2H(f,g) - H(f) - H(g) = H(f,g) - I(f;g),$$

where I(f;g) = H(f) + H(g) - H(f,g) is the Shannon mutual information. If P is the uniform probability law, then Goppa showed that the Shannon distance can be obtained as a limiting case of the **finite subgroup metric**.

In general, the **information metric** (or **entropy metric**) between two random variables (information sources) X and Y is defined by

$$H(X|Y) + H(Y|X) = H(X,Y) - I(X;Y),$$

where the *conditional entropy* H(X|Y) is defined by $\sum_{x \in X} \sum_{y \in Y} p(x,y)$ ln p(x|y), and p(x|y) = P(X = x|Y = y) is the conditional probability.

The **Rajski distance** (or normalized information metric) is defined (Rajski 1961, for discrete probability distributions X, Y) by

$$\frac{H(X|Y) + H(Y|X)}{H(X,Y)} = 1 - \frac{I(X;Y)}{H(X,Y)}.$$

It is equal to 1 if X and Y are independent. (Cf., a different one, **normalized information distance** in Chap. 11).

• Kantorovich–Mallows–Monge–Wasserstein metric Given a metric space (\mathcal{X}, d) , the Kantorovich–Mallows–Monge– Wasserstein metric is defined by

inf
$$\mathbb{E}_S[d(X,Y)]$$
,

where the infimum is taken over all joint distributions S of pairs (X, Y) of random variables X, Y such that marginal distributions of X and Y are P_1 and P_2 .

For any **separable** metric space (\mathcal{X},d) , this is equivalent to the **Lipschitz distance between measures** $\sup_f \int f d(P_1 - P_2)$, where the supremum is taken over all functions f with $|f(x) - f(y)| \leq d(x,y)$ for any $x,y \in \mathcal{X}$.

More generally, the L_p -Wasserstein distance for $\mathcal{X} = \mathbb{R}^n$ is defined by

$$(\inf \mathbb{E}_S[d^p(X,Y)])^{1/p},$$

and, for p = 1, it is also called the $\overline{\rho}$ -distance. For $(\mathcal{X}, d) = (\mathbb{R}, |x - y|)$, it is also called the L_p -metric between distribution functions (CDF), and can be written as

$$(\inf \mathbb{E}[|X - Y|^p])^{1/p} = \left(\int_{\mathbb{R}} |F_1(x) - F_2(x)|^p dx\right)^{1/p}$$
$$= \left(\int_0^1 |F_1^{-1}(x) - F_2^{-1}(x)|^p dx\right)^{1/p}$$

with $F_i^{-1}(x) = \sup_u (P_i(X \le x) < u)$.

The case p = 1 of this metric is called the Monge–Kantorovich metric or Hutchinson metric (in Fractal Theory), Wasserstein metric, Fortet–Mourier metric.

• Ornstein \bar{d} -metric

The **Ornstein** \overline{d} -metric is a metric on \mathcal{P} (for $\mathcal{X} = \mathbb{R}^n$), defined by

$$\frac{1}{n}\inf\int_{x,y}\left(\sum_{i=1}^n 1_{x_i\neq y_i}\right)dS,$$

where the infimum is taken over all joint distributions S of pairs (X,Y) of random variables X,Y such that marginal distributions of X and Y are P_1 and P_2 .

$\begin{array}{c} {\bf Part~IV} \\ {\bf Distances~in~Applied~Mathematics} \end{array}$

Chapter 15 Distances in Graph Theory

A graph is a pair G = (V, E), where V is a set, called the set of vertices of the graph G, and E is a set of unordered pairs of vertices, called the edges of the graph G. A directed graph (or digraph) is a pair D = (V, E), where V is a set, called the set of vertices of the digraph D, and E is a set of ordered pairs of vertices, called arcs of the digraph D.

A graph in which at most one edge may connect any two vertices, is called a *simple graph*. If multiple edges are allowed between vertices, the graph is called a *multi-graph*.

The graph is called *finite* (*infinite*) if the set V of its vertices is finite (infinite, respectively). The *order* of a finite graph is the number of its vertices; the *size* of a finite graph is the number of its edges.

A graph, together with a function which assigns a positive weight to each edge, is called a *weighted graph* or *network*.

A subgraph of a graph G is a graph G' whose vertices and edges form subsets of the vertices and edges of G. If G' is a subgraph of G, then G is called a supergraph of G'. An induced subgraph is a subset of the vertices of a graph G together with all edges both of whose endpoints are in this subset.

A graph G=(V,E) is called connected if, for any vertices $u,v\in V$, there exists a (u-v) path, i.e., a sequence of edges $uw_1=w_0w_1,\,w_1w_2,\,\ldots,\,w_{n-1}w_n=w_{n-1}v$ from E such that $w_i\neq w_j$ for $i\neq j,\,i,j\in\{0,1,\ldots,n\}$. A graph is called m-connected if there is no set of m-1 edges whose removal disconnects the graph; a connected graph is 1-connected. A digraph D=(V,E) is called strongly connected if, for any vertices $u,v\in V$, the directed (u-v) path and the directed (v-u) path both exist. A maximal connected subgraph of a graph G is called its connected component.

Vertices connected by an edge are called *adjacent*. The *degree* deg(v) of a vertex $v \in V$ of a graph G = (V, E) is the number of its vertices adjacent to v.

A complete graph is a graph in which each pair of vertices is connected by an edge. A bipartite graph is a graph in which the set V of vertices is decomposed into two disjoint subsets so that no two vertices within the same subset are adjacent. A path is a simple connected graph in which two vertices have degree one, and other vertices (if they exist) have degree two; the length

of a path is the number of its edges. A *cycle* is a *closed path*, i.e., a simple connected graph in which every vertex has degree two. A *tree* is a simple connected graph without cycles.

Two graphs which contain the same number of vertices connected in the same way are called *isomorphic*. Formally, two graphs G = (V(G), E(G)) and H = (V(H), E(H)) are called *isomorphic* if there is a bijection $f: V(G) \to V(H)$ such that, for any $u, v \in V(G)$, $uv \in E(G)$ if and only if $f(u)f(v) \in E(H)$.

We will consider mainly simple finite graphs and digraphs; more exactly, the equivalence classes of such isomorphic graphs.

15.1 Distances on vertices of a graph

• Path metric

The **path metric** (or **graphic metric**, shortest path metric) d_{path} is a metric on the vertex-set V of a connected graph G = (V, E), defined, for any $u, v \in V$, as the length of a shortest (u - v) path in G, i.e., a geodesic.

The path metric of the Cayley graph Γ of a finitely-generated group (G, \cdot, e) is called a **word metric**. The path metric of a graph G = (V, E), such that V can be cyclically ordered in a Hamiltonian cycle (a circuit containing each vertex exactly once), is called a **Hamiltonian metric**.

The **hypercube metric** is the path metric of a hypercube graph H(m, 2) with the vertex-set $V = \{0, 1\}^m$, and whose edges are the pairs of vectors $x, y \in \{0, 1\}^m$ such that $|\{i \in \{1, \dots, n\} : x_i \neq y_i\}| = 1$; it is equal to $|\{i \in \{1, \dots, n\} : x_i = 1\} \triangle \{i \in \{1, \dots, n\} : y_i = 1\}|$. The graphic metric space associated with a hypercube graph is called a hypercube metric space. It coincides with the metric space $(\{0, 1\}^m, d_{l_1})$.

Given an integer $n \ge 1$, the **line metric on** $\{1, ..., n\}$ in Chap. 1 is the path metric of the path $P_n = \{1, ..., n\}$.

• Weighted path metric

The weighted path metric d_{wpath} is a metric on the vertex-set V of a connected weighted graph G = (V, E) with positive edge-weights $(w(e))_{e \in E}$, defined by

$$\min_{P} \sum_{e \in P} w(e),$$

where the minimum is taken over all (u-v) paths P in G.

• Detour distance

Given a connected graph G = (V, E), the **detour distance** (or *codistance*) is a distance on the vertex-set V, defined as the length of a longest (u - v) path in G. In general, it is not a metric.

The smallest detour distance between distinct vertices is called the co-diameter of G. A graph is called a $detour\ graph$ if its detour distance coincides with its path metric.

• Path quasi-metric in digraphs

The path quasi-metric in digraphs d_{dpath} is a quasi-metric on the vertex-set V of a strongly connected directed graph D = (V, E), defined, for any $u, v \in V$, as the length of a shortest directed (u - v) path in D.

• Graph diameter

Given a connected graph G = (V, E), its **graph diameter** is the maximal length of shortest (u - v)-path in G, i.e., it is the largest value of the **path metric** between vertices of G. A connected graph is *vertex-critical* (*edge-critical*) if deleting any vertex (edge) increases its diameter.

Given a strong orientation O of a connected graph G=(V,E), i.e., a strongly connected directed graph D=(V,E') with arcs $e'\in E'$ obtained from edges $e'\in E'$ by orientation O, the oriented diameter of D is the maximal length of shortest directed (u-v)-path in it. The orientation O is tight if the diameter of G is equal to the oriented diameter of D. For example, a hypercube graph H(m,2) admits a tight orientation if $m\geq 4$ (McCanna 1988).

• Circular metric in digraphs

The **circular metric in digraphs** is a metric on the vertex-set V of a strongly connected directed graph D = (V, E), defined by

$$d_{dpath}(u,v) + d_{dpath}(v,u),$$

where d_{dpath} is the path quasi-metric in digraphs.

• Strong metric in digraphs

The strong metric in digraphs is a metric between vertices v and v of a strongly connected directed graph D = (V, E), defined (Chartrand, Erwin, Raines and Zhang 1999) as the minimum size (the number of edges) of a strongly connected subdigraph of D, containing v and v. Cf. Steiner distance.

• Y-metric

Given a class Υ of connected graphs, the metric d of a metric space (X, d) is called a Υ -metric if (X, d) is isometric to a subspace of a metric space (V, d_{wpath}) , where $G = (V, E) \in \Upsilon$, and d_{wpath} is the weighted path metric on the vertex-set V of G with positive edge-weight function w (see tree-like metric).

• Tree-like metric

A tree-like metric (or weighted tree metric) d on a set X is a Υ -metric for the class Υ of all trees, i.e., the metric space (X,d) is isometric to a subspace of a metric space (V,d_{wpath}) , where T=(V,E) is a tree, and d_{wpath} is the weighted path metric on the vertex-set V of T with a positive weight function w. A metric is a tree-like metric if and only if it satisfies the four-point inequality.

A metric d on a set X is called a **relaxed tree-like metric** if the set X can be embedded in some (not necessary positively) edge-weighted tree such that, for any $x, y \in X$, d(x, y) is equal to the sum of all edge weights along the (unique) path between corresponding vertices x and y in the tree. A metric is a relaxed tree-like metric if and only if it satisfies the **relaxed four-point inequality**.

• Katz similarity

Given a graph G = (V, E) with positive edge-weight function $w = (w(e))_{e \in E}$, let $V = \{v_1, \ldots, v_n\}$. Denote by A the $n \times n$ -matrix $((a_{ij}))$, where $a_{ij} = a_{ji} = w(ij)$ if ij is an edge, and $a_{ij} = 0$, otherwise. Let I be the identity $n \times n$ -matrix, and let $\alpha, 0 < \alpha < 1$, be a parameter with $\alpha < (\max_i |\lambda_i|)^{-1}$, where the maximum is taken over the eigenvalues λ_i of matrix A. Define the $n \times n$ -matrix $S = ((s_{ij}))$ as follows:

$$S = \sum_{1 \le k \le \infty} \lambda^k A^k = (I - \lambda A)^{-1} - I.$$

The number s_{ij} is called the **Katz similarity** between vertices v_i and v_j ; it was proposed (Katz 1953) for evaluating social status with better accounting of all paths between v_i and v_j .

• Resistance metric

Given a connected graph G=(V,E) with positive edge-weight function $w=(w(e))_{e\in E}$, let us interpret the edge-weights as resistances. For any two different vertices u and v, suppose that a battery is connected across them, so that one unit of a current flows in at u and out in v. The voltage (potential) difference, required for this is, by Ohm's law, the effective resistance between u and v in an electrical network; it is called the **resistance metric** $\Omega(u,v)$ between them (Gvishiani-Gurvich, 1987, and Klein-Randic, 1993). So, if a potential of one volt is applied across vertices u and v, a current of $\frac{1}{\Omega(u,v)}$ will flow. The number $\frac{1}{\Omega(u,v)}$ can be seen, like electrical conductance, as a measure of connectivity between u and v.

Let $r(u,v) = \frac{1}{w(v)}$ if uv is an edge, and r(u,v) = 0 otherwise. Formally,

$$\Omega(u,v) = (\sum_{w \in V} f(w)r(w,v))^{-1},$$

where $f:V\to [0,1]$ is the unique function with f(u)=1, f(v)=0 and $\sum_{z\in V}(f(w)-f(z))r(w,z)=0$ for any $w\neq u,v$.

The resistance metric is applied when the number of paths between any two vertices u and v matters; in short, it is a weighted average of the lengths of all (u-v) paths.

A probabilistic interpretation (Gobel and Jagers 1974) is: $\Omega(u,v) = (deg(u)Pr(u \to v))^{-1}$, where deg(u) is the degree of the vertex u, and $Pr(u \to v)$ is the probability for a random walk leaving u to arrive at v before returning to u. The expected commuting time between vertices u and v is $2\sum_{e\in E} w(e)\Omega(u,v)$ in general.

Then $\Omega(u,v) \leq \min_P \sum_{e \in P} \frac{1}{w(e)}$, where P is any (u-v) path, with equality if and only if such a path P is unique. So, if w(e) = 1 for all edges, the equality means that G is a **geodetic graph**, and hence the path and resistance metrics coincide.

If w(e) = 1 for all edges, then $\Omega(u, v) = (g_{uu} + g_{vv}) - (g_{uv} + g_{vu})$, where $((g_{ij}))$ is the Moore–Penrose generalized inverse of the Laplacian matrix $((l_{ij}))$ of the graph G: here l_{ii} is the degree of vertex i while, for $i \neq j$, $l_{ij} = 1$ if the vertices i and j are adjacent, and $l_{ij} = 0$ otherwise.

The distance $\sqrt{\Omega(u,v)}$ is a **Mahalanobis distance** (cf. Chap. 17) with a weighting matrix $((g_{ij}))$. This distance is called a *diffusion metric* in [CLLMNWZ05], because it (as well as diffusion) depends on a random walk.

• Hitting time quasi-metric

Let G = (V, E) be a connected graph with m edges. Consider random walks on G, where at each step the walk moves to a vertex randomly with uniform probability from the neighbors of the current vertex. The **hitting** (or *first-passage*) **time quasi-metric** H(u, v) from $u \in V$ to $v \in V$ is the expected number of steps (edges) for a random walk on G beginning at u to reach v for the first time; it is 0 for u = v.

This quasi-metric is a weightable quasi-semi-metric (cf. Chap. 1). The commuting time metric is C(u, v) = H(u, v) + H(v, u).

Then $C(u,v)=2m\Omega(u,v)$, where $\Omega(u,v)$ is the **resistance metric** (or *effective resistance*), i.e., 0 if u=v and, otherwise, $\frac{1}{\Omega(u,v)}$ is the current flowing into v, when grounding v and applying a 1V potential to u (each edge is seen as a resistor of 1Ω). Then $\Omega(u,v)=\sup_{f:V\to\mathbb{R},\ D(f)>0}\frac{(f(u)-f(v))^2}{D(f)}$, where D(f) is the *Dirichlet energy* of f, i. e., $\sum_{st\in E}(f(s)-f(t))^2$.

Above setting can be easily generalized to weighted graphs.

• Forest metric

Given $\alpha > 0$ and a connected weighted multi-graph (multiple edges are allowed) G = (V, E; w) with positive edge-weight function $w = (w(e))_{e \in E}$, the α -forest metric (Chebotarev and Shamis 2006) between vertices u and v is defined by

$$\frac{1}{2}(q_{uu} + q_{vv} - q_{uv} - q_{vu})$$

for $((g_{ij})) = (I + \alpha L)^{-1}$, where I is the identity $|V| \times |V|$ matrix, and $L = ((l_{il}))$ is the Laplacian (or Kirchhoff) matrix of G, i.e., $l_{ij} = -w(ij)$ for $i \neq j$ and $l_{ij} = -\sum_{j \neq i} l_{ij}$.

Chebotarev and Shamis showed that twice the α -forest metric of G is the **resistance distance** of the weighted multi-graph G' = (V', E'; w') with $V' = V \cup \{0\}$, $E' = E \cup \{u0 : u \in V\}$), while $w'(e) = \alpha w(e)$ for all $e \in E$ and w'(u0) = 1 for all $u \in V$.

Their forest metric (1998) is the case $\alpha = 1$ of the α -forest metric.

• Truncated metric

The **truncated metric** is a metric on the vertex-set of a graph, which is equal to 1 for any two adjacent vertices, and is equal to 2 for any non-adjacent different vertices. It is the 2-**truncated metric** for the path metric of the graph. It is the (1,2) - B-metric if the degree of any vertex is at most B.

• Hsu-Lyuu-Flandrin-Li distance

Given an m-connected graph G = (V, E) and two vertices $u, v \in V$, a container C(u, v) of width m is a set of m (u - v) paths with any two of them intersecting only in u and v. The length of a container is the length of the longest path in it.

The **Hsu-Lyuu-Flandrin-Li distance** between vertices u and v (Hsu-Lyuu 1991 and Flandrin-Li 1994) is the minimum of container lengths taken over all containers C(u, v) of width m. This generalization of the path metric is used in parallel architectures for interconnection networks.

• Multiply-sure distance

The **multiply-sure distance** is a distance on the vertex-set V of an m-connected weighted graph G = (V, E), defined, for any $u, v \in V$, as the minimum weighted sum of lengths of m disjoint (u - v) paths. This generalization of the path metric helps when several disjoint paths between two points are needed, for example, in communication networks, where m-1 of (u-v) paths are used to code the message sent by the remaining (u-v) path (see [McCa97]).

• Cut semi-metric

A cut is a partition of a set into two parts. Given a subset S of $V_n = \{1, \ldots, n\}$, we obtain the partition $\{S, V_n \setminus S\}$ of V_n . The cut semi-metric (or split semi-metric) δ_S , defined by this partition, is a semi-metric on V_n , defined by

$$\delta_S(i,j) = \begin{cases} 1, & \text{if } i \neq j, |S \cap \{i,j\}| = 1, \\ 0, & \text{otherwise.} \end{cases}$$

Usually, it is considered as a vector in $\mathbb{R}^{|E_n|}$, $E(n) = \{\{i, j\} : 1 \leq i < j \leq n\}$.

A circular cut of V_n is defined by a subset $S_{[k+1,l]} = \{k+1,\ldots,l\}$ $(mod\,n) \subset V_n$: if we consider the points $\{1,\ldots,n\}$ as being ordered along a circle in that circular order, then $S_{[k+1,l]}$ is the set of its consecutive vertices from k+1 to l. For a circular cut, the corresponding cut semi-metric is called a **circular cut semi-metric**.

An even cut semi-metric (odd cut semi-metric) is δ_S on V_n with even (odd, respectively) |S|. A k-uniform cut semi-metric is δ_S on V_n with $|S| \in \{k, n - k\}$. An equicut semi-metric (inequicut semi-metric) is δ_S on V_n with $|S| \in \{\lfloor \frac{n}{2} \rfloor, \lceil \frac{n}{2} \rceil\}$ ($|S| \notin \{\lfloor \frac{n}{2} \rfloor, \lceil \frac{n}{2} \rceil\}$, respectively); see, for example, [DeLa97].

• Decomposable semi-metric

A decomposable semi-metric is a semi-metric on $V_n = \{1, ..., n\}$ which can be represented as a non-negative linear combination of **cut semi-metrics**. The set of all decomposable semi-metrics on V_n is a *convex cone*, called the *cut cone* CUT_n .

A semi-metric on V_n is decomposable if and only if it is a **finite** l_1 -semi-metric.

A circular decomposable semi-metric is a semi-metric on $V_n = \{1, \ldots, n\}$ which can be represented as a non-negative linear combination of circular cut semi-metrics.

A semi-metric on V_n is circular decomposable if and only if it is a **Kalmanson semi-metric** with respect to the same ordering (see [ChFi98]).

• Finite l_p -semi-metric

A finite l_p -semi-metric d is a semi-metric on $V_n = \{1, \ldots, n\}$ such that (V_n, d) is a semi-metric subspace of the l_p^m -space (\mathbb{R}^m, d_{l_p}) for some $m \in \mathbb{N}$. If, instead of V_n , is taken $X = \{0, 1\}^n$, the metric space (X, d) is called the l_p^n -cube. The l_1^n -cube is called a **Hamming cube**; cf. Chap. 4.

• Kalmanson semi-metric

A Kalmanson semi-metric d is a semi-metric on $V_n = \{1, ..., n\}$ which satisfies the condition

$$\max\{d(i,j) + d(r,s), d(i,s) + d(j,r)\} \le d(i,r) + d(j,s)$$

for all $1 \leq i \leq j \leq r \leq s \leq n$. In this definition the ordering of the elements is important; so, d is a Kalmanson semi-metric with respect to the ordering $1, \ldots, n$.

Equivalently, if the points $\{1, \ldots, n\}$ are ordered along a circle C_n in that circular order, then the distance d on V_n is a Kalmanson semi-metric if the inequality

$$d(i,r) + d(j,s) \leq d(i,j) + d(r,s)$$

holds for $i, j, r, s \in V_n$ whenever the segments [i, j], [r, s] are crossing chords of C_n .

A tree-like metric is a Kalmanson metric for some ordering of the vertices of the tree. The Euclidean metric, restricted to the points that form a convex polygon in the plane, is a Kalmanson metric.

• Multi-cut semi-metric

Let $\{S_1, \ldots, S_q\}$, $q \geq 2$, be a partition of the set $V_n = \{1, \ldots, n\}$, i.e., a collection S_1, \ldots, S_q of pairwise disjoint subsets of V_n such that $S_1 \cup \cdots \cup S_q = V_n$.

The **multi-cut semi-metric** $\delta_{S_1,...,S_q}$ is a semi-metric on V_n , defined by

$$\delta_{S_1,...,S_q}(i,j) = \begin{cases} 0, & \text{if } i,j \in S_h \text{ for some } h,1 \le h \le q, \\ 1, & \text{otherwise.} \end{cases}$$

• Oriented cut quasi-semi-metric

Given a subset S of $V_n = \{1, ..., n\}$, the **oriented cut quasi-semi-metric** δ'_S is a quasi-semi-metric on V_n , defined by

$$\delta_S^{'}(i,j) = \begin{cases} 1, & \text{if} \ i \in S, j \not\in S, \\ 0, & \text{otherwise.} \end{cases}$$

Usually, it is considered as the vector of $\mathbb{R}^{|I_n|}$, $I(n) = \{(i,j) : 1 \leq i \neq j \leq n\}$. The **cut semi-metric** δ_S is $\delta_S^{'} + \delta_{V_n \setminus S}^{'}$.

• Oriented multi-cut quasi-semi-metric Given a partition $\{S_1, \ldots, S_q\}$, $q \geq 2$, of V_n , the oriented multi-cut quasi-semi-metric $\delta'_{S_1, \ldots, S_q}$ is a quasi-semi-metric on V_n , defined by

$$\delta^{'}_{S_1,...,S_n}(i,j) = \begin{cases} 1, & \text{if } i \in S_h, j \in S_m, h < m, \\ 0, & \text{otherwise} \end{cases}$$

15.2 Distance-defined graphs

Below we first give some graphs defined in terms of distances between their vertices. Then some graphs associated with metric spaces are presented.

A graph (V, E) is, say, distance-invariant or distance monotone if its metric space (V, d_{path}) is **distance invariant** or **distance monotone**, respectively (cf. Chap. 1). The definitions of such graphs, being straightforward subcases of corresponding metric spaces, will be not given below.

• k-power of a graph

The k-power of a graph G = (V, E) is the supergraph $G^k = (V, E')$ of G with edges between all pairs of vertices having path distance at most k.

• Isometric subgraph

A subgraph H of a graph G = (V, E) is called an **isometric subgraph** if the path metric between any two points of H is the same as their path metric in G.

A subgraph H is called a *convex subgraph* if it is isometric, and for any $u, v \in H$ every vertex on a shortest (u - v)-path belonging to H also belongs to H.

A subset $M \subset V$ is called *gated* if for every $u \in V \setminus M$ there exists a unique vertex $g \in M$ (called a *gate*) lying on a shortest (u - v)-path for every $v \in M$. The subgraph induced by a gated set is a convex subgraph.

• Retract subgraph

A subgraph H of a graph G=(V,E) is called a **retract subgraph** if it is induced by an idempotent **short mapping** of G into itself, i.e., $f^2=f:V\to V$ with $d_{path}(f(u),f(v))\leq d_{path}(u,v)$ for all $u,v\in V$. Any retract subgraph is **isometric**.

• Median graph

A connected graph G = (V, E) is called a **median** if, for every three vertices $u, v, w \in V$, there exists a unique vertex that lies simultaneously on a shortest (u - v)-path, a shortest (u - w)-path and a shortest (w - v)-path, i.e., (V, d_{path}) is a **median metric space**. The median graphs are exactly **retract subgraphs** of hypercubes. Also, they are exactly such **isometric subgraphs** of hypercubes that the vertex-set of any *convex subgraph* is *gated* (cf. **isometric subgraph**).

• Geodetic graph

A connected graph is called **geodetic** if there exists exactly one shortest path between any two of its vertices. Every tree is a geodetic graph.

The geodetic number of a finite connected graph (V, E) [BuHa90] is $\min |M|$ over sets $M \subset V$ of vertices such that any vertex $x \in V$ lies on a shortest (u - v)-path where $u, v \in M$.

• Interval distance monotone graph

A connected graph G = (V, E) is called **interval distance monotone** if any of its intervals $I_G(u, v)$ induces a distance monotone graph, i.e., its path-metric is **distance monotone**, cf. Chap. 1. A graph is interval distance monotone if and only if (Zhang and Wang 2007) each of its intervals is isomorphic to either a path, a cycle or a hypercube.

• Distance-regular graph

A connected graph G=(V,E) of diameter T is called **distance-regular** if, for any of its vertices u,v and any integers $0 \le i,j \le T$, the number of vertices w, such that $d_{path}(u,w)=i$ and $d_{path}(v,w)=j$, depends only on i,j and $k=d_{path}(u,v)$, but not on the choice of u and v.

A special case of it is a **distance-transitive graph**, i.e., such that its group of automorphisms is transitive, for any $0 \le i \le T$, on the pairs of vertices (u, v) with $d_{path}(u, v) = i$.

Any distance-regular graph is a **distance-balanced graph**(i.e., $|\{x \in V : d(x,u) < d(x,v)\}| = |\{x \in V : d(x,v) < d(x,u)\}|$ for any edge uv), a **distance degree regular graph** (i.e., $|\{x \in V : d(x,u) = i\}|$ depends only on i, not on $u \in V$), and a **walk-regular graph** (i.e., the number of closed walks of length i starting at u depends only on i, not on u).

A distance-regular graph is also called a **metric association scheme** or P-polynomial association scheme. A finite **polynomial metric space** (cf. Chap. 1) is a special case of it, also called a $(P \ and \ Q)$ -polynomial association scheme.

• Metrically almost transitive graph

An automorphism of a graph G = (V, E) is a map $g : V \to V$ such that u is adjacent to v if and only if g(u) is adjacent to g(v), for any vertices u and v. The set Aut(G) of all automorphisms of G is a group with respect to the composition of functions.

A graph G(V, E) is **metrically almost transitive** (Krön and Möller 2008) if there is an integer r such that, for any vertex u,

$$\cup_{g \in Aut(G)} \{ g(B(u,r)) \} = V,$$

where $B(u,r) = \{v \in V : d_{path}(u,v) \leq r\}$. The smallest such integer r is called the *covering radius* of G. Cf radii of metric space in Chap. 1.

• Graph of polynomial growth

Let G = (V, E) be a transitive locally-finite graph. For a vertex $v \in V$, the growth function is defined by

$$f(n) = |\{u \in V : d(u, v) \le n\}|,$$

and it does not depend on a particular vertex v. Cf. growth rate of metric space in Chap. 1.

The graph G is a **graph of polynomial growth** if there are some positive constants k, C such that $f(n) \leq Cn^k$ for all $n \geq 0$. It is a **graph of exponential growth** if there is a constant C > 1 such that $f(n) > C^n$ for all $n \geq 0$.

A group with a finite symmetric set of generators has polynomial growth rate if the corresponding $Cayley\ graph$ has polynomial growth. Here the metric ball consists of all elements of the group which can be expressed as products of at most n generators, i.e., it is a closed ball centered in the identity in the **word metric**, cf. Chap. 10.

• Distance-polynomial graph

Given a connected graph G = (V, E) of diameter T, for any $2 \le i \le T$ denote by G_i the graph with the same vertex-set as G, and with edges uv such that $d_{path}(u, v) = i$. The graph G is called a **distance-polynomial** if the adjacency matrix of any G_i , $2 \le i \le T$, is a polynomial in terms of the adjacency matrix of G.

Any distance-regular graph is a distance-polynomial.

• Distance-hereditary graph

A connected graph is called **distance-hereditary** if each of its connected induced subgraphs is isometric.

A graph is distance-hereditary if each of its induced paths is isometric. A graph is distance-hereditary, bipartite distance-hereditary, block graph, tree if and only if its path metric is a relaxed tree-like metric for edge-weights being, respectively, non-zero half-integers, non-zero integers, positive half-integers, positive integers.

A graph is called a **parity graph** if, for any of its vertices u and v, the lengths of all induced (u-v)-paths have the same parity. A graph G is k-distance hereditary (Meslem-Aïder, 2009) if $d_H(u,v) \leq d_G(u,v) + k$ for vertices of any its connected induced subgraph H.

• Block graph

A graph is called a **block graph** if each of its *blocks* (i.e., a maximal 2-connected induced subgraph) is a complete graph. Any tree is a block graph.

A graph is a block graph if and only if its path metric is a tree-like metric or, equivalently, satisfies the four-point inequality.

• Ptolemaic graph

A graph is called **Ptolemaic** if its path metric satisfies the **Ptolemaic** inequality

$$d(x,y)d(u,z) \le d(x,u)d(y,z) + d(x,z)d(y,u).$$

A graph is Ptolemaic if and only if it is distance-hereditary and chordal, i.e., every cycle of length greater than 3 has a chord. So, any block graph is Ptolemaic.

• t-irredundant set

A set $S \subset V$ of vertices in a connected graph G = (V, E) is called t-irredundant (Hattingh and Henning 1994) if for any $u \in S$ there exists a vertex $v \in V$ such that, for the path metric d_{path} of G,

$$d_{path}(v, x) \le t < d_{path}(v, V \setminus S) = \min_{u \notin S} d_{path}(v, u).$$

The t-irredundance number ir_t of G is the smallest cardinality |S| such that S is t-irredundant but $S \cup \{v\}$ is not, for every $v \in V \setminus S$.

The t-domination number γ_t and t-independent number α_t of G are, respectively, the cardinality of the smallest t-covering and largest $\frac{t}{2}$ packing of the metric space $(V, d_{path}(u, v))$ (cf. radius of metric space in Chap. 1). Denote by γ_t^{in} the smallest |S| such that S is $\frac{t}{2}$ -packing but $S \cup \{v\}$ is not, for every $v \in V \setminus S$; so, this non-extendible $\frac{t}{2}$ -packing is also a t-covering. Then $\frac{\gamma_t+1}{2} \leq ir_t \leq \gamma_t \leq \gamma_t^{in} \leq \alpha_t$. k-distant chromatic number

The k-distant chromatic number of a graph G = (V, E) is the minimum number of colors needed to color vertices of G so that any two vertices at distance at most k have distinct colors, i.e., it is the chromatic number of the k-power of G.

D-distance graph

Given a set D of positive numbers containing 1 and a metric space (X, d), the D-distance graph D(X,d) is a graph with the vertex-set X and the edge-set $\{uv: d(u,v) \in D\}$ (cf. **D-chromatic number** in Chap. 1).

A D-distance graph is called a **distance graph** (or unit-distance graph) if $D = \{1\}$, an ϵ -unit graph if $D = [1 - \epsilon, 1 + \epsilon]$, a unit-neighborhood graph if D=(0,1], an integral-distance graph if $D=\mathbb{Z}_+$, a rational-distance graph if $D = \mathbb{Q}_+$, and a prime-distance graph if D is the set of prime numbers (with 1).

Usually, the metric space (X,d) is a subspace of a Euclidean space \mathbb{E}^n . Moreover, every finite graph G=(V,E) can be represented by a D-distance graph in some \mathbb{E}^n . The minimum dimension of such Euclidean space is called the D-dimension of G.

• Distance-number of a graph

Given a graph G = (V, E), its degenerate drawing is a mapping $f : V \to \mathbb{R}^2$ such that |f(V)| = |V| and f(uv) is an open straight-line segment joining the vertices f(u) and f(v) for any edge $uv \in E$; it is a drawing if, moreover, $f(w) \notin f(uv)$ for any $uv \in E$ and $w \in V$.

The distance-number of a graph G = (V, E), denoted by dn(G), is (Carmi, Dujmović, Morin and Wood 2008) the minimum number of distinct edge-lengths in a drawing of G. The degenerate distance-number of G, denoted by ddn(G), is the minimum number of distinct edge-lengths in a degenerated drawing of G.

The first of the **Erdös-type distance problems** in Chap. 19 is equivalent to determining $ddn(K_n)$.

The unit-distance graph of a set $M \subset \mathbb{R}^2$ is a graph G' = (V', E') with V' = M and $xy \in E'$ if and only if points $x, y \in S$ are at unit distance. In general, ddn(G) = 1 if and only if G is isomorphic to a subgraph of a unit-distance graph.

Any *n*-vertex *m*-edge graph G satisfies (Spencer, Szemerédi and Trotter 1984) $dn(G) > ddn(G) > Cmn^{-\frac{4}{3}}$ for a constant C > 0.

Erdös, Harary and Tutte (1965) defined the dimension of a graph G as the minimum number k such that G has a degenerate drawing in \mathbb{R}^k with straight-line edges of unit length.

A graph is k-realizable if, for every mapping of its vertices to (not necessarily distinct) points of \mathbb{R}^s with $s \geq k$, there exists such a mapping in \mathbb{R}^k which preserves edge-lengths. K_3 is 2-realizable but not 1-realizable. Belk and Connely (2007) proved that a graph is 3-realizable if and only if it has no minors K_5 or $K_{2,2,2}$.

• Bar framework

A pair (G, f) is called a **bar framework** if G = (V, E) is a finite graph (no loops and multiple edges) and $f: V \to \mathbb{R}^n$ is a map with $f(u) \neq f(v)$ whenever $uv \in E$. The vertices and edges are called *joints* and *bars*, respectively, in terms of Structural Engineering. A **tensegrity** (Fuller 1948) is a bar framework in which bars are either *cables* (i.e., cannot get further apart), or *struts* (i.e., cannot get closer together).

An (infinitesimal) motion of a bar framework (G, f) is a map $m : V \to \mathbb{R}^n$ with (m(u) - m(v))(f(u) - f(v)) = 0 whenever $uv \in E$. A motion is trivial if it can be extended to an isometry of \mathbb{R}^n . A bar framework is an (infinitesimally) **rigid framework** if every motion of it is trivial.

A bar framework (G, f) is an **elastic framework** if, for any $\epsilon > 0$, there exists a $\delta > 0$ such that the following condition holds: for every edge-weighting $w: E \to \mathbb{R}_{>0}$ with $\max_{uv \in E} |w(uv) - ||f(u) - f(v)||_2| \leq \delta$, there exist a bar framework (G, f') with $\max_{v \in V} ||f(u) - f'(v)||_2 < \epsilon$.

A bar framework is *isostatic* (i.e., rigid and the deletion of any of its edges will cause loss of rigidity) if and only if (Tay and Nievergelt 1997) it is elastic and the addition of any new edge will cause loss of elasticity.

• Distance-two labelling

Given a decreasing sequence $\alpha = (\alpha_1, \ldots, \alpha_k)$ of numbers, λ_{α} -labelling of a graph G = (V, E) is an assignment of labels f(v) from the set $\{0, 1, \ldots, \lambda\}$ of integers to the vertices $v \in V$ such that, for any t with $0 \le t \le k$, $|f(v) - f(u)| \ge \alpha_t$ whenever the path distance between u and v is t. The radio frequency assignment problem, where vertices v are transmitters and labels f(v) are frequencies of (not-interfering) channels, consists of minimizing λ .

Distance-two labelling $(\lambda_{(2,1)}$ -labelling) is the main interesting case $\alpha = (2,1)$.

• Distance labelling scheme

A graph family A is said (Peleg 2000) to have an l(n) distance labelling scheme if there is a function L_G labelling the vertices of each n-vertex graph $G \in A$ with distinct labels up to l(n) bits, and there exists an algorithm, called a **distance decoder**, that decides the distance d(u, v) between any two vertices $u, v \in X$ in a graph $G \in A$, i.e., $d(u, v) = f(L_G u), L_G v)$, polynomial in time in the length of their labels L(u), L(v).

• Arc routing problems

Given a finite set X, a quasi-distance d(x,y) on it and a set $A \subseteq \{(x,y): x,y \in X\}$, consider the weighted digraph D=(X,A) with the vertex-set X and arc-weights d(x,y) for all arcs $(x,y) \in A$. For given sets V of vertices and E of arcs, the **arc routing problem** consists of finding a shortest (i.e., with minimal sum of weights of its arcs) (V,E)-tour, i.e., a circuit in D=(X,A), visiting each vertex in V and each arc in E exactly once or, in a variation, at least once.

The Asymmetric Traveling Salesman Problem corresponds to the case V = X, $E = \emptyset$; the Traveling Salesman Problem is the symmetric version of it (usually, each vertex should be visited exactly once). The Bottleneck Traveling Salesman Problem consists of finding a (V, E)-tour T with smallest $\max_{(x,y)\in T} d(x,y)$.

The Windy Postman Problem corresponds to the case $V = \emptyset$, E = A, while the Chinese Postman Problem is the symmetric version of it.

The above problems are also considered for general arc- or edge-weights; then, for example, term $Metric\ TSP$ is used when edge-weights in the Traveling Salesman Problem satisfy the triangle inequality, i.e., d is a quasi-semi-metric.

• Steiner distance of a set

The **Steiner distance of a set** $S \subset V$ of vertices in a connected graph G = (V, E) is (Chartrand, Oellermann, Tian and Zou 1989) the minimum size (number of edges) of a connected subgraph of G, containing S. Such a subgraph is, obviously, a tree, and is called a *Steiner tree* for S. Those of its vertices which are not in S are called *Steiner points*.

The Steiner distance of the set $S = \{u, v\}$ is the path metric between u and v.

• t-spanner

A spanning subgraph H=(V,E(H)) of a connected graph G=(V,E) is called a t-spanner of G if, for every $u,v\in V$, the inequality $d_{path}^H(u,v)/d_{path}^G(u,v)\leq t$ holds. The value t is called the stretch factor of H.

A spanning subgraph H=(V,E(H)) of a graph G=(V,E) is called a k-additive spanner of G if, for every $u,v\in V$, the inequality $d_{path}^H(u,v)\leq d_{path}^G(u,v)+k$ holds.

• Proximity graph

Given a finite subset V of a metric space (X,d), a **proximity graph** of V is a graph representing neighbor relationships between points of V. Such graphs are used in Computational Geometry and many real-world problems. The main examples are presented below. Cf. also **underlying** graph of a metric space in Chap. 1.

A spanning tree of V is a set T of |V|-1 unordered pairs (x,y) of different points of V forming a tree on V; the weight of T is $\sum_{(x,y)\in T} d(x,y)$. A **minimum spanning tree** MST(V) of V is a spanning tree with the minimal weight. Such a tree is unique if the edge-weights are distinct.

Nearest neighbor graph is the directed graph NNG(V) = (V, E) with vertex-set $V = v_1, \ldots, v_{|V|}$ and, for $x, y \in V$, $xy \in E$ if y is the nearest neighbor of x, i.e., $d(x,y) = \min_{v_i \in V \setminus \{x\}} d(x,v_i)$ and only v_i with maximal index i is picked. The k-nearest neighbor graph arises if k such v_i with maximal indices are picked. The indirect version of NNG(V) is a subgraph of MST(V).

Relative neighborhood graph is (Toussaint 1980) the graph RNG(V) = (V, E) with vertex-set V and, for $x, y \in V$, $xy \in E$ if there is no point $z \in V$ with $\max\{d(x, z), d(y, z)\} < d(x, y)$. Also considered, in the main case $(X, d) = (\mathbb{R}^2, ||x - y||_2)$, are the related Gabriel graph GG(V) (in general, β -skeleton) and Delaunay triangulation DT(V); then $NNG(V) \subseteq MST(V) \subseteq RNG(V) \subseteq GG(V) \subseteq DT(V)$.

For any $x \in V$, its *sphere of influence* is the open metric ball $B(x, r_x) = \{z \in X : d(x, z) < r\}$ in (X, d) centered at x with radius $r_x = \min_{z \in V \setminus \{x\}} d(x, z)$.

Sphere of influence graph is the graph SIG(V) = (V, E) with vertexset V and, for $x, y \in V$, $xy \in E$ if $B(x, r_x) \cap B(y, r_y) \neq \emptyset$; so, it is a proximity graph and an *intersection graph*. The *closed sphere of influ-* $E(x, r_x) \cap B(y, r_y) \neq \emptyset$.

15.3 Distances on graphs

• Chartrand-Kubicki-Schultz distance

The Chartrand-Kubicki-Schultz distance (or ϕ -distance 1998) between two connected graphs $G_1 = (V_1, E_1)$ and $G_2 = (V_2, E_2)$ with $|V_1| = |V_2| = n$ is

$$\min\{\sum |d_{G_1}(u,v)-d_{G_2}(\phi(u),\phi(v))|\},$$

where d_{G_1}, d_{G_2} are path metrics of graphs G_1, G_2 , the sum is taken over all unordered pairs u, v of vertices of G_1 , and the minimum is taken over all bijections $\phi: V_1 \to V_2$.

• Subgraph metric

Let $\mathbb{F} = \{F_1 = (V_1, E_1), F_2 = (V_2, E_2), \dots, \}$ be the set of isomorphism classes of finite graphs. Given a finite graph G = (V, E), denote by $s_i(G)$ the number of *injective homomorphisms* from F_i into G (i.e., the number of injections $\phi: V_i \to V$ with $\phi(x)\phi(y) \in E$ whenever $xy \in E_i$) divided by the number $\frac{|V|!}{(|V|-|V_i|)!}$ of such injections from F_i with $|V_i| \leq |V|$ into $K_{|V|}$). Set $s(G) = (s_i(G))_{i=1}^{\infty} \in [0,1]^{\infty}$.

Let d be the **Cantor metric** (cf. Chap. 18) $d(x,y) = \sum_{i=1}^{\infty} 2^{-i}|x_i - y_i|$ on $[0,1]^{\infty}$ or any metric on $[0,1]^{\infty}$ inducing the *product topology*. Then, the **subgraph metric** (Bollobás and Riordan 2007) between the graphs G_1 and G_2 is defined by

$$d(s(G_1), s(G_2)).$$

Bollobás and Riordan (2007) defined other metrics and generalized the subgraph distance on *kernels* (or *graphons*), i.e., symmetric measurable functions $k: [0,1] \times [0,1] \to \mathbb{R}_{\geq 0}$, replacing G by k and the above $s_i(G)$ by $s_i(k) = \int_{[0,1]^{|V_i|}} \prod_{st \in E_i} k(x_s x_t) \prod_{s=1}^{|V_i|} dx_s$.

• Rectangle distance on weighted graphs

Let $G = G(\alpha, \beta)$ be a complete weighted graph on $\{1, \ldots, n\}$ with vertexweights $\alpha_i > 0, \ 1 \le i \le n$, and edge-weights $\beta_{ij} \in \mathbb{R}, \ 1 \le i < j \le n$. Denote by A(G) the $n \times n$ matrix $((a_{ij}))$, where $a_{ij} = \frac{\alpha_i \alpha_j \beta_{ij}}{(\sum_{1 \le i \le n} \alpha_i)^2}$.

The **rectangle distance** (or *cut distance*) between two weighted graphs $G = G(\alpha, \beta)$ and $G' = G(\alpha', \beta')$ (with vertex-weights (α'_i) and edgeweights (β'_{ij})) is (Borgs, Chayes, Lovász, Sós and Vesztergombi 2007):

$$\max_{I,J \subset \{1,...,n\}} |\sum_{i \in I, i \in J} (a_{ij} - a'_{ij})| + \sum_{i=1}^{n} |\frac{\alpha_i}{\sum_{1 \le j \le n} \alpha_j} - \frac{\alpha'_i}{\sum_{1 \le j \le n} \alpha'_j}|,$$

where $A(G) = ((a_{ij}))$ and $A(G') = ((a'_{ij}))$.

In the case $(\alpha'_i) = (\alpha_i)$, the rectangle distance is $||A(G) - A(G')||_{cut}$, i.e., the **cut norm metric** (cf. Chap. 12) between matrices A(G) and A(G')

and the rectangle distance from Frieze and Kannan (1999). In this case, the l_1 - and l_2 -metrics between two weighted graphs G and G' are defined as $||A(G)-A(G')||_1$ and $||A(G)-A(G')||_2$, respectively. The subcase $\alpha_i=1$ for all $1\leq i\leq n$ corresponds to unweighted vertices.

Cf. the Robinson–Foulds weighted metric between phylogenetic trees.

Borgs, Chayes, Lovász, Sós and Vesztergombi (2007) defined other metrics and generalized the rectangle distance on *kernels* (or *graphons*), i.e., symmetric measurable functions $k:[0,1]\times[0,1]\to\mathbb{R}_{\geq 0}$, using the *cut* norm $||k||_{cut}=\sup_{T\subset[0,1]}|\int_{S\times T}k(x,y)dxdy|$.

norm $||k||_{cut} = \sup_{S,T \subset [0,1]} |\int_{S \times T} k(x,y) dx dy|$. A map $\phi : [0,1] \to [0,1]$ is measure-preserving if, for any measurable subset $A \subset [0,1]$, the measures of A and $\phi^{-1}(A)$ are equal. For a kernel k, define the kernel k^{ϕ} by $k^{\phi}(x,y) = k(\phi(x),\phi(y))$. The **Lovász–Szegedy semi-metric** (2007) between kernels k_1 and k_1 is defined by

$$\inf_{\phi} ||k_1^{\phi} - k_2||_{cut},$$

where ϕ ranges over all measure-preserving bijections $[0,1] \to [0,1]$. Cf. Chartrand–Kubicki–Schultz distance.

• Subgraph-supergraph distances

A common subgraph of graphs G_1 and G_2 is a graph which is isomorphic to induced subgraphs of both G_1 and G_2 . A common supergraph of graphs G_1 and G_2 is a graph which contains induced subgraphs isomorphic to G_1 and G_2 .

The **Zelinka distance** d_Z [Zeli75] on the set **G** of all graphs (more exactly, on the set of all equivalence classes of isomorphic graphs) is defined by

$$d_Z = \max\{n(G_1), n(G_2)\} - n(G_1, G_2)$$

for any $G_1, G_2 \in \mathbf{G}$, where $n(G_i)$ is the number of vertices in G_i , i = 1, 2, and $n(G_1, G_2)$ is the maximum number of vertices of a common subgraph of G_1 and G_2 .

The **Bunke–Shearer metric** (1998) on the set of non-empty graphs is defined by

$$1 - \frac{n(G_1, G_2)}{\max\{n(G_1), n(G_2)\}}.$$

Given an arbitrary set M of graphs, the common subgraph distance d_M on M is defined by

$$\max\{n(G_1), n(G_2)\} - n(G_1, G_2),$$

and the **common supergraph distance** d_M^* on M is defined by

$$N(G_1, G_2) - \min\{n(G_1), n(G_2)\}\$$

for any $G_1, G_2 \in \mathbf{M}$, where $n(G_i)$ is the number of vertices in G_i , i = 1, 2, $n(G_1, G_2)$ is the maximum number of vertices of a common subgraph $G \in \mathbf{M}$ of G_1 and G_2 , and $N(G_1, G_2)$ is the minimum number of vertices of a common supergraph $H \in \mathbf{M}$ of G_1 and G_2 .

 d_M is a metric on **M** if the following condition (1) holds:

(1) if $H \in \mathbf{M}$ is a common supergraph of $G_1, G_2 \in \mathbf{M}$, then there exists a common subgraph $G \in \mathbf{M}$ of G_1 and G_2 with $n(G) \ge n(G_1) + n(G_2) - n(H)$.

 d_M^* is a metric on **M** if the following condition (2) holds:

(2) if $G \in \mathbf{M}$ is a common subgraph of $G_1, G_2 \in \mathbf{M}$, then there exists a common supergraph $H \in \mathbf{M}$ of G_1 and G_2 with $n(H) \leq n(G_1) + n(G_2) - n(G)$.

One has $d_M \leq d_M^*$ if the condition (1) holds, and $d_M \geq d_M^*$ if the condition (2) holds.

The distance d_M is a metric on the set \mathbf{G} of all graphs, the set of all cycle-free graphs, the set of all bipartite graphs, and the set of all trees. The distance d_M^* is a metric on the set \mathbf{G} of all graphs, the set of all connected graphs, the set of all connected bipartite graphs, and the set of all trees. The Zelinka distance d_Z coincides with d_M and d_M^* on the set \mathbf{G} of all graphs. On the set \mathbf{T} of all trees the distances d_M and d_M^* are identical, but different from the Zelinka distance.

The Zelinka distance d_Z is a metric on the set $\mathbf{G}(n)$ of all graphs with n vertices, and is equal to n-k or to K-n for all $G_1, G_2 \in \mathbf{G}(n)$, where k is the maximum number of vertices of a common subgraph of G_1 and G_2 , and K is the minimum number of vertices of a common supergraph of G_1 and G_2 . On the set $\mathbf{T}(n)$ of all trees with n vertices the distance d_Z is called the **Zelinka tree distance** (see, for example, [Zeli75]).

• Fernández-Valiente metric

Given graphs G and H, let $G_1 = (V_1, E_1)$ and $G_2 = (V_2, E_2)$ be their maximum common subgraph and minimum common supergraph; cf. subgraph—supergraph distances.

The **Fernández–Valiente metric** (2001) between graphs G and H is defined by

$$(|V_2| + |E_2|) - (|V_1| + |E_1|).$$

• Editing graph metric

The **editing graph metric** (Axenovich, Kézdy and Martin 2008) between graphs $G_1 = (V_1, E_1)$ and $G_2 = (V_2, E_2)$ with the same number of vertices is defined by

$$\min_{G_3} |E_1 \Delta E_3|,$$

where $G_3 = (V_3, E_3)$ is any graph isomorphic to G_2 . It is the minimum number of edge deletions or additions (cf. the **indel metric** in Chap. 11) needed to transform G_1 into a graph isomorphic to G_2 . It corresponds to the Hamming distance between the adjacency matrices of G_1 and G_2 .

Bunke (1997) defined the **graph edit distance** between vertex- and edge-labeled graphs G_1 and G_2 as the minimal total cost of matching G_1 and G_2 , using deletions, additions and substitutions of vertices and edges. Cf. also **tree**, **top-down**, **unit cost** and **restricted edit distance** between rooted trees.

Myers, Wilson and Hancock (2000) defined the **Bayesian graph edit** distance between two *relational graphs* (i.e., triples (V, E, A), where V, E and A are the sets of vertices, edges and *vertex-attributes*) as their graph edit distance with costs defined by probabilities of operations along an editing path seen as a memoryless error process. Cf. transduction edit distances (Chap. 11) and **Bayesian distance** (Chap. 14).

• Edge distance

The edge distance is a distance on the set G of all graphs, defined by

$$|E_1| + |E_2| - 2|E_{12}| + ||V_1| - |V_2||$$

for any graphs $G_1 = (V_1, E_1)$ and $G_2 = (V_2, E_2)$, where $G_{12} = (V_{12}, E_{12})$ is a common subgraph of G_1 and G_2 with maximal number of edges. This distance has many applications in Organic and Medical Chemistry.

• Contraction distance

The **contraction distance** is a distance on the set $\mathbf{G}(n)$ of all graphs with n vertices, defined by

$$n-k$$

for any $G_1, G_2 \in \mathbf{G}(n)$, where k is the maximum number of vertices of a graph which is isomorphic simultaneously to a graph, obtained from each of G_1 and G_2 by a finite number of *edge contractions*.

To perform the *contraction* of the edge $uv \in E$ of a graph G = (V, E) means to replace u and v by one vertex that is adjacent to all vertices of $V \setminus \{u, v\}$ which were adjacent to u or to v.

• Edge move distance

The **edge move distance** is a metric on the set $\mathbf{G}(n,m)$ of all graphs with n vertices and m edges, defined, for any $G_1, G_2 \in \mathbf{G}(m,n)$, as the minimum number of *edge moves* necessary for transforming the graph G_1 into the graph G_2 . It is equal to m-k, where k is the maximum size of a common subgraph of G_1 and G_2 .

An edge move is one of the edge transformations, defined as follows: H can be obtained from G by an edge move if there exist (not necessarily distinct) vertices u, v, w, and x in G such that $uv \in E(G)$, $wx \notin E(G)$, and H = G - uv + wx.

• Edge jump distance

The **edge jump distance** is an extended metric (which in general can take the value ∞) on the set $\mathbf{G}(n,m)$ of all graphs with n vertices and m edges, defined, for any $G_1, G_2 \in \mathbf{G}(m,n)$, as the minimum number of *edge jumps* necessary for transforming G_1 into G_2 .

An edge jump is one of the edge transformations, defined as follows: H can be obtained from G by an edge jump if there exist four distinct vertices u, v, w, and x in G, such that $uv \in E(G)$, $wx \notin E(G)$, and H = G - av + wx.

• Edge flipping distance

Let $P = \{v_1, ..., v_n\}$ be a collection of points on the plane. A triangulation T of P is a partition of the convex hull of P into a set of triangles such that each triangle has a disjoint interior and the vertices of each triangle are points of P.

The **edge flipping distance** is a distance on the set of all triangulations of P, defined, for any triangulations T and T_1 , as the minimum number of edge flippings necessary for transforming T into T_1 .

An edge e of T is called *flippable* if it is the boundary of two triangles t and t' of T, and $C = t \cup t'$ is a convex quadrilateral. The *flipping* e is one of the *edge transformations*, which consists of removing e and replacing it by the other diagonal of C. The edge flipping is an special case of *edge* jump.

Edge flipping distance can be extended on pseudo-triangulations, i.e., partitions of the convex hull of P into a set of disjoint interior pseudo-triangles (simply connected subsets of the plane that lie between any three mutually tangent convex sets) whose vertices are given points.

• Edge rotation distance

The edge rotation distance is a metric on the set $\mathbf{G}(n, m)$ of all graphs with n vertices and m edges, defined, for any $G_1, G_2 \in \mathbf{G}(m, n)$, as the minimum number of edge rotations necessary for transforming G_1 into G_2 .

An edge rotation is one of the edge transformations, defined as follows: H can be obtained from G by an edge rotation if there exist distinct vertices u, v, and w in G, such that $uv \in E(G)$, $uw \notin E(G)$, and H = G - uv + uw.

• Tree edge rotation distance

The **tree edge rotation distance** is a metric on the set $\mathbf{T}(n)$ of all trees with n vertices, defined, for all $T_1, T_2 \in \mathbf{T}(n)$, as the minimum number of *tree edge rotations* necessary for transforming T_1 into T_2 . For $\mathbf{T}(n)$ the tree edge rotation distance and the edge rotation distance may differ.

A $tree\ edge\ rotation$ is an $edge\ rotation$ performed on a tree, and resulting in a tree.

• Edge shift distance

The edge shift distance (or edge slide distance) is a metric on the set $\mathbf{G}_c(n,m)$ of all connected graphs with n vertices and m edges, defined, for any $G_1, G_2 \in \mathbf{G}_c(m,n)$, as the minimum number of edge shifts necessary for transforming G_1 into G_2 .

An edge shift is one of the edge transformations, defined as follows: H can be obtained from G by an edge shift if there exist distinct vertices u, v, and w in G such that $uv, vw \in E(G)$, $uw \notin E(G)$, and H = G - uv + uw. Edge shift is a special kind of edge rotation in the case when the vertices v, w are adjacent in G.

The edge shift distance can be defined between any graphs G and H with components $G_i (1 \le i \le k)$ and $H_i (1 \le i \le k)$, respectively, such that G_i and H_i have the same order and the same size.

• F-rotation distance

The F-rotation distance is a distance on the set $\mathbf{G}_F(n,m)$ of all graphs with n vertices and m edges, containing a subgraph isomorphic to a given graph F of order at least 2, defined, for all $G_1, G_2 \in \mathbf{G}_F(m,n)$, as the minimum number of F-rotations necessary for transforming G_1 into G_2 .

An F-rotation is one of the edge transformations, defined as follows: let F' be a subgraph of a graph G, isomorphic to F, let u, v, w be three distinct vertices of the graph G such that $u \notin V(F')$, $v, w \in V(F')$, $uv \in E(G)$, and $uw \notin E(G)$; H can be obtained from G by the F-rotation of the edge uv into the position uw if H = G - uv + uw.

• Binary relation distance

Let R be a non-reflexive binary relation between graphs, i.e., $R \subset \mathbf{G} \times \mathbf{G}$, and there exists $G \in \mathbf{G}$ such that $(G, G) \notin R$.

The **binary relation distance** is an extended metric (which in general can take the value ∞) on the set **G** of all graphs, defined, for any graphs G_1 and G_2 , as the minimum number of *R-transformations* necessary for transforming G_1 into G_2 . We say, that a graph H can be obtained from a graph G by an *R-transformation* if $(H, G) \in R$.

An example is the distance between two triangular embeddings of a complete graph (i.e., its cellular embeddings in a surface with only 3-gonal faces) defined as the minimal number t such that, up to replacing t faces, the embeddings are isomorphic.

Crossing-free transformation metrics

Given a subset S of \mathbb{R}^2 , a non-crossing spanning tree of S is a tree whose vertices are points of S, and edges are pairwise non-crossing straight line segments.

The **crossing-free edge move metric** (see [AAH00]) on the set \mathbf{T}_S of all non-crossing spanning trees of a set S, is defined, for any $T_1, T_2 \in \mathbf{T}_S$, as the minimum number of *crossing-free edge moves* needed to transform T_1 into T_2 . A *crossing-free edge move* is an *edge transformation* which consists of adding some edge e in $T \in \mathbf{T}_S$ and removing some edge f from the induced cycle so that e and f do not cross.

The **crossing-free edge slide metric** is a metric on the set \mathbf{T}_S of all non-crossing spanning trees of a set S, defined, for any $T_1, T_2 \in \mathbf{T}_S$, as the minimum number of crossing-free edge slides necessary for transforming T_1 into T_2 . A crossing-free edge slide is one of the edge transformations which consists of taking some edge e in $T \in \mathbf{T}_S$ and moving one of its endpoints along some edge adjacent to e in T, without introducing edge crossings and without sweeping across points in S (that gives a new edge f instead of e). The edge slide is a special kind of crossing-free edge move: the new tree is obtained by closing with f a cycle C of length f in f and removing f from f in such a way that f avoids the interior of the triangle f.

• Traveling salesman tours distances

The Traveling Salesman Problem is the problem of finding the shortest tour that visits a set of cities. We shall consider only Traveling Salesman Problems with undirected links. For an N-city traveling salesman problem, the space \mathcal{T}_N of tours is the set of $\frac{(N-1)!}{2}$ cyclic permutations of the cities $1, 2, \ldots, N$.

The metric D on \mathcal{T}_N is defined in terms of the difference in form: if tours $T, T' \in \mathcal{T}_N$ differ in m links, then D(T, T') = m.

A k-OPT transformation of a tour T is obtained by deleting k links from T, and reconnecting. A tour T', obtained from T by a k-OPT transformation, is called a k-OPT of T. The distance d on the set T_N is defined in terms of the 2-OPT transformations: d(T,T') is the minimal i, for which there exists a sequence of i 2-OPT transformations which transforms T to T'.

In fact, $d(T, T') \leq D(T, T')$ for any $T, T' \in \mathcal{T}_N$ (see, for example, [MaMo95]).

Cf. arc routing problems.

• Orientation distance

The **orientation distance** (Chartrand, Erwin, Raines and Zhang 2001) between two orientations D and D' of a finite graph is the minimum number of arcs of D whose directions must be reversed to produce an orientation isomorphic to D'.

• Subgraphs distances

The standard distance on the set of all subgraphs of a connected graph G = (V, E) is defined by

$$\min\{d_{path}(u,v): u \in V(F), v \in V(H)\}$$

for any subgraphs F, H of G. For any subgraphs F, H of a strongly connected digraph D = (V, E), the standard quasi-distance is defined by

$$\min\{d_{dpath}(u,v): u \in V(F), v \in V(H)\}.$$

Using standard transformations (rotation, shift, etc.) on the edge-set of a graph, one gets corresponding distances between its edge-induced subgraphs with given size, which are subcases of similar distances on the set of all graphs with a given size and order.

The **edge rotation distance** on the set $\mathbf{S}^k(G)$ of all edge-induced subgraphs with k edges in a connected graph G is defined as the minimum number of *edge rotations* required to transform $F \in \mathbf{S}^k(G)$ into $H \in \mathbf{S}^k(G)$. We say that H can be obtained from F by an *edge rotation* if there exist distinct vertices u, v, and w in G such that $uv \in E(F)$, $uw \in E(G) \setminus E(F)$, and H = F - uv + uw.

The edge shift distance on the set $S^k(G)$ of all edge-induced subgraphs with k edges in a connected graph G is defined as the minimum

number of edge shifts required to transform $F \in \mathbf{S}^k(G)$ into $H \in \mathbf{S}^k(G)$. We say that H can be obtained from F by an edge shift if there exist distinct vertices u, v and w in G such that $uv, vw \in E(F), uw \in E(G) \setminus E(F)$, and H = F - uv + uw.

The edge move distance on the set $\mathbf{S}^k(G)$ of all edge-induced subgraphs with k edges of a graph G (not necessary connected) is defined as the minimum number of edge moves required to transform $F \in \mathbf{S}^k(G)$ into $H \in \mathbf{S}^k(G)$. We say that H can be obtained from F by an edge move if there exist (not necessarily distinct) vertices u, v, w, and x in G such that $uv \in E(F)$, $wx \in E(G) \setminus E(F)$, and H = F - uv + wx. The edge move distance is a metric on $\mathbf{S}^k(G)$. If F and H have s edges in common, then it is equal to k-s.

The edge jump distance (which in general can take the value ∞) on the set $\mathbf{S}^k(G)$ of all edge-induced subgraphs with k edges of a graph G (not necessary connected) is defined as the minimum number of edge jumps required to transform $F \in \mathbf{S}^k(G)$ into $H \in \mathbf{S}^k(G)$. We say that H can be obtained from F by an edge jump if there exist four distinct vertices u, v, w, and x in G such that $uv \in E(F)$, $wx \in E(G) \setminus E(F)$, and H = F - uv + wx.

15.4 Distances on trees

Let T be a rooted tree, i.e., a tree with one of its vertices being chosen as the root. The depth of a vertex v, depth(v), is the number of edges on the path from v to the root. A vertex v is called a parent of a vertex u, v = par(u), if they are adjacent, and depth(u) = depth(v) + 1; in this case u is called a child of v. A leaf is a vertex without child. Two vertices are siblings if they have the same parent. The in-degree of a vertex is the number of its children. T(v) is the subtree of T, rooted at a node $v \in V(T)$. If $w \in V(T(v))$, then v is an ancestor of w, and w is a descendant of v; nca(u, v) is the nearest common ancestor of the vertices u and v. T is called a labeled tree if a symbol from a fixed finite alphabet \mathcal{A} is assigned to each node. T is called an ordered tree if a left-to-right order among siblings in T is given.

On the set \mathbb{T}_{rlo} of all rooted labeled ordered trees there are three *editing* operations:

- Relabel change the label of a vertex v
- Deletion delete a non-rooted vertex v with parent v', making the children of v become the children of v'; the children are inserted in the place of v as a subsequence in the left-to-right order of the children of v'
- Insertion the complement of deletion; insert a vertex v as a child of a v' making v the parent of a consecutive subsequence of the children of v'

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For unordered trees the editing operations can be defined similarly, but insert and delete operations work on a subset instead of a subsequence.

We assume that there is a $cost\ function$ defined on each editing operation, and the cost of a sequence of editing operations is the sum of the costs of these operations.

The ordered edit distance mapping is a representation of the editing operations. Formally, define the triple (M, T_1, T_2) to be an ordered edit distance mapping from T_1 to T_2 , $T_1, T_2 \in \mathbb{T}_{rlo}$, if $M \subset V(T_1) \times V(T_2)$ and, for any (v_1, w_1) , $(v_2, w_2) \in M$, the following conditions hold: $v_1 = v_2$ if and only if $w_1 = w_2$ (one-to-one condition), v_1 is an ancestor of v_2 if and only if w_1 is an ancestor of w_2 (ancestor condition), v_1 is to the left of v_2 if and only if w_1 is to the left of w_2 (sibling condition).

We say that a vertex v in T_1 and T_2 is touched by a line in M if v occurs in some pair in M. Let N_1 and N_2 be the set of vertices in T_1 and T_2 , respectively, not touched by any line in M. The cost of M is given by $\gamma(M) = \sum_{(v,w)\in M} \gamma(v\to w) + \sum_{v\in N_1} \gamma(v\to \lambda) + \sum_{w\in N_2} \gamma(\lambda\to w)$, where $\gamma(a\to b) = \gamma(a,b)$ is the cost of an editing operation $a\to b$ which is a relabel if $a,b\in \mathcal{A}$, a deletion if $b=\lambda$, and an insertion if $a=\lambda$. Here $\lambda\not\in\mathcal{A}$ is a special blank symbol, and γ is a metric on the set $\mathcal{A}\cup\lambda$ (excepting the value $\gamma(\lambda,\lambda)$).

• Tree edit distance

The **tree edit distance** (see [Tai79]) on the set \mathbb{T}_{rlo} of all rooted labeled ordered trees is defined, for any $T_1, T_2 \in \mathbb{T}_{rlo}$, as the minimum cost of a sequence of editing operations (relabels, insertions, and deletions) turning T_1 into T_2 .

In terms of ordered edit distance mappings, it is equal to $\min_{(M,T_1,T_2)} \gamma(M)$, where the minimum is taken over all ordered edit distance mappings (M,T_1,T_2) .

The edit tree distance can be defined in similar way on the set of all rooted labeled unordered trees.

• Selkow distance

The Selkow distance (or top-down edit distance, degree-1 edit distance) is a distance on the set \mathbb{T}_{rlo} of all rooted labeled ordered trees, defined, for any $T_1, T_2 \in \mathbb{T}_{rlo}$, as the minimum cost of a sequence of editing operations (relabels, insertions, and deletions) turning T_1 into T_2 if insertions and deletions are restricted to leaves of the trees (see [Selk77]). The root of T_1 must be mapped to the root of T_2 , and if a node v is to be deleted (inserted), then any subtree rooted at v is to be deleted (inserted).

In terms of ordered edit distance mappings, it is equal to $\min_{(M,T_1,T_2)} \gamma(M)$, where the minimum is taken over all ordered edit distance mappings (M,T_1,T_2) satisfying the following condition: if $(v,w) \in M$, where neither v nor w is the root, then $(par(v),par(w)) \in M$.

• Restricted edit distance

The **restricted edit distance** is a distance on the set \mathbb{T}_{rlo} of all rooted labeled ordered trees, defined, for any $T_1, T_2 \in \mathbb{T}_{rlo}$, as the minimum cost

of a sequence of editing operations (relabels, insertions, and deletions) turning T_1 into T_2 with the restriction that disjoint subtrees should be mapped to disjoint subtrees.

In terms of ordered edit distance mappings, it is equal to $\min_{(M,T_1,T_2)} \gamma(M)$, where the minimum is taken over all ordered edit distance mappings (M,T_1,T_2) satisfying the following condition: for all (v_1,w_1) , (v_2,w_2) , $(v_3,w_3) \in M$, $nca(v_1,v_2)$ is a proper ancestor of v_3 if and only if $nca(w_1,w_2)$ is a proper ancestor of w_3 .

This distance is equivalent to the structure respecting edit distance, defined by $\min_{(M,T_1,T_2)} \gamma(M)$, where the minimum is taken over all ordered edit distance mappings (M,T_1,T_2) , satisfying the following condition:

for all (v_1, w_1) , (v_2, w_2) , $(v_3, w_3) \in M$, such that none of v_1, v_2 , and v_3 is an ancestor of the others, $nca(v_1, v_2) = nca(v_1, v_3)$ if and only if $nca(w_1, w_2) = nca(w_1, w_3)$.

Cf. constrained edit distance in Chap. 11.

• Unit cost edit distance

The unit cost edit distance is a distance on the set \mathbb{T}_{rlo} of all rooted labeled ordered trees, defined, for any $T_1, T_2 \in \mathbb{T}_{rlo}$, as the minimum number of editing operations (relabels, insertions, and deletions) turning T_1 into T_2 .

• Alignment distance

The alignment distance (see [JWZ94]) is a distance on the set \mathbb{T}_{rlo} of all rooted labeled ordered trees, defined, for any $T_1, T_2 \in \mathbb{T}_{rlo}$, as the minimum *cost* of an *alignment* of T_1 and T_2 . It corresponds to a restricted edit distance, where all insertions must be performed before any deletions.

Thus, one inserts spaces, i.e., vertices labeled with a $blank \ symbol \ \lambda$, into T_1 and T_2 so that they become isomorphic when labels are ignored; the resulting trees are overlayed on top of each other giving the $alignment \ T_A$ which is a tree, where each vertex is labeled by a pair of labels. The cost of T_A is the sum of the costs of all pairs of opposite labels in T_A .

• Splitting-merging distance

The splitting-merging distance (see [ChLu85]) is a distance on the set \mathbb{T}_{rlo} of all rooted labeled ordered trees, defined, for any $T_1, T_2 \in \mathbb{T}_{rlo}$, as the minimum number of vertex splittings and mergings needed to transform T_1 into T_2 .

• Degree-2 distance

The **degree-2 distance** is a metric on the set \mathbb{T}_l of all labeled trees (labeled free trees), defined, for any $T_1, T_2 \in \mathbb{T}_l$, as the minimum number of editing operations (relabels, insertions, and deletions) turning T_1 into T_2 if any vertex to be inserted (deleted) has no more than two neighbors. This metric is a natural extension of the **tree edit distance** and the **Selkow distance**.

A phylogenetic X-tree is an unordered, unrooted tree with the labeled leaf set X and no vertices of degree two. If every interior vertex has degree three, the tree is called binary (or $fully\ resolved$).

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• Robinson-Foulds metric

A $cut\ A|B$ of X is a partition of X into two subsets A and B (see cut semi-metric). Removing an edge e from a phylogenetic X-tree induces a cut of the leaf set X which is called the cut associated with e.

The Robinson–Foulds metric (or Bourque metric, bipartition distance) is a metric on the set $\mathbb{T}(X)$ of all phylogenetic X-trees, defined, for all $T_1, T_2 \in \mathbb{T}(X)$, by

$$\frac{1}{2}|\Sigma(T_1)\Delta\Sigma(T_2)| = \frac{1}{2}|\Sigma(T_1) - \Sigma(T_2)| + \frac{1}{2}|\Sigma(T_2) - \Sigma(T_1)|,$$

where $\Sigma(T)$ is the collection of all cuts of X associated with edges of T.

The Robinson–Foulds weighted metric is a metric on the set $\mathbb{T}(X)$ of all phylogenetic X-trees, defined by

$$\sum_{A|B\in\Sigma(T_1)\cup\Sigma(T_2)} |w_1(A|B) - w_2(A|B)|$$

for all $T_1, T_2 \in \mathbb{T}(X)$, where $w_i = (w(e))_{e \in E(T_i)}$ is the collection of positive weights, associated with the edges of the X-tree T_i , $\Sigma(T_i)$ is the collection of all cuts of X, associated with edges of T_i , and $w_i(A|B)$ is the weight of the edge, corresponding to the cut A|B of X, i = 1, 2.

Cf. more general cut norm metric in Chap. 12 and rectangle distance on weighted graphs.

• μ-metric

Given a phylogenetic X-tree T with n leaves and a vertex v in it, let $\mu(v) = (\mu_1(v), \ldots, \mu_n(v))$, where $\mu_i(v)$ is the number of different paths from the vertex v to the i-th leaf. Let $\mu(T)$ denote the multiset on the vertex-set of T with $\mu(v)$ being the multiplicity of the vertex v.

The μ -metric (Cardona, Roselló and Valiente 2008) is a metric on the set $\mathbb{T}(X)$ of all phylogenetic X-trees, defined, for all $T_1, T_2 \in \mathbb{T}(X)$, by

$$\frac{1}{2}|\mu(T_1)\Delta\mu(T_2)|,$$

where Δ denotes the symmetric difference of multisets. Cf. **metrics** between multisets in Chap. 1 and **Dodge–Shiode WebX quasi-distance** in Chap. 22.

• Nearest neighbor interchange metric

The nearest neighbor interchange metric (or crossover metric) is a metric on the set $\mathbb{T}(X)$ of all phylogenetic X-trees, defined, for all $T_1, T_2 \in \mathbb{T}(X)$, as the minimum number of nearest neighbor interchanges required to transform T_1 into T_2 .

A nearest neighbor interchange consists of swapping two subtrees in a tree that are adjacent to the same internal edge; the remainder of the tree is unchanged.

• Subtree prune and regraft distance

The subtree prune and regraft distance is a metric on the set $\mathbb{T}(X)$ of all phylogenetic X-trees, defined, for all $T_1, T_2 \in \mathbb{T}(X)$, as the minimum number of subtree prune and regraft transformations required to transform T_1 into T_2 .

A subtree prune and regraft transformation proceeds in three steps: one selects and removes an edge uv of the tree, thereby dividing the tree into two subtrees T_u (containing u) and T_v (containing v); then one selects and subdivides an edge of T_v , giving a new vertex w; finally, one connects u and w by an edge, and removes all vertices of degree two.

• Tree bisection-reconnection metric

The tree bisection-reconnection metric (or TBR-metric) on the set $\mathbb{T}(X)$ of all phylogenetic X-trees is defined, for all $T_1, T_2 \in \mathbb{T}(X)$, as the minimum number of tree bisection and reconnection transformations required to transform T_1 into T_2 .

A tree bisection and reconnection transformation proceeds in three steps: one selects and removes an edge uv of the tree, thereby dividing the tree into two subtrees T_u (containing u) and T_v (containing v); then one selects and subdivides an edge of T_v , giving a new vertex w, and an edge of T_u , giving a new vertex z; finally one connects w and z by an edge, and removes all vertices of degree two.

• Quartet distance

The quartet distance (see [EMM85]) is a distance of the set $\mathbb{T}_b(X)$ of all binary phylogenetic X-trees, defined, for all $T_1, T_2 \in \mathbb{T}_b(X)$, as the number of mismatched quartets (from the total number $\binom{n}{4}$ possible quartets) for T_1 and T_2 .

This distance is based on the fact that, given four leaves $\{1, 2, 3, 4\}$ of a tree, they can only be combined in a binary subtree in three different ways: (12|34), (13|24), or (14|23): a notation (12|34) refers to the binary tree with the leaf set $\{1, 2, 3, 4\}$ in which removing the inner edge yields the trees with the leaf sets $\{1, 2\}$ and $\{3, 4\}$.

• Triples distance

The **triples distance** (see [CPQ96]) is a distance of the set $\mathbb{T}_b(X)$ of all binary phylogenetic X-trees, defined, for all $T_1, T_2 \in \mathbb{T}_b(X)$, as the number of triples (from the total number $\binom{n}{3}$ possible triples) that differ (for example, by which leaf is the outlier) for T_1 and T_2 .

• Perfect matching distance

The **perfect matching distance** is a distance on the set $\mathbb{T}_{br}(X)$ of all rooted binary phylogenetic X-trees with the set X of n labeled leaves, defined, for any $T_1, T_2 \in \mathbb{T}_{br}(X)$, as the minimum number of interchanges necessary to bring the perfect matching of T_1 to the perfect matching of T_2 .

Given a set $A = \{1, ..., 2k\}$ of 2k points, a perfect matching of A is a partition of A into k pairs. A rooted binary phylogenetic tree with n labeled leaves has a root and n-2 internal vertices distinct from the root. It can be identified with a perfect matching on 2n-2, different from the

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root, vertices by following construction: label the internal vertices with numbers $n+1,\ldots,2n-2$ by putting the smallest available label as the parent of the pair of labeled children of which one has the smallest label among pairs of labeled children; now a matching is formed by peeling off the children, or sibling pairs, two by two.

• Tree rotation distance

The **tree rotation distance** is a distance on the set \mathbf{T}_n of all rooted ordered binary trees with n interior vertices, defined, for all $T_1, T_2 \in \mathbf{T}_n$, as the minimum number of *rotations*, required to transform T_1 into T_2 .

Given interior edges uv, vv', vv'' and uw of a binary tree, the rotation is replacing them by edges uv, uv'', vv' and vw.

There is a bijection between edge flipping operations in triangulations of convex polygons with n+2 vertices and rotations in binary trees with n interior vertices.

• Attributed tree metrics

An attributed tree is a triple (V, E, α) , where T = (V, E) is the underlying tree, and α is a function which assigns an attribute vector $\alpha(v)$ to every vertex $v \in V$. Given two attributed trees (V_1, E_1, α) and (V_2, E_2, β) , consider the set of all subtree isomorphisms between them, i.e., the set of all isomorphisms $f: H_1 \to H_2, H_1 \subset V_1, H_2 \subset V_2$, between their induced subtrees.

Given a similarity s on the set of attributes, the similarity between isomorphic induced subtrees is defined as $W_s(f) = \sum_{v \in H_1} s(\alpha(v), \beta(f(v)))$. Let ϕ be the isomorphism with maximal similarity $W_s(\phi) = W(\phi)$.

The following semi-metrics on the set \mathbf{T}_{att} of all attributed trees are used:

$$\begin{split} &1. \ \max\{|V_1|,|V_2|\} - W(\phi) \\ &2. \ |V_1| + |V_2| - 2W(\phi) \\ &3. \ 1 - \frac{W(\phi)}{\max\{|V_1|,|V_2|\}} \\ &4. \ 1 - \frac{W(\phi)}{|V_1| + |V_2| - W(\phi)} \end{split}$$

They become metrics on the set of equivalences classes of attributed trees: two attributed trees (V_1, E_1, α) and (V_2, E_2, β) are called *equivalent* if they are *attribute-isomorphic*, i.e., if there exists an isomorphism $g: V_1 \to V_2$ between the trees T_1 and T_2 such that, for any $v \in V_1$, we have $\alpha(v) = \beta(g(v))$. Then $|V_1| = |V_2| = W(g)$.

• Greatest agreement subtree distance

The greatest agreement subtree distance is a distance of the set **T** of all trees, defined, for all $T_1, T_2 \in \mathbf{T}$, as the minimum number of leaves removed to obtain a *(greatest) agreement subtree.*

An agreement subtree (or common pruned tree) of two trees is an identical subtree that can be obtained from both trees by pruning leaves with the same label.

Chapter 16 Distances in Coding Theory

Coding Theory deals with the design and properties of error-correcting codes for the reliable transmission of information across noisy channels in transmission lines and storage devices. The aim of Coding Theory is to find codes which transmit and decode fast, contain many valid code words, and can correct, or at least detect, many errors. These aims are mutually exclusive, however; so, each application has its own good code.

In communications, a *code* is a rule for converting a piece of information (for example, a letter, word, or phrase) into another form or representation, not necessarily of the same sort. *Encoding* is the process by which a source (object) performs this conversion of information into data, which is then sent to a receiver (observer), such as a data processing system. *Decoding* is the reverse process of converting data, which has been sent by a source, into information understandable by a receiver.

An error-correcting code is a code in which every data signal conforms to specific rules of construction so that departures from this construction in the received signal can generally be automatically detected and corrected. It is used in computer data storage, for example in dynamic RAM, and in data transmission. Error detection is much simpler than error correction, and one or more "check" digits are commonly embedded in credit card numbers in order to detect mistakes. The two main classes of error-correcting codes are block codes, and convolutional codes.

A block code (or uniform code) of length n over an alphabet \mathcal{A} , usually, over a finite field $\mathbb{F}_q = \{0, \dots, q-1\}$, is a subset $C \subset \mathcal{A}^n$; every vector $x \in C$ is called a codeword, and M = |C| is called size of the code. Given a metric d on \mathbb{F}_q^n (for example, the **Hamming metric**, **Lee metric**, **Levenstein metric**), the value $d^* = d^*(C) = \min_{x,y \in C, x \neq y} d(x,y)$ is called the **minimum distance** of the code C. The weight w(x) of a codeword $x \in C$ is defined as w(x) = d(x,0). An (n,M,d^*) -code is a q-ary block code of length q, size q, and minimum distance q^* . A binary code is a code over q.

When codewords are chosen such that the distance between them is maximized, the code is called *error-correcting*, since slightly garbled vectors can be recovered by choosing the nearest codeword. A code C is a t-error-correcting code (and a 2t-error-detecting code) if $d^*(C) \geq 2t + 1$. In this case each

neighborhood $U_t(x) = \{y \in C : d(x,y) \leq t\}$ of $x \in C$ is disjoint with $U_t(y)$ for any $y \in C, y \neq x$. A perfect code is a q-ary (n, M, 2t + 1)-code for which the M spheres $U_t(x)$ of radius t centered on the codewords fill the whole space \mathbb{F}_q^n completely, without overlapping.

A block code $C \subset \mathbb{F}_q^n$ is called *linear* if C is a vector subspace of \mathbb{F}_q^n . An [n,k]-code is a k-dimensional linear code $C \subset \mathbb{F}_q^n$ (with the minimum distance d^*); it has size q^k , i.e., it is an (n,q^k,d^*) -code. The *Hamming code* is the linear perfect one-error correcting $(\frac{q^r-1}{q-1},\frac{q^r-1}{q-1}-r,3)$ -code.

A $k \times n$ matrix G with rows that are basis vectors for a linear [n,k]-code C is called a generator matrix of C. In standard form it can be written as $(1_k|A)$, where 1_k is the $k \times k$ identity matrix. Each message (or information symbol, source symbol) $u = (u_1, \ldots, u_k) \in \mathbb{F}_q^k$ can be encoded by multiplying it (on the right) by the generator matrix: $uG \in C$. The matrix $H = (-A^T|1_{n-k})$ is called the parity-check matrix of C. The number r = n-k corresponds to the number of parity check digits in the code, and is called the redundancy of the code C. The information rate (or code rate) of a code C is the number $R = \frac{\log_2 M}{n}$. For a q-ary [n,k]-code, $R = \frac{k}{n}\log_2 q$; for a binary [n,k]-code, $R = \frac{k}{n}$

A convolutional code is a type of error-correction code in which each k-bit information symbol to be encoded is transformed into an n-bit codeword, where $R = \frac{k}{n}$ is the code rate $(n \ge k)$, and the transformation is a function of the last m information symbols, where m is the constraint length of the code. Convolutional codes are often used to improve the performance of radio and satellite links. A variable length code is a code with codewords of different lengths.

In contrast to error-correcting codes which are designed only to increase the reliability of data communications, cryptographic codes are designed to increase their security. In Cryptography, the sender uses a key to encrypt a message before it is sent through an insecure channel, and an authorized receiver at the other end then uses a key to decrypt the received data to a message. Often, data compression algorithms and error-correcting codes are used in tandem with cryptographic codes to yield communications that are efficient, robust to data transmission errors, and secure to eavesdropping and tampering. Encrypted messages which are, moreover, hidden in text, image, etc., are called steganographic messages.

16.1 Minimum distance and relatives

• Minimum distance

Given a code $C \subset V$, where V is an n-dimensional vector space equipped with a metric d, the **minimum distance** $d^* = d^*(C)$ of the code C is defined by

$$\min_{x,y \in C, x \neq y} d(x,y).$$

The metric d depends on the nature of the errors for the correction of which the code is intended. For a prescribed correcting capacity it is necessary to use codes with a maximum number of codewords. The most widely investigated such codes are the q-ary block codes in the **Hamming metric** $d_H(x,y) = |\{i : x_i \neq y_i, i = 1, \dots, n\}|$.

For a linear code C the minimum distance $d^*(C) = w(C)$, where $w(C) = \min\{w(x) : x \in C\}$ is a minimum weight of the code C. As there are $rank(H) \leq n - k$ independent columns in the parity check matrix H of an [n,k]-code C, then $d^*(C) \leq n - k + 1$ (Singleton upper bound).

• Dual distance

The dual distance d^{\perp} of a linear [n,k]-code $C \subset \mathbb{F}_q^n$ is the minimum distance of the dual code C^{\perp} of C.

The dual code C^{\perp} of C is defined as the set of all vectors of \mathbb{F}_q^n that are orthogonal to every codeword of C: $C^{\perp} = \{v \in \mathbb{F}_q^n : \langle v, u \rangle = 0 \text{ for any } u \in C\}$. The code C^{\perp} is a linear [n, n-k]-code. The $(n-k) \times n$ generator matrix of C^{\perp} is the parity-check matrix of C.

• Bar product distance

Given linear codes C_1 and C_2 of length n with $C_2 \subset C_1$, their bar product $C_1|C_2$ is a linear code of length 2n, defined by $C_1|C_2 = \{x|x+y: x \in C_1, y \in C_2\}$.

The **bar product distance** is the minimum distance $d^*(C_1|C_2)$ of the bar product $C_1|C_2$.

• Design distance

A linear code is called a cyclic code if all cyclic shifts of a codeword also belong to C, i.e., if for any $(a_0, \ldots, a_{n-1}) \in C$ the vector $(a_{n-1}, a_0, \ldots, a_{n-2}) \in C$. It is convenient to identify a codeword (a_0, \ldots, a_{n-1}) with the polynomial $c(x) = a_0 + a_1x + \cdots + a_{n-1}x^{n-1}$; then every cyclic [n, k]-code can be represented as the principal ideal $\langle g(x) \rangle = \{r(x)g(x) : r(x) \in R_n\}$ of the ring $R_n = \mathbb{F}_q(x)/(x^n - 1)$, generated by the polynomial $g(x) = g_0 + g_1x + \cdots + x^{n-k}$, called the generator polynomial of the code C.

Given an element α of order n in a finite field \mathbb{F}_{q^s} , a Bose-Chaudhuri-Hocquenghem <math>[n,k]-code of **design distance** d is a cyclic code of length n, generated by a polynomial g(x) in $\mathbb{F}_q(x)$ of degree n-k, that has roots at $\alpha, \alpha^2, \ldots, \alpha^{d-1}$. The minimum distance d^* of such a code of odd design distance d is at least d.

A Reed–Solomon code is a Bose–Chaudhuri–Hocquenghem code with s=1. The generator polynomial of a Reed–Solomon code of design distance d is $g(x)=(x-\alpha)\dots(x-\alpha^{d-1})$ with degree n-k=d-1; that is, for a Reed–Solomon code the design distance d=n-k+1, and the minimum distance $d^*\geq d$. Since, for a linear [n,k]-code the minimum distance $d^*\leq n-k+1$ (Singleton upper bound), a Reed–Solomon code has the minimum distance $d^*=n-k+1$ and achieves the Singleton upper bound. Compact disc players use a double-error correcting (255, 251, 5) Reed–Solomon code over \mathbb{F}_{256} .

• Goppa designed minimum distance

The Goppa designed minimum distance [Gopp71] is a lower bound $d^{\star}(m)$ for the minimum distance of one-point geometric Goppa codes (or algebraic geometry codes) G(m). For G(m), associated to the divisors D and mP, $m \in \mathbb{N}$, of a smooth projective absolutely irreducible algebraic curve of genus g > 0 over a finite field \mathbb{F}_q , one has $d^{\star}(m) = m + 2 - 2g$ if 2g - 2 < m < n.

In fact, for a Goppa code C(m) the structure of the gap sequence at P may allow one to give a better lower bound of the minimum distance (cf. Feng-Rao distance).

• Feng-Rao distance

The Feng–Rao distance $\delta_{FR}(m)$ is a lower bound for the minimum distance of one-point geometric Goppa codes G(m) which is better than the Goppa designed minimum distance. The method of Feng and Rao for encoding the code C(m) decodes errors up to half the Feng–Rao distance $\delta_{FR}(m)$, and gives an improvement of the number of errors that one can correct for one-point geometric Goppa codes.

Formally, the Feng–Rao distance is defined as follows. Let S be a subsemi-group S of $\mathbb{N} \cup \{0\}$ such that the *genus* $g = |\mathbb{N} \cup \{0\} \setminus S|$ of S is finite, and $0 \in S$. The **Feng–Rao distance** on S is a function $\delta_{FR}: S \to \mathbb{N} \cup \{0\}$ such that $\delta_{FR}(m) = \min\{\nu(r): r \geq m, r \in S\}$, where $\nu(r) = |\{(a,b) \in S^2: a+b=r\}|$. The generalized r-th Feng–Rao distance on S is defined by $\delta_{FR}^r(m) = \min\{\nu[m_1, \ldots, m_r]: m \leq m_1 < \cdots < m_r, m_i \in S\}$, where $\nu[m_1, \ldots, m_r] = |\{a \in S: m_i - a \in S \text{ for some } i = 1, \ldots, r\}|$. Then $\delta_{FR}(m) = \delta_{FR}^1(m)$. (See, for example, [FaMu03].)

• Free distance

The free distance is the minimum non-zero $Hamming\ weight$ of any codeword in a $convolutional\ code$ or a $variable\ length\ code$.

Formally, the k-th minimum distance d_k^* of a convolutional code or a variable length code is the smallest Hamming distance between any two initial codeword segments which are k frame long and disagree in the initial frame. The sequence $d_1^*, d_2^*, d_3^*, \ldots$ $(d_1^* \leq d_2^* \leq d_3^* \leq \ldots)$ is called the distance profile of the code. The free distance of a convolutional code or a variable length code is $\max_l d_l^* = \lim_{l \to \infty} d_l^* = d_{\infty}^*$.

• Effective free distance

A turbo code is a long block code in which there are L input bits, and each of these bits is encoded q times. In the j-th encoding, the L bits are sent through a permutation box P_j , and then encoded via an $[N_j, L]$ block encoder (code fragment encoder) which can be thought of as an $L \times N_j$ matrix. The overall turbo code is then a linear $[N_1 + \cdots + N_q, L]$ -code (see, for example, [BGT93]).

The weight-i input minimum distance $d^i(C)$ of a turbo-code C is the minimum weight among codewords corresponding to input words of weight i. The **effective free distance** of C is its weight-2 input minimum distance $d^2(C)$, i.e., the minimum weight among codewords corresponding to input words of weight 2.

• Distance distribution

Given a code C over a finite metric space (X,d) with the diameter diam(X,d) = D, the **distance distribution** of C is a (D+1)-vector (A_0,\ldots,A_D) , where $A_i = \frac{1}{|C|}|\{(c,c') \in C^2 : d(c,c') = i\}|$. That is, one considers $A_i(c)$ as the number of code words at distance i from the codeword c, and takes A_i as the average of $A_i(c)$ over all $c \in C$. $A_0 = 1$ and, if $d^* = d^*(C)$ is the minimum distance of C, then $A_1 = \cdots = A_{d^*-1} = 0$.

The distance distribution of a code with given parameters is important, in particular, for bounding the probability of decoding error under different decoding procedures from maximum likelihood decoding to error detection. It can also be helpful in revealing structural properties of codes and establishing nonexistence of some codes.

• Unicity distance

The unicity distance of a cryptosystem (Shannon 1949) is the minimal length of a cyphertext that is required in order to expect that there exists only one meaningful decryption for it. For classic cryptosystems with fixed key space, the unicity distance is approximated by the formula H(K)/D, where H(K) is the key space entropy (roughly $\log_2 N$, where N is the number of keys), and D measures the redundancy of the plaintext source language in bits per letter.

A cryptosystem offers perfect secrecy if its unicity distance is infinite. For example, the *one-time pads* offer perfect secrecy; they were used for the "red telephone" between the Kremlin and the White House.

More generally, **Pe-security distance** of a cryptosystem (Tilburg and Boekee 1987) is the minimal expected length of cyphertext that is required in order to break the cryptogram with an average error probability of at most Pe.

16.2 Main coding distances

• Arithmetic codes distance

An arithmetic code (or code with correction of arithmetic errors) is a finite subset of the set \mathbb{Z} of integers (usually, non-negative integers). It is intended for the control of the functioning of an adder (a module performing addition). When adding numbers represented in the binary number system, a single slip in the functioning of the adder leads to a change in the result by some power of 2, thus, to a single arithmetic error. Formally, a single arithmetic error on \mathbb{Z} is defined as a transformation of a number $n \in \mathbb{Z}$ to a number $n' = n \pm 2^i$, $i = 1, 2, \ldots$

The **arithmetic codes distance** is a metric on \mathbb{Z} , defined, for any $n_1, n_2 \in \mathbb{Z}$, as the minimum number of arithmetic errors taking n_1 to n_2 . It can be written as $w_2(n_1-n_2)$, where $w_2(n)$ is the arithmetic 2-weight of n, i.e., the smallest possible number of non-zero coefficients in representations

 $n = \sum_{i=0}^{k} e_i 2^i$, where $e_i = 0, \pm 1$, and k is some non-negative integer. In fact, for each n there is a unique such representation with $e_k \neq 0$, $e_i e_{i+1} = 0$ for all $i = 0, \dots, k-1$, which has the smallest number of non-zero coefficients (cf. **arithmetic** r-**norm metric** in Chap. 12).

Sharma–Kaushik distance

Let $q \geq 2$, $m \geq 2$. A partition $\{B_0, B_1, \dots B_{q-1}\}$ of \mathbb{Z}_m is called a Sharma-Kaushik partition if the following conditions hold:

- 1. $B_0 = \{0\};$
- 2. For any $i \in \mathbb{Z}_m$, $i \in \mathbb{B}_s$ if and only if $m i \in \mathbb{B}_s$, $s = 1, 2, \ldots, q 1$;
- 3. If $i \in B_s, j \in B_t$, and s > t, then $\min\{i, m i\} > \min\{j, m j\}$;
- 4. If $s \ge t$, s, t = 0, 1, ..., q 1, then $|B_s| \ge |B_t|$ except for s = q 1 in which case $|B_{q-1}| \ge \frac{1}{2} |B_{q-2}|$.

Given a Sharma–Kaushik partition of \mathbb{Z}_m , the Sharma–Kaushik weight $w_{SK}(x)$ of any element $x \in \mathbb{Z}_m$ is defined by $w_{SK}(x) = i$ if $x \in B_i$, $i \in \{0, 1, \ldots, q-1\}$.

The **Sharma–Kaushik distance** (see, for example, [ShKa97]) is a metric on \mathbb{Z}_m , defined by

$$w_{SK}(x-y)$$
.

The Sharma–Kaushik distance on Z_m^n is defined by $w_{SK}^n(x-y)$ where, for $x=(x_1,\ldots x_n)\in \mathbb{Z}_m^n$, one has $w_{SK}^n(x)=\sum_{i=1}^n w_{SK}(x_i)$.

The **Hamming metric** and the **Lee metric** arise from two specific partitions of the above type: $P_H = \{B_0, B_1\}$, where $B_1 = \{1, 2, \dots, q-1\}$, and $P_L = \{B_0, B_1, \dots, B_{\lfloor q/2 \rfloor}\}$, where $B_i = \{i, m-i\}, i = 1, \dots, \lfloor \frac{q}{2} \rfloor$.

Absolute summation distance

The absolute summation distance (or *Lee distance*) is the **Lee metric** on the set $\mathbb{Z}_m^n = \{0, 1, \dots, m-1\}^n$, defined by

$$w_{Lee}(x-y),$$

where $w_{Lee}(x) = \sum_{i=1}^{n} \min\{x_i, m - x_i\}$ is the Lee weight of $x = (x_1, \ldots, x_n) \in \mathbb{Z}_m^n$.

If \mathbb{Z}_m^n is equipped with the absolute summation distance, then a subset C of \mathbb{Z}_m^n is called a *Lee distance code*. Lee distance codes are used for phase-modulated and multilevel quantized-pulse-modulated channels, and have several applications to the toroidal interconnection networks. The most important Lee distance codes are *negacyclic codes*.

Mannheim distance

Let $\mathbb{Z}[i] = \{a+bi : a, b \in \mathbb{Z}\}$ be the set of Gaussian integers. Let $\pi = a+bi$ (a > b > 0) be a Gaussian prime, i.e., either

- (1) $(a+bi)(a-bi) = a^2+b^2 = p$, where $p \equiv 1 \pmod{4}$ is a prime number, or
- (2) $\pi = p + 0 \cdot i = p$, where $p \equiv 3 \pmod{4}$ is a prime number.

The Mannheim distance is a distance on $\mathbb{Z}[i]$, defined, for any two Gaussian integers x and y, as the sum of the absolute values of the real and imaginary parts of the difference $x-y \pmod{\pi}$. The modulo reduction, before summing the absolute values of the real and imaginary parts, is the difference between the Manhattan metric and the Mannheim distance.

The elements of the finite field $\mathbb{F}_p = \{0, 1, \dots, p-1\}$ for $p \equiv 1 \pmod{4}$, $p = a^2 + b^2$, and the elements of the finite field \mathbb{F}_{p^2} for $p \equiv 3 \pmod{4}$, p = a, can be mapped on a subset of the Gaussian integers using the complex modulo function $\mu(k) = k - [\frac{k(a-bi)}{p}](a+bi)$, $k = 0, \dots, p-1$, where [.] denotes rounding to the closest Gaussian integer. The set of the selected Gaussian integers a+bi with the minimal Galois norms $\sqrt{a^2+b^2}$ is called a constellation. This representation gives a new way to construct codes for two-dimensional signals. Mannheim distance was introduced to make QAM-like signals more susceptible to algebraic decoding methods. For codes over hexagonal signal constellations a similar metric can be introduced over the set of the Eisenstein–Jacobi integers. It is useful for block codes over tori. (See, for example, [Hube93], [Hube94].)

• Generalized Lee metric

Let \mathbb{F}_{p^m} denote the finite field with p^m elements, where p is prime number and $m \geq 1$ is an integer. Let $e_i = (0, \dots, 0, 1, 0, \dots, 0), \ 1 \leq i \leq k$, be the standard basis of \mathbb{Z}^k . Choose elements $a_i \in \mathbb{F}_{p^m}, \ 1 \leq i \leq k$, and the mapping $\phi: \mathbb{Z}^k \to \mathbb{F}_{p^m}$, sending any $x = \sum_{i=1}^k x_i e_i, \ x_i \in \mathbb{Z}^k$, to $\phi(x) = \sum_{i=1}^k a_i x_i \pmod{p}$, so that ϕ is surjective. So, for each $a \in \mathbb{F}_{p^m}$, there exists $x \in \mathbb{Z}^k$ such that $a = \phi(x)$. For each $a \in \mathbb{F}_{p^m}$, its k-dimensional Lee weight is $w_{kL}(a) = \min\{\sum_{i=1}^k |x_i| : x = (x_i) \in \mathbb{Z}, a = \phi(x)\}$.

The **generalized Lee metric** between vectors (a_j) and (b_j) of $\mathbb{F}_{p^m}^n$ is defined (Nishimura and Hiramatsu 2008) by

$$\sum_{j=1}^{n} w_{kL}(a_j - b_j).$$

It is the **Lee metric** (or **absolute summation distance**) if $\phi(e_1) = 1$ while $\phi(e_i) = 0$ for $2 \le i \le k$. It is the **Mannheim distance** if k = 2, $p \equiv 1 \pmod{4}$, $\phi(e_1) = 1$ while $\phi(e_2) = a$ is a solution in \mathbb{F}_p of the quadratic congruence $x^2 \equiv -1 \pmod{p}$.

• Poset distance

Let (V_n, \preceq) be a poset on $V_n = \{1, \ldots, n\}$. A subset I of V_n is called ideal if $x \in I$ and $y \preceq x$ imply that $y \in I$. If $J \subset V_n$, then $\langle J \rangle$ denotes the smallest ideal of V_n which contains J. Consider the vector space \mathbb{F}_q^n over a finite field \mathbb{F}_q . The P-weight of an element $x = (x_1, \ldots, x_n) \in \mathbb{F}_q^n$ is defined as the cardinality of the smallest ideal of V_n containing the support of x: $w_P(x) = |\langle supp(x) \rangle|$, where $supp(x) = \{i : x_i \neq 0\}$.

The **poset distance** (see [BGL95]) is a metric on \mathbb{F}_q^n , defined by

$$w_P(x-y)$$
.

If \mathbb{F}_q^n is equipped with a poset distance, then a subset C of \mathbb{F}_q^n is called a poset code. If V_n forms the chain $1 \leq 2 \leq \cdots \leq n$, then the linear code C of dimension k consisting of all vectors $(0, \ldots, 0, a_{n-k+1}, \ldots, a_n) \in \mathbb{F}_q^n$ is a perfect poset code with the minimum (poset) distance $d_p^n(C) = n - k + 1$.

If V_n forms an antichain, then the poset distance coincides with the **Hamming metric**. If V_n consists of finite disjoint union of chains of equal lengths, then the poset distance coincides with the **Rosenbloom–Tsfasman metric**.

• Rank distance

Let \mathbb{F}_q be a finite field, $\mathbb{K} = \mathbb{F}_{q^m}$ an extension of degree m of \mathbb{F}_q , and $\mathbb{E} = \mathbb{K}^n$ a vector space of dimension n over \mathbb{K} . For any $a = (a_1, \dots a_n) \in \mathbb{E}$ define its rank, rank(a), as the dimension of the vector space over \mathbb{F}_q , generated by $\{a_1, \dots, a_n\}$.

The rank distance is a metric on \mathbb{E} , defined by

$$rank(a-b)$$
.

Since the rank distance between two codewords is at most the Hamming distance between them, for any code $C \subset \mathbb{E}$ its minimum (rank) distance $d_{RK}^*(C) \leq \min\{m, n - \log_{q^m} |C| + 1\}$. A code C with $d_{RK}^*(C) = n - \log_{q^m} |C| + 1$, n < m, is called a Gabidulin code (see [Gabi85]). A code C with $d_{RK}^*(C) = m$, $m \leq n$, is called a full rank distance code. Such a code has at most q^n elements. A maximal full rank distance code is a full rank distance code with q^n elements; it exists if and only if m divides n.

• Gabidulin–Simonis metrics

Let \mathbb{F}_q^n be the vector space over a finite field \mathbb{F}_q and let $F = \{F_i : i \in I\}$ be a finite family of its subsets such that the minimal linear subspace of \mathbb{F}_q^n containing $\bigcup_{i \in I} F_i$ is \mathbb{F}_q^n . Without loss of generality, F can be an antichain of linear subspaces of \mathbb{F}_q^n .

The F-weight w_F of a vector $x = (x_1, \ldots, x_n) \in \mathbb{F}_q^n$ is the smallest |J| over such subsets $J \subset I$ that x belongs to the minimal linear subspace of \mathbb{F}_q^n containing $\bigcup_{i \in J} F_i$. A **Gabidulin–Simonis metric** (or F-distance, see [GaSi98]) on \mathbb{F}_q^n is defined by

$$w_F(x-y)$$
.

The **Hamming metric** corresponds to the case of F_i , $i \in I$, forming the standard basis. The **Vandermonde metric** is F-distance with F_i , $i \in I$, being the columns of a generalized Vandermonde matrix. Among other coding Gabidulin–Simonis metrics are: **rank distance**, b-burst distance, Gabidulin's combinatorial metrics (cf. **poset distance**), etc.

• Rosenbloom-Tsfasman metric

Let $M_{m,n}(\mathbb{F}_q)$ be the set of all $m \times n$ matrices with entries from a finite field \mathbb{F}_q (in general, from any finite alphabet $\mathcal{A} = \{a_1, \ldots, a_q\}$). The $Rosenbloom-Tsfasman\ norm\ ||.||_{RT}$ on $M_{m,n}(\mathbb{F}_q)$ is defined as follows:

if m = 1 and $a = (\xi_1, \xi_2, \dots, \xi_n) \in M_{1,n}(\mathbb{F}_q)$, then $||0_{1,n}||_{RT} = 0$, and $||a||_{RT} = \max\{i : \xi_i \neq 0\}$ for $a \neq 0_{1,n}$; if $A = (a_1, \dots, a_m)^T \in M_{m,n}(\mathbb{F}_q)$, $a_j \in M_{1,n}(\mathbb{F}_q), 1 \leq j \leq m$, then $||A||_{RT} = \sum_{j=1}^m ||a_j||_{RT}$.

The Rosenbloom-Tsfasman metric [RoTs96] (or ordered distance, in [MaSt99]) is a matrix norm metric (in fact, an ultrametric) on $M_{m,n}(\mathbb{F}_q)$, defined by

$$||A-B||_{RT}$$
.

For every matrix code $C \subset M_{m,n}(F_q)$ with q^k elements the minimum (Rosenbloom-Tsfasman) distance $d_{RT}^*(C) \leq mn - k + 1$. Codes meeting this bound are called maximum distance separable codes.

The most used distance between codewords of a matrix code $C \subset M_{m,n}(F_q)$ is the **Hamming metric** on $M_{m,n}(F_q)$, defined by $||A - B||_H$, where $||A||_H$ is the *Hamming weight* of a matrix $A \in M_{m,n}(\mathbb{F}_q)$, i.e., the number of its non-zero entries.

The **LRTJ-metric** (introduced as *Generalized-Lee-Rosenbloom-Tsfasman pseudo-metric* by Jain, 2008) is the **norm metric** for the following generalization of the above norm $||a||_{RT}$ in the case $a \neq 0_{1,n}$:

$$||a||_{LRTJ} = \max_{1 \leq i \leq n} \min\{\xi_i, q - \xi_i\} + \max\{i - 1 : \xi_i \neq 0\}.$$

It is the **Lee metric** for m = 1 and the Rosenbloom-Tsfasman metric for q = 2, 3.

• Interchange distance

The interchange distance (or swap metric) is a metric on the code $C \subset \mathcal{A}^n$ over an alphabet \mathcal{A} , defined, for any $x, y \in C$, as the minimum number of transpositions, i.e., interchanges of adjacent pairs of symbols, converting x into y.

• ACME distance

The **ACME distance** on a code $C \subset \mathcal{A}^n$ over an alphabet \mathcal{A} is defined by

$$\min\{d_H(x,y),d_I(x,y)\},\$$

where d_H is the **Hamming metric**, and d_I is the **interchange distance**.

• Indel distance

Let W be the set of all words over an alphabet \mathcal{A} . A deletion of a letter in a word $\beta = b_1 \dots b_n$ of the length n is a transformation of β into a word $\beta' = b_1 \dots b_{i-1} b_{i+1} \dots b_n$ of the length n-1. An insertion of a letter in a word $\beta = b_1 \dots b_n$ of the length n is a transformation of β into a word $\beta'' = b_1 \dots b_i b_{i+1} \dots b_n$, of the length n+1.

The **indel distance** (or distance of codes with correction of deletions and insertions) is a metric on W, defined, for any $\alpha, \beta \in W$, as the minimum number of deletions and insertions of letters converting α into β . (Cf. **indel metric** in Chap. 11.)

A code C with correction of deletions and insertions is an arbitrary finite subset of W. An example of such a code is the set of words $\beta = b_1 \dots b_n$ of length n over the alphabet $\mathcal{A} = \{0,1\}$ for which $\sum_{i=1}^{n} ib_i \equiv 0 \pmod{n+1}$. The number of words in this code is equal to $\frac{1}{2(n+1)} \sum_{k} \phi(k) 2^{(n+1)/k}$, where the sum is taken over all odd divisors k of n+1, and ϕ is the Euler function.

• Interval distance

The **interval distance** (see, for example, [Bata95]) is a metric on a finite group (G, +, 0), defined by

$$w_{int}(x-y)$$
,

where $w_{int}(x)$ is an interval weight on G, i.e., a group norm whose values are consecutive non-negative integers $0, \ldots, m$. This distance is used for group codes $C \subset G$.

• Fano metric

The **Fano metric** is a *decoding metric* with the goal to find the best sequence estimate used for the *Fano algorithm* of *sequential decoding* of *convolutional codes*.

A convolutional code is a type of error-correction code in which each k-bit information symbol to be encoded is transformed into an n-bit codeword, where $R = \frac{k}{n}$ is the code rate $(n \ge k)$, and the transformation is a function of the last m information symbols. The linear time-invariant decoder (fixed convolutional decoder) maps an information symbol $u_i \in \{u_1, \ldots, u_N\},\$ $u_i = (u_{i1}, \dots u_{ik}), u_{ij} \in \mathbb{F}_2$, into a codeword $x_i \in \{x_1, \dots, x_N\}, x_i =$ $(x_{i1},\ldots,x_{in}), x_{ij} \in \mathbb{F}_2$, so one has a code $\{x_1,\ldots,x_N\}$ with N codewords which occur with probabilities $\{p(x_1), \ldots, p(x_N)\}$. A sequence of l codewords forms a stream (or path) $x = x_{[1,l]} = \{x_1, \ldots, x_l\}$ which is transmitted through a discrete memoryless channel, resulting in the received sequence $y = y_{[1,l]}$. The task of a decoder which minimizes the sequence error probability is to find a sequence which maximizes the joint probability of input and output channel sequences $p(y,x) = p(y|x) \cdot p(x)$. Usually it is sufficient to find a procedure that maximizes p(y|x), and a decoder that always chooses as its estimate one of the sequences that maximizes it or, equivalently, the **Fano metric**, is called a max-likelihood decoder.

Roughly, we consider each code as a tree, where each branch represents one codeword. The decoder begins at the first vertex in the tree, and computes the branch metric for each possible branch, determining the best branch to be the one corresponding to the codeword x_j resulting in the largest branch metric, $\mu_F(x_j)$. This branch is added to the path, and the algorithm continues from the new node which represents the sum of the previous node and the number of bits in the current best codeword. Through iterating until a terminal node of the tree is reached, the algorithm traces the most likely path.

In this construction, the **bit Fano metric** is defined by

$$\log_2 \frac{p(y_i|x_i)}{p(y_i)} - R,$$

the branch Fano metric is defined by

$$\mu_F(x_j) = \sum_{i=1}^n (\log_2 \frac{p(y_i|x_{ji})}{p(y_i)} - R),$$

and the path Fano metric is defined by

$$\mu_F(x_{[1,l]}) = \sum_{j=1}^{l} \mu_F(x_j),$$

where $p(y_i|x_{ji})$ are the channel transition probabilities, $p(y_i) = \sum_{x_m} p(x_m)p(y_i|x_m)$ is the probability distribution of the output given the input symbols averaged over all input symbols, and $R = \frac{k}{n}$ is the code rate.

For a hard-decision decoder $p(y_j = 0|x_j = 1) = p(y_j = 1|x_j = 0) = p$, $0 , the Fano metric for a path <math>x_{[1,l]}$ can be written as

$$\mu_F(x_{[1,l]}) = -\alpha d_H(y_{[1,l]}, x_{[1,l]}) + \beta \cdot l \cdot n,$$

where $\alpha = -\log_2 \frac{p}{1-p} > 0$, $\beta = 1 - R + \log_2(1-p)$, and d_H is the **Hamming metric**.

The generalized Fano metric for sequential decoding is defined by

$$\mu_F^w(x_{[1,l]}) = \sum_{j=1}^{ln} \left(\log_2 \frac{p(y_j|x_j)^w}{p(y_j)^{1-w}} - wR \right),$$

 $0 \le w \le 1$. When w = 1/2, the generalized Fano metric reduces to the Fano metric with a multiplicative constant 1/2.

Metric recursion of a MAP decoding

Maximum a posteriori sequence estimation, or MAP decoding for variable length codes, used the Viterbi algorithm, and is based on the **metric recursion**

$$\Lambda_k^{(m)} = \Lambda_{k-1}^{(m)} + \sum_{n=1}^{l_k^{(m)}} x_{k,n}^{(m)} \log_2 \frac{p(y_{k,n} | x_{k,n}^{(m)} = +1)}{p(y_{k,n} | x_{k,n}^{(m)} = -1)} + 2\log_2 p(u_k^{(m)}),$$

where $\Lambda_k^{(m)}$ is the **branch metric** of branch m at time (level) k, $x_{k,n}$ is the n-th bit of the codeword having $l_k^{(m)}$ bits labeled at each branch,

 $y_{k,n}$ is the respective received soft-bit, u_k^m is the source symbol of branch m at time k and, assuming statistical independence of the source symbols, the probability $p(u_k^{(m)})$ is equivalent to the probability of the source symbol labeled at branch m, that may be known or estimated. The metric increment is computed for each branch, and the largest value, when using log-likelihood-values, of each state is used for further recursion. The decoder first computes the metric of all branches, and then the branch sequence with largest metric starting from the final state backward is selected.

Chapter 17 Distances and Similarities in Data Analysis

A data set is a finite set comprising m sequences (x_1^j, \ldots, x_n^j) , $j \in \{1, \ldots, m\}$, of length n. The values x_i^1, \ldots, x_i^m represent an attribute S_i . It can be numerical, including continuous (real numbers) and binary (presence/absence expressed by 1/0), ordinal (numbers expressing rank only), or nominal (which are not ordered).

Cluster Analysis (or Classification, Taxonomy, Pattern Recognition) consists mainly of partition of data A into a relatively small number of clusters, i.e., such sets of objects that (with respect to a selected measure of distance) are at best possible degree, "close" if they belong to the same cluster, "far" if they belong to different clusters, and further subdivision into clusters will impair the above two conditions.

We give three typical examples. In *Information Retrieval* applications, nodes of peer-to-peer database network export data (collection of text documents); each document is characterized by a vector from \mathbb{R}^n . An user query consists of a vector $x \in \mathbb{R}^n$, and the user needs all documents in the database which are relevant to it, i.e., belong to the ball in \mathbb{R}^n , center x, of fixed radius and with a convenient distance function. In Record Linkage, each document (database record) is represented by a term-frequency vector $x \in \mathbb{R}^n$ or a string, and one wants to measure semantic relevancy of syntactically different records.

In Ecology, let x, y be species abundance distributions, obtained by two sample methods (i.e., x_j, y_j are the numbers of individuals of species j, observed in a corresponding sample); one needs a measure of the distance between x and y, in order to compare two methods. Often data are organized in a **metric tree** first, i.e., in a tree indexed in a metric space.

Once a distance d between objects is selected, the **linkage metric**, i.e., a distance between clusters $A = \{a_1, \ldots, a_m\}$ and $B = \{b_1, \ldots, b_n\}$ is usually one of the following:

average linkage: the average of the distances between the all members of the clusters, i.e., $\frac{\sum_i \sum_j d(a_i,b_j)}{mn}$;

single linkage: the distance between the nearest members of the clusters, i.e., $\min_{i,j} d(a_i, b_j)$;

complete linkage: the distance between the furthest members of the clusters, i.e., $\max_{i,j} d(a_i, b_j)$;

centroid linkage: the distance between the centroids of the clusters, i.e., $||\tilde{a} - \tilde{b}||_2$, where $\tilde{a} = \frac{\sum_i a_i}{m}$, and $\tilde{b} = \frac{\sum_j b_j}{n}$; Ward linkage: the distance $\sqrt{\frac{mn}{m+n}}||\tilde{a} - \tilde{b}||_2$.

Multidimensional Scaling is a technique developed in the behavioral and social sciences for studying the structure of objects or people. Together with Cluster Analysis, it is based on distance methods. But in Multi-dimensional Scaling, as opposed to Cluster Analysis, one starts only with some $m \times m$ matrix D of distances of the objects and (iteratively) looks for a representation of objects in \mathbb{R}^n with low n, so that their Euclidean distance matrix has minimal square deviation from the original matrix D. The related MetricNearness Problem (Dhillon, Sra and Tropp 2003) is to approximate a given finite distance space (X, d) by a metric space (X, d').

There are many similarities used in Data Analysis; the choice depends on the nature of data and is not an exact science. We list below the main such similarities and distances.

Given two objects, represented by non-zero vectors $x = (x_1, \ldots, x_n)$ and $y = (y_1, \ldots, y_n)$ from \mathbb{R}^n , the following notation is used in this chapter.

 $\sum x_i$ means $\sum_{i=1}^n x_i$.

 $\overline{1_F}$ is the *characteristic function* of event $F: 1_F = 1$ if F happens, and $1_F = 0$ otherwise.

 $||x||_2 = \sqrt{\sum x_i^2}$ is the ordinary Euclidean norm on \mathbb{R}^n . \overline{x} denotes $\frac{\sum x_i}{n}$, i.e., the *mean value* of components of x. So, $\overline{x} = \frac{1}{n}$ if xis a frequency vector (discrete probability distribution), i.e., all $x_i \geq 0$, and $\sum x_i = 1$; and $\overline{x} = \frac{n+1}{2}$ if x is a ranking (permutation), i.e., all x_i are different numbers from $\{1, \ldots, n\}$.

In the binary case $x \in \{0,1\}^n$ (i.e., when x is a binary n-sequence), let $X = \{1 \le i \le n : x_i = 1\} \text{ and } \overline{X} = \{1 \le i \le n : x_i = 0\}. \text{ Let } |X \cap Y|, |X \cup Y|,$ $|X\setminus Y|$ and $|X\triangle Y|$ denote the cardinality of the intersection, union, difference and symmetric difference $(X \setminus Y) \cup (Y \setminus X)$ of the sets X and Y, respectively.

Similarities and distances for numerical data 17.1

• Ruzicka similarity The Ruzicka similarity is a similarity on \mathbb{R}^n , defined by

$$\frac{\sum \min\{x_i, y_i\}}{\sum \max\{x_i, y_i\}}.$$

The corresponding distance

$$1 - \frac{\sum \min\{x_i, y_i\}}{\sum \max\{x_i, y_i\}} = \frac{\sum |x_i - y_i|}{\sum \max\{x_i, y_i\}}$$

coincides on $\mathbb{R}^n_{>0}$ with the fuzzy polyonucleotide metric (cf. Chap. 23).

• Roberts similarity

The **Roberts similarity** is a similarity on \mathbb{R}^n , defined by

$$\frac{\sum (x_i + y_i) \frac{\min\{x_i, y_i\}}{\max\{x_i, y_i\}}}{\sum (x_i + y_i)}.$$

• Ellenberg similarity

The **Ellenberg similarity** is a similarity on \mathbb{R}^n , defined by

$$\frac{\sum (x_i + y_i) 1_{x_i \cdot y_i \neq 0}}{\sum (x_i + y_i) (1 + 1_{x_i y_i = 0})}.$$

The binary cases of Ellenberg and Ruzicka similarities coincide; it is called Tanimoto similarity (or Jaccard similarity of community, Jaccard 1908):

$$\frac{|X \cap Y|}{|X \cup Y|}.$$

The **Tanimoto distance** (or **biotope distance**, **Jaccard distance**) is a distance on $\{0,1\}^n$, defined by

$$1 - \frac{|X \cap Y|}{|X \cup Y|} = \frac{|X \Delta Y|}{|X \cup Y|}.$$

• Gleason similarity

The **Gleason similarity** is a similarity on \mathbb{R}^n , defined by

$$\frac{\sum (x_i + y_i) 1_{x_i \cdot y_i \neq 0}}{\sum (x_i + y_i)}.$$

The binary cases of Cleason, Motyka and Bray-Curtis similarities coincide; it is called **Dice similarity** 1945, (or *Sørensen similarity*, *Czekanowsky similarity*):

$$\frac{2|X\cap Y|}{|X\cup Y|+|X\cap Y|}=\frac{2|X\cap Y|}{|X|+|Y|}.$$

The **Czekanowsky–Dice distance** (or *nonmetric coefficient*, Bray and Curtis 1957) is a **near-metric** on $\{0,1\}^n$, defined by

$$1-\frac{2|X\cap Y|}{|X|+|Y|}=\frac{|X\Delta Y|}{|X|+|Y|}.$$

• Intersection distance

The **intersection distance** is a distance on \mathbb{R}^n , defined by

$$1 - \frac{\sum \min\{x_i, y_i\}}{\min\{\sum x_i, \sum y_i\}}.$$

• Motyka similarity

The **Motyka similarity** is a similarity on \mathbb{R}^n , defined by

$$\frac{\sum \min\{x_i, y_i\}}{\sum (x_i + y_i)} = n \frac{\sum \min\{x_i, y_i\}}{\overline{x} + \overline{y}}.$$

• Bray-Curtis similarity

The Bray-Curtis similarity 1957, is a similarity on \mathbb{R}^n , defined by

$$\frac{2}{n(\overline{x} + \overline{y})} \sum \min\{x_i, y_j\}.$$

It is called Renkonen % similarity (or percentage similarity) if x, y are frequency vectors.

• Sørensen distance

The Sørensen distance (or Bray-Curtis distance) is a distance on \mathbb{R}^n , defined (Sørensen 1948) by

$$\frac{\sum |x_i - y_i|}{\sum (x_i + y_i)}.$$

• Canberra distance

The **Canberra distance** (Lance and Williams 1967) is a distance on \mathbb{R}^n , defined by

$$\sum \frac{|x_i - y_i|}{|x_i| + |y_i|}.$$

• Kulczynski similarity 1

The **Kulczynski similarity** 1 is a similarity on \mathbb{R}^n , defined by

$$\frac{\sum \min\{x_i, y_i\}}{\sum |x_i - y_i|}.$$

The corresponding distance is

$$\frac{\sum |x_i - y_i|}{\sum \min\{x_i, y_i\}}.$$

The Soergel distance is

$$\frac{\sum |x_i - y_i|}{\sum \max\{x_i, y_i\}}.$$

• Kulczynski similarity 2 The Kulczynski similarity 2 is a similarity on \mathbb{R}^n , defined by

$$\frac{n}{2}\left(\frac{1}{x} + \frac{1}{y}\right) \sum \min\{x_i, y_i\}.$$

In the binary case it takes the form

$$\frac{|X \cap Y| \cdot (|X| + |Y|)}{2|X| \cdot |Y|}.$$

• Baroni–Urbani–Buser similarity
The Baroni–Urbani–Buser similarity is a similarity on \mathbb{R}^n , defined by

$$\frac{\sum \min\{x_i, y_i\} + \sqrt{\sum \min\{x_i, y_i\}} \sum (\max_{1 \le j \le n} x_j - \max\{x_i, y_i\})}{\sum \max\{x_i, y_i\} + \sqrt{\sum \min\{x_i, y_i\}} \sum (\max_{1 \le j \le n} x_j - \max\{x_i, y_i\})}.$$

In the binary case it takes the form

$$\frac{|X\cap Y|+\sqrt{|X\cap Y|\cdot|\overline{X\cup Y}|}}{|X\cup Y|+\sqrt{|X\cap Y|\cdot|\overline{X\cup Y}|}}.$$

17.2 Relatives of Euclidean distance

• Power (p, r)-distance The power (p, r)-distance is a distance on \mathbb{R}^n , defined by

$$\left(\sum |x_i - y_i|^p\right)^{\frac{1}{r}}.$$

For $p = r \ge 1$, it is the l_p -metric, including the Euclidean, Manhattan (or magnitude) and Chebyshev (or maximum-value, dominance) metrics for p = 2, 1 and ∞ , respectively.

The case (p,r)=(2,1) corresponds to the squared Euclidean distance.

The power (p,r)-distance with 0 is called the**fractional** $<math>l_p$ -distance (not a metric since the unit balls are not convex); it is used for "dimensionality-cursed" data, i.e., when there are few observations and the number n of variables is large. The case 0 , i.e., of the <math>p-th power of the fractional l_p -distance, corresponds to a metric on \mathbb{R}^n .

The weighted versions $(\sum w_i|x_i-y_i|^p)^{\frac{1}{p}}$ (with non-negative weights w_i) are also used, for p=1,2, in applications.

• Penrose size distance

The **Penrose size distance** is a distance on \mathbb{R}^n , defined by

$$\sqrt{n}\sum |x_i-y_i|.$$

It is proportional to the Manhattan metric.

The **mean character distance** (Czekanowsky 1909) is defined by $\sum |x_i-y_i|$.

The Lorentzian distance is a distance defined by $\sum \ln(1+|x_i-y_i|)$.

• Penrose shape distance

The **Penrose shape distance** is a distance on \mathbb{R}^n , defined by

$$\sqrt{\sum ((x_i - \overline{x}) - (y_i - \overline{y}))^2}.$$

The sum of squares of two above **Penrose distances** is the **squared Euclidean distance**.

• Binary Euclidean distance

The binary Euclidean distance is a distance on \mathbb{R}^n , defined by

$$\sqrt{\sum (1_{x_i>0} - 1_{y_i>0})^2}.$$

• Mean censored Euclidean distance

The **mean censored Euclidean distance** is a distance on \mathbb{R}^n , defined by

$$\sqrt{\frac{\sum (x_i - y_i)^2}{\sum 1_{x_i^2 + y_i^2 \neq 0}}}.$$

• Normalized l_p -distance

The **normalized** l_p -distance, $1 \le p \le \infty$, is a distance on \mathbb{R}^n , defined by

$$\frac{||x - y||_p}{||x||_p + ||y||_p}.$$

The only integer value p for which the normalized l_p -distance is a metric, is p=2. Moreover, in [Yian91] it is shown that, for any a,b>0, the distance $\frac{||x-y||_2}{a+b(||x||_2+||y||_2)}$ is a metric.

• Clark distance

The Clark distance (Clark 1952) is a distance on \mathbb{R}^n , defined by

$$\left(\frac{1}{n}\sum \left(\frac{x_i-y_i}{|x_i|+|y_i|}\right)^2\right)^{\frac{1}{2}}.$$

• Meehl distance

The **Meehl distance** (or *Meehl index*) is a distance on \mathbb{R}^n , defined by

$$\sum_{1 \le i \le n-1} (x_i - y_i - x_{i+1} + y_{i+1})^2.$$

• Hellinger distance

The **Hellinger distance** is a distance on \mathbb{R}^n_+ , defined by

$$\sqrt{2\sum\left(\sqrt{\frac{x_i}{\overline{x}}}-\sqrt{\frac{y_i}{\overline{y}}}\right)^2}.$$

(Cf. Hellinger metric in Chap. 14.)

The Whittaker index of association is defined by $\frac{1}{2} \sum \left| \frac{x_i}{\overline{x}} - \frac{y_i}{\overline{y}} \right|$.

• Symmetric χ^2 -measure

The **symmetric** χ^2 -measure is a distance on \mathbb{R}^n , defined by

$$\sum \frac{2}{\overline{x} \cdot \overline{y}} \cdot \frac{(x_i \overline{y} - y_i \overline{x})^2}{x_i + y_i}.$$

• Symmetric χ^2 -distance

The **symmetric** χ^2 -distance (or *chi-distance*) is a distance on \mathbb{R}^n , defined by

$$\sqrt{\sum \frac{\overline{x} + \overline{y}}{n(x_i + y_i)} (\frac{x_i}{\overline{x}} - \frac{y_i}{\overline{y}})^2} = \sqrt{\sum \frac{\overline{x} + \overline{y}}{n(\overline{x} \cdot \overline{y})^2} \cdot \frac{(x_i \overline{y} - y_i \overline{x})^2}{x_i + y_i}}.$$

Mahalanobis distance

The **Mahalanobis distance** (or *statistical distance*) is a distance on \mathbb{R}^n , defined (Mahalanobis 1936) by

$$||x-y||_A = \sqrt{(det A)^{\frac{1}{n}}(x-y)A^{-1}(x-y)^T},$$

where A is a positive-definite matrix (usually, the *covariance matrix* of a finite set consisting of *observation vectors*); cf. Mahalanobis semimetric in Chap. 14).

If A is a diagonal matrix (moreover, the identity matrix), then the Mahalanobis distance is called the **normalized Euclidean distance** (moreover, is the Euclidean distance). For heterogenous data sets (i.e., with ranges and variances of data points $x, y \in \mathbb{R}^n$ dependent on dimension $i \in \{1, ..., n\}$) the **scaled Euclidean distance** is

$$\sqrt{\sum_{i} \frac{(x_i - y_i)^2}{\sigma_i^2}},$$

where σ_i^2 is the variance in dimension *i*. The **maximum scaled difference** (used by Maxwell and Buddemeier 2002, for coastal typology) is defined by

$$\max_{i} \frac{(x_i - y_i)^2}{\sigma_i^2}.$$

17.3 Similarities and distances for binary data

Usually, such similarities s range from 0 to 1 or from -1 to 1; the corresponding distances are usually 1-s or $\frac{1-s}{2}$, respectively.

• Hamann similarity
The Hamann similarity 1961, is a similarity on $\{0,1\}^n$, defined by

$$\frac{2|\overline{X\Delta Y}|}{n} - 1 = \frac{n - 2|X\Delta Y|}{n}.$$

• Rand similarity

The **Rand similarity** (or Sokal–Michener's *simple matching*) is a similarity on $\{0,1\}^n$, defined by

$$\frac{\overline{|X\Delta Y|}}{n} = 1 - \frac{|X\Delta Y|}{n}.$$

Its square root is called the *Euclidean similarity*. The corresponding metric $\frac{|X\Delta Y|}{n}$ is called the *variance* or *Manhattan similarity*; cf. **Penrose size distance**.

• Sokal–Sneath similarity 1 The Sokal–Sneath similarity 1 is a similarity on $\{0,1\}^n$, defined by

$$\frac{2|\overline{X\Delta Y}|}{n+|\overline{X\Delta Y}|}.$$

• Sokal–Sneath similarity 2 The Sokal–Sneath similarity 2 is a similarity on $\{0,1\}^n$, defined by

$$\frac{|X\cap Y|}{|X\cup Y|+|X\Delta Y|}.$$

• Sokal–Sneath similarity 3 The Sokal–Sneath similarity 3 is a similarity on $\{0,1\}^n$, defined by

$$\frac{|X\Delta Y|}{|\overline{X\Delta Y}|}$$
.

• Russel-Rao similarity

The Russel-Rao similarity is a similarity on $\{0,1\}^n$, defined by

$$\frac{|X \cap Y|}{n}$$
.

• Simpson similarity

The **Simpson similarity** (overlap similarity) is a similarity on $\{0,1\}^n$, defined by

$$\frac{|X\cap Y|}{\min\{|X|,|Y|\}}.$$

• Forbes similarity

The **Forbes similarity** is a similarity on $\{0,1\}^n$, defined by

$$\frac{n|X\cap Y|}{|X||Y|}.$$

• Braun-Blanquet similarity

The Braun–Blanquet similarity is a similarity on $\{0,1\}^n$, defined by

$$\frac{|X \cap Y|}{\max\{|X|,|Y|\}}.$$

The average between it and the **Simpson similarity** is the **Dice** similarity.

• Roger-Tanimoto similarity

The **Roger-Tanimoto similarity** 1960, is a similarity on $\{0,1\}^n$, defined by

$$\frac{|\overline{X\Delta Y}|}{n+|X\Delta Y|}.$$

• Faith similarity

The **Faith similarity** is a similarity on $\{0,1\}^n$, defined by

$$\frac{|X \cap Y| + |\overline{X\Delta Y}|}{2n}.$$

• Tversky similarity

The **Tversky similarity** is a similarity on $\{0,1\}^n$, defined by

$$\frac{|X \cap Y|}{a|X\Delta Y| + b|X \cap Y|}.$$

It becomes the **Tanimoto**, **Dice** and (the binary case of) **Kulczynsky** 1 **similarities** for $(a,b)=(1,1), (\frac{1}{2},1)$ and (1,0), respectively.

Mountford similarity

The **Mountford similarity** 1962, is a similarity on $\{0,1\}^n$, defined by

$$\frac{2|X\cap Y|}{|X||Y\backslash X|+|Y||X\backslash Y|}.$$

• Gower-Legendre similarity

The Gower-Legendre similarity is a similarity on $\{0,1\}^n$, defined by

$$\frac{|\overline{X\Delta Y}|}{a|X\Delta Y|+|\overline{X\Delta Y}|} = \frac{|\overline{X\Delta Y}|}{n+(a-1)|X\Delta Y|}.$$

• Anderberg similarity

The **Anderberg similarity** (or *Sokal–Sneath 4 similarity*) is a similarity on $\{0,1\}^n$, defined by

$$\frac{|X\cap Y|}{4}\left(\frac{1}{|X|}+\frac{1}{|Y|}\right)+\frac{|\overline{X\cup Y}|}{4}\left(\frac{1}{|\overline{X}|}+\frac{1}{|\overline{Y}|}\right).$$

• Yule Q similarity

The Yule Q similarity (Yule 1900) is a similarity on $\{0,1\}^n$, defined by

$$\frac{|X \cap Y| \cdot |\overline{X \cup Y}| - |X \setminus Y| \cdot |Y \setminus X|}{|X \cap Y| \cdot |\overline{X \cup Y}| + |X \setminus Y| \cdot |Y \setminus X|}.$$

• Yule Y similarity of colligation

The Yule Y similarity of colligation (Yule 1912) is a similarity on $\{0,1\}^n$, defined by

$$\frac{\sqrt{|X\cap Y|\cdot |\overline{X\cup Y}|}-\sqrt{|X\backslash Y|\cdot |Y\backslash X|}}{\sqrt{|X\cap Y|\cdot |\overline{X\cup Y}|}+\sqrt{|X\backslash Y|\cdot |Y\backslash X|}}$$

• Dispersion similarity

The dispersion similarity is a similarity on $\{0,1\}^n$, defined by

$$\frac{|X\cap Y|\cdot |\overline{X\cup Y}|-|X\backslash Y|\cdot |Y\backslash X|}{n^2}.$$

• Pearson ϕ similarity

The **Pearson** ϕ **similarity** is a similarity on $\{0,1\}^n$, defined by

$$\frac{|X\cap Y|\cdot |\overline{X\cup Y}| - |X\backslash Y|\cdot |Y\backslash X|}{\sqrt{|X|\cdot |\overline{X}|\cdot |Y|\cdot |\overline{Y}|}}.$$

• Gower similarity 2

The **Gower similarity** 2 (or *Sokal–Sneath* 5 *similarity*) is a similarity on $\{0,1\}^n$, defined by

$$\frac{|X \cap Y| \cdot |\overline{X \cup Y}|}{\sqrt{|X| \cdot |\overline{X}| \cdot |Y| \cdot |\overline{Y}|}}.$$

• Pattern difference

The **pattern difference** is a distance on $\{0,1\}^n$, defined by

$$\frac{4|X\backslash Y|\cdot |Y\backslash X|}{n^2}.$$

• Q_0 -difference

The Q_0 -difference is a distance on $\{0,1\}^n$, defined by

$$\frac{|X\backslash Y|\cdot |Y\backslash X|}{|X\cap Y|\cdot |\overline{X\cup Y}|}.$$

17.4 Correlation similarities and distances

• Covariance similarity

The **covariance similarity** is a similarity on \mathbb{R}^n , defined by

$$\frac{\sum (x_i - \overline{x})(y_i - \overline{y})}{n} = \frac{\sum x_i y_i}{n} - \overline{x} \cdot \overline{y}.$$

• Correlation similarity

The **correlation similarity** (or *Pearson correlation*, or, by its full name, *Pearson product-moment correlation linear coefficient*) s is a similarity on \mathbb{R}^n , defined by

$$\frac{\sum (x_i - \overline{x})(y_i - \overline{y})}{\sqrt{(\sum (x_j - \overline{x})^2)(\sum (y_j - \overline{y})^2)}}.$$

The dissimilarities 1 - s and $1 - s^2$ are called the **Pearson correlation** distance and squared Pearson distance, respectively. Moreover,

$$\sqrt{2(1-s)} = \sqrt{\sum \left(\frac{x_i - \overline{x}}{\sqrt{\sum (x_j - \overline{x})^2}} - \frac{y_i - \overline{y}}{\sqrt{\sum (y_j - \overline{y})^2}}\right)}$$

is a normalization of the Euclidean distance (cf., a different one, **normalized** l_p -distance above in this chapter).

In the case $\overline{x} = \overline{y} = 0$, the correlation similarity becomes $\frac{\langle x, y \rangle}{||x||_2 \cdot ||y||_2}$.

• Cosine similarity

The **cosine similarity** (or *Orchini similarity*, angular similarity, normalized dot product) is a similarity on \mathbb{R}^n , defined by

$$\frac{\langle x, y \rangle}{||x||_2 \cdot ||y||_2} = \cos \phi,$$

where ϕ is the angle between vectors x and y. In the binary case, it becomes

$$\frac{|X \cap Y|}{\sqrt{|X| \cdot |Y|}}$$

and is called the Ochiai-Otsuka similarity.

In Record Linkage, cosine similarity is called **TF-IDF** similarity; it (or *tf-idf*, *TFIDF*) are used as an abbreviation of *Frequency – Inverse Document Frequency*. In Ecology, cosine similarity is often called **niche overlap similarity**, cf. Chap. 23.

The **cosine distance** is defined by $1 - \cos \phi$.

The Kumar–Hassebrook similarity is defined by

$$\frac{\langle x, y \rangle}{||x - y||_2^2 + \langle x, y \rangle}.$$

• Angular semi-metric

The **angular semi-metric** on \mathbb{R}^n is the angle (measured in radians) between vectors x and y:

$$\arccos \frac{\langle x, y \rangle}{||x||_2 \cdot ||y||_2}.$$

• Orloci distance

The **Orloci distance** (or *chord distance*) is a distance on \mathbb{R}^n , defined by

$$\sqrt{2\left(1 - \frac{\langle x, y \rangle}{||x||_2 \cdot ||y||_2}\right)}.$$

• Similarity ratio

The **similarity ratio** (or *Kohonen similarity*) is a similarity on \mathbb{R}^n , defined by

$$\frac{\langle x, y \rangle}{\langle x, y \rangle + ||x - y||_2^2}.$$

Its binary case is the **Tanimoto similarity**. Sometimes, the similarity ratio is called the *Tanimoto coefficient* or extended Jaccard coefficient.

• Morisita-Horn similarity

The Morisita-Horn similarity (Morisita 1959) is a similarity on \mathbb{R}^n , defined by

$$\frac{2\langle x,y\rangle}{||x||_2^2\cdot\frac{\overline{y}}{\overline{x}}+||y||_2^2\cdot\frac{\overline{x}}{\overline{y}}}.$$

• Spearman rank correlation

In the case when $x, y \in \mathbb{R}^n$ are rankings (or permutations), i.e., the components of each of them are different numbers $1, \ldots, n$, one has $\overline{x} = \overline{y} = \frac{n+1}{2}$. For such ordinal data, the **correlation similarity** becomes

$$1 - \frac{6}{n(n^2 - 1)} \sum_{i=1}^{\infty} (x_i - y_i)^2.$$

It is the **Spearman** ρ **rank correlation**, called also the *Spearman rho metric*, but it is not a distance. The **Spearman** ρ **distance** is the Euclidean metric on permutations.

The **Spearman footrule** is defined by

$$1 - \frac{3}{n^2 - 1} \sum |x_i - y_i|.$$

It is the l_1 -version of the **Spearman rank correlation**. The **Spearman footrule distance** is the l_1 -metric on permutations.

Another correlation similarity for rankings is the **Kendall** τ **rank correlation**, called also *Kendall* τ *metric* (but it is not a distance), which is defined by

$$\frac{2\sum_{1 \le i < j \le n} sign(x_i - x_j)sign(y_i - y_j)}{n(n-1)}.$$

The **Kendall** τ **distance** on permutations is defined by

$$|\{(i,j): 1 \le i < j \le n, (x_i - x_j)(y_i - y_j) < 0\}|.$$

• Cook distance

The **Cook distance** is a distance on \mathbb{R}^n giving a statistical measure of deciding if some *i*-th observation alone affects much regression estimates. It is a normalized **squared Euclidean distance** between estimated parameters from regression models constructed from all data and from data without *i*-th observation.

The main similar distances, used in Regression Analysis for detecting influential observations, are *DFITS distance*, *Welsch distance*, and *Hadi distance*.

• Distance-based machine learning

The following setting is used for many real-world applications (neural networks, etc.), where data are incomplete and have both continuous

and nominal attributes. Given an $m \times (n+1)$ matrix $((x_{ij}))$, its row $(x_{i0}, x_{i1}, \ldots, x_{in})$ denotes an instance input vector $x_i = (x_{i1}, \ldots, x_{in})$ with output class x_{i0} ; the set of m instances represents a training set during learning. For any new input vector $y = (y_1, \ldots, y_n)$, the closest (in terms of a selected distance d) instance x_i is sought, in order to classify y, i.e., predict its output class as x_{i0} .

The distance [WiMa97] $d(x_i, y)$ is defined by

$$\sqrt{\sum_{j=1}^{n} d_j^2(x_{ij}, y_j)}$$

with $d_j(x_{ij}, y_j) = 1$ if x_{ij} or y_j is unknown. If the attribute j (i.e., the range of values x_{ij} for $1 \le i \le m$) is nominal, then $d_j(x_{ij}, y_j)$ is defined, for example, as $1_{x_{ij} \ne y_j}$, or as

$$\sum_{a} \left| \frac{|\{1 \le t \le m : x_{t0} = a, x_{ij} = x_{ij}\}|}{|\{1 \le t \le m : x_{tj} = x_{ij}\}|} - \frac{|\{1 \le t \le m : x_{t0} = a, x_{ij} = y_{j}\}|}{|\{1 \le t \le m : x_{tj} = y_{j}\}|} \right|^{q}$$

for q=1 or 2; the sum is taken over all output classes, i.e., values a from $\{x_{t0}: 1 \leq t \leq m\}$. For continuous attributes j, the number d_j is taken to be $|x_{ij}-y_j|$ divided by $\max_t x_{tj} - \min_t x_{tj}$, or by $\frac{1}{4}$ of the standard deviation of values x_{tj} , $1 \leq t \leq m$.

Chapter 18 Distances in Mathematical Engineering

In this chapter we group the main distances used in Robot Motion, Cellular Automata, Feedback Systems and Multi-objective Optimization.

18.1 Motion planning distances

Automatic motion planning methods are applied in Robotics, Virtual Reality Systems and Computer Aided Design. A motion planning metric is a metric used in automatic motion planning methods.

Let a robot be a finite collection of rigid links organized in a kinematic hierarchy. If the robot has n degrees of freedom, this leads to an n-dimensional manifold C, called the configuration space (or C-space) of the robot. The workspace W of the robot is the space in which the robot moves. Usually, it is modeled as the Euclidean space \mathbb{E}^3 . The obstacle region CB is the set of all configurations $q \in C$ that either cause the robot to collide with obstacles B, or cause different links of the robot to collide among themselves. The closure $cl(C_{free})$ of $C_{free} = C \setminus \{CB\}$ is called the space of collision-free configurations. A motion planning algorithm must find a collision-free path from an initial configuration to a goal configuration.

A **configuration metric** is a motion planning metric on the configuration space C of a robot.

Usually, the configuration space C consists of six-tuples $(x, y, z, \alpha, \beta, \gamma)$, where the first three coordinates define the position, and the last three the orientation. The orientation coordinates are the angles in radians. Intuitively, a good measure of the distance between two configurations is a measure of the workspace region swept by the robot as it moves between them (the **swept volume distance**). However, the computation of such a metric is prohibitively expensive.

The simplest approach has been to consider the C-space as a Cartesian space and to use Euclidean distance or its generalizations. For such **configuration metrics**, one normalizes the orientation coordinates so that they

get the same magnitude as the position coordinates. Roughly, one multiplies the orientation coordinates by the maximum x, y or z range of the workspace bounding box. Examples of such metrics are given below.

More generally, the configuration space of a three-dimensional rigid body can be identified with the Lie group ISO(3): $C \cong \mathbb{R}^3 \times \mathbb{R}P^3$. The general form of a matrix in ISO(3) is given by

$$\begin{pmatrix} R & X \\ 0 & 1 \end{pmatrix}$$
,

where $R \in SO(3) \cong \mathbb{R}P^3$, and $X \in \mathbb{R}^3$. If X_q and R_q represent the translation and rotation components of the configuration $q = (X_q, R_q) \in ISO(3)$, then a configuration metric between configurations q and r is given by $w_{tr}||X_q-X_r||+w_{rot}f(R_q,R_r)$, where the **translation distance** $||X_q-X_r||$ is obtained using some norm ||.|| on \mathbb{R}^3 , and the **rotation distance** $f(R_q,R_r)$ is a positive scalar function which gives the distance between the rotations $R_q, R_r \in SO(3)$. The rotation distance is scaled relative to the translation distance via the weights w_{tr} and w_{rot} .

A workspace metric is a motion planning metric in the workspace \mathbb{R}^3 . There are many other types of metrics used in motion planning methods, in particular, the **Riemannian metrics**, the **Hausdorff metric** and, in Chap. 9, the **separation distance**, the **penetration depth distance** and the **growth distances**.

• Weighted Euclidean distance The weighted Euclidean distance is a configuration metric on \mathbb{R}^6 , defined by

$$\left(\sum_{i=1}^{3} |x_i - y_i|^2 + \sum_{i=4}^{6} (w_i |x_i - y_i|)^2\right)^{\frac{1}{2}}$$

for any $x, y \in \mathbb{R}^6$, where $x = (x_1, \dots, x_6)$, x_1, x_2, x_3 are the position coordinates, x_4, x_5, x_6 are the orientation coordinates, and w_i is the normalization factor. It gives the same importance to both position and orientation.

• Scaled weighted Euclidean distance The scaled weighted Euclidean distance is a configuration metric on \mathbb{R}^6 , defined by

$$\left(s\sum_{i=1}^{3}|x_i-y_i|^2+(1-s)\sum_{i=4}^{6}(w_i|x_i-y_i|)^2\right)^{\frac{1}{2}}$$

for any $x, y \in \mathbb{R}^6$. The scaled weighted Euclidean distance changes the relative importance of the position and orientation components through the scale parameter s.

• Weighted Minkowskian distance

The weighted Minkowskian distance is a configuration metric on \mathbb{R}^6 , defined by

$$\left(\sum_{i=1}^{3} |x_i - y_i|^p + \sum_{i=4}^{6} (w_i|x_i - y_i|)^p\right)^{\frac{1}{p}}$$

for any $x, y \in \mathbb{R}^6$. It uses a parameter $p \geq 1$; as with Euclidean, both position and orientation have the same importance.

• Modified Minkowskian distance

The modified Minkowskian distance is a configuration metric on \mathbb{R}^6 , defined by

$$\left(\sum_{i=1}^{3} |x_i - y_i|^{p_1} + \sum_{i=4}^{6} (w_i |x_i - y_i|)^{p_2}\right)^{\frac{1}{p_3}}$$

for all $x, y \in \mathbb{R}^6$. It distinguishes between position and orientation coordinates using the parameters $p_1 \geq 1$ (for the position) and $p_2 \geq 1$ (for the orientation).

• Weighted Manhattan distance

The weighted Manhattan distance is a configuration metric on \mathbb{R}^6 , defined by

$$\sum_{i=1}^{3} |x_i - y_i| + \sum_{i=4}^{6} w_i |x_i - y_i|$$

for any $x, y \in \mathbb{R}^6$. It coincides, up to a normalization factor, with the usual l_1 -metric on \mathbb{R}^6 .

• Robot displacement metric

The **robot displacement metric** (or *DISP distance*, Latombe 1991, and LaValle 2006) is a **configuration metric** on a configuration space C of a robot, defined by

$$\max_{a \in A} ||a(q) - a(r)||$$

for any two configurations $q, r \in C$, where a(q) is the position of the point a in the workspace \mathbb{R}^3 when the robot is at configuration q, and ||.|| is one of the norms on \mathbb{R}^3 , usually the Euclidean norm. Intuitively, this metric yields the maximum amount in workspace that any part of the robot is displaced when moving from one configuration to another (cf. **bounded box metric**).

• Euler angle metric

The Euler angle metric is a rotation metric on the group SO(3) (for the case of using roll-pitch-yaw Euler angles for rotation), defined by

$$w_{rot}\sqrt{\Delta(\theta_1, \theta_2)^2 + \Delta(\phi_1, \phi_2)^2 + \Delta(\eta_1, \eta_2)^2}$$

for all $R_1, R_2 \in SO(3)$, given by Euler angles $(\theta_1, \phi_1, \eta_1)$, $(\theta_2, \phi_2, \eta_2)$, respectively, where $\Delta(\theta_1, \theta_2) = \min\{|\theta_1 - \theta_2|, 2\pi - |\theta_1 - \theta_2|\}$, $\theta_i \in [0, 2\pi]$, is the **metric between angles**, and w_{rot} is a scaling factor.

• Unit quaternions metric

The unit quaternions metric is a rotation metric on the unit quaternion representation of SO(3), i.e., a representation of SO(3) as the set of points (unit quaternions) on the unit sphere S^3 in \mathbb{R}^4 with identified antipodal points $(q \sim -q)$.

This representation of SO(3) suggested a number of possible metrics on it, for example, the following ones:

- 1. $\min\{||q-r||, ||q+r||\}$
- 2. $||\ln(q^{-1}r)||$
- 3. $w_{rot}(1 |\lambda|)$
- 4. $\arccos |\lambda|$

where $q = q_1 + q_2 i + q_3 j + q_4 k$, $\sum_{i=1}^{4} q_i^2 = 1$, ||.|| is a norm on \mathbb{R}^4 , $\lambda = \langle q, r \rangle = \sum_{i=1}^{4} q_i r_i$, and w_{rot} is a scaling factor.

• Center of mass metric

The **center of mass metric** is a **workspace metric**, defined as the Euclidean distance between the *centers of mass* of the robot in the two configurations. The center of mass is approximated by averaging all object vertices.

Bounded box metric

The **bounded box metric** is a **workspace metric**, defined as the maximum Euclidean distance between any vertex of the *bounding box* of the robot in one configuration and its corresponding vertex in the other configuration. Cf. unrelated **box metric** in Chap. 4.

• Pose distance

A **pose distance** provides a measure of dissimilarity between actions of *agents* (including robots and humans) for Learning by Imitation in Robotics.

In this context, agents are considered as kinematic chains, and are represented in the form of a kinematic tree, such that every link in the kinematic chain is represented by a unique edge in the corresponding tree. The configuration of the chain is represented by the pose of the corresponding tree which is obtained by an assignment of the pair (n_i, l_i) to every edge e_i . Here n_i is the unit normal, representing the orientation of the corresponding link in the chain, and l_i is the length of the link.

The pose class consists of all poses of a given kinematic tree. One of the possible pose distances is a distance on a given pose class which is the sum of measures of dissimilarity for every pair of compatible segments in the two given poses.

Another way is to view a pose D(m) in the context of the a precedent and a subsequent frames as a 3D point cloud $\{D^j(i): m-a \le i \le m+a, j \in J\}$, where J is the joint set. The set D(m) contains k = |J|(2a+1) points

(joint positions) $p_i = (x_i, y_i, z_i), 1 \le i \le k$. Let $T_{\theta, x, z}$ denote the linear transformation which simultaneously rotates all points of a point cloud about the y axis by an angle $\theta \in [0.2\pi]$ and then shifts the resulting points in the xz plane by a vector $(x, 0, z) \in \mathbb{R}^3$. Then the **3D point cloud distance** (Kover and Gleicher 2002) between the poses $D(m) = (p_i)_{i \in [1,k]}$ and $D(n) = (q_i)_{i \in [1,k]}$ is defined as

$$\min_{\theta, x, z} \{ \sum_{i=1}^{k} ||p_i - T_{\theta, x, z}(q_i)||_2^2 \}.$$

Cf. Procrustes distance in Chap. 21.

• Millibot train metrics

In Microbotics (the field of miniature mobile robots), nanorobot, microrobot, millirobot, minirobot, and $small\ robot$ are terms for robots with characteristic dimensions at most $1\,\mu\mathrm{m}$, mm , cm , dm , and m , respectively. A $millibot\ train$ is a team of heterogeneous, resource-limited millirobots which can collectively share information. They are able to fuse range information from a variety of different platforms to build a global occupancy map that represents a single collective view of the environment. In the motion planning of millibot trains for the construction of a $motion\ planning\ metric$, one casts a series of random points about a robot and pose each point as a candidate position for movement. The point with the highest overall utility is then selected, and the robot is directed to that point.

Thus, the **free space metric**, determined by free space contour, only allows candidate points that do not drive the robot through obstructions; **obstacle avoidance metric** penalizes for moves that get too close to obstacles; **frontier metric** rewards for moves that take the robot towards open space; **formation metric** rewards for moves that maintain formation; **localization metric**, based on the separation angle between one or more localization pairs, rewards for moves that maximize localization (see [GKC04]); cf. **collision avoidance distance**, **piano movers distance** in Chap. 19.

18.2 Cellular automata distances

Let $S, 2 \leq |S| < \infty$, denote a finite set (alphabet), and let S^{∞} denote the set of bi-infinite sequences $\{x_i\}_{i=-\infty}^{\infty}$ (configurations) of elements (letters) of S. A (one-dimensional) cellular automaton is a continuous mapping $f: S^{\infty} \to S^{\infty}$ that commutes with the translation map $g: S^{\infty} \to S^{\infty}$, defined by $g(x_i) = x_{i+1}$. Once a metric on S^{∞} is defined, the resulting metric space (X, d) together with the self-mapping f form a **dynamical system**, cf. Chap. 1.

A cellular automaton can be defined as any discrete dynamical system on the finite state space X. Cellular automata (generally, bi-infinite arrays instead of sequences) are used in Symbolic Dynamics, Computer Science and, as models, in Physics and Biology. The main distances between configurations $\{x_i\}_i$ and $\{y_i\}_i$ from S^{∞} (see [BFK99]) follow.

• Cantor metric

The **Cantor metric** is a metric on S^{∞} defined, for $x \neq y$, by

$$2^{-\min\{i\geq 0:|x_i-y_i|+|x_{-i}-y_{-i}|\neq 0\}}$$

It corresponds to the case $a = \frac{1}{2}$ of the **generalized Cantor metric** in Chap. 11. The corresponding metric space is compact.

• Besicovitch semi-metric

The **Besicovitch semi-metric** is a semi-metric on S^{∞} defined, for $x \neq y$, by

$$\overline{\lim}_{l\to\infty}\frac{|-l\le i\le l: x_i\ne y_i|}{2l+1}.$$

Cf. Besicovitch distance on measurable functions in Chap. 13.

The corresponding semi-metric space is complete.

• Weyl semi-metric

The **Weyl semi-metric** is a semi-metric on S^{∞} , defined by

$$\overline{\lim}_{l \to \infty} \max_{k \in \mathbb{Z}} \frac{|k+1 \le i \le k+l : x_i \ne y_i|}{l}.$$

This and the above semi-metric are **translation invariant**, but are neither separable nor locally compact. Cf. **Weyl distance** in Chap. 13.

18.3 Distances in Control Theory

Control Theory considers the feedback loop of a plant P (a function representing the object to be controlled, a system) and a controller C (a function to design). The output y, measured by a sensor, is fed back to the reference value r. Then the controller takes the error e = r - y to make inputs u = Ce. Subject to zero initial conditions, the input and output signals to the plant are related by y = Pu, where r, u, v and P, C are functions of the frequency variable s. So, $y = \frac{PC}{1+PC}r$ and $y \approx r$ (i.e., one controls the output by simply setting the reference) if PC is large for any value of s. If the system is modeled by a system of linear differential equations, then its transfer function $\frac{PC}{1+PC}$ is a rational function. The plant P is stable if it has no poles in the closed right half-plane $\mathbb{C}_+ = \{s \in \mathbb{C} : \Re s \geq 0\}$.

The robust stabilization problem is: given a nominal plant (a model) P_0 and some metric d on plants, find the open ball of maximal radius which is centered in P_0 , such that some controller (rational function) C stabilizes every element of this ball.

The graph G(P) of the plant P is the set of all bounded input-output pairs (u, y = Pu). Both u and y belong to the Hardy space $H^2(\mathbb{C}_+)$ of the right half-plane; the graph is a closed subspace of $H^2(\mathbb{C}_+) + H^2(\mathbb{C}_+)$. In fact, G(P) is a closed subspace of $H^2(\mathbb{C}^2)$, and $G(P) = f(P) \cdot H^2(\mathbb{C}^2)$ for some function f(P), called the graph symbol.

All metrics below are *gap-like metrics*; they are topologically equivalent, and the stabilization is a robust property with respect of each of them.

Gap metric

The **gap metric** between plants P_1 and P_2 (Zames and El-Sakkary 1980) is defined by

$$gap(P_1, P_2) = ||\Pi(P_1) - \Pi(P_2)||_2,$$

where $\Pi(P_i)$, i = 1, 2, is the orthogonal projection of the graph $G(P_i)$ of P_i seen as a closed subspace of $H^2(\mathbb{C}^2)$. We have

$$gap(P_1, P_2) = \max\{\delta_1(P_1, P_2), \delta_1(P_2, P_1)\},\$$

where $\delta_1(P_1, P_2) = \inf_{Q \in H^{\infty}} ||f(P_1) - f(P_2)Q||_{H^{\infty}}$, and f(P) is a graph symbol.

Here H^{∞} is the space of matrix-valued functions that are analytic and bounded in the open right half-plane $\{s \in \mathbb{C} : \Re s > 0\}$; the H^{∞} -norm is the maximum singular value of the function over this space.

If A is an $m \times n$ matrix with m < n, then its n columns span an n-dimensional subspace, and the matrix B of the orthogonal projection onto the column space of A is $A(A^TA)^{-1}A^T$. If the basis is orthonormal, then $B = AA^T$. In general, the **gap metric** between two subspaces of the same dimension is the l_2 -norm of the difference of their orthogonal projections; see also the definition of this distance as an **angle distance** between subspaces.

In some applications, when subspaces correspond to autoregressive models, the *Frobenius norm* is used instead of the l_2 -norm. Cf. **Frobenius distance** in Chap. 12.

• Vidyasagar metric

The Vidyasagar metric (or graph metric) between plants P_1 and P_2 is defined by

$$\max\{\delta_2(P_1, P_2), \delta_2(P_2, P_1)\},\$$

where
$$\delta_2(P_1, P_2) = \inf_{||Q|| \le 1} ||f(P_1) - f(P_2)Q||_{H^{\infty}}$$
.

The **behavioral distance** is the gap between *extended* graphs of P_1 and P_2 ; a term is added to the graph G(P), in order to reflect all possible initial conditions (instead of the usual setup with the initial conditions being zero).

• Vinnicombe metric

The Vinnicombe metric (ν -gap metric) between plants P_1 and P_2 is defined by

$$\delta_{\nu}(P_1, P_2) = ||(1 + P_2 P_2^*)^{-\frac{1}{2}} (P_2 - P_1) (1 + P_1^* P_1)^{-\frac{1}{2}}||_{\infty}$$

if $wno(f^*(P_2)f(P_1)) = 0$, and it is equal to 1 otherwise. Here f(P) is the graph symbol function of plant P. See [Youn98] for the definition of the winding number wno(f) of a rational function f and for a good introduction to Feedback Stabilization.

18.4 MOEA distances

Most optimization problems have several objectives but, for simplicity, only one of them is optimized, and the others are handled as constraints. *Multi-objective optimization* considers (besides some inequality constraints) an objective vector function $f: X \subset \mathbb{R}^n \to \mathbb{R}^k$ from the search (or genotype, decision variables) space X to the objective (or phenotype, decision vectors) space $f(X) = \{f(x) : x \in X\} \subset \mathbb{R}^k$.

A point $x^* \in X$ is Pareto optimal if, for every other $x \in X$, the decision vector f(x) does not Pareto dominate $f(x^*)$, i.e., $f(x) \leq f(x^*)$. The Pareto optimal front is the set $PF^* = \{f(x) : x \in X^*\}$, where X^* is the set of all Pareto optimal points.

Multi-objective evolutionary algorithms (MOEA) produce, at each generation, an approximation set (the found Pareto front PF_{known} approximating the desired Pareto front PF^*) in objective space in which no element Pareto dominates another element. Examples of **MOEA metrics**, i.e., measures evaluating how close PF_{known} is to PF^* , follow.

• Generational distance

The **generational distance** is defined by

$$\frac{(\sum_{j=1}^{m} d_j^2)^{\frac{1}{2}}}{m},$$

where $m = |PF_{known}|$, and d_j is the Euclidean distance (in the objective space) between $f^j(x)$ (i.e., j-th member of PF_{known}) and the nearest member of PF^* . This distance is zero if and only if $PF_{known} = PF^*$.

The term **generational distance** (or *rate of turnover*) is also used for the minimal number of branches between two positions in any system of ranked descent represented by an hierarchical tree. Examples are: **phylogenetic distance** on a phylogenetic tree, the number of generations

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separating a photocopy from the original block print, and the number of generations separating the audience at a memorial from the commemorated event.

Spacing

The **spacing** is defined by

$$\left(\frac{\sum_{j=1}^{m}(\overline{d}-d_j)^2}{m-1}\right)^{\frac{1}{2}},$$

where $m = |PF_{known}|$, d_j is the L_1 -metric (in the objective space) between $f^j(x)$ (i.e., j-th member of PF_{known}) and the nearest other member of PF_{known} , while \overline{d} is the mean of all d_j .

• Overall non-dominated vector ratio

The overall non-dominated vector ratio is defined by

$$\frac{|PF_{known}|}{|PF^*|}.$$

$\begin{array}{c} & \text{Part V} \\ \text{Computer-related Distances} \end{array}$

Chapter 19

Distances on Real and Digital Planes

19.1 Metrics on real plane

In the plane \mathbb{R}^2 we can use many different metrics. In particular, any L_p -metric (as well as any norm metric for a given norm ||.|| on \mathbb{R}^2) can be used on the plane, and the most natural is the L_2 -metric, i.e., the Euclidean metric $d_E(x,y) = \sqrt{(x_1-y_1)^2 + (x_2-y_2)^2}$ which gives the length of the straight line segment [x,y], and is the intrinsic metric of the plane. However, there are other, often "exotic," metrics on \mathbb{R}^2 . Many of them are used for the construction of generalized Voronoi diagrams on \mathbb{R}^2 (see, for example, Moscow metric, network metric, nice metric). Some of them are used in Digital Geometry.

Erdős-type distance problems (given, usually, for the Euclidean metric on \mathbb{R}^2) are of interest for \mathbb{R}^n and for other metrics on \mathbb{R}^2 . Examples of such problems are to find out:

the least number of different distances (or largest occurrence of a given distance) in an m-subset of \mathbb{R}^2 ; the largest size of a subset of \mathbb{R}^2 determining at most m distances;

the minimum diameter of an m-subset of \mathbb{R}^2 with only integral distances (or, say, without a pair (d_1, d_2) of distances with $0 < |d_1 - d_2| < 1$);

existence of an m-subset of \mathbb{R}^2 with, for each $1 \leq i \leq m$, a distance occurring exactly i times (examples are known for $m \leq 8$);

existence of a dense subset of \mathbb{R}^2 with rational distances (Ulam problem); existence of m, m > 7, non-collinear points of \mathbb{R}^2 with integral distances; forbidden distances of a partition of \mathbb{R}^2 , i.e., distances not occurring within each part.

• City-block metric

The **city-block metric** is the L_1 -metric on \mathbb{R}^2 , defined by

$$||x - y||_1 = |x_1 - y_1| + |x_2 - y_2|.$$

This metric is also called the **taxicab metric**, **Manhattan metric**, **rectilinear metric**, **right-angle metric**; on \mathbb{Z}^2 it is called the **grid metric** and 4-metric.

• Chebyshev metric

The Chebyshev metric (or lattice metric, chessboard metric, kingmove metric, 8-metric) is the L_{∞} -metric on \mathbb{R}^2 , defined by

$$||x - y||_{\infty} = \max\{|x_1 - y_1|, |x_2 - y_2|\}.$$

On \mathbb{Z}^n , this metric is called also the **uniform metric**, **sup metric** and box metric.

• (p,q)-relative metric

Let $0 < q \le 1$, $p \ge \max\{1 - q, \frac{2 - q}{3}\}$, and let $||.||_2$ be the Euclidean norm on \mathbb{R}^2 (in general, on \mathbb{R}^n).

The (p,q)-relative metric is a metric on \mathbb{R}^2 (in general, on \mathbb{R}^n and even on any *Ptolemaic* space (V,||.||)), defined by

$$\frac{||x-y||_2}{\left(\frac{1}{2}(||x||_2^p + ||y||_2^p)\right)^{\frac{q}{p}}}$$

for x or $y \neq 0$ (and is equal to 0 otherwise). In the case of $p = \infty$ it has the form

$$\frac{||x-y||_2}{(\max\{||x||_2,||y||_2\})^q}.$$

For q=1 and any $1 \leq p < \infty$, one obtains the *p*-relative metric (or *Klamkin-Meir metric*); for q=1 and $p=\infty$, one obtains the relative metric. The original (1,1)-relative metric is called the **Schattschneider** metric.

• *M*-relative metric

Let $f:[0,\infty)\to(0,\infty)$ be a convex increasing function such that $\frac{f(x)}{x}$ is decreasing for x>0. Let $||.||_2$ be the Euclidean norm on \mathbb{R}^2 (in general, on \mathbb{R}^n).

The M-relative metric is a metric on \mathbb{R}^2 (in general, on \mathbb{R}^n and even on any Ptolemaic space (V, ||.||)), defined by

$$\frac{||x-y||_2}{f(||x||_2) \cdot f(||y||_2)}.$$

In particular, the distance

$$\frac{||x-y||_2}{\sqrt[p]{1+||x||_p^2}\sqrt[p]{1+||y||_p^2}}$$

is a metric on \mathbb{R}^2 (on \mathbb{R}^n) if and only if $p \geq 1$. A similar metric on $\mathbb{R}^2 \setminus \{0\}$ (on $\mathbb{R}^n \setminus \{0\}$) can be defined by

$$\frac{||x - y||_2}{||x||_2 \cdot ||y||_2}.$$

• MBR metric

The MBR metric (Schönemann 1982, for bounded response scales in Psychology) is a metric on \mathbb{R}^2 , defined by

$$\frac{||x-y||_1}{1+|x_1-y_1||x_2-y_2|} = \tanh(\operatorname{arctanh}(|x_1-y_1|) + \operatorname{arctanh}(|x_2-y_2|)).$$

• Moscow metric

The Moscow metric (or Karlsruhe metric) is a metric on \mathbb{R}^2 , defined as the minimum Euclidean length of all *admissible* connecting curves between x and $y \in \mathbb{R}^2$, where a curve is called *admissible* if it consists only of segments of straight lines passing through the origin and segments of circles centered at the origin (see, for example, [Klei88]).

If the polar coordinates for points $x, y \in \mathbb{R}^2$ are (r_x, θ_x) , (r_y, θ_y) , respectively, then the distance between them is equal to $\min\{r_x, r_y\}\Delta(\theta_x - \theta_y) + |r_x - r_y| \text{ if } 0 \leq \Delta(\theta_x, \theta_y) < 2$, and is equal to $r_x + r_y \text{ if } 2 \leq \Delta(\theta_x, \theta_y) < \pi$, where $\Delta(\theta_x, \theta_y) = \min\{|\theta_x - \theta_y|, 2\pi - |\theta_x - \theta_y|\}, \theta_x, \theta_y \in [0, 2\pi)$, is the **metric between angles**.

• French Metro metric

Given a norm ||.|| on \mathbb{R}^2 , the **French Metro metric** is a metric on \mathbb{R}^2 , defined by

$$||x-y||$$

if x = cy for some $c \in \mathbb{R}$, and by

$$||x|| + ||y||,$$

otherwise. For the Euclidean norm $||.||_2$, it is called the **Paris metric**, **hedgehog metric**, **radial metric**, or **enhanced SNCF metric**. In this case it can be defined as the minimum Euclidean length of all *admissible* connecting curves between two given points x and y, where a curve is called *admissible* if it consists only of segments of straight lines passing through the origin.

In graph terms, this metric is similar to the **path metric** of the tree consisting of a point from which radiate several disjoint paths.

The Paris metric is an example of an \mathbb{R} -tree T which is *simplicial*, i.e., the set of points x with $T \setminus \{x\}$ not having two components is discrete and closed.

• Lift metric

The lift metric (or raspberry picker metric or metric "river") is a metric on \mathbb{R}^2 , defined by

$$|x_1-y_1|$$

if $x_2 = y_2$, and by

$$|x_1| + |x_2 - y_2| + |y_1|$$

if $x_2 \neq y_2$ (see, for example, [Brya85]). It can be defined as the minimum Euclidean length of all admissible connecting curves between two given points x and y, where a curve is called admissible if it consists only of segments of straight lines parallel to the x_1 -axis and segments of the x_2 -axis.

The lift metric is an example of an *non-simplicial* (cf. **French Metro** metric) \mathbb{R} -tree.

• British Rail metric

Given a norm ||.|| on \mathbb{R}^2 (in general, on \mathbb{R}^n), the **British Rail metric** is a metric on \mathbb{R}^2 (in general, on \mathbb{R}^n), defined by

$$||x|| + ||y||$$

for $x \neq y$ (and it is equal to 0, otherwise).

It is also called the **Post Office metric**, **caterpillar metric** and **shuttle metric**.

• Flower-shop metric

Let d be a metric on \mathbb{R}^2 , and let f be a fixed point (a flower-shop) in the plane.

The flower-shop metric (sometimes called SNCF metric) is a metric on \mathbb{R}^2 (in general, on any metric space), defined by

$$d(x, f) + d(f, y)$$

for $x \neq y$ (and is equal to 0 otherwise). So, a person living at point x, who wants to visit someone else living at point y, first goes to f, to buy some flowers. In the case d(x,y) = ||x - y|| and the point f being the origin, it is the **British Rail metric**.

If k > 1 flower-shops f_1, \ldots, f_k are available, one buys the flowers, where the detour is a minimum, i.e., the distance between distinct points x, y is equal to $\min_{1 \le i \le k} \{d(x, f_i) + d(f_i, y)\}.$

• Radar screen metric

Given a norm ||.|| on \mathbb{R}^2 (in general, on \mathbb{R}^n), the **radar screen metric** is a metric on \mathbb{R}^2 (in general, on \mathbb{R}^n), defined by

$$\min\{1,||x-y||\}.$$

It is a special case of the *t*-truncated metric from Chap. 4.

• Rickman's rug metric

Given a number $\alpha \in (0,1)$, the **Rickman's rug metric** on \mathbb{R}^2 is defined by

$$|x_1 - y_1| + |x_2 - y_2|^{\alpha}$$
.

It is the case n=2 of the **parabolic distance** in Chap. 6; see there other metrics on \mathbb{R}^n , $n \geq 2$.

• Burago-Ivanov metric

The Burago-Ivanov metric [BuIv01] is a metric on \mathbb{R}^2 , defined by

$$|||x||_2 - ||y||_2| + \min\{||x||_2, ||y||_2\} \cdot \sqrt{\angle(x, y)},$$

where $\angle(x,y)$ is the angle between vectors x and y, and $||.||_2$ is the Euclidean norm on \mathbb{R}^2 . The corresponding **internal metric** on \mathbb{R}^2 is equal to $|||x||_2 - ||y||_2|$ if $\angle(x,y) = 0$, and is equal to $||x||_2 + ||y||_2$, otherwise.

• 2*n*-gon metric

Given a centrally symmetric regular 2n-gon K on the plane, the 2n-gon **metric** is a metric on \mathbb{R}^2 , defined, for any $x, y \in \mathbb{R}^2$, as the shortest Euclidean length of a polygonal line from x to y with each of its sides parallel to some edge of K.

If K is a square with the vertices $\{(\pm 1, \pm 1)\}$, one obtains the Manhattan metric. The Manhattan metric arises also as the Minkowskian metric with the unit ball being the diamond, i.e., a square with the vertices $\{(1,0),(0,1),(-1,0),(0,-1)\}$.

• Fixed orientation metric

Given a set A, $|A| \geq 2$, of distinct orientations (i.e., angles with fixed x-axis) on the plane \mathbb{R}^2 , the A-distance (Widmayer, Wu and Wong 1987) is Euclidean length of the shortest (zig-zag) path of line segments with orientations from A. Any A-distance is a metric; it is called also a **fixed** orientation metric.

A fixed orientation metric with $A = \{\frac{i\pi}{n} : 1 \leq i \leq n\}$ for fixed $n \in [2, \infty]$, is called a **uniform orientation metric**; cf. 2n-gon metric above. It is the L_1 -metric, hexagonal metric, L_2 -metric for $n = 2, 3, \infty$, respectively.

• Central Park metric

The **Central Park metric** is a metric on \mathbb{R}^2 , defined as the length of a shortest L_1 -path (*Manhattan path*) between two points $x, y \in \mathbb{R}^2$ in the presence of a given set of areas which are traversed by a shortest Euclidean path (for example, Central Park in Manhattan).

• Collision avoidance distance

Let $\mathcal{O} = \{O_1, \dots, O_m\}$ be a collection of pairwise disjoint polygons on the Euclidean plane representing a set of obstacles which are neither transparent nor traversable.

The collision avoidance distance (or piano movers distance, shortest path metric with obstacles) is a metric on the set $\mathbb{R}^2 \setminus \{\mathcal{O}\}$, defined, for any $x, y \in \mathbb{R}^2 \setminus \{\mathcal{O}\}$, as the length of the shortest path among all possible continuous paths, connecting x and y, that do not intersect obstacles $O_i \setminus \partial O_i$ (a path can pass through points on the boundary ∂O_i of O_i), $i = 1, \ldots m$.

• Rectilinear distance with barriers

Let $\mathcal{O} = \{O_1, \dots, O_m\}$ be a set of pairwise disjoint open polygonal barriers on \mathbb{R}^2 . A rectilinear path (or Manhattan path) P_{xy} from x to y is a collection of horizontal and vertical segments in the plane, joining x and y. The path P_{xy} is called feasible if $P_{xy} \cap (\bigcup_{i=1}^m B_i) = \emptyset$.

The rectilinear distance with barriers (or rectilinear distance in the presence of barriers) is a metric on $\mathbb{R}^2 \setminus \{\mathcal{O}\}$, defined, for any $x, y \in \mathbb{R}^2 \setminus \{\mathcal{O}\}$, as the length of the shortest feasible rectilinear path from x to y.

The rectilinear distance in the presence of barriers is a restriction of the **Manhattan metric**, and usually it is considered on the set $\{q_1, \ldots, q_n\} \subset \mathbb{R}^2$ of *n origin-destination points*: the problem to find such a path arises, for example, in Urban Transportation, or in Plant and Facility Layout (see, for example, [LaLi81]).

• Link distance

Let $P \subset \mathbb{R}^2$ be a polygonal domain (on n vertices and h holes), i.e., a closed multiply-connected region whose boundary is a union of n line segments, forming h+1 closed polygonal cycles. The **link distance** (Suri 1986) is a metric on P, defined, for any $x, y \in P$, as the minimum number of edges in a polygonal path from x to y within the polygonal domain P.

If the path is restricted to be rectilinear, one obtains the rectilinear link distance. If each line segment of the path is parallel to one from a set A of fixed orientations, one obtains the A-oriented link distance; cf. fixed orientation metric above.

• Facility layout distances

A *layout* is a partition of a rectangular plane region into smaller rectangles, called *departments*, by lines parallel to the sides of original rectangle. All interior vertices should be three-valent, and some of them, at least one on the boundary of each department, are *doors*, i.e., input-output locations.

The problem is to design a convenient notion of distance d(x,y) between departments x and y which minimizes the cost function $\sum_{x,y} F(x,y)d(x,y)$, where F(x,y) is some material flow between x and y. The main distances used are:

The **centroid distance**, i.e., the shortest Euclidean or **Manhattan** distance between *centroids* (the intersections of the diagonals) of x and y;

The **perimeter distance**, i.e., the shortest rectilinear distance between doors of x and y, but going only along the walls, i.e., department perimeters.

• Quickest path metric

A quickest path metric (or network metric, time metric) is a metric on \mathbb{R}^2 (or on a subset of \mathbb{R}^2) in the presence of a given transportation network, i.e., a finite graph G=(V,E) with $V\subset\mathbb{R}^2$ and edge-weight function w(e)>1: the vertices and edges are stations and roads. For any $x,y\in\mathbb{R}^2$, it is the time needed for a quickest path (i.e., a path minimizing thetravel duration) between them when using, eventually, the network.

Movement takes place, either off the network with unit speed, or along its roads $e \in E$ with fixed speeds w(e) >> 1, with respect to a given metric d on the plane (usually, the Euclidean metric, or the **Manhattan metric**). The network G can be accessed or exited only at stations (usual discrete model) or at any point of roads (the continuous model).

The **heavy luggage metric** (Abellanas, Hurtado and Palop 2005) is a quickest path metric on \mathbb{R}^2 in the presence of a network with speed 1 outside of the network and speed ∞ (so, travel time 0) inside of it.

The **airlift metric** is a quickest path metric on \mathbb{R}^2 in the presence of an airports network, i.e., a planar graph G = (V, E) on n vertices (airports) with positive edge weights $(w_e)_{e \in E}$ (flight durations). The graph may be entered and exited only at the airports. Movement off the network takes place with unit speed with respect to the Euclidean metric. We assume that going by car takes time equal to the Euclidean distance d, whereas the flight along an edge e = uv of G takes time w(e) < d(u, v). In the simplest case, when there is an airlift between two points $a, b \in \mathbb{R}^2$, the distance between x and y is equal to

$$\min\{d(x,y), d(x,a) + w + d(b,y), d(x,b) + w + d(a,y)\},\$$

where w is the flight duration from a to b.

The **city metric** is a quickest path metric on \mathbb{R}^2 in the presence of a city public transportation network, i.e., a planar straight line graph G with horizontal or vertical edges. G may be composed of many connected components, and may contain cycles. One can enter/exit G at any point, be it at a vertex or on an edge (it is possible to postulate fixed entry points, too). Once having accessed G, one travels at fixed speed v > 1 in one of the available directions. Movement off the network takes place with unit speed with respect to the **Manhattan metric** (as in a large modern-style city with streets arranged in north-south and east-west directions).

The **subway metric** is a quickest path metric on \mathbb{R}^2 which is a variant of the city metric: a subway (in the form of a line in the plane) is used within a city grid.

• Shantaram metric

For any positive numbers a, b with $b \le 2a \le 2b$, the **Shantaram metric** between two points $x, y \in \mathbb{R}^2$ is 0, a or b if x and y coincide in exactly 2, 1 or no coordinates, respectively.

• Periodic metric

A metric d on \mathbb{R}^2 is called **periodic** if there exist two linearly independent vectors v and u such that the *translation* by any vector w = mv + nu, $m, n \in \mathbb{Z}$, preserves distances, i.e., d(x, y) = d(x + w, y + w) for any $x, y \in \mathbb{R}^2$ (cf. **translation invariant metric** in Chap. 5).

• Nice metric

A metric d on \mathbb{R}^2 with the following properties is called **nice** (Klein and Wood 1989):

- 1. d induces the Euclidean topology;
- 2. The d-circles are bounded with respect to the Euclidean metric;
- 3. If $x, y \in \mathbb{R}^2$ and $x \neq y$, then there exists a point $z, z \neq x, z \neq y$, such that d(x, y) = d(x, z) + d(z, y);
- 4. If $x,y \in \mathbb{R}^2$, $x \prec y$ (where \prec is a fixed order on \mathbb{R}^2 , the lexicographic order, for example), $C(x,y) = \{z \in \mathbb{R}^2 : d(x,z) \leq d(y,z)\}$, $D(x,y) = \{z \in \mathbb{R}^2 : d(x,z) < d(y,z)\}$, and D(x,y) is the closure of D(x,y), then $J(x,y) = C(x,y) \cap D(x,y)$ is a curve homeomorphic to (0,1). The intersection of two such curves consists of finitely many connected components.

Every **norm metric** fulfills 1, 2, and 3 Property 2 means that the metric d is continuous at infinity with respect to the Euclidean metric. Property 4 is to ensure that the boundaries of the correspondent *Voronoi diagrams* are curves, and that not too many intersections exist in a neighborhood of a point, or at infinity.

A nice metric d has a nice Voronoi diagram: in the Voronoi diagram $V(P, d, \mathbb{R}^2)$ (where $P = \{p_1, \dots, p_k\}, k \geq 2$, is the set of generator points) each Voronoi region $V(p_i)$ is a path-connected set with a non-empty interior, and the system $\{V(p_1), \dots, V(p_k)\}$ forms a partition of the plane.

• Contact quasi-distances

The **contact quasi-distances** are the following variations of the **distance convex function** (cf. Chap. 1) defined on \mathbb{R}^2 (in general, on \mathbb{R}^n) for any $x, y \in \mathbb{R}^2$.

Given a set $B \subset \mathbb{R}^2$, the first contact quasi-distance d_B is defined by

$$\inf\{\alpha>0:y-x\in\alpha B\}$$

(cf. sensor network distances in Chap. 29).

Given, moreover, a point $b \in B$ and a set $A \subset \mathbb{R}^2$, the **linear contact quasi-distance** is a **point-set distance** defined by $d_b(x, A) = \inf\{\alpha \geq 0 : \alpha b + x \in A\}.$

The intercept quasi-distance is, for a finite set B, defined by $\frac{\sum_{b \in B} d_b(x,y)}{|B|}$.

• Radar discrimination distance

The radar discrimination distance is a distance on \mathbb{R}^2 , defined by

$$|\rho_x - \rho_y + \theta_{xy}|$$

if $x, y \in \mathbb{R}^2 \setminus \{0\}$, and by

$$|\rho_x - \rho_y|$$

if x = 0 or y = 0, where, for each "location" $x \in \mathbb{R}^2$, ρ_x denotes the radial distance of x from $\{0\}$ and, for any $x, y \in \mathbb{R}^2 \setminus \{0\}$, θ_{xy} denotes the radian angle between them.

• Ehrenfeucht-Haussler semi-metric

Let S be a subset of \mathbb{R}^2 such that $x_1 \geq x_2 - 1 \geq 0$ for any $x = (x_1, x_2) \in S$. The **Ehrenfeucht-Haussler semi-metric** (see [EhHa88]) on S is defined by

$$\log_2\left(\left(\frac{x_1}{y_2}+1\right)\left(\frac{y_1}{x_2}+1\right)\right).$$

• Toroidal metric

The **toroidal metric** is a metric on $T = [0,1) \times [0,1) = \{x = (x_1, x_2) \in \mathbb{R}^2 : 0 \le x_1, x_2 < 1\}$, defined for any $x, y \in \mathbb{R}^2$ by

$$\sqrt{t_1^2 + t_2^2},$$

where $t_i = \min\{|x_i - y_i|, |x_i - y_i + 1|\}$ for i = 1, 2 (cf. **torus metric**).

• Circle metric

The **circle metric** is the **intrinsic metric** on the *unit circle* S^1 in the plane. As $S^1 = \{(x,y): x^2 + y^2 = 1\} = \{e^{i\theta}: 0 \le \theta < 2\pi\}$, it is the length of the shorter of the two arcs joining the points $e^{i\theta}$, $e^{i\vartheta} \in S^1$, and can be written as

$$\min\{|\theta-\vartheta|, 2\pi-|\theta-\vartheta|\} = \left\{ \begin{aligned} |\vartheta-\theta|, & \text{if} \ \ 0 \leq |\vartheta-\theta| \leq \pi, \\ 2\pi-|\vartheta-\theta|, & \text{if} \ \ |\vartheta-\theta| > \pi. \end{aligned} \right.$$

(Cf. metric between angles.)

• Angular distance

The **angular distance** traveled around a circle is the number of radians the path subtends, i.e.,

$$\theta = \frac{l}{r},$$

where l is the length of the path, and r is the radius of the circle.

Metric between angles

The **metric between angles** Δ is a metric on the set of all angles in the plane, defined for any $\theta, \vartheta \in [0, 2\pi)$ (cf. **circle metric**) by

$$\min\{|\theta-\vartheta|, 2\pi-|\theta-\vartheta|\} = \begin{cases} |\vartheta-\theta|, & \text{if } 0 \leq |\vartheta-\theta| \leq \pi, \\ 2\pi-|\vartheta-\theta|, & \text{if } |\vartheta-\theta| > \pi. \end{cases}$$

• Metric between directions

On \mathbb{R}^2 , a direction \hat{l} is a class of all straight lines which are parallel to a given straight line $l \subset \mathbb{R}^2$. The **metric between directions** is a metric on the set \mathcal{L} of all directions on the plane, defined, for any directions $\hat{l}, \hat{m} \in \mathcal{L}$, as the angle between any two representatives.

• Circular-railroad quasi-metric

The circular-railroad quasi-metric on the unit circle $S^1 \subset \mathbb{R}^2$ is defined, for any $x, y \in S^1$, as the length of the counterclockwise circular arc from x to y in S^1 .

• Inversive distance

The **inversive distance** between two non-intersecting circles in the plane is defined as the natural logarithm of the ratio of the radii (the larger to the smaller) of two concentric circles into which the given circles can be inverted.

Let c be the distance between the centers of two non-intersecting circles of radii a and b < a. Then their inversive distance is given by

$$\cosh^{-1} \left| \frac{a^2 + b^2 - c^2}{2ab} \right|.$$

The *circumcircle* and *incircle* of a triangle with *circumradius* R and *inradius* r are at the inversive distance $2\sinh^{-1}(\frac{1}{2}\sqrt{\frac{r}{R}})$.

Given three non-collinear points, construct three tangent circles such that one is centered at each point and the circles are pairwise tangent to one another. Then there exist exactly two non-intersecting circles that are tangent to all three circles. These are called the inner and outer *Soddy circles*. The inversive distance between the Soddy circles is $2 \cosh^{-1} 2$.

19.2 Digital metrics

Here we list special metrics which are used in *Computer Vision* (or *Pattern Recognition*, *Robot Vision*, *Digital Geometry*).

A computer picture (or computer image) is a subset of \mathbb{Z}^n which is called a digital nD space. Usually, pictures are represented in the digital plane (or image plane) \mathbb{Z}^2 , or in the digital space (or image space) \mathbb{Z}^3 . The points of \mathbb{Z}^2 and \mathbb{Z}^3 are called pixels and voxels, respectively. An nD m-quantized space is a scaling $\frac{1}{m}\mathbb{Z}^n$.

A digital metric (see, for example, [RoPf68]) is any metric on a digital nD space. Usually, it should take integer values.

The metrics on \mathbb{Z}^n that are mainly used are the L_1 - and L_{∞} -metrics, as well as the L_2 -metric after rounding to the nearest greater (or lesser) integer. In general, a given list of neighbors of a pixel can be seen as a list of permitted one-step moves on \mathbb{Z}^2 . Let us associate a **prime distance**, i.e., a positive weight, to each type of such move.

Many digital metrics can be obtained now as the minimum, over all admissible paths (i.e., sequences of permitted moves), of the sum of corresponding prime distances.

In practice, the subset $(\mathbb{Z}_m)^n = \{0, 1, \dots, m-1\}^n$ is considered instead of the full space \mathbb{Z}^n . $(\mathbb{Z}_m)^2$ and $(\mathbb{Z}_m)^3$ are called the m-grill and m-framework, respectively. The most used metrics on $(\mathbb{Z}_m)^n$ are the **Hamming metric** and the **Lee metric**.

• Grid metric

The **grid metric** is the L_1 -metric on \mathbb{Z}^n . The L_1 -metric on \mathbb{Z}^n can be seen as the path metric of an infinite graph: two points of \mathbb{Z}^n are adjacent if their L_1 -distance is equal to one. For \mathbb{Z}^2 this graph is the usual *grid*. Since each point has exactly four closest neighbors in \mathbb{Z}^2 for the L_1 -metric, it is also called the 4-metric.

For n = 2, this metric is the restriction on \mathbb{Z}^2 of the **city-block metric** which is also called the **taxicab metric**, **rectilinear metric**, or **Manhattan metric**.

• Lattice metric

The lattice metric is the L_{∞} -metric on \mathbb{Z}^n . The L_{∞} -metric on \mathbb{Z}^n can be seen as the path metric of an infinite graph: two points of \mathbb{Z}^n are adjacent if their L_{∞} -distance is equal to one. For \mathbb{Z}^2 , the adjacency corresponds to the king move in chessboard terms, and this graph is called the L_{∞} -grid, while this metric is also called the **chessboard metric**, **king-move metric**, or **king metric**. Since each point has exactly eight closest neighbors in \mathbb{Z}^2 for the L_{∞} -metric, it is also called the 8-metric.

This metric is the restriction on \mathbb{Z}^n of the **Chebyshev metric** which is also called the **sup metric**, or **uniform metric**.

• Hexagonal metric

The **hexagonal metric** is a metric on \mathbb{Z}^2 with a unit sphere $S^1(x)$ (centered at $x \in \mathbb{Z}^2$), defined by $S^1(x) = S^1_{L_1}(x) \cup \{(x_1 - 1, x_2 - 1), (x_1 - 1, x_2 + 1)\}$ for x even (i.e., with even x_2), and by $S^1(x) = S^1_{L_1}(x) \cup \{(x_1 + 1, x_2 - 1), (x_1 + 1, x_2 + 1)\}$ for x odd (i.e., with odd x_2). Since any unit sphere $S^1(x)$ contains exactly six integral points, the hexagonal metric is also called the 6-metric (see [LuRo76]).

For any $x, y \in \mathbb{Z}^2$, this metric can be written as

$$\max \left\{ |u_2|, \frac{1}{2} (|u_2| + u_2) + \left\lfloor \frac{x_2 + 1}{2} \right\rfloor - \left\lfloor \frac{y_2 + 1}{2} \right\rfloor - u_1, \right.$$
$$\left. \frac{1}{2} (|u_2| - u_2) - \left\lfloor \frac{x_2 + 1}{2} \right\rfloor + \left\lfloor \frac{y_2 + 1}{2} \right\rfloor + u_1 \right\},$$

where $u_1 = x_1 - y_1$, and $u_2 = x_2 - y_2$.

The hexagonal metric can be defined as the path metric on the hexagonal grid of the plane. In hexagonal coordinates (h_1, h_2) (in which the h_1 -and h_2 -axes are parallel to the grid's edges) the hexagonal distance between points (h_1, h_2) and (i_1, i_2) can be written as $|h_1 - i_1| + |h_2 - i_2|$ if $(h_1 - i_1)(h_2 - i_2) \ge 0$, and as $\max\{|h_1 - i_1|, |h_2 - i_2|\}$ if $(h_1 - i_1)(h_2 - i_2) \le 0$. Here the hexagonal coordinates (h_1, h_2) of a point x are related to its

Cartesian coordinates (x_1, x_2) by $h_1 = x_1 - \lfloor \frac{x_2}{2} \rfloor$, $h_2 = x_2$ for x even, and by $h_1 = x_1 - \lfloor \frac{x_2+1}{2} \rfloor$, $h_2 = x_2$ for x odd.

The hexagonal metric is a better approximation to the Euclidean metric than either L_1 -metric or L_{∞} -metric.

• Neighborhood sequence metric

On the digital plane \mathbb{Z}^2 , consider two types of motions: the *city-block motion*, restricting movements only to the horizontal or vertical directions, and the *chessboard motion*, also allowing diagonal movements.

The use of both these motions is determined by a neighborhood sequence $B = \{b(1), b(2), \ldots, b(l)\}$, where $b(i) \in \{1, 2\}$ is a particular type of neighborhood, with b(i) = 1 signifying unit change in 1 coordinate (city-block neighborhood), and b(i) = 2 meaning unit change also in 2 coordinates (chessboard neighborhood). The sequence B defines the type of motion to be used at every step (see [Das90]).

The **neighborhood sequence metric** is a metric on \mathbb{Z}^2 , defined as the length of a shortest path between x and $y \in \mathbb{Z}^2$, determined by a given neighborhood sequence B. It can be written as

$$\max\{d_B^1(u),d_B^2(u)\},$$

where $u_1 = x_1 - y_1$, $u_2 = x_2 - y_2$, $d_B^1(u) = \max\{|u_1|, |u_2|\}$, $d_B^2(u) = \sum_{j=1}^{l} \lfloor \frac{|u_1| + |u_2| + g(j)}{f(l)} \rfloor$, f(0) = 0, $f(i) = \sum_{j=1}^{i} b(j)$, $1 \le i \le l$, g(j) = f(l) - f(j-1) - 1, $1 \le j \le l$.

For $B = \{1\}$ one obtains the **city-block metric**, for $B = \{2\}$ one obtains the **chessboard metric**. The case $B = \{1, 2\}$, i.e., the alternative use of these motions, results in the **octagonal metric**, introduced in [RoPf68].

A proper selection of the B-sequence can make the corresponding metric very close to the Euclidean metric. It is always greater than the chessboard distance, but smaller than the city-block distance.

• nD-neighborhood sequence metric

The nD-neighborhood sequence metric is a metric on \mathbb{Z}^n , defined as the length of a shortest path between x and $y \in \mathbb{Z}^n$, determined by a given nD-neighborhood sequence B (see [Faze99]).

Formally, two points $x,y\in\mathbb{Z}^n$ are called m-neighbors, $0\leq m\leq n$, if $0\leq |x_i-y_i|\leq 1, 1\leq i\leq n$, and $\sum_{i=1}^n |x_i-y_i|\leq m$. A finite sequence $B=\{b(1),\ldots,b(l)\},\,b(i)\in\{1,2,\ldots,n\},$ is called an nD-neighborhood sequence with period l. For any $x,y\in\mathbb{Z}^n$, a point sequence $x=x^0,x^1,\ldots,x^k=y$, where x^i and $x^{i+1},\,0\leq i\leq k-1$, are r-neighbors, $r=b((i\ mod\ l)+1)$, is called a path from x to y determined by B with length k. The distance between x and y can be written as

$$\max_{1 \le i \le n} d_i(u) \text{ with } d_i(x, y) = \sum_{i=1}^l \left\lfloor \frac{a_i + g_i(j)}{f_i(l)} \right\rfloor,$$

where $u = (|u_1|, |u_2|, ..., |u_n|)$ is the non-increasing ordering of $|u_m|, u_m = x_m - y_m, m = 1, ..., n$, that is, $|u_i| \le |u_j|$ if i < j; $a_i = \sum_{j=1}^{n-i+1} u_j$; $b_i(j) = b(j)$ if b(j) < n-i+2, and is n-i+1 otherwise; $f_i(j) = \sum_{k=1}^{j} b_i(k)$ if $1 \le j \le l$, and is 0 if j = 0; $g_i(j) = f_i(l) - f_i(j-1) - 1, 1 \le j \le l$.

The set of 3D-neighborhood sequence metrics forms a complete distributive lattice under the natural comparison relation.

• Strand-Nagy distances

The face-centered cubic lattice is $A_3 = \{(a_1, a_2, a_3) \in \mathbb{Z}^3 : a_1 + a_2 + a_3 \equiv 0 \pmod{2}\}$, and the body-centered cubic lattice is its dual $A_3^* = \{(a_1, a_2, a_3) \in \mathbb{Z}^3 : a_1 \equiv a_2 \equiv a_3 \pmod{2}\}$.

Let $L \in \{A_3, A_3^*\}$. For any points $x, y \in L$, let $d_1(x, y) = \sum_{j=1}^3 |x_j - y_j|$ denote the L_1 -metric and $d_{\infty}(x, y) = \max_{j \in \{1, 2, 3\}} |x_j - y_j|$ denote the L_{∞} -metric between them. Two points $x, y \in L$ are called 1-neighbors if $d_1(x, y) \leq 3$ and $0 < d_{\infty}(x, y) \leq 1$; they are called 2-neighbors if $d_1(x, y) \leq 3$ and $1 < d_{\infty}(x, y) \leq 2$. Given a sequence $B = \{b(i)\}_{i=1}^{\infty}$ over the alphabet $\{1, 2\}$, a B-path in L is a point sequence $x = x^0, x^1, \ldots, x^k = y$, where x^i and x^{i+1} , $0 \leq i \leq k-1$, are 1-neighbors if b(i) = 1 and 2-neighbors if b(i) = 2.

The **Strand–Nagy distance** between two points $x, y \in L$ (called the *B-distance* by Strand and Nagy 2007) is the length of a shortest *B*-path between them. For $L = A_3$, it is

$$\min\{k: k \ge \max\{\frac{d_1(x,y)}{2}, d_{\infty}(x,y) - |\{1 \le i \le k: b(i) = 2\}|\}.$$

The Strand–Nagy distance is a metric, for example, for the periodic sequence $B = (1, 2, 1, 2, 1, 2, \dots)$ but not for the periodic sequence $B = (2, 1, 2, 1, 2, 1, \dots)$.

• Path-generated metric

Consider the l_{∞} -grid, i.e., the graph with the vertex-set \mathbb{Z}^2 , and two vertices being neighbors if their l_{∞} -distance is 1. Let \mathcal{P} be a collection of paths in the l_{∞} -grid such that, for any $x, y \in \mathbb{Z}^2$, there exists at least one path from \mathcal{P} between x and y, and if \mathcal{P} contains a path Q, then it also contains every path contained in Q. Let $d_{\mathcal{P}}(x,y)$ be the length of the shortest path from \mathcal{P} between x and $y \in \mathbb{Z}^2$. If $d_{\mathcal{P}}$ is a metric on \mathbb{Z}^2 , then it is called a **path-generated metric** (see, for example, [Melt91]).

Let G be one of the sets: $G_1 = \{\uparrow, \rightarrow\}$, $G_{2A} = \{\uparrow, \nearrow\}$, $G_{2B} = \{\uparrow, \nwarrow\}$, $G_{2C} = \{\nearrow, \nwarrow\}$, $G_{2D} = \{\rightarrow, \nwarrow\}$, $G_{3A} = \{\rightarrow, \uparrow, \nearrow\}$, $G_{3B} \{\rightarrow, \uparrow, \nwarrow\}$, $G_{4A} = \{\rightarrow, \nearrow, \nwarrow\}$, $G_{4B} \{\uparrow, \nearrow, \nwarrow, \}$, $G_5 = \{\rightarrow, \uparrow, \nearrow, \nwarrow\}$. Let $\mathcal{P}(G)$ be the set of paths which are obtained by concatenation of paths in G and the corresponding paths in the opposite directions. Any path-generated metric coincides with one of the metrics $d_{\mathcal{P}(G)}$. Moreover, one can obtain the following formulas:

1.
$$d_{\mathcal{P}(G_1)}(x, y) = |u_1| + |u_2|$$

2. $d_{\mathcal{P}(G_{2A})}(x, y) = \max\{|2u_1 - u_2|, |u_2|\}$

```
3. d_{\mathcal{P}(G_{2B})}(x,y) = \max\{|2u_1 + u_2|, |u_2|\}

4. d_{\mathcal{P}(G_{2C})}(x,y) = \max\{|2u_2 + u_1|, |u_1|\}

5. d_{\mathcal{P}(G_{2D})}(x,y) = \max\{|2u_2 - u_1|, |u_1|\}

6. d_{\mathcal{P}(G_{3A})}(x,y) = \max\{|u_1|, |u_2|, |u_1 - u_2|\}

7. d_{\mathcal{P}(G_{3B})}(x,y) = \max\{|u_1|, |u_2|, |u_1 + u_2|\}

8. d_{\mathcal{P}(G_{4A})}(x,y) = \max\{2\lceil(|u_1| - |u_2|)/2\rceil, 0\} + |u_2|

9. d_{\mathcal{P}(G_{4B})}(x,y) = \max\{2\lceil(|u_2| - |u_1|)/2\rceil, 0\} + |u_1|

10. d_{\mathcal{P}(G_n)}(x,y) = \max\{|u_1|, |u_2|\}
```

where $u_1 = x_1 - y_1$, $u_2 = x_2 - y_2$, and $\lceil . \rceil$ is the *ceiling function*: for any real x the number $\lceil x \rceil$ is the least integer greater than or equal to x.

The metric spaces obtained from G-sets which have the same numerical index are isometric. $d_{\mathcal{P}(G_1)}$ is the **city-block metric**, and $d_{\mathcal{P}(G_5)}$ is the **chessboard metric**.

• Knight metric

The **knight metric** is a metric on \mathbb{Z}^2 , defined as the minimum number of moves a chess knight would take to travel from x to $y \in \mathbb{Z}^2$. Its *unit sphere* S^1_{knight} , centered at the origin, contains exactly 8 integral points $\{(\pm 2, \pm 1), (\pm 1, \pm 2)\}$, and can be written as $S^1_{knight} = S^3_{L_1} \cap S^2_{l_{\infty}}$, where $S^3_{L_1}$ denotes the L_1 -sphere of radius 3, and $S^2_{L_{\infty}}$ denotes the L_{∞} -sphere of radius 2, both centered at the origin (see [DaCh88]).

The distance between x and y is 3 if (M, m) = (1, 0), is 4 if (M, m) = (2, 2) and is equal to $\max\{\lceil \frac{M}{2} \rceil, \lceil \frac{M+m}{3} \rceil\} + (M+m) - \max\{\lceil \frac{M}{2} \rceil, \lceil \frac{M+m}{3} \rceil\}$ $(mod \ 2)$ otherwise, where $M = \max\{|u_1|, |u_2|\}, m = \min\{|u_1|, |u_2|\}, u_1 = x_1 - y_1, u_2 = x_2 - y_2.$

• Super-knight metric

Let $p, q \in \mathbb{N}$ such that p + q is odd, and (p, q) = 1.

A (p,q)-super-knight (or (p,q)-leaper) is a (variant) chess piece whose move consists of a leap p squares in one orthogonal direction followed by a 90° direction change, and q squares leap to the destination square.

Chess-variant terms exist for a (p, 1)-leaper with p = 0, 1, 2, 3, 4 (Wazir, Ferz, usual Knight, Camel, Giraffe), and for a (p, 2)-leaper with p = 0, 1, 2, 3 (Dabbaba, usual Knight, Alfil, Zebra).

A (p,q)-super-knight metric (or (p,q)-leaper metric) is a metric on \mathbb{Z}^2 , defined as the minimum number of moves a chess (p,q)-super-knight would take to travel from x to $y \in \mathbb{Z}^2$. Thus, its unit sphere $S^1_{p,q}$, centered at the origin, contains exactly 8 integral points $\{(\pm p, \pm q), (\pm q, \pm p)\}$. (See [DaMu90].)

The **knight metric** is the (1,2)-super-knight metric. The **city-block metric** can be considered as the *Wazir metric*, i.e., (0,1)-super-knight metric.

Rook metric

The **rook metric** is a metric on \mathbb{Z}^2 , defined as the minimum number of moves a chess rook would take to travel from x to $y \in \mathbb{Z}^2$. This metric can take only the values $\{0,1,2\}$, and coincides with the **Hamming metric** on \mathbb{Z}^2 .

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• Chamfer metric

Given two positive numbers α, β with $\alpha \leq \beta < 2\alpha$, consider the (α, β) -weighted l_{∞} -grid, i.e., the infinite graph with the vertex-set \mathbb{Z}^2 , two vertices being adjacent if their l_{∞} -distance is one, while horizontal/vertical and diagonal edges have weights α and β , respectively.

A chamfer metric (or (α, β) -chamfer metric, [Borg86]) is the weighted path metric in this graph. For any $x, y \in \mathbb{Z}^2$ it can be written as

$$\beta m + \alpha (M - m),$$

where $M = \max\{|u_1|, |u_2|\}$, $m = \min\{|u_1|, |u_2|\}$, $u_1 = x_1 - y_1$, $u_2 = x_2 - y_2$. If the weights α and β are equal to the Euclidean lengths 1, $\sqrt{2}$ of horizontal/vertical and diagonal edges, respectively, then one obtains the Euclidean length of the shortest chessboard path between x and y. If $\alpha = \beta = 1$, one obtains the **chessboard metric**. The (3, 4)-chamfer metric is the most used one for digital images; it is called simply the (3, 4)-metric.

A **3D-chamfer metric** is the weighted path metric of the graph with the vertex-set \mathbb{Z}^3 of *voxels*, two voxels being adjacent if their l_{∞} -distance is one, while weights α, β , and γ are associated, respectively, to the distance from 6 face neighbors, 12 edge neighbors, and 8 corner neighbors.

• Weighted cut metric

Consider the weighted l_{∞} -grid, i.e., the graph with the vertex-set \mathbb{Z}^2 , two vertices being adjacent if their l_{∞} -distance is one, and each edge having some positive weight (or cost). The usual weighted path metric between two pixels is the minimal cost of a path connecting them. The weighted cut metric between two pixels is the minimal cost (defined now as the sum of costs of crossed edges) of a cut, i.e., a plane curve connecting them while avoiding pixels.

• Digital volume metric

The **digital volume metric** is a metric on the set K of all bounded subsets (*pictures*, or *images*) of \mathbb{Z}^2 (in general, of \mathbb{Z}^n), defined by

$$vol(A\triangle B)$$
,

where vol(A) = |A|, i.e., the number of pixels contained in A, and $A \triangle B$ is the *symmetric difference* between sets A and B.

This metric is a digital analog of the **Nikodym metric**.

• Hexagonal Hausdorff metric

The **hexagonal Hausdorff metric** is a metric on the set of all bounded subsets (*pictures*, or *images*) of the *hexagonal grid* on the plane, defined by

$$\inf\{p,q:A\subset B+qH,\,B\subset A+pH\}$$

for any pictures A and B, where pH is the regular hexagon of size p (i.e., with p+1 pixels on each edge), centered at the origin and including its interior, and + is the Minkowski addition: $A+B=\{x+y:x\in A,y\in B\}$ (cf. Pompeiu–Hausdorff–Blaschke metric in Chap. 9). If A is a pixel x, then the distance between x and B is equal to $\sup_{y\in B}d_6(x,y)$, where d_6 is the hexagonal metric, i.e., the path metric on the hexagonal grid.

Chapter 20 Voronoi Diagram Distances

Given a finite set A of objects A_i in a space S, computing the *Voronoi diagram* of A means partitioning the space S into *Voronoi regions* $V(A_i)$ in such a way that $V(A_i)$ contains all points of S that are "closer" to A_i than to any other object A_j in A.

Given a generator set $P = \{p_1, \dots p_k\}, k \geq 2$, of distinct points (generators) from \mathbb{R}^n , $n \geq 2$, the ordinary Voronoi polygon $V(p_i)$ associated with a generator p_i is defined by

$$V(p_i) = \{x \in \mathbb{R}^n : d_E(x, p_i) \le d_E(x, p_i) \text{ for any } j \ne i\},$$

where d_E is the ordinary Euclidean distance on \mathbb{R}^n . The set

$$V(P, d_E, \mathbb{R}^n) = \{V(p_1), \dots, V(p_k)\}$$

is called the n-dimensional ordinary Voronoi diagram, generated by P.

The boundaries of (n-dimensional) Voronoi polygons are called ((n-1)-dimensional) Voronoi facets, the boundaries of Voronoi facets are called (n-2)-dimensional Voronoi faces, ..., the boundaries of two-dimensional Voronoi faces are called Voronoi edges, and the boundaries of Voronoi edges are called Voronoi vertices.

A generalization of the ordinary Voronoi diagram is possible in the following three ways:

- 1. The generalization with respect to the generator set $A = \{A_1, \ldots, A_k\}$ which can be a set of lines, a set of areas, etc.
- 2. The generalization with respect to the space S which can be a sphere (spherical Voronoi diagram), a cylinder (cylindrical Voronoi diagram), a cone (conic Voronoi diagram), a polyhedral surface (polyhedral Voronoi diagram), etc.
- 3. The generalization with respect to the function d, where $d(x, A_i)$ measures the "distance" from a point $x \in S$ to a generator $A_i \in A$.

This generalized distance function d is called the **Voronoi generation** distance (or *Voronoi distance*, V-distance), and allows many more functions

than an ordinary metric on S. If F is a strictly increasing function of a V-distance d, i.e., $F(d(x,A_i)) \leq F(d(x,A_j))$ if and only if $d(x,A_i) \leq d(x,A_j)$, then the generalized Voronoi diagram V(A,F(d),S) and V(A,d,S) coincide, and one says that the V-distance F(d) is transformable to the V-distance d, and that the generalized Voronoi diagram V(A,F(d),S) is a trivial triv

In applications, one often uses for trivial generalizations of the ordinary Voronoi diagram $V(P, d, \mathbb{R}^n)$ the **exponential distance**, the **logarithmic distance**, and the **power distance**. There are generalized Voronoi diagrams $V(P, D, \mathbb{R}^n)$, defined by V-distances, that are not transformable to the Euclidean distance d_E : the **multiplicatively weighted Voronoi distance**, the **additively weighted Voronoi distance**, etc.

For additional information see, for example, [OBS92], [Klei89].

20.1 Classical Voronoi generation distances

• Exponential distance

The exponential distance is the Voronoi generation distance

$$D_{exp}(x, p_i) = e^{d_E(x, p_i)}$$

for the trivial generalization $V(P, D_{exp}, \mathbb{R}^n)$ of the ordinary Voronoi diagram $V(P, d_E, \mathbb{R}^n)$, where d_E is the Euclidean distance.

• Logarithmic distance

The logarithmic distance is the Voronoi generation distance

$$D_{\ln}(x, p_i) = \ln d_E(x, p_i)$$

for the trivial generalization $V(P, D_{ln}, \mathbb{R}^n)$ of the ordinary Voronoi diagram $V(P, d_E, \mathbb{R}^n)$, where d_E is the Euclidean distance.

• Power distance

The **power distance** is the Voronoi generation distance

$$D_{\alpha}(x, p_i) = d_E(x, p_i)^{\alpha}, \ \alpha > 0,$$

for the trivial generalization $V(P, D_{\alpha}, \mathbb{R}^n)$ of the ordinary Voronoi diagram $V(P, d_E, \mathbb{R}^n)$, where d_E is the Euclidean distance.

• Multiplicatively weighted distance

The multiplicatively weighted distance d_{MW} is the Voronoi generation distance of the generalized Voronoi diagram $V(P, d_{MW}, \mathbb{R}^n)$ (multiplicatively weighted Voronoi diagram), defined by

$$d_{MW}(x, p_i) = \frac{1}{w_i} d_E(x, p_i)$$

for any point $x \in \mathbb{R}^n$ and any generator point $p_i \in P = \{p_1, \dots, p_k\}, k \geq 2$, where $w_i \in w = \{w_i, \dots, w_k\}$ is a given positive multiplicative weight of the generator p_i , and d_E is the ordinary Euclidean distance.

A Möbius diagram (Boissonnat and Karavelas 2003) is a diagram the **midsets** (bisectors) of which are hyperspheres. It generalizes the Euclidean Voronoi and power diagrams, and it is equivalent to power diagrams in \mathbb{R}^{n+1} .

For \mathbb{R}^2 , the multiplicatively weighted Voronoi diagram is called a *circular Dirichlet tessellation*. An edge in this diagram is a circular arc or a straight line.

In the plane \mathbb{R}^2 , there exists a generalization of the multiplicatively weighted Voronoi diagram, the *crystal Voronoi diagram*, with the same definition of the distance (where w_i is the speed of growth of the crystal p_i), but a different partition of the plane, as the crystals can grow only in an empty area.

• Additively weighted distance

The additively weighted distance d_{AW} is the Voronoi generation distance of the generalized Voronoi diagram $V(P, d_{AW}, \mathbb{R}^n)$ (additively weighted Voronoi diagram), defined by

$$d_{AW}(x, p_i) = d_E(x, p_i) - w_i$$

for any point $x \in \mathbb{R}^n$ and any generator point $p_i \in P = \{p_1, \dots, p_k\}$, $k \geq 2$, where $w_i \in w = \{w_i, \dots, w_k\}$ is a given additive weight of the generator p_i , and d_E is the ordinary Euclidean distance.

For \mathbb{R}^2 , the additively weighted Voronoi diagram is called a *hyperbolic Dirichlet tessellation*. An edge in this Voronoi diagram is a hyperbolic arc or a straight line segment.

• Additively weighted power distance

The additively weighted power distance d_{PW} is the Voronoi generation distance of the generalized Voronoi diagram $V(P, d_{PW}, \mathbb{R}^n)$ (additively weighted power Voronoi diagram), defined by

$$d_{PW}(x, p_i) = d_E^2(x, p_i) - w_i$$

for any point $x \in \mathbb{R}^n$ and any generator point $p_i \in P = \{p_1, \dots, p_k\}, k \geq 2$, where $w_i \in w = \{w_i, \dots, w_k\}$ is a given additive weight of the generator p_i , and d_E is the ordinary Euclidean distance.

This diagram can be seen as a Voronoi diagram of circles or as a Voronoi diagram with the *Laguerre geometry*.

The multiplicatively weighted power distance $d_{MPW}(x, p_i) = \frac{1}{w_i} d_E^2(x, p_i)$, $w_i > 0$, is transformable to the multiplicatively weighted distance, and gives a trivial extension of the multiplicatively weighted Voronoi diagram.

• Compoundly weighted distance

The **compoundly weighted distance** d_{CW} is the Voronoi generation distance of the generalized Voronoi diagram $V(P, d_{CW}, \mathbb{R}^n)$ (compoundly weighted Voronoi diagram), defined by

$$d_{CW}(x, p_i) = \frac{1}{w_i} d_E(x, p_i) - v_i$$

for any point $x \in \mathbb{R}^n$ and any generator point $p_i \in P = \{p_1, \dots, p_k\}$, $k \geq 2$, where $w_i \in w = \{w_i, \dots, w_k\}$ is a given positive multiplicative weight of the generator $p_i, v_i \in v = \{v_1, \dots, v_k\}$ is a given additive weight of the generator p_i , and d_E is the ordinary Euclidean distance.

An edge in the two-dimensional compoundly weighted Voronoi diagram is a part of a fourth-order polynomial curve, a hyperbolic arc, a circular arc, or a straight line.

20.2 Plane Voronoi generation distances

• Shortest path distance with obstacles

Let $\mathcal{O} = \{O_1, \dots, O_m\}$ be a collection of pairwise disjoint polygons on the Euclidean plane, representing a set of obstacles which are neither transparent nor traversable.

The shortest path distance with obstacles d_{sp} is the Voronoi generation distance of the generalized Voronoi diagram $V(P, d_{sp}, \mathbb{R}^2 \setminus \{\mathcal{O}\})$ (shortest path Voronoi diagram with obstacles), defined, for any $x, y \in \mathbb{R}^2 \setminus \{\mathcal{O}\}$, as the length of the shortest path among all possible continuous (x-y)-paths that do not intersect obstacles $O_i \setminus \partial O_i$ (a path can pass through points on the boundary ∂O_i of O_i), $i = 1, \ldots m$.

The shortest path is constructed with the aid of the visibility polygon and the visibility graph of $V(P, d_{sp}, \mathbb{R}^2 \setminus \{\mathcal{O}\})$.

• Visibility shortest path distance

Let $\mathcal{O} = \{O_1, \ldots, O_m\}$ be a collection of pairwise disjoint line segments $O_l = [a_l, b_l]$ in the Euclidean plane, with $P = \{p_1, \ldots, p_k\}, k \geq 2$, the set of generator points,

$$VIS(p_i) = \{x \in \mathbb{R}^2 : [x, p_i] \cap]a_l, b_l [= \emptyset \text{ for all } l = 1, \dots, m\}$$

the *visibility polygon* of the generator p_i , and d_E the ordinary Euclidean distance.

The **visibility shortest path distance** d_{vsp} is the Voronoi generation distance of the generalized Voronoi diagram $V(P, d_{vsp}, \mathbb{R}^2 \setminus \{\mathcal{O}\})$ (visibility shortest path Voronoi diagram with line obstacles), defined by

$$d_{vsp}(x, p_i) = \begin{cases} d_E(x, p_i), & \text{if } x \in VIS(p_i), \\ \infty, & \text{otherwise} . \end{cases}$$

• Network distances

A network on \mathbb{R}^2 is a connected planar geometrical graph G = (V, E) with the set V of vertices and the set E of edges (links).

Let the generator set $P = \{p_1, \ldots, p_k\}$ be a subset of the set $V = \{p_1, \ldots, p_l\}$ of vertices of G, and let the set L be given by points of links of G.

The **network distance** d_{netv} on the set V is the Voronoi generation distance of the *network Voronoi node diagram* $V(P, d_{netv}, V)$, defined as the shortest path along the links of G from $p_i \in V$ to $p_j \in V$. It is the weighted path metric of the graph G, where w_e is the Euclidean length of the link $e \in E$.

The **network distance** d_{netl} on the set L is the Voronoi generation distance of the *network Voronoi link diagram* $V(P, d_{netl}, L)$, defined as the shortest path along the links from $x \in L$ to $y \in L$.

The access network distance d_{accnet} on \mathbb{R}^2 is the Voronoi generation distance of the network Voronoi area diagram $V(P, d_{accnet}, \mathbb{R}^2)$, defined by

$$d_{accnet}(x,y) = d_{netl}(l(x), l(y)) + d_{acc}(x) + d_{acc}(y),$$

where $d_{acc}(x) = \min_{l \in L} d(x, l) = d_E(x, l(x))$ is the access distance of a point x. In fact, $d_{acc}(x)$ is the Euclidean distance from x to the access point $l(x) \in L$ of x which is the nearest to x point on the links of G.

• Airlift distance

An airports network is an arbitrary planar graph G on n vertices (airports) with positive edge weights (flight durations). This graph may be entered and exited only at the airports. Once having accessed G, one travels at fixed speed v>1 within the network. Movement off the network takes place with the unit speed with respect to the ordinary Euclidean distance.

The **airlift distance** d_{al} is the Voronoi generation distance of the *airlift Voronoi diagram* $V(P, d_{al}, \mathbb{R}^2)$, defined as the time needed for a *quickest path* between x and y in the presence of the airports network G, i.e., a path minimizing the travel time between x and y.

City distance

A city public transportation network, like a subway or a bus transportation system, is a planar straight line graph G with horizontal or vertical edges. G may be composed of many connected components, and may contain cycles. One is free to enter G at any point, be it at a vertex or on an edge (it is possible to postulate fixed entry points, too). Once having accessed G, one travels at a fixed speed v > 1 in one of the available directions. Movement off the network takes place with the unit speed with respect to the **Manhattan metric** (we imagine a large modern-style city with streets arranged in north-south and east-west directions).

The **city distance** d_{city} is the Voronoi generation distance of the *city Voronoi diagram* $V(P, d_{city}, \mathbb{R}^2)$, defined as the time needed for the *quick-est path* between x and y in the presence of the network G, i.e., a path minimizing the travel time between x and y.

The set $P = \{p_1, \dots, p_k\}$, $k \geq 2$, can be seen as a set of some city facilities (say, post offices or hospitals): for some people several facilities of the same kind are equally attractive, and they want to find out which facility is reachable first.

• Distance in a river

The distance in a river d_{riv} is the Voronoi generation distance of the generalized Voronoi diagram $V(P, d_{riv}, \mathbb{R}^2)$ (Voronoi diagram in a river), defined by

$$d_{riv}(x,y) = \frac{-\alpha(x_1 - y_1) + \sqrt{(x_1 - y_1)^2 + (1 - \alpha^2)(x_2 - y_2)^2}}{v(1 - \alpha^2)},$$

where v is the speed of the boat on still water, w > 0 is the speed of constant flow in the positive direction of the x_1 -axis, and $\alpha = \frac{w}{v}$ (0 < α < 1) is the relative flow speed.

• Boat-sail distance

Let $\Omega \subset \mathbb{R}^2$ be a *domain* in the plane (water surface), let $f: \Omega \to \mathbb{R}^2$ be a continuous vector field on Ω , representing the velocity of the water flow (flow field); let $P = \{p_1, \dots, p_k\} \subset \Omega, k \geq 2$, be a set of k points in Ω (harbors).

The **boat-sail distance** [NiSu03] d_{bs} is the Voronoi generation distance of the generalized Voronoi diagram $V(P, d_{bs}, \Omega)$ (boat-sail Voronoi diagram), defined by

$$d_{bs}(x,y) = \inf_{\gamma} \delta(\gamma, x, y)$$

for all $x, y \in \Omega$, where $\delta(\gamma, x, y) = \int_0^1 \left| F \frac{\gamma'(s)}{|\gamma'(s)|} + f(\gamma(s)) \right|^{-1} ds$ is the time necessary for the boat with the maximum speed F on still water to move from x to y along the curve $\gamma: [0,1] \to \Omega, \ \gamma(0) = x, \ \gamma(1) = y$, and the infimum is taken over all possible curves γ .

• Peeper distance

Let $S = \{(x_1, x_2) \in \mathbb{R}^2 : x_1 > 0\}$ be the half-plane in \mathbb{R}^2 , let $P = \{p_1, \ldots, p_k\}, k \geq 2$, be a set of points contained in the half-plane $\{(x_1, x_2) \in \mathbb{R}^2 : x_1 < 0\}$, and let the *window* be the interval (a, b) with a = (0, 1) and b = (0, -1).

The **peeper distance** d_{pee} is the Voronoi generation distance of the generalized Voronoi diagram $V(P, d_{pee}, S)$ (peeper's Voronoi diagram), defined by

$$d_{pee}(x, p_i) = \begin{cases} d_E(x, p_i), & \text{if } [x, p] \cap]a, b \neq \emptyset, \\ \infty, & \text{otherwise}, \end{cases}$$

where d_E is the ordinary Euclidean distance.

• Snowmobile distance

Let $\Omega \subset \mathbb{R}^2$ be a *domain* in the x_1x_2 -plane of the space \mathbb{R}^3 (a *two-dimensional mapping*), and let $\Omega^* = \{(q, h(q)) : q = (x_1(q), x_2(q)) \in \Omega, h(q) \in \mathbb{R}\}$ be the three-dimensional *land surface* associated with the mapping Ω . Let $P = \{p_1, \ldots, p_k\} \subset \Omega, k \geq 2$, be a set of k points in Ω (snowmobile stations).

The **snowmobile distance** d_{sm} is the Voronoi generation distance of the generalized Voronoi diagram $V(P, d_{sm}, \Omega)$ (snowmobile Voronoi diagram), defined by

$$d_{sm}(q,r) = \inf_{\gamma} \int_{\gamma} \frac{1}{F\left(1 - \alpha \frac{dh(\gamma(s))}{ds}\right)} ds$$

for any $q, r \in \Omega$, and calculating the minimum time necessary for the snowmobile with the speed F on flat land to move from (q, h(q)) to (r, h(r)) along the land path $\gamma^* : \gamma^*(s) = (\gamma(s), h(\gamma(s)))$ associated with the domain path $\gamma : [0, 1] \to \Omega$, $\gamma(0) = q$, $\gamma(1) = r$ (the infimum is taken over all possible paths γ , and α is a positive constant).

A snowmobile goes uphill more slowly than downhill. The situation is opposite for a forest fire: the frontier of the fire goes uphill faster than downhill. This situation can be modeled using a negative value of α . The resulting distance is called the **forest-fire distance**, and the resulting Voronoi diagram is called the *forest-fire Voronoi diagram*.

• Skew distance

Let T be a *tilted plane* in \mathbb{R}^3 , obtained by rotating the x_1x_2 -plane around the x_1 -axis through the angle α , $0 < \alpha < \frac{\pi}{2}$, with the coordinate system obtained by taking the coordinate system of the x_1x_2 -plane, accordingly rotated. For a point $q \in T$, $q = (x_1(q), x_2(q))$, define the *height* h(q) as its x_3 -coordinate in \mathbb{R}^3 . Thus, $h(q) = x_2(q) \cdot \sin \alpha$. Let $P = \{p_1, \ldots, p_k\} \subset T$, k > 2.

The **skew distance** [AACL98] d_{skew} is the Voronoi generation distance of the generalized Voronoi diagram $V(P, d_{skew}, T)$ (skew Voronoi diagram), defined by

$$d_{skew}(q,r) = d_E(q,r) + (h(r) - h(q)) = d_E(q,r) + \sin \alpha (x_2(r) - x_2(q))$$

or, more generally, by

$$d_{skew}(q,r) = d_E(q,r) + k(x_2(r) - x_2(q))$$

for all $q, r \in T$, where d_E is the ordinary Euclidean distance, and $k \geq 0$ is a constant.

20.3 Other Voronoi generation distances

• Voronoi distance for line segments

The Voronoi distance for (a set of) line segments d_{sl} is the Voronoi generation distance of the generalized Voronoi diagram $V(A, d_{sl}, \mathbb{R}^2)$ (line Voronoi diagram, generated by straight line segments), defined by

$$d_{sl}(x, A_i) = \inf_{y \in A_i} d_E(x, y),$$

where the generator set $A = \{A_1, \ldots, A_k\}$, $k \geq 2$, is a set of pairwise disjoint straight line segments $A_i = [a_i, b_i]$, and d_E is the ordinary Euclidean distance. In fact,

$$d_{sl}(x,A_i) = \begin{cases} d_E(x,a_i), & \text{if} & x \in R_{a_i}, \\ d_E(x,b_i), & \text{if} & x \in R_{b_i}, \\ d_E(x-a_i, \frac{(x-a_i)^T(b_i-a_i)}{d_E^2(a_i,b_i)}(b_i-a_i)), & \text{if} & x \in \mathbb{R}^2 \backslash \{R_{a_i} \cup R_{b_i}\}, \end{cases}$$

where $R_{a_i} = \{x \in \mathbb{R}^2 : (b_i - a_i)^T (x - a_i) < 0\}, R_{b_i} = \{x \in \mathbb{R}^2 : (a_i - b_i)^T (x - b_i) < 0\}.$

• Voronoi distance for arcs

The Voronoi distance for (a set of circle) arcs d_{ca} is the Voronoi generation distance of the generalized Voronoi diagram $V(A, d_{ca}, \mathbb{R}^2)$ (line Voronoi diagram, generated by circle arcs), defined by

$$d_{ca}(x, A_i) = \inf_{y \in A_i} d_E(x, y),$$

where the generator set $A = \{A_1, \ldots, A_k\}, k \geq 2$, is a set of pairwise disjoint circle arcs A_i (less than or equal to a semicircle) with radius r_i centered at x_{c_i} , and d_E is the ordinary Euclidean distance. In fact,

$$d_{ca}(x, A_i) = \min\{d_E(x, a_i), d_E(x, b_i), |d_E(x, x_{c_i}) - r_i|\},\$$

where a_i and b_i are the end points of A_i .

• Voronoi distance for circles

The Voronoi distance for (a set of) circles d_{cl} is the Voronoi generation distance of a generalized Voronoi diagram $V(A, d_{cl}, \mathbb{R}^2)$ (line Voronoi diagram, generated by circles), defined by

$$d_{cl}(x, A_i) = \inf_{y \in A_i} d_E(x, y),$$

where the generator set $A = \{A_1, \ldots, A_k\}$, $k \geq 2$, is a set of pairwise disjoint circles A_i with radius r_i centered at x_{c_i} , and d_E is the ordinary Euclidean distance. In fact,

$$d_{cl}(x, A_i) = |d_E(x, x_{c_i}) - r_i|.$$

There exist different distances for the line Voronoi diagram, generated by circles. For example, $d_{cl}^{\star}(x, A_i) = d_E(x, x_{c_i}) - r_i$, or $d_{cl}^{\star}(x, A_i) = d_E^2(x, x_{c_i}) - r_i^2$ (the Laguerre Voronoi diagram).

• Voronoi distance for areas

The Voronoi distance for areas d_{ar} is the Voronoi generation distance of the generalized Voronoi diagram $V(A, d_{ar}, \mathbb{R}^2)$ (area Voronoi diagram), defined by

$$d_{ar}(x, A_i) = \inf_{y \in A_i} d_E(x, y),$$

where $A = \{A_1, \ldots, A_k\}, k \geq 2$, is a collection of pairwise disjoint connected closed sets (areas), and d_E is the ordinary Euclidean distance.

Note, that for any generalized generator set $A = \{A_1, \ldots, A_k\}, k \geq 2$, one can use as the Voronoi generation distance the **Hausdorff distance** from a point x to a set A_i : $d_{Haus}(x, A_i) = \sup_{y \in A_i} d_E(x, y)$, where d_E is the ordinary Euclidean distance.

Cylindrical distance

The **cylindrical distance** d_{cyl} is the **intrinsic metric** on the surface of a cylinder S which is used as the Voronoi generation distance in the *cylindrical Voronoi diagram* $V(P, d_{cyl}, S)$. If the axis of a cylinder with unit radius is placed at the x_3 -axis in \mathbb{R}^3 , the cylindrical distance between any points $x, y \in S$ with the cylindrical coordinates $(1, \theta_x, z_x)$ and $(1, \theta_y, z_y)$ is given by

$$d_{cyl}(x,y) = \begin{cases} \sqrt{(\theta_x - \theta_y)^2 + (z_x - z_y)^2}, & \text{if } \theta_y - \theta_x \le \pi, \\ \sqrt{(\theta_x + 2\pi - \theta_y)^2 + (z_x - z_y)^2}, & \text{if } \theta_y - \theta_x > \pi. \end{cases}$$

• Cone distance

The **cone distance** d_{con} is the **intrinsic metric** on the surface of a cone S which is used as the Voronoi generation distance in the *conic Voronoi diagram* $V(P, d_{con}, S)$. If the axis of the cone S is placed at the x_3 -axis in \mathbb{R}^3 , and the radius of the circle made by the intersection of the cone S with the x_1x_2 -plane is equal to one, then the cone distance between any points $x, y \in S$ is given by

$$d_{con}(x,y) = \begin{cases} \sqrt{r_x^2 + r_y^2 - 2r_x r_y \cos(\theta_y' - \theta_x')}, \\ \text{if } \theta_y' \le \theta_x' + \pi \sin(\alpha/2), \\ \sqrt{r_x^2 + r_y^2 - 2r_x r_y \cos(\theta_x' + 2\pi \sin(\alpha/2) - \theta_y')}, \\ \text{if } \theta_y' > \theta_x' + \pi \sin(\alpha/2), \end{cases}$$

where (x_1, x_2, x_3) are the Cartesian coordinates of a point x on S, α is the top angle of the cone, θ_x is the counterclockwise angle from the x_1 -axis

to the ray from the origin to the point $(x_1, x_2, 0)$, $\theta'_x = \theta_x \sin(\alpha/2)$, $r_x = \sqrt{x_1^2 + x_2^2 + (x_3 - \coth(\alpha/2))^2}$ is the straight line distance from the top of the cone to the point (x_1, x_2, x_3) .

• Voronoi distances of order m

Given a finite set A of objects in a metric space (S,d), and an integer $m \geq 1$, consider the set of all m-subsets M_i of A (i.e., $M_i \subset A$, and $|M_i| = m$). The Voronoi diagram of order m of A is a partition of S into Voronoi regions $V(M_i)$ of m-subsets of A in such a way that $V(M_i)$ contains all points $s \in S$ which are "closer" to M_i than to any other m-set M_j : d(s,x) < d(s,y) for any $x \in M_i$ and $y \in S \setminus M_i$. This diagram provides first, second, ..., m-th closest neighbors of a point in S.

Such diagrams can be defined in terms of some "distance function" $D(s, M_i)$, in particular, some m-hemi-metric (cf. Chap. 3) on S. For $M_i = \{a_i, b_i\}$, there were considered the functions $|d(s, a_i) - d(s, b_i)|$, $d(s, a_i) + d(s, b_i)$, $d(s, a_i) \cdot d(s, b_i)$, as well as 2-metrics $d(s, a_i) + d(s, b_i) + d(a_i, b_i)$ and the area of triangle (s, a_i, b_i) .

Chapter 21 Image and Audio Distances

21.1 Image distances

Image Processing treats signals such as photographs, video, or tomographic output. In particular, Computer Graphics consists of image synthesis from some abstract models, while Computer Vision extracts some abstract information: say, the 3D (i.e., three-dimensional) description of a scene from video footage of it. From about 2000, analog image processing (by optical devices) gave way to digital processing, and, in particular, digital image editing (for example, processing of images taken by popular digital cameras).

Computer graphics (and our brains) deals with vector graphics images, i.e., those represented geometrically by curves, polygons, etc. A raster graphics image (or digital image, bitmap) in 2D is a representation of a 2D image as a finite set of digital values, called pixels (short for picture elements) placed on a square grid \mathbb{Z}^2 or a hexagonal grid. Typically, the image raster is a square $2^k \times 2^k$ grid with k = 8, 9 or 10.

Video images and tomographic or MRI (obtained by cross-sectional slices) images are 3D (2D plus time); their digital values are called voxels (volume elements). The **spacing distance** between two pixels in one slice is referred to as the *interpixel distance*, while the **spacing distance** between slices is the *interslice distance*.

A digital binary image corresponds to only two values 0, 1 with 1 being interpreted as logical "true" and displayed as black; so, such image is identified with the set of black pixels. The elements of a binary 2D image can be seen as complex numbers x + iy, where (x, y) are coordinates of a point on the real plane \mathbb{R}^2 . A continuous binary image is a (usually, compact) subset of a locally compact metric space (usually, Euclidean space \mathbb{E}^n with n = 2, 3).

The gray-scale images can be seen as point-weighted binary images. In general, a fuzzy set is a point-weighted set with weights (membership values); see **metrics between fuzzy sets** in Chap. 1. For the gray-scale images, xyi-representation is used, where plane coordinates (x, y) indicate shape, while the weight i (short for intensity, i.e., brightness) indicates texture (intensity est pattern). Sometimes, the matrix ((ixy)) of gray-levels is used. The

brightness histogram of a gray-scale image provides the frequency of each brightness value found in that image. If an image has m brightness levels (bins of gray-scale), then there are 2^m different possible intensities. Usually, m=8 and numbers $0,1,\ldots,255$ represent the intensity range from black to white; other typical values are m=10,12,14,16. Humans can differ between around 350,000 different colors but between only 30 different gray-levels; so, color has much higher discriminatory power.

For color images, (RGB)-representation is the b known, where space coordinates R, G, B indicate red, green and blue levels; a 3D histogram provides brightness at each point. Among many other 3D color models (spaces) are: (CMY) cube (Cyan, Magenta, Yellow colors), (HSL) cone (Hue-color type given as an angle, Saturation in %, Luminosity in %), and (YUV), (YIQ) used, respectively, in PAL, NTSC television. CIE-approved conversion of (RGB) into luminance (luminosity) of gray-level is 0.299R + 0.587G + 0.114B.

The *color histogram* is a feature vector of length n (usually, n = 64, 256) with components representing either the total number of pixels, or the percentage of pixels of a given color in the image.

Images are often represented by feature vectors, including color histograms, color moments, textures, shape descriptors, etc. Examples of feature spaces are: raw intensity (pixel values), edges (boundaries, contours, surfaces), salient features (corners, line intersections, points of high curvature), and statistical features (moment invariants, centroids). Typical video features are in terms of overlapping frames and motions.

Image Retrieval (similarity search) consists of (as for other data: audio recordings, DNA sequences, text documents, time-series, etc.) finding images whose features have values either mutual similarity, or similarity to a given query or in a given range.

There are two methods to compare images directly: intensity-based (color and texture histograms), and geometry-based (shape representations by *medial axis*, *skeletons*, etc.). The imprecise term *shape* is used for the extent (silhouette) of the object, for its local geometry or geometrical pattern (conspicuous geometric details, points, curves, etc.), or for that pattern modulo a similarity transformation group (translations, rotations, and scalings). The imprecise term *texture* means all that is left after color and shape have been considered, or it is defined via structure and randomness.

The similarity between vector representations of images is measured by the usual practical distances: l_p -metrics, weighted editing metrics, Tanimoto distance, cosine distance, Mahalanobis distance and its extension, Earth Mover distance. Among probabilistic distances, the following ones are most used: Bhattacharya 2, Hellinger, Kullback–Leibler, Jeffrey and (especially, for histograms) χ^2 -, Kolmogorov–Smirnov, Kuiper distances.

The main distances applied for compact subsets X and Y of \mathbb{R}^n (usually, n = 2, 3) or their digital versions are: **Asplund metric**, **Shephard metric**,

symmetric difference semi-metric $Vol(X\Delta Y)$ (see Nykodym metric, area deviation, digital volume metric and their normalizations) and variations of the Hausdorff distance (see below).

For Image Processing, the distances below are between "true" and approximated digital images, in order to assess the performance of algorithms. For Image Retrieval, distances are between feature vectors of a query and reference.

• Color distances

A color space is a 3-parameter description of colors. The need for exactly three parameters comes from the existence of three kinds of receptors (cells on the retina) in the human eye: for short, middle and long wavelengths, corresponding to blue, green, and red. In fact, their respective sensitivity peaks are situated around 570, 543 and 442 nm, while wavelength limits of extreme violet and red are about 700 and 390 nm, respectively. Some women are tetrachromats, i.e., they have a fourth type of color receptor. The zebrafish Danio rerio has cone cells sensitive to red, green, blue, and ultraviolet light.

The CIE (International Commission on Illumination) derived (XYZ) color space in 1931 from the (RGB)-model and measurements of the human eye. In the CIE (XYZ) color space, the values X, Y and Z are also roughly red, green and blue, respectively.

The basic assumption of Colorimetry, supported experimentally (Indow 1991), is that the perceptual color space admits a metric, the true **color distance**. This metric is expected to be locally Euclidean, i.e., a **Riemannian metric**. Another assumption is that there is a continuous mapping from the metric space of *photic* (light) stimuli to this metric space.

Cf. probability-distance hypothesis in Chap. 28 that the probability with which one stimulus is discriminated from another is a (continuously increasing) function of some subjective quasi-metric between these stimuli.

Such a uniform color scale, where equal distances in the color space correspond to equal differences in color, is not obtained yet and existing color distances are various approximations of it. A first step in this direction was given by MacAdam ellipses, i.e., regions on a chromaticity (x,y) diagram which contains all colors looking indistinguishable to the average human eye; cf. JND (just-noticeable difference) video quality metric. Those 25 ellipses define, for any $\epsilon > 0$, the MacAdam metric in a color space as the metric for which those ellipses are circles of radius ϵ . Here $x = \frac{X}{X+Y+Z}$ and $y = \frac{Y}{X+Y+Z}$ are projective coordinates, and the colors of the chromaticity diagram occupy a region of the real projective plane.

The CIE $(L^*a^*b^*)$ (CIELAB) is an adaptation of CIE 1931 (XYZ) color space; it gives a partial linearization of the MacAdam color metric. The parameters L^*, a^*, b^* of the most complete model are derived from L, a, b which are: the luminance L of the color from black L = 0 to white L = 100, its position a between green a < 0 and red a > 0, and its position b between green b < 0 and yellow b > 0.

• Average color distance

For a given 3D color space and a list of n colors, let (c_{i1}, c_{i2}, c_{i3}) be the representation of the i-th color of the list in this space. For a color histogram $x = (x_1, \ldots, x_n)$, its average color is the vector $(x_{(1)}, x_{(2)}, x_{(3)})$, where $x_{(j)} = \sum_{i=1}^{n} x_i c_{ij}$ (for example, the average red, blue and green values in (RGB)).

The average color distance between two color histograms [HSEFN95] is the Euclidean distance of their average colors.

• Color component distance

Given an image (as a subset of \mathbb{R}^2), let p_i denote the area percentage of this image occupied by the color c_i . A color component of the image is a pair (c_i, p_i) .

The **color component distance** (Ma, Deng and Manjunath 1997) between color components (c_i, p_i) and (c_j, p_j) is defined by

$$|p_i - p_j| \cdot d(c_i, c_j),$$

where $d(c_i, c_j)$ is the distance between colors c_i and c_j in a given color space. Mojsilović, Hu and Soljanin (2002) developed an **Earth Mover distance**-like modification of this distance.

• Riemannian color space

The proposal to measure perceptual dissimilarity of colors by a *Riemannian metric* (cf. Chap. 7) on a strictly convex cone $C \subset \mathbb{R}^3$ comes from von Helmholtz (1892) and Luneburg (1947).

Roughly, it was shown in [Resn74] that the only such GL-homogeneous cones C (i.e., the group of all orientation preserving linear transformations of \mathbb{R}^3 , carrying C into itself, acts transitively on C) are either $C_1 = \mathbb{R}_{>0} \times (\mathbb{R}_{>0} \times \mathbb{R}_{>0})$, or $C_2 = \mathbb{R}_{>0} \times C'$, where C' is the set of 2×2 real symmetric matrices with determinant 1. The first factor $\mathbb{R}_{>0}$ can be identified with variation of brightness and the other with the set of lights of a fixed brightness. Let α_i be some positive constants.

The **Stiles color metric** (Stiles 1946) is the *GL*-invariant Riemannian metric on $C_1 = \{(x_1, x_2, x_3) \in \mathbb{R}^3 : x_i > 0\}$ given by the *line element*

$$ds^2 = \sum_{i=1}^3 \alpha_i (\frac{dx_i}{x_i})^2.$$

The **Resnikoff color metric** (Resnikoff 1974) is the GL-invariant Riemannian metric on $C_2 = \{(x, u) : x \in \mathbb{R}_{>0}, u \in C'\}$ given by the *line element*

$$ds^2 = \alpha_1 \left(\frac{dx}{x}\right)^2 + \alpha_2 ds_{C'}^2,$$

where $ds_{C'}^2$, is the **Poincare metric** (cf. Chap. 6) on C'; so, C_2 with this metric is not isometric to a Euclidean space.

Histogram intersection quasi-distance

Given two color histograms $x = (x_1, \ldots, x_n)$ and $y = (y_1, \ldots, y_n)$ (with x_i, y_i representing the number of pixels in the bin i), the **histogram** intersection quasi-distance between them (cf. intersection distance in Chap. 17) is (Swain and Ballard 1991) defined by

$$1 - \frac{\sum_{i=1}^{n} \min\{x_i, y_i\}}{\sum_{i=1}^{n} x_i}.$$

For normalized histograms (total sum is 1) the above quasi-distance becomes the usual l_1 -metric $\sum_{i=1}^n |x_i - y_i|$. The normalized cross correlation (Rosenfeld and Kak 1982) between x and y is a similarity, defined by $\frac{\sum_{i=1}^{n} x_i, y_i}{\sum_{i=1}^{n} x_i^2}$. Histogram quadratic distance

Given two color histograms $x = (x_1, \ldots, x_n)$ and $y = (y_1, \ldots, y_n)$ (usually, n=256 or n=64) representing the color percentages of two images, their histogram quadratic distance (used in IBM's Query By Image Content system) is the **Mahalanobis distance**, defined by

$$\sqrt{(x-y)^T A(x-y)},$$

where $A = ((a_{ij}))$ is a symmetric positive-definite matrix, and the weight a_{ij} is some, perceptually justified, similarity between colors i and j. For example (see [HSEFN95]), $a_{ij} = 1 - \frac{d_{ij}}{\max_{1 \le p,q \le n} d_{pq}}$, where d_{ij} is the Euclidean distance between 3-vectors representing i and j in some color space. Another definition is given by $a_{ij} = 1 - \frac{1}{\sqrt{5}}((v_i - v_j)^2 + (s_i \cos h_i - v_j)^2)$ $(s_j \cos h_j)^2 + (s_i \sin h_i - s_j \sin h_j)^2)^{\frac{1}{2}}$, where (h_i, s_i, v_i) and (h_j, s_j, v_j) are the representations of the colors i and j in the color space (HSV).

Histogram diffusion distance

Given two histogram-based descriptors $x = (x_1, \dots, x_n)$ and y = (y_1,\ldots,y_n) , their **histogram diffusion distance** (Ling and Okada 2006) is defined by

$$\int_0^T ||u(t)||_1 dt,$$

where T is a constant, and u(t) is a heat diffusion process with initial condition u(0) = x - y. In order to approximate the diffusion, the initial condition is convoluted with a Gaussian window; then the sums of l_1 -norms after each convolution approximate the integral.

This distance was generalized in Yan, Wang, Liu, Lu and Ma (2007).

Gray-scale image distances

Let f(x) and g(x) denote the brightness values of two digital gray-scale images f and q at the pixel $x \in X$, where X is a raster of pixels. Any distance between point-weighted sets (X, f) and (X, g) (for example, the Earth Mover distance) can be applied for measuring distances between f and g. However, the main used distances (called also *errors*) between the images f and g are:

- 1. The root mean-square error $RMS(f,g) = \left(\frac{1}{|X|}\sum_{x \in X}(f(x) g(x))^2\right)^{\frac{1}{2}}$ (a variant is to use the l_1 -norm |f(x) g(x)| instead of the l_2 -norm)
- 2. The signal-to-noise ratio $SNR(f,g) = \left(\frac{\sum_{x \in X} g(x)^2}{\sum_{x \in X} (f(x) g(x))^2}\right)^{\frac{1}{2}}$
- 3. The pixel misclassification error rate $\frac{1}{|X|}|\{x \in X : f(x) \neq g(x)\}|$ (normalized **Hamming distance**)
- 4. The frequency root mean-square error $\left(\frac{1}{|U|^2}\sum_{u\in U}(F(u)-G(u))^2\right)^{\frac{1}{2}}$, where F and G are the discrete Fourier transforms of f and g, respectively, and U is the frequency domain
- 5. The Sobolev norm of order δ error $\left(\frac{1}{|U|^2}\sum_{u\in U}(1+|\eta_u|^2)^{\delta}(F(u)-G(u))^2\right)^{\frac{1}{2}}$, where $0<\delta<1$ is fixed (usually, $\delta=\frac{1}{2}$), and η_u is the 2D frequency vector associated with position u in the frequency domain U

Cf. metrics between fuzzy sets in Chap. 1.

• Image compression L_p -metric

Given a number r, $0 \le r < 1$, the **image compression** L_p -**metric** is the usual L_p -**metric** on $\mathbb{R}^{n^2}_{\ge 0}$ (the set of gray-scale images seen as $n \times n$ matrices) with p being a solution of the equation $r = \frac{p-1}{2p-1} \cdot e^{\frac{p}{2p-1}}$. So, p = 1, 2, or ∞ for, respectively, $r = 0, r = \frac{1}{3}e^{\frac{2}{3}} \approx 0.65$, or $r \ge \frac{\sqrt{e}}{2} \approx 0.82$. Here r estimates the *informative* (i.e., filled with non-zeros) part of the image. According to [KKN02], it is the best quality metric to select a lossy compression scheme.

• Chamfering distances

The **chamfering distances** are distances approximating Euclidean distance by a weighted path distance on the graph $G = (\mathbb{Z}^2, E)$, where two pixels are neighbors if one can be obtained from another by an *one-step move* on \mathbb{Z}^2 . The list of permitted moves is given, and a **prime distance**, i.e., a positive weight (see Chap. 19), is associated to each type of such move.

An (α, β) -chamfer metric corresponds to two permitted moves – with l_1 -distance 1 and with l_{∞} -distance 1 (diagonal moves only) – weighted α and β , respectively.

The main applied cases are $(\alpha, \beta) = (1, 0)$ (the **city-block metric**, or 4-metric), (1,1) (the **chessboard metric**, or 8-metric), $(1, \sqrt{2})$ (the Montanari metric), (3,4) (the (3,4)-metric), (2,3) (the Hilditch–Rutovitz metric), (5,7) (the Verwer metric).

The Borgefors metric corresponds to three permitted moves – with l_1 -distance 1, with l_{∞} -distance 1 (diagonal moves only), and knight moves – weighted 5, 7 and 11, respectively.

An 3D-chamfer metric (or (α, β, γ) -chamfer metric) is the weighted path metric of the infinite graph with the vertex-set \mathbb{Z}^3 of voxels, two vertices being adjacent if their l_{∞} -distance is one, while weights α, β and γ are associated to 6 face, 12 edge and 8 corner neighbors, respectively. If $\alpha = \beta = \gamma = 1$, we obtain the l_{∞} -metric. The (3,4,5)- and (1, 2, 3)-chamfer metrics are the most used ones for digital 3D images.

The Chaudhuri-Murthy-Chaudhuri metric between sequences $x = (x_1, \ldots, x_m)$ and $y = (y_1, \ldots, y_n)$ is defined by

$$|x_{i(x,y)} - y_{i(x,y)}| + \frac{1}{1 + \lceil \frac{n}{2} \rceil} \sum_{1 \le i \le n, i \ne i(x,y)} |x_i - y_i|,$$

where the maximum value of $x_i - y_i$ is attained for i = i(x, y). For n = 2it is the $(1, \frac{3}{2})$ -chamfer metric.

Earth Mover distance

The Earth Mover distance is a discrete form of the Monge-Kantorovich distance. Roughly, it is the minimal amount of work needed to transport earth or mass from one position (properly spread in space) to the other (a collection of holes). For any two finite sequences (x_1, \ldots, x_m) and (y_1, \ldots, y_n) over a metric space (X, d), consider signatures, i.e., point-weighted sets $P_1 = (p_1(x_1), \dots, p_1(x_m))$ and $P_2 =$ $(p_2(y_1), \ldots, p_2(y_n))$. For example [RTG00], signatures can represent clustered color or texture content of images: elements of X are centroids of clusters, and $p_1(x_i), p_2(y_i)$ are sizes of corresponding clusters. The ground distance d is a color distance, say, the Euclidean distance in 3D CIE $(L^*a^*b^*)$ color space.

Let $W_1 = \sum_i p_1(x_i)$ and $W_2 = \sum_i p_2(y_i)$ be the total weights of P_1 and P_2 , respectively. Then the **Earth Mover distance** (or transport distance) between signatures P_1 and P_2 is defined as the function

$$\frac{\sum_{i,j} f_{ij}^* d(x_i, y_j)}{\sum_{i,j} f_{ij}^*},$$

where the $m \times n$ matrix $S^* = ((f_{ij}^*))$ is an optimal, i.e., minimizing $\sum_{i,j} f_{ij} d(x_i, y_j)$, flow. A flow (of the weight of the earth) is an $m \times n$ matrix $S = ((f_{ij}))$ with following constraints:

- 1. All $f_{ij} \geq 0$.
- 2. $\sum_{i,j} f_{ij} = \min\{W_1, W_2\}.$ 3. $\sum_i f_{ij} \leq p_2(y_j)$, and $\sum_j f_{ij} \leq p_1(x_i).$

So, this distance is the average ground distance d that weights travel during an optimal flow.

In the case $W_1 = W_2$, the above two inequalities 3 became equalities. Normalizing signatures to $W_1 = W_2 = 1$ (which not changes the distance) allows us to see P_1 and P_2 as probability distributions of random variables, say, X and Y. Then $\sum_{i,j} f_{ij} d(x_i, y_j)$ is just $\mathbb{E}_S[d(X, Y)]$, i.e., the Earth Mover distance coincides, in this case, with the **Kantorovich–Mallows–Monge–Wasserstein metric**.

For, say, $W_1 < W_2$, it is not a metric in general. However, replacing the inequalities 3 in the above definition by equalities:

3'. $\sum_{i} f_{ij} = p_2(y_j)$, and $\sum_{j} f_{ij} = \frac{p_1(\tilde{x_i})W_1^{\dagger}}{W_2}$, produces the Giannopoulos–Veltkamp's **proportional transport semi-metric**.

• Parameterized curves distance

The shape can be represented by a parametrized curve on the plane. Usually, such a curve is *simple*, i.e., it has no self-intersections. Let X = X(x(t)) and Y = Y(y(t)) be two parametrized curves, where the (continuous) parametrization functions x(t) and y(t) on [0,1] satisfy x(0) = y(0) = 0 and x(1) = y(1) = 1.

The most used **parametrized curves distance** is the minimum, over all monotone increasing parametrizations x(t) and y(t), of the maximal Euclidean distance $d_E(X(x(t)), Y(y(t)))$. It is the Euclidean special case of the **dogkeeper distance** which is, in turn, the **Fréchet metric** for the case of curves. Among variations of this distance are dropping the monotonicity condition of the parametrization, or finding the part of one curve to which the other has the smallest such distance [VeHa01].

• Non-linear elastic matching distances

Consider a digital representation of curves. Let $r \geq 1$ be a constant, and let $A = \{a_1, \ldots, a_m\}$, $B = \{b_1, \ldots, b_n\}$ be finite ordered sets of consecutive points on two closed curves. For any order-preserving correspondence f between all points of A and all points of B, the $stretch\ s(a_i, b_j)$ of $(a_i, f(a_i) = b_j)$ is r if either $f(a_{i-1}) = b_j$ or $f(a_i) = b_{j-1}$, or zero otherwise.

The **relaxed non-linear elastic matching distance** is the minimum, over all such f, of $\sum (s(a_i, b_j) + d(a_i, b_j))$, where $d(a_i, b_j)$ is the difference between the tangent angles of a_i and b_j . It is a **near-metric** for some r. For r = 1, it is called the **non-linear elastic matching distance**.

• Turning function distance

For a plane polygon P, its turning function $T_P(s)$ is the angle between the counterclockwise tangent and the x-axis as a function of the arc length s. This function increases with each left hand turn and decreases with right hand turns.

Given two polygons of equal perimeters, their turning function distance is the L_p -metric between their turning functions.

• Size function distance

For a plane graph G=(V,E) and a measuring function f on its vertexset V (for example, the distance from $v\in V$ to the center of mass of V), the size function $S_G(x,y)$ is defined, on the points $(x,y) \in \mathbb{R}^2$, as the number of connected components of the restriction of G on vertices $\{v \in V : f(v) \leq y\}$ which contain a point v' with $f(v') \leq x$.

Given two plane graphs with vertex-sets belonging to a raster $R \subset \mathbb{Z}^2$, their Uras-Verri's **size function distance** is the normalized l_1 -distance between their size functions over raster pixels.

• Reflection distance

For a finite union A of plane curves and each point $x \in \mathbb{R}^2$, let V_A^x denote the union of intervals (x, a), $a \in A$, which are visible from x, i.e., $(x, a) \cap A = \emptyset$. Denote by ρ_A^x the area of the set $\{x + v \in V_A^x : x - v \in V_A^x\}$.

The Hagedoorn-Veltkamp's **reflection distance** between finite unions A and B of plane curves is the normalized l_1 -distance between the corresponding functions ρ_A^x and ρ_B^x , defined by

$$\frac{\int_{\mathbb{R}^2} |\rho_A^x - \rho_B^x| dx}{\int_{\mathbb{R}^2} \max\{\rho_A^x, \rho_B^x\} dx}.$$

• Distance transform

Given a metric space $(X = \mathbb{Z}^2, d)$ and a binary digital image $M \subset X$, the **distance transform** is a function $f_M : X \to \mathbb{R}_{\geq 0}$, where $f_M(x) = \inf_{u \in M} d(x, u)$ is the **point-set distance** d(x, M). Therefore, a distance transform can be seen as a gray-scale digital image where each pixel is given a label (a gray-level) which corresponds to the distance to the nearest pixel of the background. Distance transforms, in Image Processing, are also called distance fields and, especially, **distance maps**; but we reserve the last term only for this notion in any metric space. A distance transform of a shape is the distance transform with M being the boundary of the image. For $X = \mathbb{R}^2$, the graph $\{(x, f(x)) : x \in X\}$ of d(x, M) is called the Voronoi surface of M.

• Medial axis and skeleton

Let (X,d) be a metric space, and let M be a subset of X. The **medial axis** of X is the set $MA(X) = \{x \in X : |\{m \in M : d(x,m) = d(x,M)\}| \geq 2\}$, i.e., all points of X which have in M at least two **elements of best approximation**. MA(X) consists of all points of boundaries of *Voronoi regions* of points of M. The cut locus of X is the closure $\overline{MA(X)}$ of the medial axis. The medial axis transform MAT(X) is the point-weighted set MA(X) (the restriction of the **distance transform** on MA(X)) with d(x,M) being the weight of $x \in X$.

If (as usual in applications) $X \subset \mathbb{R}^n$ and M is the boundary of X, then the **skeleton** Skel(X) of X is the set of the centers of all d-balls inscribed in X and not belonging to any other such ball; so, Skel(X) = MA(X). The skeleton with M being continuous boundary is a limit of Voronoi diagrams as the number of the generating points becomes infinite. For 2D

binary images X, the skeleton is a curve, a single-pixel thin one, in the digital case. The *exoskeleton* of X is the skeleton of the complement of X, i.e., of the background of the image for which X is the foreground.

• Procrustes distance

The shape of a form (configuration of points in \mathbb{R}^2), seen as expression of translation-, rotation- and scale-invariant properties of form, can be represented by a sequence of landmarks, i.e., specific points on the form, selected accordingly to some rule. Each landmark point a can be seen as an element $(a', a'') \in \mathbb{R}^2$ or an element $a' + a'' i \in \mathbb{C}$.

Consider two shapes x and y, represented by their landmark vectors (x_1, \ldots, x_n) and (y_1, \ldots, y_n) from \mathbb{C}^n . Suppose that x and y are corrected for translation by setting $\sum_t x_t = \sum_t y_t = 0$. Then their **Procrustes distance** is defined by

$$\sqrt{\sum_{t=1}^{n} |x_t - y_t|^2},$$

where two forms are, first, optimally (by least squares criterion) aligned to correct for scale, and their **Kendall shape distance** is defined by

$$\arccos\sqrt{\frac{(\sum_t x_t \overline{y}_t)(\sum_t y_t \overline{x}_t)}{(\sum_t x_t \overline{x}_t)(\sum_t y_t \overline{y}_t)}},$$

where $\overline{\alpha} = a' - a''i$ is the *complex conjugate* of $\alpha = a' + a''i$.

• Tangent distance

For any $x \in \mathbb{R}^n$ and a family of transformations $t(x,\alpha)$, where $\alpha \in \mathbb{R}^k$ is the vector of k parameters (for example, the scaling factor and rotation angle), the set $M_x = \{t(x,\alpha) : \alpha \in \mathbb{R}^k\} \subset \mathbb{R}^n$ is a manifold of dimension at most k. It is a curve if k=1. The minimum Euclidean distance between manifolds M_x and M_y would be a useful distance since it is invariant with respect to transformations $t(x,\alpha)$. However, the computation of such a distance is too difficult in general; so, M_x is approximated by its tangent subspace at x: $\{x + \sum_{i=1}^k \alpha_k x^i : \alpha \in \mathbb{R}^k\} \subset \mathbb{R}^n$, where the tangent vectors x^i , $1 \le i \le k$, spanning it are the partial derivatives of $t(x,\alpha)$ with respect to α . The **one-sided** (or directed) **tangent distance** between elements x and y of \mathbb{R}^n is a quasi-distance d, defined by

$$\sqrt{\min_{\alpha} ||x + \sum_{i=1}^{k} \alpha_k x^i - y||^2}.$$

The Simard-Le Cun-Denker's **tangent distance** is defined by $\min\{d(x,y), d(y,x)\}.$

Cf. metric cone structure, tangent metric cone in Chap. 1.

• Pixel distance

Consider two digital images, seen as binary $m \times n$ matrices $x = ((x_{ij}))$ and $y = ((y_{ij}))$, where a pixel x_{ij} is black or white if it is equal to 1 or 0, respectively.

For each pixel x_{ij} , the fringe distance map to the nearest pixel of opposite color $D_{BW}(x_{ij})$ is the number of fringes expanded from (i,j) (where each fringe is composed by the pixels that are at the same distance from (i,j)) until the first fringe holding a pixel of opposite color is reached.

The **pixel distance** (Smith, Bourgoin, Sims and Voorhees 1994) is defined by

$$\sum_{1 \le i \le m} \sum_{1 \le j \le n} |x_{ij} - y_{ij}| (D_{BW}(x_{ij}) + D_{BW}(y_{ij})).$$

• Figure of merit quasi-distance

Given two binary images, seen as non-empty finite subsets A and B of a finite metric space (X, d), their Pratt's figure of merit quasi-distance is defined by

$$\left(\max\{|A|, |B|\} \sum_{x \in B} \frac{1}{1 + \alpha d(x, A)^2}\right)^{-1},$$

where α is a scaling constant (usually, $\frac{1}{9}$), and $d(x, A) = \min_{y \in A} d(x, y)$ is the **point-set distance**.

Similar quasi-distances are Peli-Malah's mean error distance $\frac{1}{|B|} \sum_{x \in B} d(x, A)$, and the mean square error distance $\frac{1}{|B|} \sum_{x \in B} d(x, A)^2$.

• p-th order mean Hausdorff distance

Given $p \ge 1$ and two binary images, seen as non-empty subsets A and B of a finite metric space (say, a raster of pixels) (X,d), their p-th order mean Hausdorff distance is [Badd92] a normalized L_p -Hausdorff distance, defined by

$$\left(\frac{1}{|X|}\sum_{x\in X}|d(x,A)-d(x,B)|^p\right)^{\frac{1}{p}},$$

where $d(x, A) = \min_{y \in A} d(x, y)$ is the **point-set distance**. The usual Hausdorff metric is proportional to the ∞ -th order mean Hausdorff distance.

Venkatasubraminian's Σ -Hausdorff distance $d_{dHaus}(A, B) + d_{dHaus}(B, A)$ is equal to $\sum_{x \in A \cup B} |d(x, A) - d(x, B)|$, i.e., it is a version of L_1 -Hausdorff distance.

Another version of the first order mean Hausdorff distance is Lindstrom-Turk's mean geometric error (1998) between two images, seen as surfaces A and B. It is defined by

$$\frac{1}{Area(A) + Area(B)} \left(\int_{x \in A} d(x,B) dS + \int_{x \in B} d(x,A) dS \right),$$

where Area(A) denotes the area of A. If the images are seen as finite sets A and B, their mean geometric error is defined by

$$\frac{1}{|A|+|B|} \left(\sum_{x \in A} d(x,B) + \sum_{x \in B} d(x,A) \right).$$

Modified Hausdorff distance

Given two binary images, seen as non-empty finite subsets A and B of a finite metric space (X, d), their Dubuisson-Jain's modified Hausdorff distance (1994) is defined as the maximum of point-set distances averaged over A and B:

$$\max \left\{ \frac{1}{|A|} \sum_{x \in A} d(x, B), \frac{1}{|B|} \sum_{x \in B} d(x, A) \right\}.$$

• Partial Hausdorff quasi-distance

Given two binary images, seen as non-empty subsets A, B of a finite metric space (X,d), and integers k,l with $1 \leq k \leq |A|$, $1 \leq l \leq |B|$, their Huttenlocher-Rucklidge's partial (k, l)-Hausdorff quasi-distance (1992) is defined by

$$\max\{k_{x\in A}^{th}d(x,B), l_{x\in B}^{th}d(x,A)\},\$$

where $k_{x\in A}^{th}d(x,B)$ means the k-th (rather than the largest |A|-th ranked one) among |A| distances d(x, B) ranked in increasing order. The case $k = \lfloor \frac{|A|}{2} \rfloor, \ l = \lfloor \frac{|B|}{2} \rfloor$ corresponds to the modified median Hausdorff quasi-distance.

• Bottleneck distance

Given two binary images, seen as non-empty subsets A, B with |A| =|B| = m, of a metric space (X, d), their **bottleneck distance** is defined by

$$\min_{f} \max_{x \in A} d(x, f(x)),$$

where f is any bijective mapping between A and B.

Variations of the above distance are:

- 1. The minimum weight matching: $\min_f \sum_{x \in A} d(x, f(x))$ 2. The uniform matching: $\min_f \{\max_{x \in A} d(x, f(x)) \min_{x \in A} d(x, f(x)) \}$ f(x)
- 3. The **minimum deviation matching**: $\min_{f} \{ \max_{x \in A} d(x, f(x)) \max_{x \in A} d(x, f(x)) \}$ $\frac{1}{|A|} \sum_{x \in A} d(x, f(x))$

Given an integer t with $1 \le t \le |A|$, the t-bottleneck distance between A and B [InVe00] is the above minimum but with f being any mapping from A to B such that $|\{x \in A : f(x) = y\}| \le t$.

The cases t = 1 and t = |A| correspond, respectively, to the bottleneck distance, and the **directed Hausdorff distance** $d_{dHaus}(A, B) = \max_{x \in A} \min_{y \in B} d(x, y)$.

• Hausdorff distance up to G

Given a group (G, \cdot, id) acting on the Euclidean space \mathbb{E}^n , the **Hausdorff** distance up to G between two compact subsets A and B (used in Image Processing) is their **generalized** G-Hausdorff distance (see Chap. 1), i.e., the minimum of $d_{Haus}(A, g(B))$ over all $g \in G$. Usually, G is the group of all isometries or all translations of \mathbb{E}^n .

• Hyperbolic Hausdorff distance

For any compact subset A of \mathbb{R}^n , denote by MAT(A) its Blum's medial axis transform, i.e., the subset of $X = \mathbb{R}^n \times \mathbb{R}_{\geq 0}$, whose elements are all pairs $x = (x', r_x)$ of the centers x' and the radii r_x of the maximal inscribed (in A) balls, in terms of the Euclidean distance d_E in \mathbb{R}^n . (Cf. **medial axis and skeleton** transforms for the general case.)

The hyperbolic Hausdorff distance [ChSe00] is the Hausdorff metric on non-empty compact subsets MAT(A) of the metric space (X,d), where the hyperbolic distance d on X is defined, for its elements $x=(x',r_x)$ and $y=(y',r_y)$, by

$$\max\{0, d_E(x', y') - (r_v - r_x)\}.$$

• Non-linear Hausdorff metric

Given two compact subsets A and B of a metric space (X,d), their **non-linear Hausdorff metric** (or $Szatm\acute{a}ri-Rekeczky-Roska$ wave distance) is the **Hausdorff distance** $d_{Haus}(A \cap B, (A \cup B)^*)$, where $(A \cup B)^*$ is the subset of $A \cup B$ which forms a closed contiguous region with $A \cap B$, and the distances between points are allowed to be measured only along paths wholly in $A \cup B$.

• Video quality metrics

These metrics are between test and reference color video sequences, usually aimed at optimization of encoding/compression/decoding algorithms. Each of them is based on some perceptual model of the human vision system, the simplest ones being RMSE (root-mean-square error) and PSNR (peak signal-to-noise ratio) error measures. Among others, threshold metrics estimate the probability of detecting in video an artifact (i.e., a visible distortion that gets added to a video signal during digital encoding). Examples are: Sarnoff's JND (just-noticeable differences) metric, Winkler's PDM (perceptual distortion metric), and Watson's DVQ (digital video quality) metric. DVQ is an l_p -metric between feature vectors representing two video sequences. Some metrics measure special artifacts in the video: the appearance of block structure, blurriness, added "mosquito" noise (ambiguity in the edge direction), texture distortion, etc.

• Time series video distances

The time series video distances are objective wavelet-based spatial-temporal video quality metrics. A video stream x is processed into a time series x(t) (seen as a curve on coordinate plane) which is then (piecewise linearly) approximated by a set of n contiguous line segments that can be defined by n+1 endpoints (x_i, x_i') , $0 \le i \le n$, in the coordinate plane. In [WoPi99] are given the following (cf. Meehl distance) distances between video streams x and y:

1.
$$Shape(x,y) = \sum_{i=0}^{n-1} |(x'_{i+1} - x'_i) - (y'_{i+1} - y'_i)|.$$

2. $Offset(x,y) = \sum_{i=0}^{n-1} |\frac{x'_{i+1} + x'_i}{2} - \frac{y'_{i+1} + y'_i}{2}|.$

• Handwriting spatial gap distances

Automatic recognition of unconstrained handwritten texts (for example, legal amounts on bank checks or pre-hospital care reports) require measuring the spatial gaps between connected components in order to extract words.

Three most used ones, among handwriting spatial gap distances between two adjacent connected components x and y of text line, are:

Seni and Cohen (1994): the run-length (minimum horizontal Euclidean distance) between points of x and y;

Seni and Cohen (1994): the horizontal distance between the bounding boxes of x and y;

Mahadevan and Nagabushnam (1995): Euclidean distance between the convex hulls of x and y, on the line linking hull centroids.

21.2 Audio distances

Sound is the vibration of gas or air particles that causes pressure variations within our eardrums. Audio (speech, music, etc.) Signal Processing is the processing of analog (continuous) or, mainly, digital representation of the air pressure waveform of the sound. A sound spectrogram (or sonogram) is a visual three-dimensional representation of an acoustic signal. It is obtained either by a series of bandpass filters (an analog processing), or by application of the short-time Fourier transform to the electronic analog of an acoustic wave. Three axes represent time, frequency and intensity (acoustic energy). Often this three-dimensional curve is reduced to two dimensions by indicating the intensity with more thick lines or more intense gray or color values.

Sound is called *tone* if it is periodic (the lowest *fundamental* frequency plus its multiples, *harmonics* or *overtones*) and *noise*, otherwise. The frequency is measured in *cps* (the number of complete cycles per second) or Hz (Hertz). The range of audible sound frequencies to humans is typically 20 Hz–18 kHz. In fact, it is up to 20 kHz for most young adults, while 8 kHz in the elderly.

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The power P(f) of a signal is energy per unit of time; it is proportional to the square of signal's amplitude A(f). Decibel dB is the unit used to express the relative strength of two signals. One tenth of 1 dB is bel, the original outdated unit.

The amplitude of an audio signal in dB is $20 \log_{10} \frac{A(f)}{A(f')} = 10 \log_{10} \frac{P(f)}{P(f')}$, where f' is a reference signal selected to correspond to 0 dB (usually, the threshold of human hearing). The threshold of pain is about 120–140 dB.

Pitch and *loudness* are auditory subjective terms for frequency and amplitude.

The $mel\ scale$ is a perceptual frequency scale, corresponding to the auditory sensation of tone height and based on mel, a unit of perceived frequency (pitch). It is connected to the acoustic frequency f hertz scale by $Mel(f)=1,127\ln(1+\frac{f}{700})$ (or, simply, $Mel(f)=1,000\log_2(1+\frac{f}{1,000})$) so that $1,000\ \mathrm{Hz}$ correspond to $1,000\ \mathrm{mels}$.

The *Bark scale* (named after Barkhausen) is a psycho-acoustic scale of frequency: it ranges from 1 to 24 corresponding to the first 24 critical bands of hearing

$$(0, 100, 200, \dots, 1,270, 1,480, 1,720, \dots, 950, 12,000, 15,500 \text{ Hz}).$$

Those bands correspond to spatial regions of the basilar membrane (of the inner ear), where oscillations, produced by the sound of given frequency, activate the hair cells and neurons. The Bark scale is connected to the acoustic frequency f kilohertz scale by $Bark(f) = 13 \arctan(0.76f) + 3.5 \arctan(\frac{f}{0.75})^2$.

The main way that humans control their phonation (speech, song, laughter) is by control over the vocal tract (the throat and mouth) shape. This shape, i.e., the cross-sectional profile of the tube from the closure in the glottis (the space between the vocal cords) to the opening (lips), is represented by the cross-sectional area function Area(x), where x is the distance to the glottis. The vocal tract acts as a resonator during vowel phonation, because it is kept relatively open. These resonances reinforce the source sound (ongoing flow of lung air) at particular resonant frequencies (or formants) of the vocal tract, producing peaks in the spectrum of the sound.

Each vowel has two characteristic formants, depending on the vertical and horizontal position of the tongue in the mouth. The source sound function is modified by the frequency response function for a given area function. If the vocal tract is approximated as a sequence of concatenated tubes of constant cross-sectional area (of equal length, or epilarynx-pharynx-oral cavity, etc.), then the area ratio coefficients are the ratios $\frac{Area(x_{i+1})}{Area(x_i)}$ for consecutive tubes; those coefficients can be computed by LPC (see below).

The *spectrum* of a sound is the distribution of magnitude (dB) (and sometimes the phases in frequency (kHz)) of the components of the wave. The *spectral envelope* is a smooth contour that connects the spectral peaks. The estimation of the spectral envelopes is based on either LPC (linear predictive coding), or FTT (fast Fourier transform using real cepstrum, i.e., the log amplitude spectrum of the sound).

FT (Fourier transform) maps time-domain functions into frequency-domain representations. The cepstrum of the signal f(t) is $FT(\ln(FT(f(t) + 2\pi mi)))$, where m is the integer needed to unwrap the angle or imaginary part of the complex logarithm function. The complex and real cepstrum use, respectively, complex and real log function. The real cepstrum uses only the magnitude of the original signal f(t), while the complex cepstrum uses also phase of f(t). The FFT method is based on linear spectral analysis. The FFT performs the Fourier transform on the signal and samples the discrete transform output at the desired frequencies usually in the mel scale.

Parameter-based distances used in recognition and processing of speech data are usually derived by LPC, modeling the speech spectrum as a linear combination of the previous samples (as in autoregressive processes). Roughly, LPC processes each word of the speech signal in the following six steps: filtering, energy normalization, partition into frames, windowing (to minimize discontinuities at the borders of frames), obtaining LPC parameters by the autocorrelation method and conversion to the LPC-derived cepstral coefficients. LPC assumes that speech is produced by a buzzer at the glottis (with occasionally added hissing and popping sounds), and it removes the formants by filtering.

The majority of distortion measures between sonograms are variations of squared Euclidean distance (including a covariance-weighted one, i.e., Mahalanobis, distance) and probabilistic distances belonging to following general types: generalized total variation metric, f-divergence of Csizar and Chernoff distance.

The distances for sound processing below are between vectors x and y representing two signals to compare. For recognition, they are a template reference and input signal, while for noise reduction they are the original (reference) and distorted signal (see, for example, [OASM03]). Often distances are calculated for small segments, between vectors representing short-time spectra, and then averaged.

• Segmented signal-to-noise ratio

The segmented signal-to-noise ratio $SNR_{seg}(x, y)$ between signals $x = (x_i)$ and $y = (y_i)$ is defined by

$$\frac{10}{m} \sum_{m=0}^{M-1} \left(\log_{10} \sum_{i=n+1}^{n+n} \frac{x_i^2}{(x_i - y_i)^2} \right),$$

where n is the number of frames, and M is the number of segments. The usual signal-to-noise ratio SNR(x, y) between x and y is given by

$$10\log_{10}\frac{\sum_{i=1}^{n}x_{i}^{2}}{\sum_{i=1}^{n}(x_{i}-y_{i})^{2}}.$$

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Another measure, used to compare two waveforms x and y in the time-domain, is their **Czekanovski–Dice distance**, defined by

$$\frac{1}{n} \sum_{i=1}^{n} \left(1 - \frac{2 \min\{x_i, y_i\}}{x_i + y_i} \right).$$

• Spectral magnitude-phase distortion

The spectral magnitude-phase distortion between signals $x = x(\omega)$ and $y = y(\omega)$ is defined by

$$\frac{1}{n} \left(\lambda \sum_{i=1}^{n} (|x(w)| - |y(w)|)^2 + (1 - \lambda) \sum_{i=1}^{n} (\angle x(w) - \angle y(w))^2 \right),$$

where |x(w)|, |y(w)| are magnitude spectra, and $\angle x(w)$, $\angle y(w)$ are phase spectra of x and y, respectively, while the parameter $\lambda, 0 \le \lambda \le 1$, is chosen in order to attach commensurate weights to the magnitude and phase terms. The case $\lambda = 0$ corresponds to the **spectral phase distance**.

Given a signal $f(t) = ae^{-bt}u(t)$, a, b > 0, which has Fourier transform $x(w) = \frac{a}{b+iw}$, its magnitude (or amplitude) spectrum is $|x| = \frac{a}{\sqrt{b^2+w^2}}$, and its phase spectrum (in radians) is $\alpha(x) = \tan^{-1}\frac{w}{b}$, i.e., $x(w) = |x|e^{i\alpha} = |x|(\cos\alpha + i\sin\alpha)$.

• RMS log spectral distance

The RMS log spectral distance (or root-mean-square distance, mean quadratic distance) LSD(x,y) between discrete spectra $x = (x_i)$ and $y = (y_i)$ is the following Euclidean distance:

$$\sqrt{\frac{1}{n}\sum_{i=1}^{n}(\ln x_i - \ln y_i)^2}.$$

The corresponding l_1 - and l_{∞} -distances are called *mean absolute distance* and *maximum deviation*. These three distances are related to decibel variations in the log spectral domain by the multiple $\frac{10}{\log 10}$.

The square of the RMS log spectral distance, via the cepstrum representation $\ln x(\omega) = \sum_{j=-\infty}^{\infty} c_j e^{-j\omega i}$ (where $x(\omega)$ is the power spectrum, i.e., magnitude-squared Fourier transform) becomes, in the complex cepstral space, the **cepstral distance**.

The log area ratio distance LAR(x, y) between x and y is defined by

$$\sqrt{\frac{1}{n} \sum_{i=1}^{n} 10(\log_{10} Area(x_i) - \log_{10} Area(y_i))^2},$$

where $Area(z_i)$ denotes the cross-sectional area of the segment of the vocal tract tube corresponding to z_i .

Bark spectral distance

The Bark spectral distance (Wang-Sekey-Gersho 1992) is a perceptual distance, defined by

$$BSD(x,y) = \sum_{i=1}^{n} (x_i - y_i)^2,$$

i.e., is the squared Euclidean distance between Bark spectra (x_i) and (y_i) of x and y, where the i-th component corresponds to the i-th auditory critical band in the Bark scale.

A modification of the Bark spectral distance excludes critical bands i on which the loudness distortion $|x_i - y_i|$ is less than the noise masking threshold.

• Itakura-Saito quasi-distance

The **Itakura–Saito guasi-distance** (or maximum likelihood distance) IS(x,y) between LPC-derived spectral envelopes $x=x(\omega)$ and $y=y(\omega)$ (1968) is defined by

$$\frac{1}{2\pi} \int_{-\pi}^{\pi} \left(\ln \frac{x(w)}{y(w)} + \frac{y(w)}{x(w)} - 1 \right) dw.$$

The **cosh distance** is defined by IS(x,y) + IS(y,x), i.e., is equal to

$$\frac{1}{2\pi} \int_{-\pi}^{\pi} \left(\frac{x(w)}{y(w)} + \frac{y(w)}{x(w)} - 2 \right) dw = \frac{1}{2\pi} \int_{-\pi}^{\pi} 2 \cosh\left(\ln\frac{x(w)}{y(w)} - 1\right) dw,$$

where $\cosh(t)=\frac{e^t+e^{-t}}{2}$ is the hyperbolic cosine function. • Log likelihood ratio quasi-distance

The log likelihood ratio quasi-distance (or Kullback-Leibler distance) KL(x,y) between LPC-derived spectral envelopes $x=x(\omega)$ and $y = y(\omega)$ is defined by

$$\frac{1}{2\pi} \int_{-\pi}^{\pi} x(w) \ln \frac{x(w)}{y(w)} dw.$$

The **Jeffrey divergence** KL(x,y) + KL(y,x) is also used.

The weighted likelihood ratio distance between spectral envelopes $x = x(\omega)$ and $y = y(\omega)$ is defined by

$$\frac{1}{2\pi} \int_{-\pi}^{\pi} \left(\frac{\left(\ln\left(\frac{x(w)}{y(w)}\right) + \frac{y(w)}{x(w)} - 1\right)x(w)}{p_x} + \frac{\left(\ln\left(\frac{y(w)}{x(w)}\right) + \frac{x(w)}{y(w)} - 1\right)y(w)}{p_y} \right) dw,$$

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where P(x) and P(y) denote the power of the spectra x(w) and y(w), respectively.

• Cepstral distance

The **cepstral distance** (or squared Euclidean cepstrum metric) CEP(x,y) between the LPC-derived spectral envelopes $x=x(\omega)$ and $y=y(\omega)$ is defined by

$$\frac{1}{2\pi} \int_{-\pi}^{\pi} \left(\ln \frac{x(w)}{y(w)} \right)^2 dw = \frac{1}{2\pi} \int_{-\pi}^{\pi} \left(\ln x(w) - \ln y(w) \right)^2 dw$$
$$= \sum_{j=-\infty}^{\infty} (c_j(x) - c_j(y))^2,$$

where $c_j(z) = \frac{1}{2\pi} \int_{-\pi}^{\pi} e^{jwi} \ln |z(w)| dw$ is j-th cepstral (real) coefficient of z derived from the Fourier transform or LPC.

The quefrency-weighted cepstral distance (or Yegnanarayana distance, weighted slope distance) between x and y is defined by

$$\sum_{i=-\infty}^{\infty} i^2 (c_i(x) - c_i(y))^2.$$

"Quefrency" and "cepstrum" are anagrams of "frequency" and "spectrum," respectively.

The Martin cepstrum distance between two AR (autoregressive) models is defined, in terms of their cepstra, by

$$\sqrt{\sum_{i=0}^{\infty} i(c_i(x) - c_i(y))^2}.$$

(Cf. general Martin distance in Chap. 12, defined as an angle distance between subspaces, and Martin metric in Chap. 11 between strings which is an l_{∞} -analog of it.)

The **Klatt slope metric** (1982) between discrete spectra $x = (x_i)$ and $y = (y_i)$ with n channel filters is defined by

$$\sqrt{\sum_{i=1}^{n} ((x_{i+1} - x_i) - (y_{i+1} - y_i))^2}.$$

• Pitch distance

Pitch is a subjective correlate of the fundamental frequency; cf. the above Bark scale of loudness (perceived intensity) and Mel scale of perceived tone height. A musical scale is, usually, a linearly ordered collection of

pitches (notes). A **pitch distance** (or **interval**, **musical distance**) is the size of the section of the linearly-perceived pitch-continuum bounded by those two pitches, as modeled in a given scale. So, an interval describes the difference in pitch between two notes. The pitch distance between two successive notes in a scale is called a *scale step*.

In Western music now, the most used scale is the *chromatic scale* (octave of 12 notes) of *equal temperament*, i.e., divided into 12 equal steps with the ratio between any two adjacent frequencies being $\sqrt[12]{2}$. The scale step here is a *semitone*, i.e., the distance between two adjacent keys (black and white) on a piano. The **distance between notes** whose frequencies are f_1 and f_2 is $12 \log_2(\frac{f_1}{f_2})$ semitones.

A MIDI (Musical Instrument Digital Interface) number of fundamental frequency f is defined by $p(f) = 69 + 12\log_2\frac{f}{440}$. The distance between notes, in terms of MIDI numbers, becomes the **natural metric** $|m(f_1) - m(f_2)|$ on \mathbb{R} . It is a convenient pitch distance since it corresponds to physical distance on keyboard instruments, and psychological distance as measured by experiments and understood by musicians.

A distance model in Music, is the alternation of two different intervals to create a non-diatonic musical mode, for example, 1:2 (the octatonic scale), 1:3 (alternation of semitones and minor thirds) and 1:5.

• Distances between rhythms

A rhythm timeline (music pattern) is represented, besides the standard music notation, in the following ways, used in computational music analysis.

- 1. By a binary vector $x = (x_1, \ldots, x_m)$ of m time intervals (equal in a metric timeline), where $x_i = 1$ denotes a beat, while $x_i = 0$ denotes a rest interval (silence). For example, the five 12/8 metric timelines of Flamenco music are represented by five binary sequences of length 12.
- 2. By a pitch vector $q = (q_1, \ldots, q_n)$ of absolute pitch values q_i and a pitch difference vector $p = (p_1, \ldots, p_{n-1})$ where $p_i = q_{i+1} q_i$ represents the number of semitones (positive or negative) from q_i to q_{i+1} .
- 3. By an inter-onset interval vector $t = (t_1, \ldots, t_n)$ of n time intervals between consecutive onsets.
- 4. By a chronotonic representation which is a histogram visualizing t as a sequence of squares of sides t_1, \ldots, t_n ; such a display can be seen as a piece-wise linear function.
- 5. By a rhythm difference vector $r = (r_1, \ldots, r_{n-1})$, where $r_i = \frac{t_{i+1}}{t_i}$.

Examples of general distances between rhythms are the Hamming distance, swap metric (cf. Chap. 11) and Earth Mover distance between their given vector representations.

The Euclidean interval vector distance is the Euclidean distance between two inter-onset interval vectors. The Gustafson **chronotonic** distance is a variation of l_1 -distance between these vectors using the chronotonic representation.

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Coyle–Shmulevich **interval-ratio distance** is defined by

$$1 - n + \sum_{i=1}^{n-1} \frac{\max\{r_i, r_i'\}}{\min\{r_i, r_i'\}},$$

where r and r' are rhythm difference vectors of two rhythms (cf. the reciprocal of **Ruzicka similarity** in Chap. 17).

• Acoustics distances

The wavelength is the distance the sound wave travels to complete one cycle. This distance is measured perpendicular to the wavefront in the direction of propagation between one peak of a sine wave (sinusoid) and the next corresponding peak. The wavelength of any frequency may be found by dividing the speed of sound $(331.4\,\mathrm{m\,s^{-1}}$ at sea level) in the medium by the fundamental frequency.

The far field (cf. Rayleigh distance in Chap. 24) is the part of a sound field in which sound waves can be considered planar and the sound intensity decreases as $\frac{1}{d^2}$, where d is the distance from the source. It corresponds to a reduction of $\approx 6 \,\mathrm{dB}$ in the sound level for each doubling of distance and to halving of loudness (subjective response) for each reduction of $\approx 10 \,\mathrm{dB}$. Humans have the innate ability to adjust their vocal output to compensate for sound propagation losses to a listener's position.

The near field is the part of a sound field (usually within about two wavelengths from the source) where there is no simple relationship between sound level and distance. A sound measurement is in free field if it is made in open space at a large distance from the source.

The **critical distance** is the distance from the source at which the direct sound (produced by the source) and reverberant sound (reflected echo produced by the direct sound bouncing off, say, walls, floor, etc.) are equal in amplitude.

The **blanking distance** is the minimum sensing range of an ultrasonic **proximity sensor**.

The proximity effect (audio) is the anomaly of low frequencies being enhanced when a directional microphone is very close to the source.

The acoustic metric is the term used occasionally for some distances between vowels; for example, the Euclidean distance between vectors of formant frequencies of pronounced and intended vowel. (Not to be confused with acoustic metrics in General Relativity and Quantum Gravity; cf. Chap. 24.)

Chapter 22 Distances in Internet and Similar Networks

22.1 Scale-free networks

A **network** is a graph, directed or undirected, with a positive number (weight) assigned to each of its arcs or edges. Real-world complex networks usually have a gigantic number N of vertices and are sparse, i.e., with relatively few edges.

Interaction networks (Internet, the Web, social networks, etc.) tend to be **small-world** [Watt99], i.e., interpolate between regular geometric lattices and random graphs in the following sense: they have a large *clustering* coefficient (i.e., the probability that two distinct neighbors of a vertex are neighbors), as lattices in a local neighborhood, while the average path distance between two vertices is small, about $\ln N$, as in a random graph.

The main subcase of a small-world network is a **scale-free network** [Bara01] in which the probability distribution, say, for a vertex, to have degree k is similar to $k^{-\gamma}$ for some positive constant γ which usually belongs to the segment [2, 3].

This *power law* implies that very few vertices, called *hubs* (connectors, super-spreaders), are far more connected than other vertices.

The power law (or **long range dependent**, *heavy-tail*) distributions, in space or time, has been observed in many natural phenomena (both, physical and sociological).

• Collaboration distance

The collaboration distance is the path metric (see http://www.oakland. edu/enp/) of the Collaboration graph, having about 0.4 million vertices (authors in Mathematical Reviews database) with xy being an edge if authors x and y have a joint publication among about 2 million papers itemized in this database. The vertex of largest degree 1416, corresponds to Paul Erdős; the Erdős number of a mathematician is his collaboration distance to Paul Erdős.

The Barr's collaboration metric (http://www.oakland.edu/enp/barr. pdf) is the resistance metric from Chap. 15 in the following extension of the Collaboration graph. First, put a 1- Ω resistor between any two

authors for every joint two-authors paper. Then, for each n-authors paper, n>2, add a new vertex and connect it by a $\frac{n}{4}$ - Ω resistor to each of its co-authors.

• Co-starring distance

The **co-starring distance** is the path metric of the $Hollywood\ graph$, having about 250,000 vertices (actors in the Internet Movie database) with xy being an edge if the actors x and y appeared in a feature film together. The vertices of largest degree are Christopher Lee and Kevin Bacon; the trivia game $Six\ degrees\ of\ Kevin\ Bacon\ uses$ the $Bacon\ number$, i.e., the co-starring distance to this actor.

Similar popular examples of such social scale-free networks are graphs of musicians (who played in the same rock band), baseball players (as teammates), scientific publications (who cite each other), chess-players (who played each other), mail exchanges, acquaintances among classmates in a college, business board membership, sexual contacts among members of a given group. The path metric of the last network is called the **sexual distance**.

Among other such studied networks are air travel connections, word cooccurrences in human language, US power grid, sensor networks, worm neuronal network, gene co-expression networks, protein interaction networks and metabolic networks (with two substrates forming an edge if a reaction occurs between them via enzymes).

• Forward quasi-distance

In a directed network, where edge-weights correspond to a point in time, the **forward quasi-distance** (backward quasi-distance) is the length of the shortest directed path, but only among paths on which consecutive edge-weights are increasing (decreasing, respectively).

The forward quasi-distance is useful in epidemiological networks (disease spreading by contact, or, say, heresy spreading within a church), while the backward quasi-distance is appropriated in P2P (i.e., peer-to-peer) file-sharing networks.

• Betweenness centrality

For a **geodesic** metric space (X, d) (in particular, for the path metric of a graph), the **stress centrality** of a point $x \in X$ is defined (Shimbel 1953) by

$$\sum_{y,z\in X,\,y\neq x\neq z} \text{Number of shortest } (y-z) \text{ paths through } x,$$

the **betweenness centrality** of a point $x \in X$ is defined (Freeman 1977) by

$$g(x) = \sum_{y,z \in X, \ y \neq x \neq z} \frac{\text{Number of shortest } (y-z) \text{ paths through } x}{\text{Number of shortest } (y-z) \text{ paths}},$$

and the **distance-mass function** is a function $M: \mathbb{R}_{>0} \to \mathbb{Q}$, defined by

$$M(a) = \frac{|\{y \in X : d(x,y) + d(y,z) = a \text{ for some } x,y \in X\}|}{|\{(x,z) \in X \times X : d(x,z) = a\}|}.$$

It was conjectured in [GOJKK02] that many scale-free networks satisfy to power law $g^{-\gamma}$ (for the probability that a vertex has betweenness centrality g), where γ is either 2 or ≈ 2.2 with the distance-mass function M(a) being either linear or non-linear, respectively. In the linear case, for example, $\frac{M(a)}{a} \approx 4.5$ for the **Internet AS metric**, and ≈ 1 for the **Web** hyperlink quasi-metric.

Distance centrality

Given a finite metric space (X, d) (usually, the **path metric** on the graph of a network) and a point $x \in X$, we give here examples of metric functionals used to measure distance centrality, i.e., the amount of centrality of the point x in X expressed in terms of its distances d(x, y) to other points:

- 1. The **eccentricity** (or *Koenig number*) $\max_{y \in X} d(x, y)$ was given in
- Chap. 1; Hage and Harary (1995) considered $\frac{1}{\max_{y \in X} d(x,y)}$.

 2. The closeness centrality (Sabidussi 1966) $\frac{1}{\sum_{y \in X} d(x,y)}$ and the mean
- distance $\frac{\sum_{y \in X} d(x,y)}{|X|-1}$.

 3. Dangalchev (2006) introduced $\sum_{y \in X} y \neq x} 2^{-d(x,y)}$, which allows the case $d(x,y) = \infty$ (disconnected graphs).

Drift distance

The **drift distance** is the absolute value of the difference between observed and actual coordinates of a node in a NVE (Networked Virtual Environment). In models of such large-scale peer-to-peer NVE (for example, Massively Multiplayer On-line Games), the users are represented as coordinate points on the plane (nodes) which can move at discrete timesteps, and each has a visibility range called the Area of Interest. NVE creates a synthetic 3D world where each user assumes avatar (a virtual identity) to interact with other users or computer AI.

The term **drift distance** is also used for the current going through a material, in tire production, etc.

• Semantic proximity

For the words in a document, there are short range syntactic relations and long range semantic correlations, i.e., meaning correlations between concepts.

The main document networks are Web and bibliographic databases (digital libraries, scientific databases, etc.); the documents in them are related by, respectively, hyperlinks and citation or collaboration.

Also, some semantic tags (keywords) can be attached to the documents in order to index (classify) them: terms selected by author, title words, journal titles, etc.

The **semantic proximity** between two keywords x and y is their **Tanimoto similarity** $\frac{|X \cap Y|}{|X \cup Y|}$, where X and Y are the sets of documents indexed by x and y, respectively. Their **keyword distance** is defined by $\frac{|X \Delta Y|}{|X \cap Y|}$; it is not a metric.

22.2 Network-based semantic distances

Among the main lexical networks (such as WordNet, Framenet, Medical Search Headings, Roget's Thesaurus) a semantic lexicon WordNet is the most popular lexical resource used in Natural Language Processing and Computational Linguistics.

WordNet (see http://wordnet.princeton.edu) is an on-line lexical database in which English nouns, verbs, adjectives and adverbs are organized into synsets (synonym sets), each representing one underlying lexical concept.

Two synsets can be linked semantically by one of the following links: upwards x (hyponym) IS-A y (hypernym) link, downwards x (meronym) CONTAINS y (holonym) link, or a horizontal link expressing frequent co-occurrence (antonymy), etc. IS-A links induce a partial order, called IS-A taxonomy. The version 2.0 of WordNet has 80,000 noun concepts and 13,500 verb concepts, organized into 9 and 554 separate IS-A hierarchies, respectively.

In the resulting DAG (directed acyclic graph) of concepts, for any two synsets (or concepts) x and y, let l(x,y) denote the length of the shortest path between them, using only IS-A links, and let LPS(x,y) denote their least common subsumer (ancestor) by IS-A taxonomy. Let d(x) denote the depth of x (i.e., its distance from the root in IS-A taxonomy) and let $D = \max_x d(x)$.

The semantic relatedness of two noons can be estimated by their **ancestral** path distance (cf. Chap. 23), i.e., the length of the shortest *ancestral* path (directed path through a common ancestor) connecting them. A list of the other main semantic similarities and distances follows.

• Path similarity

The path similarity between synsets x and y is defined by

$$path(x, y) = (l(x, y))^{-1}.$$

Leacock-Chodorow similarity

The **Leacock–Chodorow similarity** between synsets x and y is defined by

$$lch(x,y) = -\ln\frac{l(x,y)}{2D},$$

and the **conceptual distance** between them is defined by $\frac{l(x,y)}{D}$.

• Wu-Palmer similarity

The Wu-Palmer similarity between synsets x and y is defined by

$$wup(x,y) = \frac{2d(LPS(x,y))}{d(x) + d(y)}.$$

• Resnik similarity

The **Resnik similarity** between synsets x and y is defined by

$$res(x, y) = -\ln p(LPS(x, y)),$$

where p(z) is the probability of encountering an instance of concept z in a large corpus, and $-\ln p(z)$ is called the *information content* of z.

• Lin similarity

The Lin similarity between synsets x and y is defined by

$$lin(x,y) = \frac{2 \ln p(LPS(x,y))}{\ln p(x) + \ln p(y)}.$$

• Jiang-Conrath distance

The **Jiang–Conrath distance** between synsets x and y is defined by

$$jcn(x, y) = 2 \ln p(LPS(x, y)) - (\ln p(x) + \ln p(y)).$$

• Lesk similarities

A gloss of a synonym set z is the member of this set giving a definition or explanation of an underlying concept. The **Lesk similarities** are those defined by a function of the overlap of glosses of corresponding concepts; for example, the **gloss overlap** is

$$\frac{2t(x,y)}{t(x)+t(y)},$$

where t(z) is the number of words in the synset z, and t(x, y) is the number of common words in x and y.

• Hirst-St-Onge similarity

The Hirst-St-Onge similarity between synsets x and y is defined by

$$hso(x,y) = C - L(x,y) - ck,$$

where L(x,y) is the length of a shortest path between x and y using all links, k is the number of changes of direction in that path, and C,c are constants.

The Hirst-St-Onge distance is defined by $\frac{L(x,y)}{k}$.

• Semantic biomedical distances

The **semantic biomedical distances** are the distances used in biomedical lexical networks. The main clinical terminologies are UMLS (United Medical Language System) and SNOMED CT (Systematized Nomenclature of Medicine – Clinical Terms).

The *conceptual distance* between two biomedical concepts in UMLS is (Caviedes and Cimino 2004) the minimum number of *IS-A* parent links between them in the directed acyclic graph of *IS-A* taxonomy of concepts.

An example of semantic biomedical distances used in SNOMED and presented in Melton, Parsons, Morrisin, Rothschild, Markatou and Hripsak (2006) is given by the *inter-patient distance* between two *medical cases* (sets X and Y of patient data). It is the **Tanimoto distance** (cf. Chap. 1) $\frac{|X \triangle Y|}{|X \cup Y|}$ between them.

22.3 Distances in Internet and Web

Let us consider in detail the graphs of the Web and of its hardware substrate, Internet, which are small-world and scale-free.

The *Internet* is the largest WAN (wide area network), spanning the Earth. This publicly available worldwide computer network came from ARPANET (started in 1969 by US Department of Defense), NSFNet, Usenet, Bitnet, and other networks. In 1995, the National Science Foundation in the US gave up the stewardship of the Internet.

Its nodes are *routers*, i.e., devices that forward packets of data along networks from one computer to another, using IP (Internet Protocol relating names and numbers), TCP and UDP (for sending data), and (built on top of them) HTTP, Telnet, FTP and many other *protocols* (i.e., technical specifications of data transfer). Routers are located at *gateways*, i.e., places where at least two networks connect.

The links that join the nodes together are various physical connectors, such as telephone wires, optical cables and satellite networks. The Internet uses *packet switching*, i.e., data (fragmented if needed) are forwarded not along a previously established path, but so as to optimize the use of available *bandwidth* (bit rate, in million bits per second) and minimize the *latency* (the time, in milliseconds, needed for a request to arrive).

Each computer linked to the Internet is usually given a unique "address," called its *IP address*. The number of possible IP addresses is $2^{32} \approx 4.3$ billion only. The most popular applications supported by the Internet are e-mail, file transfer, Web, and some multimedia as Internet TV and YouTube. In 2006, 161 EB (161 billion gigabytes, $1,288 \times 10^{18}$ bites, $\approx 3 \times 10^6$ times the information in all the books ever written) of digital information was created and copied. Internet traffic more than doubles each year.

The *Internet IP graph* has, as the vertex-set, the IP addresses of all computers linked to the Internet; two vertices are adjacent if a router connects them directly, i.e., the passing datagram makes only one *hop*.

The Internet also can be partitioned into ASs (administratively Autonomous Systems or domains). Within each AS the intra-domain routing is done by IGP (Interior Gateway Protocol), while inter-domain routing is done by BGP (Border Gateway Protocol) which assigns an ASN (16-bit number) to each AS. The *Internet AS graph* has ASs (about 25.000 in 2007) as vertices and edges represent the existence of a BGP peer connection between corresponding ASs.

The World Wide Web (WWW or Web, for short) is a major part of Internet content consisting of interconnected documents (resources). It corresponds to HTTP (Hyper Text Transfer Protocol) between browser and server, HTML (Hyper Text Markup Language) of encoding information for a display, and URLs (Uniform Resource Locators), giving unique "addresses" to web pages. The Web was started in 1989 in CERN which gave it for public use in 1993.

The Web digraph is a virtual network, the nodes of which are documents (i.e., static HTML pages or their URLs) which are connected by incoming or outcoming HTML hyperlinks, i.e., hypertext links.

The number of nodes in the Web digraph in 2007 was, by different estimation, between 15 and 30 billion.

The number of web sites (collections of related web pages found at a single address) reached 182 million in 2008, from 18,957 in 1995. Along with the Web lies the Deep or Invisible Web, i.e., searchable databases (about 300.000) with the number of pages (if not actual content) estimated as being about 500 times more than on static Web pages. Those pages are not indexed by search engines; they have dynamic URL and so can be retrieved only by a direct query in real time.

On 30 June 2007, 1,173,109,925 users (17.8% of the global population 6,574,666,417) were online, including 69.5% in North America and 39.8% in Europe. The top six languages on the Internet, at 30 June 2007, were: English, Chinese, Spanish, Japanese, French, German with, respectively, 31%, 16%, 9%, 7%, 5%, 5% of all Internet users with corresponding Internet penetration 18%, 14%, 23%, 67%, 15%, 61% by languages.

There are several hundred thousand cyber-communities, i.e., clusters of nodes of the Web digraph, where the link density is greater among members than between members and the rest. The cyber-communities (a customer group, a social network, a concept in a technical paper, etc.) are usually focused around a definite topic and contain a bipartite hubs-authorities subgraph, where all hubs (guides and resource lists) point to all authorities (useful and relevant pages on the topic). Examples of new media, created by the Web are (we)blogs (digital diaries posted on the Web), Skype (telephone calls), social site Facebook and Wikipedia (the collaborative encyclopedia). The project Semantic Web by WWW Consortium aims at linking to metadata, merging social data and transformation of WWW into GGG (Giant Global Graph) of users.

On average, nodes of the Web digraph are of size $10\,\mathrm{kB}$, out-degree 7.2, and probability k^{-2} to have out-degree or in-degree k. A study in [BKMR00] of over 200 million web pages gave, approximatively, the largest connected component "core" of 56 million pages, with another 44 million of pages connected to the core (newcomers?), 44 million to which the giant core is connected (corporations?) and 44 million connected to the core only by undirected paths or disconnected from it. For randomly chosen nodes x and y, the probability of the existence of a directed path from x to y was 0.25 and the average length of such a shortest path (if it exists) was 16, while maximal length of a shortest path was over 28 in the core and over 500 in the whole digraph.

A study in [CHKSS06] of Internet AS graphs revealed the following Medusa structure of the Internet: "nucleus" (diameter 2 cluster of ≈ 100 nodes), "fractal" ($\approx 15{,}000$ nodes around it), and "tentacles" ($\approx 5{,}000$ nodes in isolated subnetworks communicating with the outside world only via the nucleus).

The distances below are examples of host-to-host **routing metrics**, i.e., values used by routing algorithms in the Internet, in order to compare possible routes. Examples of other such measures are: bandwidth consumption, communication cost, reliability (probability of packet loss). Also, the main computer-related *quality metrics* are mentioned.

• Internet IP metric

The Internet IP metric (or hop count, RIP metric, IP path length) is the path metric in the Internet IP graph, i.e., the minimal number of hops (or, equivalently, routers, represented by their IP addresses) needed to forward a packet of data. RIP imposes a maximum distance of 15 and advertises by 16 non-reachable routes.

• Internet AS metric

The Internet AS metric (or *BGP-metric*) is the path metric in the *Internet AS graph*, i.e., the minimal number of ISPs (Independent Service Providers), represented by their ASs, needed to forward a packet of data.

• Geographic distance

The **geographic distance** is the **great circle distance** on the Earth from the client x (destination) to the server y (source). However, for economical reasons, the data often do not follow such geodesics; for example, most data from Japan to Europe transits via US.

• RTT-distance

The **RTT-distance** is the RTT (Round Trip Time: to send a packet and receive an acknowledgement back) of transmission between x and y, measured in milliseconds (usually, by the ping command) during the previous day.

See [HFPMC02] for variations of this distance and connections with the above three metrics. Fraigniaud, Lebbar and Viennot (2008) found that RTT is a C-inframetric (Chap. 1) with $C \approx 7$. Sinha, Raz and Choudhuri (2006) asserted that the average RTT-distance from x to the nearest backbone (i.e., of Class 1) network, coupled with its **geographic distance**, predicts the network distance better than on-line metrics.

• Administrative cost distance

The administrative cost distance is the nominal number (rating the trustworthiness of a routing information), assigned by the network to the route between x and y. For example, Circe assigns values $0, 1, \ldots, 200, 255$ for the Connected Interface, Static Route, ..., Internal BGP, Unknown, respectively.

• DRP-metrics

The DD (Distributed Director) system of Cisco uses (with priorities and weights) the **administrative cost distance**, the **random metric** (selecting a random number for each IP address) and the **DRP** (Direct Response Protocol) metrics. DRP-metrics ask from all DRP-associated routers one of the following distances:

- The DRP-external metric, i.e., the number of BGP (Border Gateway Protocol) hops between the client requesting service and the DRP server agent
- 2. The **DRP-internal metric**, i.e., the number of IGP hops between the DRP server agent and the closest border router at the edge of the autonomous system
- 3. The **DRP-server metric**, i.e., the number of IGP hops between the DRP server agent and the associated server

• Network tomography metrics

Consider a network with fixed routing protocol, i.e., a strongly connected digraph D=(V,E) with a unique directed path T(u,v) selected for any pair (u,v) of vertices. The routing protocol is described by a binary routing matrix $A=((a_{ij}))$, where $a_{ij}=1$ if the arc $e\in E$, indexed i, belongs to the directed path T(u,v), indexed j. The **Hamming distance** between two rows (columns) of A is called the **distance between corresponding arcs** (directed paths) of the network.

Consider two networks with the same digraph, but different routing protocols with routing matrices A and A', respectively. Then a **routing protocol semi-metric** [Var04] is the smallest Hamming distance between A and a matrix B, obtained from A' by permutations of rows and columns (both matrices are seen as strings).

• Web hyperlink quasi-metric

The **Web hyperlink quasi-metric** (or *click count*) is the length of the shortest directed path (if it exists) between two web pages (vertices in the Web digraph), i.e., the minimal number of necessary mouse-clicks in this digraph.

• Average-clicks Web quasi-distance

The average-clicks Web quasi-distance between two web pages x and y in the Web digraph [YOI03] is the minimum $\sum_{i=1}^{m} \ln p \frac{z_i^+}{\alpha}$ over all directed paths $x = z_0, z_1, \ldots, z_m = y$ connecting x and y, where z_i^+ is the outdegree of the page z_i . The parameter α is 1 or 0.85, while p (the average out-degree) is 7 or 6.

• Dodge-Shiode WebX quasi-distance

The **Dodge–Shiode WebX quasi-distance** between two web pages x and y of the Web digraph is the number $\frac{1}{h(x,y)}$, where h(x,y) is the number of shortest directed paths connecting x and y.

• Web similarity metrics

Web similarity metrics form a family of indicators used to quantify the extent of relatedness (in content, links or/and usage) between two web pages x and y.

Some examples are: topical resemblance in overlap terms, co-citation (the number of pages, where both are given as hyperlinks), bibliographical coupling (the number of hyperlinks in common) and co-occurrence frequency $\min\{P(x|y), P(y|x)\}$, where P(x|y) is the probability that a visitor of the page y will visit the page x.

In particular, **search-centric change metrics** are metrics used by search engines on the Web, in order to measure the degree of change between two versions x and y of a web page. If X and Y are the set of all words (excluding HTML markup) in x and y, respectively, then the **word page distance** is the **Dice distance**

$$\frac{|X\triangle Y|}{|X|+|Y|}=1-\frac{2|X\cup Y|}{|X|+|Y|}.$$

If v_x and v_y are weighted vector representations of x and y, then their cosine page distance is given by

$$1 - \frac{\langle v_x, v_y \rangle}{||v_x||_2 \cdot ||v_y||_2}.$$

Cf. **TF-IDF** similarity in Chap. 17.

• Web quality control distance function

Let P be a query quality parameter and X its domain. For example, P can be query *response time*, or accuracy, relevancy, size of result.

The Web quality control distance function (Chen, Zhu and Wang 1998) for evaluating the relative goodness of two values, x and y, of parameter P is a function $\rho: X \times X \to \mathbb{R}$ (not a **distance**) such that, for all $x, y, z \in X$:

- 1. $\rho(x,y) = 0$ if and only if x = y.
- 2. $\rho(x,y) > 0$ if and only if $\rho(y,x) < 0$.
- 3. If $\rho(x,y) > 0$ and $\rho(y,z) > 0$, then $\rho(x,z) > 0$.

The inequality $\rho(x,y) > 0$ means that x is better than y; so, it defines a partial order (reflexive, antisymmetric and transitive binary relation) on X.

• Lostness metric

Users navigating within hypertext systems often experience disorientation (the tendency to lose sense of location and direction in a non-linear document) and cognitive overhead (the additional effort and concentration needed to maintain several tasks/trails at the same time). Users miss the global view of document structure and their working space.

Smith's **lostness metric** measures it by

$$(\frac{n}{s}-1)^2+(\frac{r}{n}-1)^2,$$

where s is the total number of nodes visited while searching, n is the number of different nodes among them, and r is the number of nodes which need to be visited to complete a task.

• Trust metrics

A **trust metric** is, in Computer Security, a measure to evaluate a set of peer certificates resulting in a set of accounts accepted and, in Sociology, a measure of how a member of the group is trusted by the others in the group.

For example, the UNIX access metric is a combination of only read, write and execute kinds of access to a resource. The much finer Advogato trust metric (used in the community of open source developers to rank them) is based on bonds of trust formed when a person issues a certificate about someone else. Other examples are: Technorati, TrustFlow, Richardson et al's., Mui et al's., eBay trust metrics.

• Software metrics

A **software metric** is a measure of software quality which indicates the complexity, understandability, description, testability and intricacy of code. Managers use mainly **process metrics** which help in monitoring the processes that produce the software (say, the number of times the program failed to rebuild overnight).

An **architectural metric** is a measure of software architecture (development of large software systems) quality which indicates the coupling (inter-connectivity of composites), cohesion (intraconnectivity), abstractness, instability, etc.

• Locality metric

The **locality metric** is a physical metric measuring globally the locations of the program components, their calls, and the depth of nested calls by

$$\frac{\sum_{i,j} f_{ij} d_{ij}}{\sum_{i,j} f_{ij}},$$

where d_{ij} is a distance between calling components i and j, while f_{ij} is the frequency of calls from i to j. If the program components are of about same size, $d_{ij} = |i - j|$ is taken. In the general case, Zhang and Gorla (2000) proposed to distinguish forward calls, which are placed before the called

component, and backward (other) calls. Define $d_{ij} = d'_i + d''_{ij}$, where d'_i is the number of lines of code between the calling statement and the end of i if call is forward, and between the beginning of i and the call, otherwise, while $d''_{ij} = \sum_{k=i+1}^{j-1} L_k$ if the call is forward, and $d''_{ij} = \sum_{k=j+1}^{i-1} L_k$ otherwise. Here L_k is the number of lines in component k.

• Reuse distance

In a computer, the microprocessor (or processor) is the chip doing all the computations, and the memory usually refers to RAM (random access memory). A (processor) cache stores small amounts of recently used information right next to the processor where it can be accessed much faster than memory. The following distance estimates the cache behavior of programs.

The **reuse distance** (Mattson, Gecsei, Slutz and Treiger 1970; and Ding and Zhong 2003) of a memory location x is the number of distinct memory references between two accesses of x. Each memory reference is counted only once because after access it is moved in the cache. The reuse distance from the current access to the previous one or to the next one is called the *backward* or *forward* reuse distance, respectively.

• Action at a distance (in Computing)

In Computing, the **action at a distance** is a class of programming problems in which the state in one part of a program's data structure varies wildly because of difficult-to-identify operations in another part of the program.

In Software Engineering, Holland's *Law of Demeter* is a style guideline: an unit should "talk only to immediate friends" (closely related units) and have limited knowledge about other units; cf. **principle of locality** in Chap. 24.

Part VI Distances in Natural Sciences

Chapter 23 Distances in Biology

Distances are mainly used in *Biology* to pursue basic classification tasks, for instance, for reconstructing the evolutionary history of organisms in the form of phylogenetic trees. In the classical approach those distances were based on comparative morphology, physiology, mating studies, paleontology and immunodiffusion. The progress of modern *Molecular Biology* also allowed the use of nuclear- and/or amino-acid sequences to estimate distances between genes, proteins, genomes, organisms, species, etc. The importance of distance can be seen, for example, from the list of 23 Mathematical Challenges funded by US Department of Defense since DARPA-BAA Tech 2007; the 15-th one is "The Geometry of Genome Space:" what notion of distance is needed to incorporate biological utility?

DNA is a sequence of nucleotides (or nuclei acids) A, T, G and C, and it can be seen as a word over this alphabet of four letters. The (single ring) nucleotides A, G (short for adenine and guanine) are called purines, while (double ring) T, C (short for thymine and cytosine) are called pyrimidines (in RNA, it is uracil U instead of T). Two strands of DNA are held together and in the opposite orientation (forming a double helix) by weak hydrogen bonds between corresponding nucleotides (necessarily, a purine and a pyrimidine) in the strands alignment. These pairs are called base pairs.

A transition mutation is a substitution of a base pair, so that a purine/pyrimidine is replaced by another purine/pyrimidine; for example, GC is replaced by AT. A transversion mutation is a substitution of a base pair, so that a purine/pyrimidine is replaced by a pyrimidine/purine base pair, or vice versa; for example, GC is replaced by TA.

DNA molecules occur (in the nuclei of eukaryote cells) in the form of long chaines called *chromosomes*. Most human cells contain 23 pairs of chromosomes, one set of 23 from each parent; human *gamete* (sperm or egg) is a *haploid*, i.e., contains only one set of 23 chromosomes. The (normal) males and females differ only in the 23-rd pair: XY for males, and XX for female. The DNA from one human cell has length $\approx 1.8 \,\mathrm{m}$ but width of $\approx 2.4 \,\mathrm{nm}$.

A gene is a functionally complete segment of DNA which encodes (via transcription, information flow to RNA, and then translation, information flow from RNA to enzymes) a protein or an RNA molecule. The location of

a gene on its specific chromosome is called the *gene locus*. Different versions (states) of a gene are called its *alleles*. Only $\approx 1.5\%$ of human DNA are in protein-coding genes.

A protein is a large molecule which is a chain of amino acids; among them are hormones, catalysts (enzymes), antibodies, etc. The **protein length** is the number of amino acids in the chain; average protein length is around 300. There are 20 standard amino acids; the three-dimensional shape of a protein is defined by the (linear) sequence of amino acids, i.e., by a word in this alphabet of 20 letters.

The genetic code is universal to all organisms and is a correspondence between some codons (i.e., ordered triples of nucleotides) and 20 amino acids. It express the genotype (information contained in genes, i.e., in DNA) as the phenotype (proteins). Three stop codons (UAA, UAG, and UGA) signify the end of a protein; any two, among 61 remaining codons, are called synonymous if they correspond to the same amino acid. Slight variations of the code (codon reassignments selected, perhaps, for antiviral defense) were observed for some mitochondria, ciliates, yeasts, etc. In certain enzymes, non-standard amino acids (21st, selenocysteine and 22nd, pyrrolysine) are substituted for standard stop codons: UGA and UAG, respectively. On the other hand, more than 60 amino acids were identified in the Murchison meteorite.

A genome is entire genetic constitution of a species or of a living organism. For example, the human genome is the set of 23 chromosomes consisting of ≈ 3.2 billion base pairs of DNA and organized into 20,000–25,000 genes.

IAM (infinite-alleles model of evolution) assumes that an allele can change from any given state into any other given state. It corresponds to a primary role for genetic drift (i.e., random variation in gene frequencies from one generation to another), especially in small populations, over natural selection (stepwise mutations). IAM is convenient for allozyme (a form of a protein which is encoded by one allele at a specific gene locus) data.

SMM (for step-wise mutation model of evolution) is more convenient for (recently, most popular) micro-satellite data. A repeat is a stretch of base pairs that is repeated with a high degree of similarity in the same sequence. Micro-satellites are highly variable repeating short sequences of DNA; their mutation rate is 1 per 1,000–10,000 replication events, while it is 1/1,000,000 for allozymes. It turns out that micro-satellites alone contain enough information to plot the lineage tree of an organism. Micro-satellite data (for example, for DNA fingerprinting) consist of numbers of repeats of micro-satellites for each allele. Another popular molecular marker is SSU rRNA (small subunit ribosomal RNA) data because rRNA genes are essential for the survival of any organism and their sequences change relatively little.

Examples of distances, representing general schemes of measurement in Biology, follow.

The term **taxonomic distance** is used for every distance between two *taxa*, i.e., entities or groups, which are arranged into an hierarchy (in the form of a tree designed to indicate degrees of relationship).

Linnean taxonomic hierarchy is arranged in ascending series of ranks: Zoology (seven ranks: Kingdom, Phylum, Class, Order, Family Genus, Species) and Botany (12 ranks). A phenogram is an hierarchy expressing phenetic relationship, i.e., unweighted overall similarity. A cladogram is a strictly genealogical (by ancestry) hierarchy in which no attempt is made to estimate/depict rates or amount of genetic divergence between taxa.

A phylogenetic tree is an hierarchy representing a hypothesis of phylogeny, i.e., evolutionary relationships within and between taxonomic levels, especially the patterns of lines of descent. **Phenetic distance** is a measure of the difference in phenotype between any two nodes on a phylogenetic tree. **Phylogenetic distance** (or **cladistic distance**, **genealogical distance**) between two taxa is the branch length, i.e., the minimum number of edges, separating them in a phylogenetic tree.

Evolutionary distance (or patristic distance, general *genetic distance*) between two taxa is a measure of genetic divergence estimating the *divergence time*, i.e., the time that has passed since those populations existed as a single population. General **immunological distance** between two taxa is a measure of the strength of antigen-antibody reactions, indicating the evolutionary distance separating them.

The next three sections list the main genetic distances for different molecular data (allele frequencies and nucleotide or amino acid sequences). The main way to estimate the genetic distance between DNA, RNA or proteins is to compare their nucleotide or amino acid, respectively, sequences. Besides sequencing, the main techniques used are immunological ones, annealing (cf. hybridization metric) and comparing their gel electrophoresis (separation through an electric charge) banding patterns. In fact, chromosomes stained by some dyes show a 2D pattern of traverse bands of light and heavy staining.

23.1 Genetic distances for gene-frequency data

In this section, a **genetic distance** between populations is a way of measuring the amount of evolutionary divergence by counting the number of allelic substitutions by loci. Among the three most commonly used distances below, **Nei standard genetic distance** 1972, assumes that differences arise due to mutation and genetic drift, while **Cavalli-Sforza–Edvards chord distance** 1967, and **Reynolds–Weir–Cockerham distance** 1983, assume genetic drift only.

A population is represented by a double-indexed vector $x = (x_{ij})$ with $\sum_{j=1}^{n} m_j$ components, where x_{ij} is the frequency of *i*-th allele (the label for a state of a gene) at the *j*-th gene locus (the position of a gene on a chromosome), m_j is the number of alleles at the *j*-th locus, and n is the number of considered loci.

Denote by \sum summation over all i and j. Since x_{ij} is the frequency, $x_{ij} \ge 0$ and $\sum_{i=1}^{m_j} x_{ij} = 1$.

• Stephens et al. shared allele distance
The Stephens et al. shared allele distance (Stephens, Gilbert, Yuhki and O'Brien 1992) between populations is defined by

$$1 - \frac{\overline{SA(x,y)}}{\overline{SA(x)} + \overline{SA(y)}},$$

where, for two individuals a and b, SA(a,b) denotes the number of shared alleles summed over all n loci and divided by 2n, while $\overline{SA(x)}$, $\overline{SA(y)}$, and $\overline{SA(x,y)}$ are SA(a,b) averaged over all pairs (a,b) with individuals a,b being in populations, represented by x, by y and, respectively, between them.

• Dps distance.

The **Thorpe similarity** between populations is defined by $\sum \min\{x_{ij}, y_{ij}\}$.

The **Dps distance** between populations is defined by

$$-\ln\frac{\sum\min\{x_{ij},y_{ij}\}}{\sum_{j=1}^n m_j}.$$

Prevosti-Ocana-Alonso distance
 The Prevosti-Ocana-Alonso distance between populations is defined
 (cf. Manhattan metric) by

$$\frac{\sum |x_{ij} - y_{ij}|}{2n}.$$

• Roger distance

The **Roger distance** is a metric between populations, defined by

$$\frac{1}{\sqrt{2}n} \sum_{j=1}^{n} \sqrt{\sum_{i=1}^{m_j} (x_{ij} - y_{ij})^2}.$$

Cavalli-Sforza-Edvards chord distance
 The Cavalli-Sforza-Edvards chord distance between populations is defined by

$$\frac{2\sqrt{2}}{\pi n} \sum_{j=1}^{n} \sqrt{1 - \sum_{i=1}^{m_j} \sqrt{x_{ij} y_{ij}}}.$$

It is a metric (cf. **Hellinger distance** in Chap. 17).

• Cavalli-Sforza arc distance

The Cavalli-Sforza arc distance between populations is defined by

$$\frac{2}{\pi}\arccos\left(\sum\sqrt{x_{ij}y_{ij}}\right).$$

(Cf. Fisher distance in Chap. 14.)

• Nei-Tajima-Tateno distance

The Nei-Tajima-Tateno distance between populations is defined by

$$1 - \frac{1}{n} \sum \sqrt{x_{ij}y_{ij}}.$$

• Nei minimum genetic distance

The Nei minimum genetic distance between populations is defined by

$$\frac{1}{2n}\sum (x_{ij}-y_{ij})^2.$$

• Nei standard genetic distance

The Nei standard genetic distance between populations is defined by

$$-\ln I$$
.

where I is Nei normalized identity of genes, defined by $\frac{\langle x,y\rangle}{||x||_2 \cdot ||y||_2}$ (cf. **Bhattacharya distances** in Chap. 14 and **angular semi-metric** in Chap. 17).

• Sangvi χ^2 distance

The **Sangvi** χ^2 **distance** between populations is defined by

$$\frac{2}{n} \sum \frac{(x_{ij} - y_{ij})^2}{x_{ij} + y_{ij}}.$$

• Fuzzy set distance

The Dubois-Prade's **fuzzy set distance** between populations is defined by

$$\frac{\sum 1_{x_{ij} \neq y_{ij}}}{\sum_{j=1}^{n} m_j}.$$

• Goldstein and al. distance

The Goldstein and al. distance between populations is defined by

$$\frac{1}{n}\sum (ix_{ij}-iy_{ij})^2.$$

• Average square distance
The average square distance between populations is defined by

$$\frac{1}{n} \sum_{k=1}^{n} \left(\sum_{1 \le i < j \le m_j} (i-j)^2 x_{ik} y_{jk} \right).$$

• Shriver–Boerwinkle stepwise distance
The Shriver–Boerwinkle stepwise distance between populations is defined by

$$\frac{1}{n} \sum_{k=1}^{n} \sum_{1 \le i, j \le m_k} |i - j| (2x_{ik}y_{jk} - x_{ik}x_{jk} - y_{ik}y_{jk}).$$

• **Kinship distance**The **kinship distance** between populations is defined by

$$-\ln\langle x,y\rangle$$
,

and $\langle x, y \rangle$ is called the kinship coefficient.

• Latter F-statistics distance
The Latter F-statistics distance between populations is defined by

$$\frac{\sum (x_{ij} - y_{ij})^2}{2(n - \sum x_{ij}y_{ij})}.$$

• Reynolds-Weir-Cockerham distance
The Reynolds-Weir-Cockerham distance (or co-ancestry distance)
between populations is defined by

$$-\ln(1-\theta),$$

where $\theta = \frac{\sum (x_{ij} - y_{ij})^2}{2(1 - \sum x_{ij}y_{ij})}$ (cf. Latter F-statistics distance) is an estimation of their co-ancestry coefficient.

This coefficient of two populations (or individuals) is the probability that a randomly picked allele from the genetic pool of one population (or from one individual) is *identical by descent* (i.e., corresponding genes are physical copies of the same ancestral gene) to a randomly picked allele in another. Two genes can be *identical by state*, i.e., with the same allele label, but not identical by descent. The co-ancestry coefficient of two individuals is the *inbreeding coefficient* of their following generation.

• Ancestral path distance

Hereditary trees (or family trees, pedigree graphs) are used to represent ancestry relations and, in particular, to identify inbreeding loops and genes associated with genetic diseases. In such a directed tree, every vertex (person) has in-degree at most two (known parents).

Generally, given a directed acyclic graph, Bender, Farach, Colton, Pemmasani, Skiena and Sumazin 2001, defined, for any two vertices x, y, the **ancestral path distance** as the length of the shortest ancestral path (directed path through a common ancestor vertex) and the **LCA** ancestral path distance as the length of the shortest directed path through LCA (the least common ancestor) vertex.

The smallest *inbreeding loop* containing vertices x and y is formed by concatenating ancestral and descending paths connecting them. The ancestral path distance also measures semantic noun relatedness in WorldNet (cf. Chap. 22).

The unrelated *ancestral distance* of an extant taxon (Hearn and Huber 2006) is the time (or the number of speciation events, node depth) separating it from its most recent ancestor with at least one extant descendant having an independent character (trait). Cf., also unrelated, **co-ancestry distance**.

• Lasker distance

The Lasker distance (Rodrigues and Larralde 1989) between two human populations x and y, characterized by surname frequency vectors (x_i) and (y_i) , is the number $-\ln 2R_{x,y}$, where $R_{x,y} = \frac{1}{2} \sum_i x_i y_i$ is Lasker's coefficient of relationship by isonymy. Surname structure is related to inbreeding and (in patrilinear societies) to random genetic drift, mutation and migration. Surnames can be considered as alleles of one locus, and their distribution can be analyzed by Kimura's theory of neutral mutations; an isonymy points to a common ancestry.

23.2 Distances for DNA/RNA data

Distances between nucleotide (DNA/RNA) or protein sequences are usually measured in terms of substitutions, i.e., mutations, between them.

A *DNA sequence* will be seen as a sequence $x = (x_1, ..., x_n)$ over the four-letter alphabet of four nucleotides A, T, C, G (or two-letter alphabet purine/pyramidine); \sum denotes $\sum_{i=1}^{n}$.

Protein-coding nucleotide sequences are called codon sequences.

• No. of DNA differences

The **No. of DNA differences** between DNA sequences is the number of mutations, i.e., their **Hamming metric**:

$$\sum 1_{x_i \neq y_i}.$$

• p-distance

The p-distance d_p between DNA sequences is defined by

$$\frac{\sum 1_{x_i \neq y_i}}{n}.$$

• Jukes-Cantor nucleotide distance

The Jukes–Cantor nucleotide distance between DNA sequences is defined by

 $-\frac{3}{4}\ln\left(1-\frac{4}{3}d_p(x,y)\right),\,$

where d_p is the *p*-distance, subject to $d_p \leq \frac{3}{4}$. If the rate of substitution varies with the gamma distribution, and a is the parameter describing the shape of this distribution, then the **gamma distance for the Jukes–Cantor model** is defined by

$$\frac{3a}{4}\left(\left(1-\frac{4}{3}d_p(x,y)\right)^{-1/a}-1\right).$$

• Tajima-Nei distance

The Tajima–Nei distance between DNA sequences is defined by

$$-b \ln \left(1 - \frac{d_p(x, y)}{b} \right), \text{ where}$$

$$b = \frac{1}{2} \left(1 - \sum_{j=A, T, C, G} \left(\frac{1_{x_i = y_i = j}}{n} \right)^2 + \frac{1}{c} \sum_{c} \left(\frac{1_{x_i \neq y_i}}{n} \right)^2 \right), \text{ and}$$

$$c = \frac{1}{2} \sum_{i, k \in \{A, T, G, C\}, j \neq k} \frac{\left(\sum_{j=1}^{n} 1_{(x_i, y_i) = (j, k)} \right)^2}{\left(\sum_{j=1}^{n} 1_{x_i = y_i = j} \right) \left(\sum_{j=1}^{n} 1_{x_i = y_i = k} \right)}.$$

Let $P = \frac{1}{n}|\{1 \le i \le n : \{x_i,y_i\} = \{A,G\} \text{ or } \{T,C\}\}|$, and $Q = \frac{1}{n}|\{1 \le i \le n : \{x_i,y_i\} = \{A,T\} \text{ or } \{G,C\}\}|$, i.e., P and Q are the frequencies of, respectively, transition and transversion mutations between x and y.

The following four distances are given in terms of P and Q.

• Jin-Nei gamma distance

The Jin-Nei gamma distance between DNA sequences is defined by

$$\frac{a}{2}\left((1-2P-Q)^{-1/a}+\frac{1}{2}(1-2Q)^{-1/a}-\frac{3}{2}\right),$$

where the rate of substitution varies with the gamma distribution, and a is the parameter describing the shape of this distribution.

• Kimura 2-parameter distance

The Kimura 2-parameter distance between DNA sequences is defined by

$$-\frac{1}{2}\ln(1-2P-Q) - \frac{1}{2}\ln\sqrt{1-2Q}.$$

Tamura 3-parameter distance

The **Tamura** 3-parameter distance between DNA sequences is defined by

$$-b \ln \left(1 - \frac{P}{b} - Q\right) - \frac{1}{2}(1 - b) \ln(1 - 2Q),$$

where $f_x = \frac{1}{n} |\{1 \le i \le n : x_i = G \text{ or } C\}|, f_y = \frac{1}{n} |\{1 \le i \le n : y_i = G\}|$

G or C}|, and $b = f_x + f_y - 2f_x f_y$. In the case $f_x = f_y = \frac{1}{2}$ (so, $b = \frac{1}{2}$), it is the **Kimura 2-parameter** distance.

• Tamura-Nei distance

The Tamura-Nei distance between DNA sequences is defined by

$$-\frac{2f_{A}f_{G}}{f_{R}}\ln\left(1 - \frac{f_{R}}{2f_{A}f_{G}}P_{AG} - \frac{1}{2f_{R}}P_{RY}\right)$$

$$-\frac{2f_{T}f_{C}}{f_{Y}}\ln\left(1 - \frac{f_{Y}}{2f_{T}f_{C}}P_{TC} - \frac{1}{2f_{Y}}P_{RY}\right)$$

$$-2\left(f_{R}f_{Y} - \frac{f_{A}f_{G}f_{Y}}{f_{R}} - \frac{f_{T}f_{C}f_{R}}{f_{Y}}\right)\ln\left(1 - \frac{1}{2f_{R}f_{Y}}P_{RY}\right),$$

where $f_j = \frac{1}{2n} \sum (1_{x_i=j} + 1_{y_i=j})$ for j = A, G, T, C, and $f_R = f_A + f_G$, $f_Y = f_T + f_C$, while $P_{RY} = \frac{1}{n} | \{1 \le i \le n : |\{x_i, y_i\} \cap \{A, G\}| = |\{x_i, y_i\} \cap \{T, C\}| = 1\}|$ (the proportion of transversion differences), $P_{AG} = \frac{1}{n} | \{1 \le i \le n : \{x_i, y_i\} = \{A, G\}\}|$ (the proportion of transitions within purines), and $P_{TC} = \frac{1}{n} | \{1 \le i \le n : \{x_i, y_i\} = \{T, C\}\}|$ (the proportion of transitions within pyrimidines).

• Lake paralinear distance

Given two DNA sequences $x = (x_1, \ldots, x_n)$ and $y = (y_1, \ldots, y_n)$, denote by det(J) the determinant of the 4×4 matrix $J = ((J_{ij}))$, where $J_{ij} =$ $\frac{1}{n}|\{1 \le t \le n : x_t = i, y_t = j\}|$ (joint probability) and indices i, j = j1, 2, 3, 4 represent nucleotides A, T, C, G, respectively. Let $f_i(x)$ denote the frequency of i-th nucleotide in the sequence x (marginal probability), and let $f(x) = f_1(x)f_2(x)f_3(x)f_4(x)$. The Lake paralinear distance (1994) between sequences x and y is defined by

$$-\frac{1}{4}\ln\frac{\det(J)}{\sqrt{f(x)f(y)}}.$$

It is a **four-point inequality metric**, and it generalizes trivially for sequences over any alphabet. Related are the **LogDet distance** (Lockhart, Steel, Hendy and Penny 1994) $-\frac{1}{4}\ln\det(J)$ and the symmetrization $\frac{1}{2}(d(x,y)+d(y,x))$ of the **Barry–Hartigan quasi-metric** (1987) $d(x,y)=-\frac{1}{4}\ln\frac{\det(J)}{\sqrt{f(x)}}$.

• Eigen-McCaskill-Schuster distance

The **Eigen–McCaskill–Schuster distance** between DNA sequences $x = (x_1, \ldots, x_n)$ and $y = (y_1, \ldots, y_n)$ is defined by

$$|\{1 \le i \le n : \{x_i, y_i\} \ne \{A, G\}, \{T, C\}\}|.$$

It is the number of transversions, i.e., positions i with one of x_i, y_i denoting a purine and another one denoting a pyrimidine. It is applied to virus or cancer proliferation under control of drugs or the immune system.

• Watson-Crick distance

The **Watson-Crick distance** between DNA sequences $x = (x_1, \dots, x_n)$ and $y = (y_1, \dots, y_n)$ is defined, for $x \neq y$, by

$$|\{1 \le i \le n : \{x_i, y_i\} \ne \{A, T\}, \{G, C\}\}|,$$

i.e., it is the **Hamming metric** (cf. **No. of differences**) $\sum 1_{x_i \neq \overline{y}_i}$ between x and the Watson-Crick complement $\overline{y} = (\overline{y}_1, \dots, \overline{y}_n)$ of y. Here $\overline{y}_i = A, T, G, C$ if $y_i = T, A, C, G$, respectively.

• Hybridization metric

Hybridization is the process of combining complementary, single-stranded nucleic acids into a single molecule. Annealing is the binding of two strands by Watson-Crick complementation, i.e., interchange of all A, T, G, C by T, A, C, G, respectively. Denaturation is the reverse process of separating two strands of the double stranded DNA/RNA molecule (heating breaks the hydrogen bonds between bases).

Cf. the proximity effect in the production of chromosome aberrations among **proximity effects** in Chap. 24.

The rate of annealing of two strands (or the temperature at which denaturation occurs) measures similarity between their base sequences.

H-measure between two DNA n-sequences x and y is defined by

$$H(x,y) = \min_{-n \le k \le n} \sum 1_{x_i \ne y_{i+k}^*},$$

where the indices i + k are modulo n, and y^* is the reversal of y followed by Watson–Crick complementation.

A DNA cube is any maximal set of DNA n-sequences, such that H(x,y) = 0 for any two of them. The **hybridization metric** (Garzon, Neathery, Deaton, Murphy, Franceschetti and Stevens 1997) between DNA cubes A and B is defined by

$$\min_{x \in A, y \in B} H(x, y).$$

• Interspot distance.

A *DNA microarray* is a technology consisting of an arrayed series of thousands of *features* (microscopic spots of DNA oligonucleotides, each containing picomoles of a specific DNA sequence) that are used as probes to hybridize a *target* (cRNA sample) under high-stringency conditions. Probe-target hybridization is quantified by fluorescence-based detection of fluorophore-labeled targets to determine the relative abundance of nucleic acid sequences in the target.

The **interspot distance** is the **spacing distance** between features. Typical values are 375, 750, 1,500 μ m (1 μ m = 10⁻⁶ m)

• RNA structural distances

An RNA sequence is a string over the alphabet $\{A, C, G, U\}$ of nucleotides (bases). Inside a cell, such a string folds in 3D space, because of pairing of nucleotide bases (usually, by bonds A-U, G-C and G-U). The secondary structure of an RNA is, roughly, the set of helices (or the list of paired bases) making up the RNA. Such structure can be represented as a planar graph and further, as a rooted tree.

The *tertiary structure* is the geometric form the RNA takes in space; the secondary structure is its simplified model. The *quaternary structure* describes the arrangement of multiple protein molecules into larger complexes.

An RNA structural distance between two RNA sequences is a distance between their secondary structures. These distances are given in terms of their selected representation. For example, **tree edit distance** (and other distances on rooted trees given in Chap. 15) are based on rooted tree representation.

Let an RNA secondary structure be represented by a simple graph (V, E) with vertex-set $V = \{1, \ldots, n\}$ such that, for every $1 \le i \le n$, $(i, i+1) \notin E$ and $(i, j), (i, k) \in E$ imply j = k. Let $E = \{(i_1, j_1), \ldots, (i_k, j_k)\}$, and let (ij) denote the transposition of i and j. Then $\pi(G) = \prod_{t=1}^k (i_t j_t)$ is an involution.

Let G=(V,E) and G'=(V,E') be such planar graph representations of two RNA secondary structures. The **base pair distance** between G and G' is the number $|E\Delta E'|$, i.e., the **symmetric difference metric** between secondary structures seen as sets of paired bases.

The **Zuker distance** between G and G' is the smallest number k such that, for every edge $(i,j) \in E$, there is an edge $(i',j') \in E'$ with $\max\{|i-i'|,|j-j'|\} \le k$ and, for every edge $(k',l') \in E'$, there is an edge $(k,l) \in E$ with $\max\{|k-k'|,|l-l'|\} \le k$.

The Reidys–Stadler–Roselló metric between G and G' is defined by

$$|E\Delta E'| - 2T,$$

where T is the number of cyclic orbits of length greater than 2 induced by the action on V of the subgroup $\langle \pi(G), \pi(G') \rangle$ of the group Sym_n of permutations on V. It is the number of transpositions needed to represent $\pi(G)\pi(G')$.

Let $I_G = \langle x_i x_j : (x_i, x_j) \in E \rangle$ be the monomial ideal (in the ring of polynomials in the variables x_1, \ldots, x_n with coefficients 0, 1), and let $M(I_G)_m$ denotes the set of all monomials of total degree $\leq m$ that belong to I_G . For every $m \geq 3$, a Liabrés-Roselló **monomial metric** between G = (V, E) and G' = (V', E') is defined by

$$|M(I_G)_{m-1}\Delta M(I_{G'})_{m-1}|.$$

The secondary structure of a protein much depends on its backbone configuration, i.e., the sequence of dihedral angles defining backbone. Wang and Zheng (2007) presented a variation of **Lempel–Ziv distance** between two such sequences.

• Fuzzy polynucleotide metric

The fuzzy polynucleotide metric (or NTV-metric) is the metric introduced by Nieto, Torres and Valques-Trasande in 2003 on the 12-dimensional unit cube I^{12} . Four nucleotides U, C, A and G of RNA alphabet being coded as (1,0,0,0), (0,1,0,0), (0,0,1,0) and (0,0,0,1), respectively, 64 possible triplet codons of the genetic code can be seen as vertices of I^{12} .

So, any point $(x_1, \ldots, x_{12}) \in I^{12}$ can be seen as a fuzzy polynucleotide codon with each x_i expressing the grade of membership of element $i, 1 \le i \le 12$, in the fuzzy set x. The vertices of the cube are called the *crisp* sets.

The **NTV-metric** between different points $x, y \in I^{12}$ is defined by

$$\frac{\sum_{1 \le i \le 12} |x_i - y_i|}{\sum_{1 \le i \le 12} \max\{x_i, y_i\}}.$$

Dress and Lokot showed that $\frac{\sum_{1 \leq i \leq n} |x_i - y_i|}{\sum_{1 \leq i \leq n} \max\{|x_i|, |y_i|\}}$ is a metric on the whole of \mathbb{R}^n . On $\mathbb{R}^n_{\geq 0}$ this metric is equal to 1 - s(x, y), where $s(x, y) = \frac{\sum_{1 \leq i \leq n} \min\{x_i, y_i\}}{\sum_{1 \leq i \leq n} \max\{x_i, y_i\}}$ is the **Ruzicka similarity** (cf. Chap. 17).

• tRNA interspecies distance

An ensemble of tRNA molecules is necessary to translate triplet codons into amino acids; eukaryotes have up to 80 different tRNAs. Two tRNA molecules are called *isoacceptor tRNAs* if they bind the same amino acid.

tRNA interspecies distance between species m and n is (Xue, Tong, Marck, Grosjean and Wong 2003), averaged for all 20 amino acids, tRNA

distance for given amino acid aa_i , which is, averaged for all pairs, **Jukes–Cantor protein distance** between each of the one or more isoacceptor tRNAs of aa_i from species m and each of the one or more isoacceptor tRNAs of the same amino acid from species n.

• Whole genome composition distance

Let A_k denote the set of all $\sum_{i=1}^k 4^i$ non-empty words of length at most k over the alphabet of four RNA nucleotides. For an RNA sequence $x = (x_1, \ldots, x_n)$ and any $a \in A_k$, let $f_a(x)$ denote the number of occurrences of a as a block (contiguous subsequence) in x divided by the number of blocks of the same length in x.

The whole genome composition distance (or, WGCD, Wu, Goebel, Wan and Lin 2006) between RNA sequences x and y (of two strains of HIV-1 virus) is defined as the Euclidean distance

$$\sqrt{\sum_{a \in A_k} (f_a(x) - f_a(y))^2}.$$

Cf. k-mer distance and, in Chap. 11, q-gram similarity.

• Genome rearrangement distances

The genomes of related unichromosomal species or single chromosome organelles (such as small viruses and mitochondria) are represented by the order of genes along chromosomes, i.e., as permutations (or rankings) of a given set of n homologous genes. If one takes into account the directionality of the genes, a chromosome is described by a signed permutation, i.e., by a vector $x = (x_1, \ldots, x_n)$, where $|x_i|$ are different numbers $1, \ldots, n$, and any x_i can be positive or negative. The circular genomes are represented by circular (signed) permutations $x = (x_1, \ldots, x_n)$, where $x_{n+1} = x_1$ and so on.

Given a set of considered mutation moves, the corresponding *genomic distance* between two such genomes is the **editing metric** (cf. Chap. 11) with the editing operations being these moves, i.e., the minimal number of moves needed to transform one (signed) permutation into another.

In addition to (and, usually, instead of) local mutation events, such as character indels or replacements in the DNA sequence, the *large* (i.e., happening on a large portion of the chromosome) mutations are considered, and the corresponding genomic editing metrics are called **genome rearrangement distances**. In fact, such rearrangement mutations being rarer, these distances estimate better the true genomic evolutionary distance. The main genome (chromosomal) rearrangements are *inversions* (block reversals), *transpositions* (exchanges of two adjacent blocks) in a permutation, and also *inverted transposition* (inversion combined with transposition) and, for signed permutations only, *signed reversals* (sign reversal combined with inversion).

The main genome rearrangement distances between two unichromosomal genomes are:

reversal metric and signed reversal metric (cf. Chap. 11);

transposition distance: the minimal number of transpositions needed to transform (permutation representing) one of them into another;

ITT-distance: the minimal number of inversions, transpositions and inverted transpositions needed to transform one of them into another.

Given two circular signed permutations $x = (x_1, \ldots, x_n)$ and $y = (y_1, \ldots, y_n)$ (so, $x_{n+1} = x_1$ etc.), a breakpoint is a number $i, 1 \le i \le n$, such that $y_{i+1} \ne x_{j(i)+1}$, where the number $j(i), 1 \le j(i) \le n$, is defined by the equality $y_i = x_{j(i)}$. The **breakpoint distance** (Watterson, Ewens, Hall and Morgan 1982) between genomes, represented by x and y, is the number of breakpoints.

This distance and the **permutation editing metric** (the **Ulam metric** from Chap. 11: the minimal needed number of character moves, i.e., 1-character transpositions) are used for the approximation of genome rearrangement distances.

• Syntenic distance

This is a genomic distance between multichromosomal genomes, seen as unordered collections of synteny sets of genes, where two genes are syntenic if they appear in the same chromosome. The syntenic distance (Ferretti, Nadeau and Sankoff 1996) between two such genomes is the minimal number of mutation moves – translocations (exchanges of genes between two chromosomes), fusions (merging of two chromosomes into one) and fissions (splitting of one chromosome into two) – needed to transfer one genome into another. All (input and output) chromosomes of these mutations should be non-empty and not duplicated. The above three mutation moves correspond to interchromosomal genome rearrangements, which are rarer than intrachromosomal ones; so, they give information about deeper evolutionary history.

• Strand length

A single strand of nucleic acid (DNA or RNA sequence) is oriented down-stream, i.e., from 5' end toward 3' end (sites terminating at fifth and third carbon in the sugar-ring; 5'-phosphate binds covalently to the 3'-hydroxyl of another nucleotide). So, the structures along it (genes, transcription factors, polymerases) are either downstream or upstream. The strand length is the distance from its 5' to 3' end.

For a molecule of messenger RNA (mRNA), the **gene length** is the distance from the cap site 5', where post-translational stability is ensured, to the polyadenylation site 3', where a poly(A) tail of 50–250 adenines is attached after translation.

• Genome distance

The **genome distance** between two loci on a chromosome is the number of base pairs (bp) separating them on the chromosome.

In particular, the **intragenic distance** of two neighboring genes is the smallest distance in base pairs separating them on the chromosome. Sometimes, it is defined as the genome distance between the transcription start sites of those genes.

Nelson, Hersh and Carrol (2004) defined the *intergenic distance* of a gene as the amount of non-coding DNA between the gene and its nearest neighbors, i.e., the sum of upstream and downstream distances, where *upstream distance* is the genome distance between the start of a gene's first exon and the boundary of the closest upstream neighboring exon (irrespective of DNA strand) and *downstream distance* is the distance between the end of a gene's last exon and the boundary of the closest downstream neighboring exon. If exons overlap, the intergenic distance is 0.

• Map distance

The **map distance** between two loci on a genetic map is the recombination frequency expressed as a percentage; it is measured in *centimorgans* cM (or *map units*), where 1 cM corresponds to their statistically corrected recombination frequency 1%. It may be either recombination of genes (chromosome map) or within genes (gene map).

Typically, a map distance of 1 cM (genetic scale) corresponds to a **genome distance** (physical scale) of about one megabase (million base pairs) Mp.

• Action at a distance along a DNA

An action at a distance along a DNA happens when an event at one location on a DNA molecule affects an event at a distant (say, more than 2,500 base pairs) location on the same molecule.

Many genes are regulated by distant (up to a million base pairs away and, possibly, located on another chromosome) or short (30–200 base pairs) regions of DNA, *enhancers*. Enhancers increase the probability of such a gene to be transcribed in a manner independent of distance and position (the same or opposite strand of DNA) relative to the transcription initiation site (the promoter). The enhancer function can be preserved even if it is moved on the chromosome or its orientation is reversed.

DNA supercoiling is the coiling of a DNA double helix on itself (twisting around the helical axis once every 10.4 base pairs of sequence, forming circles and figures of eight) because it has been bent, overwound or underwound. Such folding puts a long-range enhancer, which is far from a regulated gene in **genome distance**, geometrically closer to the promoter. The genomic radius of regulatory activity of a genome is the genome distance of the most distant known enhancer from the corresponding promoter; in the human genome it is $\approx 1 \,\mathrm{Mp}$ (for enhancer of SSH, Sonic Hedgehog gene).

There is evidence that genomes are organized into enhancer-promoter loops. But the long-range enhancer function is not fully understood yet. Akbari, Bae, Johnsen, Villaluz, Wong and Drewell (2008) explain it by the action of a *tether*, i.e., a sequence next to a promoter that, as a kind of postal code, specifically attracts the enhancer.

Similarly, some viral RNA elements interact across thousands of intervening nucleotides to control translation, genomic RNA synthesis, and subgenomic mRNA transcription.

23.3 Distances for protein data

A protein sequence (or primary protein structure) is a sequence $x = (x_1, \ldots, x_n)$ over a 20-letter alphabet of 20 amino acids; \sum denotes $\sum_{i=1}^n$.

There are many notions of similarity/distance (20×20 scoring matrices) on the set of 20 amino acids, based on genetic codes, physico-chemical properties, observed frequency of mutations, secondary structural matching and structural properties. The most important is the 20×20 Dayhoff PAM250 matrix which expresses the relative mutability of 20 amino acids.

• PAM distance

The **PAM** distance (or **Dayhoff–Eck** distance, *PAM* value) between protein sequences is defined as the minimal number of accepted (i.e., fixed) point mutations per 100 amino acids needed to transform one protein into another.

1 PAM is a unit of evolution: it corresponds to 1 point mutation per 100 amino acids. PAM values 80, 100, 200, 250 correspond to the distance (in %) 50, 60, 75, 92 between proteins.

• Genetic code distance

The **genetic code distance** (Fitch and Margoliash 1967) between amino acids x and y is the minimum number of nucleotides that must be changed to obtain x from y. In fact, it is 1, 2 or 3, since each amino acid corresponds to three bases.

• Miyata-Miyazawa-Yasanaga distance

The Miyata-Masada-Yasanaga distance (or Miyata's biochemical distance 1979) between amino acids x, y with polarities p_x, p_y and volumes v_x, v_y , respectively, is

$$\sqrt{(\frac{|p_x - p_y|}{\sigma_p})^2 + (\frac{|v_x - v_y|}{\sigma_v})^2},$$

where σ_p and σ_v are standard deviations of $|p_x - p_y|$ and $|v_x - v_y|$, respectively.

This distance is derived from the similar *Grantam's chemical distance* (Grantam 1974) based on polarity, volume and carbon-composition of amino acids.

• Polar distance

The following three physico-chemical distances between amino acids x and y were defined in Hughes, Ota and Nei (1990).

Dividing amino acids into two groups – polar(C, D, E, H, K, N, Q, R, S, T, W, Y) and non-polar (the rest) – the **polar distance** is 1 if x, y belong to different groups, and 0 otherwise. The second polarity distance is the absolute difference between the polarity indices of x and y.

Dividing amino acids into three groups – positive (H, K, R), negative (D, E) and neutral (the rest) – the **charge distance** is 1 if x, y belong to different groups, and 0 otherwise.

• Feng-Wang distance

In [FeWa08], 20 amino acids are ordered linearly by their rank-scaled functions CI, NI of pK_a values for the terminal amino acid groups COOH and NH $_3^+$, respectively. 17CI is 1,2,3,4,5,6,7,7,8,9,10,11,12,13,14, 14,15,15,16,17 for C,H,F,P,N,D,R,Q,K,E,Y,S,M,V,G,A,L,I,W,T, while 18NI is 1,2,3,4,5,5,6,7,8,9,10,10,11,12,13,14,15,16,17,18 for N,K,R,Y,F,Q,S,H,M,W,G,L,V,E,I,A,D,T,P,C.

Given a protein sequence $x = (x_1, \ldots, x_m)$, define $x_i < x_j$ if i < j, $CI(x_i) < CI(x_i)$ and $NI(x_i) < NI(x_i)$ hold. Represent the sequence x by the augmented $m \times m$ Hasse matrix $((a_{ij}(x)))$, where $a_{ii}(x) = \frac{CI(x_i) + NI(x_i)}{2}$ and, for $i \neq j$, $a_{ij}(x) = -1$, 1 or 0 if, respectively, $x_i < x_j$, $x_i \geq x_j$ or otherwise.

The **Feng–Wang distance** between protein sequences $x = (x_1, \ldots, x_m)$ and $y = (y_1, \ldots, y_n)$ is $||\frac{\lambda(x)}{\sqrt{m}} - \frac{\lambda(y)}{\sqrt{n}}||_2$, where $\lambda(z)$ denotes the largest eigenvalue (possibly, a complex number) of the matrix $((a_{ij}(z)))$.

• No. of protein differences

The **No. of protein differences** is just the **Hamming metric** between protein sequences:

$$\sum 1_{x_i \neq y_i}.$$

• Amino p-distance

The amino p-distance (or uncorrected distance) d_p between protein sequences is defined by

$$\frac{\sum 1_{x_i \neq y_i}}{n}.$$

• Amino Poisson correction distance

The amino Poisson correction distance between protein sequences is defined, via the amino p-distance d_p , by

$$-\ln(1-d_p(x,y)).$$

Amino gamma distance

The amino gamma distance (or Poisson correction gamma distance) between protein sequences is defined, via the amino p-distance d_p , by

$$a((1 - d_p(x, y))^{-1/a} - 1),$$

where the substitution rate varies with i = 1, ..., n according to the gamma distribution, and a is the parameter describing the shape of the

distribution. For a = 2.25 and a = 0.65, it estimates the **Dayhoff** and **Grishin distances**, respectively. In some applications, this distance with a = 2.25 is called simply the **Dayhoff distance**.

• Jukes-Cantor protein distance

The Jukes-Cantor protein distance between protein sequences is defined, via amino p-distance d_p , by

$$-\frac{19}{20}\ln\left(1-\frac{20}{19}d_p(x,y)\right).$$

• Kimura protein distance

The **Kimura protein distance** between protein sequences is defined, via the **amino** p-distance d_p , by

$$-\ln\left(1-d_p(x,y)-\frac{d_p^2(x,y)}{5}\right).$$

• Grishin distance

The **Grishin distance** d between protein sequences can be obtained, via the **amino** p-distance d_p , from the formula

$$\frac{\ln(1+2d(x,y))}{2d(x,y)} = 1 - d_p(x,y).$$

• k-mer distance

The k-mer distance (Edgar 2004) between sequences $x = (x_1, \ldots, x_m)$ and $y = (y_1, \ldots, y_n)$ over a compressed amino acid alphabet is defined by

$$\ln\left(\frac{1}{10} + \frac{\sum_{a} \min\{x(a), y(a)\}}{\min\{m, n\} - k + 1}\right),\,$$

where a is any k-mer (a word of length k over the alphabet), while x(a) and y(a) are the number of times a occurs in x and y, respectively, as a block (contiguous subsequence). Cf. q-gram similarity in Chap. 11.

• Immunological distance

An antigen (or immunogen, pathogen) is any molecule eliciting immune response. Once it gets into the body, the immune system either neutralizes its pathogenic effect or destroys the infected cells. the most important cells in this response are white blood cells: *T-cells* and *B-cells* responsible for the production and secretion of antibodies (specific proteins that bind to the antigen). When an antibody strongly matches an antigen, the corresponding B-cell is stimulated to divide, produce clones of itself that then produce more antibodies, and then differentiate into a plasma or memory cell. A secreted antibody binds to antigen, and antigen—antibody complexes are removed.

A mammal (usually a rabbit) when injected with an antigen will produce immunoglobulins (antibodies) specific for this antigen. Then *antiserum* (blood serum containing antibodies) is purified from the mammal's serum. The produced antiserum is used to pass on passive immunity to many diseases.

Immunological distance procedures (*immunodiffusion* and, the main now, *micro-complement fixation*) measure relative strengths of the immunological responses to antigens from different taxa. This strength is dependent upon the similarity of the proteins, and the dissimilarity of the proteins is related to the evolutionary distance between the taxa concerned.

The index of dissimilarity id(x, y) between two taxa x and y is the factor $\frac{r(x,x)}{r(x,y)}$ by which the heterologous (reacting with an antibody not induced by it) antigen concentration must be raised to produce a reaction as strong as that to the homologous (reacting with its specific antibody) antigen.

The **immunological distance** between two taxa is given by

$$100(\log id(x,y) + \log id(y,x)).$$

It can be 0 between two closely related species.

Earlier immunodiffusion procedures compared the amount of precipitate when heterologous bloods were added in similar amounts as homologous ones, or compared the highest dilution giving a positive reaction.

The name of applied antigen (target protein) can be used to specify immunological distance, say, albumin, transferring lysozyme distances. Proponents of the *molecular clock hypothesis* estimate that 1 unit of albumin distance between two taxa corresponds to $\approx 540,000$ years of their divergence time, and that 1 unit of **Nei standard genetic distance** corresponds to 18–20 million years.

Adams and Boots (2006) call the *immunological distance* between two immunologically similar pathogen strains (actually, serotypes of dengue virus) their *cross-immunity*, i.e., 1 minus the probability that primary infection with one strain prevents secondary infection with the other. Lee and Chen (2004) define the *antigenetic distance* between two influenza viruses to be the reciprocal of their *antigenetic relatedness* which is (presented as a percentage) geometric mean $\sqrt{\frac{r(x,y)}{r(x,x)}\frac{r(y,x)}{r(y,y)}}$ of two ratios between the heterologous and homologous antibody *titers*.

An antiserum *titer* is a measurement of concentration of antibodies found in a serum. Titers are expressed in their highest positive dilution; for example, the antiserum dilution required to obtain a reaction curve with given peak height (say, 75% microcomplement fixed), or the reciprocal of the dilution consistently showing a twofold increase in absorbency over that obtained with the pre-bleed serum sample.

• Pharmacological distance

The *protein kinases* are enzymes which transmit signals and control cells using transfer of *phosphate groups* from high-energy donor molecules

to specific target proteins. So, many drug molecules (against cancer, inflammation, etc.) are kinase inhibitors (blockers). But their high cross-reactivity often leads to toxic side effects. Hence, designed drugs should be specific (say, not to bind to $\geq 95\%$ of other proteins).

Given a set $\{a_1, \ldots, a_n\}$ of drugs in use, the affinity vector of kinase x is defined as $(-\ln B_1(x), \ldots, -\ln B_n(x))$, where $B_i(x)$ is the binding constant for the reaction of x with drug a_i , and $B_i(x) = 1$ if no interaction was observed. The binding constants are the average of several experiments where the concentration of binded kinase is measured at equilibrium. The **pharmacological distance** (Fabian et al. 2005) between kinases x and y is the Euclidean distance $(\sum_{i=1}^n (\ln B_i(x) - \ln B_i(y))^2)^{\frac{1}{2}}$ between their affinity vectors.

The secondary structure of a protein is given by the hydrogen bonds between its residues. A dehydron in a solvable protein is a hydrogen bond which is solvent-accessible. The dehydron matrix of kinase x with residue-set $\{R_1, \ldots, R_m\}$ is the $m \times m$ matrix $((D_{ij}(x)))$, where $D_{ij}(x)$ is 1 if residues R_i and R_j are paired by a dehydron and is 0 otherwise. The **packing distance** (Maddipati and Fernándes 2006) between kinases x and y is the Hamming distance $\sum_{1 \leq i,j \leq m} |D_{ij}(x) - D_{ij}(y)|$ between their dehydron matrices; cf. base pair distance among RNA structural distances. The environmental distance (Chen, Zhang and Fernándes 2007) between kinases is a normalized variation of their packing distance.

• Förster distance

FRET (fluorescence resonance energy transfer; Főrster 1948) is a distance-dependent quantum mechanical property of a fluorophore (molecule component causing its fluorescence) resulting in direct non-radiative energy transfer between the electronic excited states of two dye molecules, the donor fluorophore and a suitable acceptor fluorophore, via a dipole. In FRET microscopy, fluorescent proteins are used as non-invasive probes in living cells since they fuse genetically to proteins of interest. The efficiency of FRET transfer depends on the square of the donor electric field magnitude, and this field decays as the inverse sixth power of intermolecular separation (the physical donor–acceptor distance). The distance at which this energy transfer is 50% efficient, i.e., 50% of excited donors are deactivated by FRET, is called the **Főrster distance** of these two fluorophores.

Measurable FRET occurs only if the donor–acceptor distance is less than $\approx 10 \,\mathrm{nm}$, the mutual orientation of the molecules is favorable, and the spectral overlap of the donor emission with acceptor absorption is sufficient.

23.4 Other biological distances

Here we collect the main examples of other notions of distance and distancerelated models used in Biology.

• Migration distance (in Biomotility)

Migration distance (or penetration distance), in Cattle Reproduction and human infertility diagnosis, is the distance (in millimeters) traveled by the vanguard spermatozoon during sperm displacement in vitro through a capillary tube filled with homologous cervical mucus or a gel mimicking it. Sperm can swim about 10–20 body lengths per second.

Such measurements, under different specifications (duration, temperature, etc.) of incubation, estimate the ability of spermatozoa to colonize the oviduct in vivo.

In general, the term **migration distance** is used in biological measurements of directional motility using controlled migration; for example, determining the molecular weight of an unknown protein via its migration distance through a gel, or comparing the migration distance of mast cells in different peptide media.

• Penetration distance

Penetration distance is, similarly to migration distance, a general term used in (especially, biological) measurements for the distance from the given surface to the point where the concentration of the penetrating substance (say, a drug) in the medium (say, a tissue) had dropped to the given level. Several examples follow.

During penetration of a macromolecular drug into the tumor interstitium, tumor interstitial penetration is the distance that the drug carrier moved away from the source at a vascular surface; it is measured in 3D to the nearest vascular surface.

During the intraperitoneal delivery of cisplatin and heat to tumor metastases in tissues adjacent to the peritoneal cavity, the *penetration distance* is the depth to which drug diffuses directly from the cavity into tissues. Specifically, it is the distance beyond which such delivery is not preferable to intravenous delivery. It can be the distance from the cavity surface into the tissues within which drug concentration is, for example, a) greater, at a given time point, than that in control cells distant from the cavity, or b) is much higher than in equivalent intravenous delivery, or c) has first peak approaching its plateau value within 1% deviation.

The *penetration distance* of a drug in the brain is the distance from the probe surface to the point where the concentration is roughly half its far-field value.

The penetration distance of chemicals into wood is the distance between the point of application and the 5 mm cut section in which the contaminants concentration is at least 3% of the total.

The forest edge-effect penetration distance is the distance to the point where invertebrate abundance ceased to be different to forest interior abundance. Cf. **penetration depth distance** in Chap. 9 and **penetration depth** in Chap. 24.

• Capillary diffusion distance

One of diffusion processes is osmosis, i.e., the net movement of water through a permeable membrane to a region of lower solvent potential. In the respiratory system (the alveoli of mammalian lungs), oxygen O_2 diffuses into the blood and carbon dioxide CO_2 diffuses out.

Capillary diffusion distance is, similarly to penetration distance, a general term used in biological measurements for the distance, from the capillary blood through the tissues to the mitochondria, to the point where the concentration of oxygen has dropped to the given low level.

This distance is measured as, say, the average distance from the capillary wall to the mitochondria, or the distance between the closest capillary endothelial cell to the epidermis, or in percentage terms. For example, it can be the distance where a given percentage (95% for maximal, 50% for average) of the fiber area is served by a capillary. Or the percent cumulative frequency of fiber area within a given distance of the capillary when the capillary to fiber ratio is increased, say, from 0.5 to 4.0.

Another practical example: the *effective diffusion distance* of nitric oxide NiO in microcirculation in vivo is the distance within which Ni concentration is greater than the equilibrium dissociation constant of the target enzyme for oxide action.

Cf. the **immunological distance** for immunodiffusion and, in Chap. 29, the **diffusion tensor distance** among **distances in Medicine**.

• Gendron et al. distance

The **Gendron et al. distance** (Gendron, Lemieux and Major 2001) between two base–base interactions, represented by 4×4 homogeneous transformation matrices X and Y, is defined by

$$\frac{S(XY^{-1}) + S(X^{-1}Y)}{2},$$

where $S(M) = \sqrt{l^2 + (\theta/\alpha)^2}$, l is the length of translation, θ is the angle of rotation, and α represents a scaling factor between the translation and rotation contributions.

• Metabolic distance

The **metabolic distance** (or *pathway distance*) between enzymes is the minimum number of metabolic steps separating two enzymes in the metabolic pathways.

• Spike train distances

A human brain has $\approx 10^{11}$ of *neurons* (nerve cells) among $\approx 10^{14}$ cells in the human body. Neuronal response to a stimulus is a continuous time series. It can be reduced, by a threshold criterion, to a much simpler discrete series of *spikes* (short electrical pulses).

A spike train is a sequence $x = (t_1, \ldots, t_s)$ of s events (neuronal spikes, or heart beats, etc.) listing absolute spike times or inter-spike time intervals.

The main distances between spike trains $x = x_1, ..., x_m$ and $y = y_1, ..., y_n$ follow:

1. The **spike count distance** is defined by

$$\frac{|n-m|}{\max\{m,n\}}.$$

2. The **firing rate distance** is defined by

$$\sum_{1 \le i \le s} (x_i' - y_i')^2,$$

where $x' = x'_1, \ldots, x'_s$ is the sequence of local firing rates of train $x = x_1, \ldots, x_m$ partitioned in s time intervals of length T_{rate} .

3. Let $\tau_{ij} = \frac{1}{2}\min\{x_{i+1} - x_i, x_i - x_{i-1}, y_{i+1} - y_i, y_i - y_{i-1}\}$ and $c(x|y) = \sum_{i=1}^{m} \sum_{j=1}^{n} J_{ij}$, where $J_{ij} = 1, \frac{1}{2}, 0$ if $0 < x_i - y_i \le \tau_{ij}, x_i = y_i$, and otherwise, respectively. The **event synchronization distance** (Quiroga, Kreuz and Grassberger 2002) is defined, respectively, by

$$1 - \frac{c(x|y) + c(y|x)}{\sqrt{mn}}.$$

4. Let $x_{isi}(t) = \min\{x_i : x_i > t\} - \max\{x_i : x_i < t\}$ for $x_1 < t < x_m$, and let $I(t) = \frac{x_{isi}(t)}{y_{isi}(t)} - 1$ if $x_{isi}(t) \le x_{isi}(t)$ and $I(t) = 1 - \frac{y_{isi}(t)}{x_{isi}(t)}$ otherwise. The time-weighted and spike-weighted variants of **ISI distances** (Kreuz, Haas, Morelli, Abarbanel and Politi 2007) are defined by

$$\int_0^T |I(t)|dt \text{ and } \sum_{i=1}^m |I(x_i)|.$$

- 5. Various *information distances* were applied to spike trains: **Kullback–Leibler distance**, **Chernoff distance** (cf. Chap. 14). Also, if x and y are mapped into binary sequences, the **Lempel–Ziv distance** and a version of the **normalized information distance** (cf. Chap. 11.) were used.
- 6. The Victor-Purpura distance is a cost-based editing metric (i.e., the minimal cost of transforming x into y) by the following operations with their associated costs: insert a spike (cost 1), delete a spike (cost 1), shift a spike by time t (cost qt); here q > 0 is a parameter. Victor and Purpura introduced this distance in 1996; the fuzzy Hamming distance (cf. Chap. 11), introduced in 2001, identifies cost functions of shift preserving the triangle inequality.

7. The van Rossum distance 2001, is defined by

$$\sqrt{\int_0^\infty (f_t(x) - f_t(y))^2 dt},$$

where x is convoluted with $h_t = \frac{1}{\tau}e^{-t/\tau}$ and $\tau \approx 12 \,\text{ms}$ (best); $f_t(x) = \sum_{i=0}^{m} h(t-x_i)$. The Victor-Purpura distance and van Rossum distance are the most commonly used metrics.

8. If components of x, y are seen as the samples of two 0-mean random variables, then the **cross-correlation distances** between x and y are defined by

$$1 - \frac{\langle f(x), f(y) \rangle}{||f(x)||_2 \cdot ||f(y)||_2},$$

where f(x) is the train x filtered by convolution with a kernel function $f(\cdot)$. This function is exponential (Haas and White 2002) or Gaussian (Schreiber, Fellous, Whitmer, Tiesinga and Sejnowski 2004).

9. Given two sets of spike trains labeled by neurons firing them, the **Aronov et al. distance** (Aronov, Reich, Mechler and Victor 2003) between them is a cost-based **editing metric** (i.e., the minimal cost of transforming one into the other) by the following operations with their associated costs: insert or delete a spike (cost 1), shift a spike by time t (cost qt), relabel a spike (cost k), where q and k are positive parameters.

• Prototype distance

Given a finite metric space (X,d) (usually, a Euclidean space) and a selected, as typical by some criterion, vertex $x_0 \in X$, called the *prototype* (or *centroid*), the **prototype distance** of every $x \in X$ is the number $d(x,x_0)$. Usually, the elements of X represent phenotypes or morphological traits. The average of $d(x,x_0)$ over $x \in X$ estimates the corresponding variability.

• Biotope distance

The biotopes here are represented as binary sequences $x = (x_1, ..., x_n)$, where $x_i = 1$ means the presence of the species i. The **biotope distance** (or **Tanimoto distance**) between biotopes x and y is defined by

$$\frac{|\{1 \le i \le n : x_i \ne y_i\}|}{|\{1 \le i \le n : x_i + y_i > 0\}|} = \frac{|X\Delta Y|}{|X \cup Y|},$$

where $X = \{1 \le i \le n : x_i = 1\}$ and $Y = \{1 \le i \le n : y_i = 1\}$.

• Niche overlap similarity

Let $p(x) = (p_1(x), \dots, p_n(x))$ be a frequency vector (i.e., all $p_i(x) \ge 0$ and $\sum_i p_i(x) = 1$) representing an ecological niche of species x, for instance, the proportion of resource $i, i \in \{1, \dots, n\}$ used by species x.

Niche overlap similarity (or $Pianka's index O_{xy}$) of species x and y is the term often (starting with Pianka 1973) used in Ecology for the cosine similarity (cf. Chap. 17)

$$\frac{\langle p(x), p(y) \rangle}{||p(x)||_2 \cdot ||p(y)||_2}.$$

• Ecological distance

Let a given species be distributed in subpopulations over a given landscape, i.e., a textured mosaic of patches (homogeneous areas of land use, as fields, lakes, forest) and linear frontiers (river shores, hedges and road sides). The individuals move across the landscape, preferentially by frontiers, until they reach a different subpopulation or they exceed a maximum dispersal distance.

The **ecological distance** between two subpopulations (patches) x and y is (Vuilleumier and Fontanillas 2007):

$$\frac{D(x,y) + D(y,x)}{2},$$

where D(x, y) is the distance an individual covers to reach patch y from patch x, averaged over all successful dispersers from x to y. If no such dispersers exist, put $D(x, y) = \min_{z} (D(x, z) + D(z, x))$.

The ecological distance in some heterogeneous landscapes depends more on the genetic than the geographic (Euclidean) distance. The term *distance* is used also to compare the species composition of two samples; cf. **biotope** distance.

• Dispersal distance

In Biology, the **dispersal distance** is a **range distance** to which a species maintains or expands the distribution of a population. It refers, for example, to seed dispersal by pollination and to natal, breeding and migration dispersal.

• Long-distance dispersal

Long-distance dispersal (or *LDD*, Kot, Lewis and van den Driessche 1996) refers to the rare events of biological dispersal (especially, plants) on distances an order of magnitude greater than the median **dispersal** distance.

Together with vicarience theory (dispersal via land bridges) based on continental drift, LDD emerged in Biogeography as the main factor of biodiversity and species migration patterns. It explained the fast spread of different organisms in new habitats, for example in paleocolonization events, plant pathogens, and invasive species. For the regional survival of some plants, LDD is more important than local (median-distance) dispersal.

Also, cancer invasion (spread from primary tumors invading new tissues) can be thought as an invasive species spread via LDD, followed by localized dispersal.

Transoceanic LDD by wind currents is a probable source of the strong floristic similarities among landmasses in the southern hemisphere.

Examples of other LDD vehicles are: rafting by water (corals can traverse 40,000 km during their lifetime), migrating birds, human transport, extreme climatic events.

• Island distance effect

An *island*, in Biogeography, is any area of habitat surrounded by areas unsuitable for the species on the island: true islands surrounded by ocean, mountains surrounded by deserts, lakes surrounded by dry land, forest fragments surrounded by human-altered landscapes. The **island distance effect** is that the number of species found on an island is smaller when the degree of isolation (distance to nearest neighbor and mainland) is larger. The second main factor of island species diversity is its size: the chance of extinction is greater on smaller islands.

• Migration distance (in Biogeography)

Migration distance, in Biogeography, is the distance between regular breeding and non-breeding areas within annual large-scale return movement of birds, fish and insects. The longest such distance recorded electronically is 64,000 km (sooty shearwaters flying from New Zealand to the North Pacific Ocean); its rival in migration distance, the Arctic tern, is too small to be fit with electronic tags.

Migration differs from ranging, i.e., a movement of an animal beyond its home range which ceases when suitable resource (food, mates, shelter) is found; for example, wandering albatrosses make several seasonal foraging round trips of up to 3,000 km. So, Kennedy (1985) defined migratory behavior as persistent and straightened-out movement effected by the animal's own locomotory exertions or by its active embarkation upon a vehicle (say, wind or water currents).

• Isolation-by-distance

Isolation-by-distance is a biological model predicting that the genetic distance between populations increases exponentially with respect to geographic distance. So, emergence of regional differences (races) and new species is explained by restricted gene flow and adaptive variations. Isolation-by-distance for humans was studied, for example, via the distribution of surnames (cf. **Lasker distance**).

Speciation by force of distance is a speciation despite gene flow between populations. It was observed in *ring species*, i.e., two reproductively isolated populations connected by gene flow through a chain of intergrading populations. For example, Irwin, Bensch, Irwin and Price (2005) showed gradual genetic change between greenish warblers in west and east Siberia; they coexist without interbreeding in central Siberia.

Malecot's distance model

The Malecot's distance model is a migratory model of isolation by distance, expressed by the following Malecot's equation for the dependency ρ_d of alleles at two loci at distance d (allelic association, linkage disequilibrium, polymorphism distance):

$$\rho_d = (1 - L)Me^{\epsilon d} + L,$$

where d is the distance between loci along the chromosome (either **genome distance** on the physical scale in kilobases, or **map distance** on the genetic scale in centimorgans), ϵ is a constant for a specified region, $M \leq 1$ is a parameter expressing mutation rate and L is the parameter predicting association between unlinked loci.

• Distances in Animal Behavior

The first definitions of such distances were derived in Zoo Biology by Hediger (1950) namely, fight distance (run boundary), critical distance (attack boundary), personal distance (separating members of non-contact species) and social distance (for intraspecies communication). Cf. Hall's distances between people in Chap. 28. The main distances in Animal Behavior follow.

The **individual distance**: the distance which an animal attempts to maintain between itself and other animals. It ranges between "proximity" and "far apart" (for example, $\leq 8 \,\mathrm{m}$ and $\geq 61 \,\mathrm{m}$ in elephant social calls).

The **group distance**: the distance which a group of animals attempts to maintain between it and other groups.

The **nearest-neighbor distance**: the more or less constant distance which an animal maintains, in directional movement of large groups (such as schools of fish or flocks of birds), from its immediate neighbors. The mechanism of *allelomimesis* ("do what your neighbor does") prevents the structural breakdown of a group and can generate seemingly intelligent evasive maneuvers in the presence of predators. When this distance decreases, the mode of movement of the group can change: marching locusts align, ants build bridges, etc.

The **flight initiation distance** (FID): the distance from the predator when escape begins. The **alert distance**: the distance between a predator and the prey when the prey turns towards the predator.

The **escape distance**: the distance at which the animal reacts on the appearance of a predator or dominating animal of the same species.

The **reaction distance**: the distance at which the animal reacts to the appearance of prey; *catching distance*: the distance at which the predator can strike a prey.

In general, the **detection distance**: the maximal distance from the observer at which the individual or cluster of individuals is seen or heard. For example, the approximate visual detection distance is 2,000 m for an eagle searching for displaying sage-grouse, 200 m for a female sage-grouse searching for a male, and 1,450 m for a sage-grouse scanning for a flying eagle.

An example of distance estimation (for prey recognition) by some insects: the velocity of the mantid's head movement is kept constant during peering; so, the distance to the target is inversely proportional to the velocity of the retinal image.

Example of unexplained distance prediction by animals is given (Vannini, Lori, Coffa and Fratini 2008) by snails Cerithidea decollata migrating up and down in mangrove shore in synchrony with tidal phases. In the absence of visual cues and chemical marks, snails cluster just above the high water line, and the distance from the ground correlates better with the incoming tide level than with previous tide levels.

The gaze following: great apes, ravens and canids follow another's gaze direction (head and eye orientation) into distant space, moreover, geometrically behind an obstacle. It could be, more than a mere co-orienting reflex, an understanding that other has different perception and knowledge, i.e., a precursor to *Theory of Mind* (the ability to attribute mental states to oneself and others). But such interpretation, as well as *foresight* (mental time travel) in non-human animals is still controversial.

Humans and, perhaps, chimpanzees posess, besides landmark-based representation of space, more flexible Euclidean mental map.

The interpupillary distance of mammals in non-leafy environments increases as $d \approx M^{\frac{1}{3}}$ (M is the body mass); their eyes face sideways in order to get the panoramic vision. In leafy environments, this distance is constrained by the maximum leaf size. Changizi and Shimojo (2008) suggested that the degree of binocular convergence is selected to maximize how much the mammal can see. So, in cluttered (say, leafy) environments, forward-facing eyes (and smaller distance d) are better.

The **distance-to-shore**: the distance to the coastline used to study clustering of whale strandings (by distorted echolocation, anomalies of magnetic field, etc.).

Daily distance traveled and feeding time are much greater in larger groups of primates, according to a meta-analysis in Majolo et al. (2008). Also, larger groups spent slightly more time grooming and less time resting than smaller groups.

In the main non-resource-based mating system, *lek mating*, females in estrous visit a congregation of displaying males, the *lek*, for fertilization, and mate preferentially with males of higher **lekking distance rank**, i.e., relative distance from male *territory* (the median of his positions) to the center of the lek. Dominance rank often influences *space use*: high-ranking individuals have smaller, centrally located (so, less to travel and more secure) home ranges.

A distance pheromone, in animal olfactory communication, is a soluble (for example, in the urine) and/or evaporable substance emitted by an animal, as a chemosensory cue, in order to send a message (on alarm, sex, food trail, recognition, etc.) to other members of the same species. In contrast, a *contact pheromone* is such an insoluble non-evaporable substance; it coats the animal's body and is a contact cue. The *action radius* of a distance pheromone is its attraction range, the maximum distance over which animals can be shown to direct their movement to a source.

Distance effect avoiding refers to observed selection (contrary to typical decision-making of a central place forager) of some good distant source of

interest over a poor but nearer one in the same direction. For example, females at a chorusing lek of anurians or arthropods may use the lower pitch of a bigger/better distant male's call to select it over louder to her weaker call nearby. High-quality males help them by placing their calls to precede or follow those of inferior males. Franks et al. (2007) showed that ant colonies are able to select a good distant nest over a poorer one in the way, even nine times closer. Ants might compensate for distance effect by increasing recruitment latencies and quorum thresholds at nearby poor nests; also their scouts, founding a low-quality nest, start to look for a new one.

Matters of relevance at a distance (a distant food source) are communicated mainly by body language; for example, honeybees dance and wolves, before a hunt, howl to rally the pack, become tense and have their tails pointing straight. In Animal Communication were observed: conceptual generalizations (dolphins can transmit identity information independent of the caller's voice and location), syntax (putty-nosed monkeys build alarm calls as "word sequences") and metacommunication (the "play face" in dogs signals that subsequent aggressive signal is a play).

The **communication distance**, in animal vocal communication, is the maximal distance at which the receiver can still get the signal; animals can vary signal amplitude and visual display with *receiver distance* in order to ensure signal transmission. The frequency and sound power of a maximal vocalization by an air-breathing animal with body mass M is, usually, proportional to $M^{0.4}$ and $M^{0.6}$, respectively.

Another example of distance-dependent communication is the protective coloration of some aposematic animals: it switches from conspicuousness (signaling non-edibility) to crypsis (camouflage) with increasing distance from a predator.

• Animal long-distance communication

The main modes of animal communication are infrasound ($<20\,\mathrm{Hz}$), sound, ultrasound ($>20\,\mathrm{kHz}$), vision (light), chemical (odor), tactile and electrical. Infrasound, low-pitched sound (as territorial calls) and light in air can be long-distance.

A blue whale infrasound can travel thousands of kilometers through the ocean water using SOFAR channel (a layer where the speed of sound is at a minimum, because water pressure, temperature, and salinity cause a minimum of water density; cf. **distances in Oceanography** in Chap. 25). On the other hand, Janik (2000) estimated that unmodulated dolphin whistles at 12 kHz in a habitat having a uniform water depth of 10 m would be detectable by conspecifics at distances of 1.5–4 km.

Most elephant communication is in the form of infrasonic rumbles which may be heard by other elephants at 5 km away and, in optimum atmospheric conditions, at 10 km. The resulting seismic waves can travel 16–32 km through the ground. But non-fundamental harmonics of elephant calls are sonic. McComb, Reby, Baker, Moss and Sayialel (2003) found that, for female African elephants, the peak of social call frequency

is $\approx 115\,\mathrm{Hz}$ and the *social recognition distance* (over which a contact call can be identified as belonging to a family) is usually 1.5 km and at most 2.5 km.

Many animals hear infrasound generated by earthquakes, tsunami and hurricanes before they strike. For example, elephants can hear storms $160-240\,\mathrm{km}$ away.

High-frequency sounds attenuate more rapidly with distance; they are more directional and vulnerable to scattering. But ultrasounds are used by bats (echolocation) and antropods. Rodents use them to communicate selectively to nearby receivers without alerting predators and competitors. Some anurans shift to ultrasound signals in the presence of continuous background noise (as waterfall, human traffic).

• Plant long-distance communication

Long-distance signaling was observed from roots and mature leaves, exposed to an environmental stress, to newly developing leaves of a higher plant. For example, flooding of the soil induces (in a few hours for some dryland species) bending of leaves and slowing of their expansion.

This communication is done cell-to-cell through the plant vascular transpiration system. In this system, macromolecules (except for water, ions and hormones) carry nutrients and signals, via *phloem* and *xylem* conducting tissues, only in one direction: from lower mature regions to shoots. The identity of long-distance signals in plants is still unknown but the existence of information macromolecules is expected.

Insecticide distance effect.

The main means of pest (termites, ants, etc.) control are chemical liquid insecticides and repellents. The efficiency of an insecticide can be measured by its *all dead distance*, i.e., the maximum distance from the standard toxicant source within which no targeted insects are found alive after a fixed period. The **insecticide distance effect** (or *transfer effect*) is that the toxicant is spread through the colony because insects groom and feed each other. The newer *bait systems* concentrate on this effect.

The toxicant (usually, a growth inhibitor) should act slowly in order to maximize distance effect and minimize secondary repellency created by the presence of dying, dead and decaying insects. A bait system should be reapplied until insects come to it by chance, eat the toxic bait and go back to the colony, creating a chemical trail. It acts slowly, but it completely eliminates a colony and is safer to environment.

• Marital distance

The **marital distance** is the distance between birthplaces of spouses (or zygotes).

• Ontogenetic depth

Nelson's **ontogenetic depth** is the distance, in number of cell divisions, from the unicellular state (fertilized egg) to the adult metazoan capable of reproduction (production of viable gametes).

• Telomere length

The telomeres are repetitive DNA sequences $((TTAGGG)_n)$ in vertebrates cells) at both ends of each linear chromosome in the cell nucleus. They are long stretches of noncoding DNA protecting coding DNA. The number n of TTAGGG repeats is called the **telomere length**; it is $\approx 2,000$ in humans. A cell can divide if each of its telomeres has positive length; otherwise, it becomes senescent and dies, or tries to self-replicate and, eventually, creates cancer. The Hayflick limit is the maximal number of divisions beneath which a normal differentiated cell will stop dividing because of shortened telomeres or DNA damage; for humans it is about 52.

Human telomeres are 3–20 kilobases in length, and they lose ≈ 100 base pairs, i.e., 16 repeats, at each mitosis (happening each 20–180 min). The mean leukocyte telomere length, for example, decreases with age by 9% per decade. There is correlation between telomere length and longevity in humans and, for example, between chronic emotional stress in women and telomere shortening.

But telomere length can increase: by transfer of repeats between daughter telomers or by action of enzyme telomerase. In humans, telomerase acts only in germ, stem or proliferating tumor cells. These cells, unicellular eukaryotes and hydra species are biologically immortal, i.e., there is no aging (sustained increase in rate of mortality with age) since the Hayflick limit does not apply.

The telomere shortening is one of the main proposed mechanisms of aging. The other ones are stem cell senescence, oxidative damage, evolutionary accumulation of late-acting harmful genes, and general transition of a biological network from plasticity (childhood) via adaptation (adolescence) to steady rigid state (aging). In Gerontology, aging is (vital functions) redundancy decay and contemporary risk.

• Gerontologic distance

The **gerontologic distance** between individuals of ages x and y from a population with *survival fraction distributions* $S_1(t)$ and $S_2(t)$, respectively, is defined by

$$|\ln \frac{S_2(y)}{S_1(x)}|.$$

A function S(t) can be either an empirical distribution, or a parametric one based on modeling. The main survival functions S(t) are: $\frac{N(t)}{N(0)}$ (where N(t) is the number of survivors, from an initial population N(0), at time t), e^{kt} (exponential model), $e^{\frac{a}{b}(1-e^{bt})}$ (Gompertz model), $e^{-\frac{at^{b+1}}{b+1}}$ (Weibull model); here a and b are, respectively, age independent and age dependent mortality rate coefficients.

Distances are used in Human Gerontology also to model the relationship between geographical distance and contact between adult children and their elderly parents. A surprising phenomenon, the *late-life mortality deceleration* (even plateau) was observed for humans and fruit flies: the probability that the somatic cells of an organism become senescent tends to be independent of its age in the long-time limit. In fact, the existence of such a plateau is typical for many Markov processes.

Also, Fukuda, Taki, Sato, Kinomura, Goteau and Kawashima (2008) found that gray matter volume linearly decreases with age, and the loss is slower in women. The presence of gene FOX03A GG triples the chance of living to 95 years.

• Body size rules

Body size, measured as mass or length, is one of the most important traits of an organism. Payne et al. (2008) claim that the maximum size of Earth's organisms increased in two great leaps (about 1,600 and 600 million years ago: appearance of eukaryotic cells and multicellularity) due leaps in the oxygen level, and each time it jumped up by a factor of about a million. Below are given the main rules of large-scale Ecology involving body size.

Island rule is a principle that, on islands, small mammal species evolve to larger ones while larger ones evolve to smaller. Damuth (1993) suggested that in mammals there is an optimum body size $\approx 1 \,\mathrm{kg}$ for energy acquisition, and so island species should, in the absence of the usual competitors and predators, evolve to that size.

Insular dwarfism is an evolutionary trend of the reduction in size of large mammals when their gene pool is limited to a very small environment (say, islands). One explanation is that food decline activates processes where only the smaller of the animals survive since they need fewer resources, and so are more likely to get past the breakpoint where population decline allows food sources to replenish.

Island gigantism is a biological phenomenon where the size of animals isolated on an island increases dramatically over generations due the removal of constraints. It is a form of natural selection in which increased size provides a survival advantage.

Abyssal gigantism is a tendency of deep-sea species to be larger than their shallow-water counterparts. It can be adaptation for scarcer food resources (delaying sexual maturity results in greater size), greater pressure and lower temperature.

The Galileo's square-cube law states that as an object increases in size its volume V (and mass) increases as the cube of its linear dimensions while surface area SA increases as the square; so the ratio $\frac{SA}{V}$ decreases. For materials, high $\frac{SA}{V}$ speeds up chemical reactions and thermodynamic processes that minimize free energy. This ratio is the main compactness measure for 3D shapes in Biology. Higher $\frac{SA}{V}$ permits smaller cells to gather nutrients and reproduce very rapidly. Also, smaller animals in hot and dry climates better lose heat through the skin and cool the body.

But lower $\frac{SA}{V}$ (and so larger size) improves temperature control in unfavorable environments: a smaller proportion of the body being exposed results in slower heat loss or gain. **Bergmann–Mayr's rule** is a principle

that, within a species, the body size increases with colder climate. For example, Northern Europeans on average are taller than Southern ones. **Allen's rule**: animals from colder climates usually have shorter limbs than the equivalent ones from warmer climates.

Cope's rule is a macro-evolutionary trend (common among mammals): the tendency of body size to increase over geological time. Large size enhances reproductive success, ability to avoid predators and capture prey, and improves thermal efficiency. In large carnivores, bigger species dominate better smaller competitors.

Cope's rule can be an evolutionary manifestation of Bergmann's rule: species and lineages that conform to Bergmann's rule should evolve toward larger sizes during episodes of climatic cooling. Large body size favors the individual but renders the clade more susceptible to extinction via, for example, dietary specialization.

An allometric law is a relation between the size of an organism and the size of any of its parts or attributes; for example, **Rensch's rule** is that, in groups of related species, sexual size dimorphism is more pronounced in larger species.

Examples of allometric power laws are, in terms of body mass M (or, assuming constant density of biomass, body size) of an animal, proportionalities of metabolic rate to $M^{0.75}$ (**Kleber's law**) and of breathing time to $M^{0.25}$.

Many of the allometric 0.25 scaling laws can be explained by the *WBE* model (West, Brown and Enquist 1997) positing that biological rates are limited by the rate at which energy, materials, and waste can be distributed or removed, and that the process requires an hierarchic space filling network which minimizes needed time and energy.

A cellular organism (for example, bacteria) of linear size (say, diameter) S has, roughly, internal metabolic activity proportional to cell volume (so, to S^3) and flux of nutrient and energy dissipation proportional to cell envelope area (so, to S^2). Therefore, this size is within the possible (nanometers) range of the ratio flux/metabolic activity. For viral particles, there is no metabolism, and their size is, roughly, proportional to the third root of the genome size.

A time period t° correlates roughly with the size of its cold-blooded organisms. Also, a rapid average decline of $\approx 20\%$ in size-reated traits was observed in human-harvested species.

• Distance running model

The distance running model is a model of antropogenesis proposed in [BrLi04]. Bipedality is a key derived behavior of hominids which appeared 4.5–6 million years ago. However, australopithecines were still animals.

The genus Homo which emerged about 2 million years ago could already produce rudimentary tools. The Bramble–Lieberman model attributes this transition to a suite of adaptations specific to running long distances in the savanna (in order to compete with other scavengers in reaching carcasses).

They specify how endurance running, a derived capability of Homo, defined the human body form, producing balanced head, low/wide shoulders, narrow chest, short forearms and heels, large hip, etc.

• Distance coercion model

The distance coercion model is a model, proposed in [OkBi08], of the origin of uniquely human extensive/intensive kinship-independent conspecific social cooperation in spite of conflicts of interest. All the unique properties of humans (complex symbolic speech, cognitive virtuosity, manipulation-proof transmission of fitness-relevant information, etc.) can be seen as elements and effects of this cooperation.

The model argues that such non-kin cooperation can arise only as a result of the instantaneous pursuit of individual self-interest by animals who possess a capacity for synchronous (remote) projection of coercive threat.

The individually adaptive advantages of cooperation come as a byproduct of an ongoing individually self-interested coercive threat conjointly with other group members (preemptive or compensated coercion). So, each individual will display public behaviors that can be construed as beneficial to other coalition members.

Humans are the only animal with an innate biological capacity to project coercive threat remotely: to kill adult conspecifics with thrown projectiles from a distance of many body diameters (at least $10\,\mathrm{m}$). The model posits that the human throwing capacity briefly preceded the emergence of brain expansion and so, of needed (in late pregnancy and child-rearing) social support.

Historical increases in the scale of human social cooperation could be associated with prior acquisition of a new coercive technology; for instance, the bow and agricultural civilizations, gunpowder weaponry and the modern state.

• Distance model of altruism

In Evolutionary Ecology, altruism is explained by kin selection and group selection, and it is supposed to be a driving force of the transition from unicellular organisms to multicellularity. The **distance model of altruism** (see [Koel00]) suggests that altruists spread locally, i.e., with small interaction distance and offspring dispersal distance, while the evolutionary response of egoists is to invest in increasing of those distances. The intermediate behaviors are not maintained, and evolution will lead to a stable bimodal spatial pattern.

Chapter 24 Distances in Physics and Chemistry

24.1 Distances in Physics

Physics studies the behavior and properties of matter in a wide variety of contexts, ranging from the sub-microscopic particles from which all ordinary matter is made (*Particle Physics*) to the behavior of the material Universe as a whole (*Cosmology*).

Physical forces which act at a distance (i.e., a push or pull which acts without "physical contact") are nuclear and molecular attraction and, beyond the atomic level, gravity (completed, perhaps, by anti-gravity), static electricity, and magnetism. Last two forces can be both push and pull.

Distances on a small scale are treated in this chapter, while large distances (in Astronomy and Cosmology) are the subject of Chaps. 25 and 26.

In fact, the distances having physical meaning range from 1.6×10^{-35} m (*Planck length*) to 4.3×10^{26} m (the estimated size of the observable Universe). The world appears Euclidean at distances less than about 10^{25} m (if gravitational fields are not too strong).

At present, the Theory of Relativity, Quantum Theory and Newtonian laws permit us to describe and predict the behavior of physical systems in the range $10^{-15} - 10^{25}$ m.

Gigantic accelerators are able to register particles measuring 10^{-18} m. Relativity and Quantum Theory effects, governing Physics on very large and small scales, are already accounted for in technology, for instance in GPS satellites and nanocrystals of solar cells.

Moment

In Physics and Engineering, **moment** is the product of a quantity and a distance (or some power of the distance) to some point associated with that quantity.

• Displacement

A displacement is a special kind of quasi-metric (directed Euclidean distance) defined in Mechanics; it is the distance along a straight line from x_1 to x_2 , where x_1 and x_2 are positions occupied by the same moving

particle at two instants t_1 and t_2 , $t_2 \ge t_1$, of time. So, a **displacement** is a vector $\overrightarrow{x_1x_2}$ of length $||x_1 - x_2||_2$ specifying the position x_2 of a particle in reference to its previous position x_1 .

• Mechanic distance

The **mechanic distance** is the position of a particle as a function of time t. For a particle with initial position x_0 and initial speed v_0 , which is acted upon by a constant acceleration a, it is given by

$$x(t) = x_0 + v_0 t + \frac{1}{2}at^2.$$

The distance fallen under uniform acceleration a, in order to reach a speed v, is given by $x=\frac{v^2}{2a}$.

A free falling body is a body which is falling subject only to acceleration

A free falling body is a body which is falling subject only to acceleration by gravity g. The distance fallen by it, after a time t, is $\frac{1}{2}gt^2$; it is called the **free fall distance**.

• Terminal distance

The **terminal distance** is the distance of an object, moving in a resistive medium, from an initial position to a stop.

Given an object of mass m moving in a resistive medium (where the drag per unit mass is proportional to speed with constant of proportionality β , and there is no other force acting on a body), the position x(t) of a body with initial position x_0 and initial velocity v_0 is given by

$$x(t) = x_0 + \frac{v_0}{\beta}(1 - e^{-\beta t}).$$

The speed of the body $v(t) = x'(t) = v_0 e^{-\beta t}$ decreases to zero over time, and the body reaches a **maximum terminal distance**

$$x_{terminal} = \lim_{t \to \infty} x(t) = x_0 + \frac{v_0}{\beta}.$$

For a projectile, moving from initial position (x_0, y_0) and velocity (v_{x_0}, v_{y_0}) , the position (x(t), y(t)) is given by $x(t) = x_0 + \frac{v_{x_0}}{\beta}(1 - e^{-\beta t})$, $y(t) = (y_0 + \frac{v_{y_0}}{\beta} - \frac{g}{\beta^2}) + \frac{v_{y_0}\beta - g}{\beta^2}e^{-\beta t}$. The horizontal motion ceases at a maximum terminal distance

$$x_{terminal} = x_0 + \frac{v_{x_0}}{\beta}.$$

• Acceleration distance

The acceleration distance is the minimum distance at which an object (or, say, flow, flame), accelerating in given conditions, reaches a given speed.

• Ballistics distances

Ballistics is the study of the motion of projectiles, i.e., bodies which are propelled (or thrown) with some initial velocity, and then allowed to be acted upon by the forces of gravity and possible drag.

The horizontal distance traveled by a projectile is called the **range**, the maximum upward distance reached by it is the **height**, and the path of the object is the **trajectory**.

The range of a projectile launched with a velocity v_0 at an angle θ to the horizontal is

$$x(t) = v_0 t \cos \theta,$$

where t is the time of motion. On a level plane, where the projectile lands at the same altitude as it was launched, the full range is

$$x_{max} = \frac{v_0^2 \sin 2\theta}{g},$$

which is maximized when $\theta = \pi/4$. If the altitude of the landing point is Δh higher that of the launch point, then

$$x_{max} = \frac{v_0^2 \sin 2\theta}{2g} \left(1 + \left(1 - \frac{2\Delta hg}{v_0^2 \sin^2 \theta} \right)^{1/2} \right).$$

The height is given by $\frac{v_0 \sin^2 \theta}{2g}$, and is maximized when $\theta = \pi/2$.

The arc length of the trajectory is given by $\frac{v_0^2}{g}(\sin\theta + \cos^2\theta g d^{-1}(\theta))$, where $gd(x) = \int_0^x \frac{dt}{\cosh t}$ is the Gudermannian function. The arc length is maximized when $gd^{-1}(\theta)\sin\theta = (\int_0^\theta \frac{dt}{\cos t})\sin\theta = 1$, and the approximate solution is $\theta \approx 0.9855$.

• Interaction distance

The **impact parameter** is the perpendicular distance between the velocity vector of a projectile and the center of the object it is approaching.

The **interaction distance** between two particles is the farthest distance of their approach at which it is discernable that they will not pass at the impact parameter, i.e., their distance of closest approach if they had continued to move in their original direction at their original speed.

• Mean free path (length)

The mean free path (length) of a particle (photon, atom or molecule) in a medium measures its probability to undergo a situation of a given kind K; it is the average of an exponential distribution of distances until the situation K occurs. In particular, this average distance d is called:

nuclear collusion length if K is a nuclear reaction;

interaction length if K is an interaction which is neither elastic, nor quasi-elastic;

scattering length if K is a scattering event;

attenuation length (or absorption length) if K means that the probability P(d), that a particle has not been absorbed, drops to $\frac{1}{e}$ (cf. Beer-Lambert law);

radiation length (or cascade unit) if K means that the energy of (high energy electromagnetic-interacting) charged particles drops by the factor $\frac{1}{e} \approx 0.368$.

In Gamma-ray Radiography, the mean free path of a beam of photons is the average distance a photon travels between collisions with atoms of the target material. It is $\frac{1}{\alpha\rho}$, where α is the material opacity and ρ is its density.

• Neutron scattering length

In Physics, *scattering* is the random deviation or reflection of a beam of radiation or a stream of particles by the particles in the medium.

In Neutron Interferometry, the **scattering length** a is the zero-energy limit of the scattering amplitude $f = -\frac{\sin \delta}{k}$. Since the *total scattering cross section* (the likelihood of particle interactions) is $4\pi |f|^2$, it can be seen as the radius of a hard sphere from which a point neutron is scattered.

The spin-independent part of the scattering length is the *coherent* scattering length.

In order to expand the scattering formalism to absorption, the scattering length is made complex a = a' - ia''.

Thomson scattering length is the classical electron radius

$$\approx 2.81794 \times 10^{-15} \, m$$

• Inelastic mean free path

In Electron Microscopy, the **inelastic mean free path** (or IMFP) is the average total distance that an electron traverses between events of inelastic scattering, while the **effective attenuation length** (or EAL) is an experimental parameter reflecting the average net distance traveled.

The EAL is the thickness in the material through which electron can pass with probability $\frac{1}{e}$ that it survives without inelastic scattering. It is about 20% less than the IMFP due to the elastic scatterings which deflect the electron trajectories.

Both are smaller than the total electron range which may be 10–100 times greater.

• Range of a charged particle

The range of a charged particle, passing through a medium and ionizing, is the distance to the point where its energy drops to almost zero.

• Gyroradius

The **gyroradius** (or *cyclotron radius*, *Larmor radius*) is the radius of the circular orbit of a charged particle (for example, an energetic electron that is ejected from Sun) gyrating around its gliding center.

• Debye screening distance

The **Debye screening distance** (or *Debye length*, *Debye–Hückel length*) is the distance over which a local electric field affects the distribution of

mobile charge carriers (for example, electrons) present in the material (plasmas and other conductors).

Its order increases with decreasing concentration of free charge carriers, from 10^{-4} m in gas discharge to 10^{5} m in intergalactic medium.

• Range of fundamental forces

The fundamental forces (or interactions) are gravity and electromagnetic, weak nuclear and strong nuclear forces. The **range** of a force is considered short if it decays (approaches 0) exponentially as the distance d increases.

Both electromagnetic force and gravity are forces of infinite range which obey **inverse-square distance laws**. The shorter the range, the higher the energy. Both weak and strong forces are very short range (about 10^{-18} and 10^{-15} m, respectively) which is limited by the uncertainty principle.

At subatomic distances, Quantum Field Theory describes electromagnetic, weak and strong interactions with the same formalism but different constants; they almost coincide at very large energy.

• Inverse-square distance laws

Any law stating that some physical quantity is inversely proportional to the square of the distance from the source that quantity.

Law of universal gravitation (Newton–Bullialdus): the gravitational attraction between two point-like objects with masses m_1 , m_2 at distance d is given by

$$G\frac{m_1m_2}{d^2}$$
,

where G is the Newton universal gravitational constant.

The existence of extra dimensions, postulated by M-theory, will be experimentally checked by LHC (Large Hadron Collider opened 10 September 2008 at CERN, near Geneva) based on the inverse proportionality of the gravitational attraction in n-dimensional space to the (n-1)-th degree of the distance between objects; if the Universe has a fourth dimension, LHC will find out the inverse proportionality to the cube of the small inter-particle distance.

Coulomb law: the force of attraction or repulsion between two point-like objects with charges e_1 , e_2 at distance d is given by

$$\kappa \frac{e_1 e_2}{d^2}$$
,

where κ is the *Coulomb constant* depending upon the medium that the charged objects are immersed in. The gravitational and electrostatic forces of two bodies with *Planck mass* $m_P \approx 2.176 \times 10^{-8}$ kg and unity electrical charge have equal strength.

The *intensity* (power per unit area in the direction of propagation) of a spherical wavefront (light, sound, etc.) radiating from a point source decreases (assuming that there are no losses caused by absorption or scattering) inversely proportional to the square d^2 of the distance from the source (cf. **distance decay** in Chap. 29). However, for a radio wave, it decrease like $\frac{1}{d}$.

• EM radiation wavelength range

The wavelength is the distance the wave travels to complete one cycle.

Electromagnetic (EM) radiation wavelength range is: <0.01 nm for gamma rays, 0.01–10 nm for X-rays, 100–400 nm for ultraviolet, 400–780 nm for visible light, 0.78–1,000 μ m for infrared (in lasers), 1–330 mm for microwave, 0.33–3,000 m for radio frequency radiation, >3 km for low frequency, and ∞ for static field.

Besides gamma rays, X-rays and far ultraviolet, the EM radiation is non-ionizing, i.e., passing through matter, it only excites electrons: moves them to a higher energy state, instead of removing them completely from an atom or molecule.

Rayleigh distance

In non-ionizing energy radiation (such as sound and much of electromagnetic radiation), the **Rayleigh distance** is the minimum of the distance d from the antenna source, from which the field strength decreases, up to a given error, as d^{-1} . This *Rayleigh limit* can be, say, the point where the phase error is $\frac{1}{16}$ of a wavelength λ .

Beyond this point, about from $d = \frac{2D^2}{\lambda}$, where D is the maximum overall dimension of the antenna, the **far field** starts: the energy radiates only in the radial direction, its angular distribution does not change with distance, the wave front is considered planar and the rays approximately parallel.

The Maxwell equations, governing the field strength decay, can be approximated as d^{-3} , d^{-2} and d^{-1} for three regions: the reactive near field, the radiating near field and the far field. Approximate outer edges of reactive and radiating near fields are given by $\frac{\lambda}{2\pi}$ and, say, $0.62(\frac{D^3}{\lambda})^{\frac{1}{2}}$, where large with respect to λ . Cf. the **acoustic distances** in Chap. 21.

In Laser Science, beam divergence is defined by its radius, i.e., (for a Gaussian beam) the distance from the beam propagation axis where intensity drops to $\frac{1}{e^2} \approx 13.5\%$ of the maximal value. The waist (or focus) of the beam is the position on its axis where the beam radius is at its minimum and the phase profile is flat.

The **Rayleigh length** (or *Rayleigh range*) of the beam is the distance along its propagation direction from the waist to the place where the beam radius increases by a factor $\sqrt{2}$, i.e., the beam can propagate without significantly diverging.

The Rayleigh length divides the *near-field* and *mid-field*; it is the distance from the waist at which the wavefront curvature is at a maximum. The divergence really starts in the *far field* where the beam radius is at least 10 times its Rayleigh length.

The Rayleigh length is the natural defocusing distance of laser beams. The *confocal parameter* (or *depth of focus*) of the beam is twice its Rayleigh length. Cf. the **lens distances** in Chap. 28.

Half-value layer

Ionizing radiation consists of highly-energetic particles or waves (especially, X-rays, gamma rays and far ultraviolet light), which are

progressively absorbed during propagation through the surrounding medium, via *ionization*, i.e., removing an electron from some of its atoms or molecules. The **half-value layer** is the depth within a material where half of the incident radiation is absorbed.

A basic rule of protection against ionizing radiation exposure: doubling of distance from its source decreases this exposure to a quarter.

In Maxwell Render light simulation software, the attenuation distance (or transparency) is the thickness of object that absorbs 50% of light energy.

• Radiation attenuation with distance

Radiation is the process by which energy is emitted from a source and propagated through the surrounding medium. Radiant energy described in wave terms includes sound and electromagnetic radiation, as light, X-rays and gamma rays.

The incident radiation partially changes its direction, gets absorbed, and the remainder transmitted. The change of direction is *reflection*, diffraction, or scattering if the direction of the outgoing radiation is reversed, split into separate rays, or randomised (diffused), respectively. Scattering occurs in non-homogeneous media.

In Physics, *attenuation* is any process in which the flux density, power amplitude or intensity of a wave, beam or signal decreases with increasing distance from the energy source, as a result of absorption of energy and scattering out of the beam by the transmitting medium. It comes in addition to the divergence of flux caused by distance alone as described by the **inverse-square distance laws**.

Attenuation of light is caused primarily by scattering and absorption of photons. The primary causes of attenuation in matter are the *photoelectric* effect (emission of electrons), Compton scattering (wavelength increase of an interacting X-ray or gamma ray photon) and pair production (creation of an elementary particle and its antiparticle from a high-energy photon).

In Physics, absorption is a process in which atoms, molecules, or ions enter some bulk phase – gas, liquid or solid material; in adsorption, the molecules are taken up by the surface, not by the volume. Absorption of EM radiation is the process by which the energy of a photon is taken up (and destroyed) by, for example, an atom whose valence electrons make the transition between two electronic energy levels. The absorbed energy may be re-emitted or transformed into heat.

Attenuation is measured in units of decibels (dB) or nepers ($\approx 8.7 \, \mathrm{dB}$) per length unit of the medium and is represented by the medium attenuation coefficient α . When possible, specific absorption or scattering coefficient is used instead.

Attenuation of signal (or loss) is the reduction of its strength during transmission. In Signal Propagation, attenuation of a propagating EM wave is called the path loss (or path attenuation). Path loss may be due to free-space loss, refraction, diffraction, reflection,

absorption, aperture-medium coupling loss, etc. Path loss in decibels is $L = 10n \log_{10} d + C$, where n is the path loss exponent, d is the transmitter-receiver distance in m, and C is a constant accounting for system losses.

The free-space path loss (FSPL) is the loss in signal strength of an electromagnetic wave that would result from a line-of-sight path through free space, with no obstacles to cause reflection or diffraction. FSPL is $(\frac{4\pi d}{\lambda})^2$, where d is the distance from the transmitter and λ is the signal wavelength (both in meters), i.e., in decibels it is $10\log_{10}(FSPL) = 20\log_{10}d + 20\log_{10}f - 147.56$, where f is the frequency in hertz.

• Beer-Lambert law

The **Beer–Lambert law** is an empirical relationship for the *absorbance* Ab of a substance when a radiation beam of given frequency goes through it:

$$Ab = \alpha d = -log_a T$$

where a=e or (for liquids) 10, d is the path length (distance the beam travels through the medium), $T=\frac{I_d}{I_0}$ is the transmittance (I_d and I_0 are the intensity of the transmitted and incident radiation), and α is the medium opacity (or linear attenuation coefficient, absorption coefficient); α is the fraction of radiation lost to absorption and/or scattering per unit length of the medium.

The extinction coefficient is $\frac{\lambda_w}{4\pi}\alpha$, where λ_w is the same frequency wavelength in a vacuum. In Chemistry, α is given as ϵC , where C is the absorber concentration, and ϵ is the molar extinction coefficient.

The **optical depth** is $\tau = -\ln \frac{I_d}{I_0}$, measured along the true (slant) optical path.

The **penetration depth** (or **attenuation length**, mean free path, optical extinction length) is the thickness d in the medium where the intensity I_d has decreased to $\frac{1}{e}$ of I_0 ; so, it is $\frac{1}{\alpha}$. Cf. half-value layer.

Also, in Helioseismology, the (meridional flow) penetration depth is the distance from the base of the solar convection zone to the location of the first reversal of the meridional velocity. In an information network, the message penetration distance is the maximum distance from the event message traverses in the valid routing region.

The **skin depth** is the thickness d where the amplitude A_d of a propagating wave (say, alternating current in a conductor) has decreased to $\frac{1}{e}$ of its initial value A_0 ; it is twice the penetration depth. The *propagation* constant is $\gamma = -\ln \frac{A_d}{A_0}$.

The Beer–Lambert law is also applied to describe the attenuation of solar or stellar radiation. The main components of the atmospheric light attenuation are: absorption and scattering by aerosols, Rayleigh scattering (from molecular oxygen O_2 and nitrogen Ni_2) and (only absorption) by carbon dioxide CO_2 , O_2 , nitrogen dioxide NiO_2 , water vapor, ozone O_3 . Cf. atmospheric visibility distances in Chap. 25.

The sea is nearly opaque to light: less than 1% penetrates 100 m deep. In Oceanography, attenuation of light is the decrease in its intensity with depth due to absorption (by water molecules) and scattering (by suspended fine particles).

In Astronomy, attenuation of EM radiation is called *extinction* (or reddening). It arises from the absorption and scattering by the interstellar medium, the Earth's atmosphere and dust around an observed object. The *photosphere* of a star is the surface where its optical depth is $\frac{2}{3}$. The optical depth of a planetary ring is the proportion of light blocked by the ring when it lies between the source and the observer.

• Sound attenuation with distance

Vibrations propagate through elastic solids and liquids, including the Earth, and consist of two types of elastic (or seismic, body) waves and two types of surface waves. Elastic waves are: primary (P) wave moving in the propagation direction of the wave and secondary (S) wave moving in this direction and perpendicular to it. Also, because the surface acts as an interface between solid and gas, surface waves occur: the Love wave moving perpendicular to the direction of the wave and the Rayleigh (R) wave moving in the direction of the wave and circularly within the vertical surface perpendicular to it. The geometric attenuation of P- and S-waves is proportional to $\frac{1}{d^2}$, when propagated by the surface of an infinite elastic body, and it is proportional to $\frac{1}{\sqrt{d}}$.

Sound propagates through gas (say, air) as a P-wave. It attenuates geometrically over a distance, normally at a rate of $\frac{1}{d^2}$: the inverse-square distance law relating the growing radius d of a wave to its decreasing intensity. The **far field** (cf. **Rayleigh distance**) is the part of a sound field in which sound pressure decreases as $\frac{1}{d}$ (but sound intensity decreases as $\frac{1}{d^2}$).

In natural media, further weakening occurs from attenuation, i.e., scattering (reflection of the sound in other directions) and absorption (conversion of the sound energy to heat). Cf. critical distance among acoustics distances in Chap. 21.

The **sound extinction distance** is the distance over which its intensity falls to $\frac{1}{e}$ of its original value. For sonic boom intensities (say, supersonic flights), the lateral *extinction distance* is the distance where in 99% of cases the sound intensity is lower than 0.1–0.2 mbar (10–20 Pa) of atmospheric pressure. The earthquake **extinction length** is the distance (in kilometers) over which the primary S-wave energy is decreased by $\frac{1}{e}$; cf. site-source **distances in Seismology** in Chap. 25.

Water is transparent to sound. Sound energy is absorbed (due to viscosity) and $\approx 6\%$ of it scattered (due to water inhomogeneities). Sound attenuation by zooplankton is used in hydroacoustic measurement of fish and zooplankton abundance.

Absorbed less in liquids and solids, low frequency sounds can propagate in these media over much greater distances along lines of minimum sound speed. Cf. **SOFAR** channel among distances in Oceanography in Chap. 25.

On the other hand, high frequency waves attenuate more rapidly. So, low frequency waves are dominant further from the source (say, a musical band or earthquake).

Attenuation of ultrasound waves with frequency f MHz at a given distance r cm is $\alpha f r$ decibels, where α dB MHz⁻¹ cm⁻¹ is the attenuation coefficient of the medium. It is used in Ultrasound Biomicroscopy; in a homogeneous medium (so, without scattering) α is 0.0022, 0.18, 0.85, 20, 41 for water, blood, brain, bone, lung, respectively.

• Optical distance

The **optical distance** (or *optical path length*) is a distance dn traveled by light, where d is the physical distance in a medium and $n = \frac{c}{v}$ is the refractive index of the medium (c and v are the speeds of an EM wave in a vacuum and in the medium). By Fermat's principle light follows the shortest optical path. Cf. **optical depth**.

The **light extinction distance** is the distance where light propagating through a given medium reaches its *steady-state speed*, i.e., a characteristic speed that it can maintain indefinitely. It is proportional to $\frac{1}{\rho\lambda}$, where ρ is the density of the medium and λ is the wavelength, and it is very small for most common media.

• Proximity effects

In Electronic Engineering, an alternating current flowing through an electric conductor induces (via the associated magnetic field) eddy currents within the conductor. The *electromagnetic proximity effect* is the "current crowding" which occurs when such currents are flowing through several nearby conductors such as within a wire. It increases the alternating current resistance (so, electrical losses) and generates undesirable heating.

In Nanotechnology, the quantum $\frac{1}{f}$ proximity effect is that the $\frac{1}{f}$ fundamental noise in a semiconductor sample is increased by the presence of another similar current-carrying sample placed in the close vicinity.

The superconducting proximity effect is the propagation of superconductivity through a NS (normal-superconductor) interface, i.e., a very thin layer of "normal" metal behaves like a superconductor (that is, with no resistance) when placed between two thicker superconductor slices.

In E-beam Lithography, if a material is exposed to an electronic beam, some molecular chains break and many electron scattering events occur. Any pattern written by the beam on the material can be distorted by this *E-beam proximity effect*.

In LECD (localized electrochemical deposition) technique for fabrication of miniature devices, the microelectrode (anode) is placed close to the tip of a fabricated microstructure (cathode). Voltage is applied and the structure is grown by deposition. The *LECD proximity effect*: at small cathode-anode distances, migration overcomes diffusion, the deposition rate increases greatly and the products are porous.

In Atomic Physics, the **proximity effect** refers to the intramolecular interaction between two (or more) functional groups (in terms of group contributions models of a molecule) that affects their properties and those of the groups located nearby.

Cf. also proximity effect (audio) among acoustics distances in Chap. 21.

The term proximity effect is also used more abstractly, to describe some undesirable proximity phenomena. For example, the proximity effect in the production of chromosome aberrations (when ionizing radiation breaks double-stranded DNA) is that DNA strands can misrejoin if separated by less than $\frac{1}{3}$ of the diameter of a cell nucleus. The proximity effect in innovation process is the tendency to the geographic agglomeration of innovation activity.

• Hopping distance

Hopping is atomic-scale long range dynamics that controls diffusivity and conductivity. For example, oxidation of DNA (loss of an electron) generates a radical cation which can migrate a long (more that 20 nm) distance, called the **hopping distance**, from site to site (to "hop" from one aggregate to another) before it is trapped by reaction with water.

Atomic jump distance

In the solid state the atoms are about closely packed on a rigid lattice. The atoms of some elements (carbon, hydrogen, nitrogen), being too small to replace the atoms of metallic elements on the lattice, are located in the interstices between metal atoms and they diffuse by squeezing between the host atoms.

Interstitial diffusion is the only mechanism by which atoms can be transported through a solid substance while, in a gas or liquid, mass transport is possible by both diffusion and the flow of fluid (for example, convection currents).

The **jump distance** is the distance an atom is moved through the lattice in a given direction by one exchange of its position with an adjacent vacant or occupied lattice site.

The mean square diffusion distance d_t from the starting point which a molecule will have diffused in time t, satisfies $d_t^2 = r^2N = r^2\nu t = 2nDt$, where r is the jump distance; N is the number of jumps (equal to νt assuming a fixed jump rate ν); n=1,2,3 for one-, two- and three-dimensional diffusion; and $D=\frac{\nu r^2}{2n}$ is the diffusivity in square centimeters per second. For example, $D=1-1.5\times 10^{-5}, 10^{-6}$ and 10^{-10} for small molecules in water, small protein in water and proteins in a membrane, respectively.

In diffusion alloy bonding, a **characteristic diffusion distance** is the distance between the joint interface and the site wherein the concentration of the diffusing substance (say, aluminum in high carbon-steel) falls to zero up to a given error.

• Diffusion length

Diffusion is a process of spontaneous spreading of matter, heat, momentum, or light: particles move to lower chemical potential implying a change in concentration.

In Microfluidics, the **diffusion length** is the distance from the point of initial mixing to the complete mixing point where the equilibrium composition is reached.

In semiconductors, electron-hole pairs are generated and recombine; the (minority carrier) diffusion length of a material is the average distance a minority carrier can move from the point of generation until it recombines with majority carriers. Also, the diffusion length, in electron transport by diffusion, is the distance over which concentration of free charge carriers injected into semiconductor falls to $\frac{1}{e}$ of its original value.

Cf. jump distance and, in Chap. 23, capillary diffusion distance.

• Thermal diffusion length

The heat propagation into material is represented by the **thermal diffusion length**, i.e., the propagation distance of the thermal wave producing an attenuation of the peak temperature to about 0.1 of the maximum surface value.

For lasers with femtosecond pulse duration, it is so small that the energy of the beam, not being absorbed by laser-induced plasma, is fully deposited into the target.

The propagation of the laser-generated shock wave is approximated as *blast wave* (instantaneous, massless point explosion). The **expansion distance** is the distance between the surface of the target and the position of a blast wave; it depends on the energy converted into the plasma state.

• Hydrodynamic radius

The **hydrodynamic radius** (or *Stokes radius*, *Stokes–Einstein radius*) of a molecule, undergoing diffusion in a *solution* (homogeneous mixture composed of two or more substances), is the hypothetical radius of a hard sphere which diffuses with the same rate as the molecule.

• Solvent migration distance

In Chromatography, the **solvent migration distance** is the distance traveled by the front line of the liquid or gas entering a chromatographic bed for *elution* (the process of using a solvent to extract an absorbed substance from a solid medium).

• Healing length

For a superfluid, the **healing length** is a length over which the wave function can vary while still minimizing energy.

For Bose–Einstein condensates, the healing length is the width of the bounding region over which the probability density of the condensate drops to zero.

• Coupling length

In optical fibre devices mode coupling occurs during transmission by multimode fibres (mainly because of random bending of the fibre axis).

Between two modes, a and b, the **coupling length** l_c is the length for which the complete power transfer cycle (from a to b and back) take place, and the **beating length** z is the length along which the modes accumulate a 2π phase difference. The resonant coupling effect is *adiabatic* (no heat is transferred) if and only if $l_c > z$.

Furuya, Suematsu and Tokiwa (1978) define the *coupling length* of modes a and b as the length of transmission at which the ratio $\frac{I_a}{I_b}$ of mode intensities reach e^2 .

• Localization length

Generally, the **localization length** is the average distance between two obstacles in a given scale. The localization scaling theory of metal-insulator transitions predicts that, in zero magnetic field, electronic wave functions are always localized in disordered 2D systems over a length scale called the **localization length**.

• Long range order

A physical system has **long range order** if remote portions of the same sample exhibit correlated behavior. For example, in crystals and some liquids, the positions of an atom and its neighbors define the positions of all other atoms.

Examples of long range ordered states are: superfluidity and, in solids, magnetism, charge density wave, superconductivity. Most strongly correlated systems develop long-range order in their ground state.

Short range refers to the first- or second-nearest neighbors of an atom. More precisely, the system has **long range order**, *quasi-long range order* or is *disordered* if the corresponding correlation function decays at large distances to a constant, to zero polynomially, or to zero exponentially (cf. **long range dependency** in Chap. 29).

• Correlation length

The **correlation length** is the distance from a point beyond which there is no further correlation of a physical property associated with that point. It is used mainly in statistical mechanics as a measure of the order in a system for phase transitions (fluid, ferromagnetic, nematic).

For example, in a spin system at high temperature, the correlation length is $-\frac{\ln d \cdot C(d)}{d}$ where d is the distance between spins and C(d) is the correlation function.

In particular, the percolation correlation length is an average distance between two sites belonging to the same cluster, while the thermal correlation length is an average diameter of spin clusters in thermal equilibrium at a given temperature. In second-order phase transitions, the correlation length diverges at the critical point.

• Magnetic length

The **magnetic length** (or *effective magnetic length*) is the distance between the effective magnetic poles of a magnet.

The magnetic correlation length is a magnetic-field dependent correlation length.

• Spatial coherence length

The spatial coherence length is the propagation distance from a coherent source to the farthest point where an electromagnetic wave still maintains a specific degree of coherence. This notion is used in Telecommunication Engineering (usually, for the optical regime) and in synchrotron X-ray Optics (the advanced characteristics of synchrotron sources provide highly coherent X-rays).

The spatial coherence length is about $20\,\mathrm{cm}$, $100\,\mathrm{m}$, and $100\,\mathrm{km}$ for helium–neon, semiconductor and fiber lasers, respectively. Cf. temporal coherence length which describes the correlation between signals observed at different moments of time.

For vortex-loop phase transitions (superconductors, superfluid, etc.), **coherence length** is the diameter of the largest loop which is thermally excited. Besides coherence length, the second **characteristic length** (cf. Chap. 29) in a superconductor is its **penetration depth**. If the ratio of these values (the *Ginzburg–Landau parameter*) is $<\sqrt{2}$, then the phase transition to superconductivity is of second-order.

• Decoherence length

In disordered media, the **decoherence length** is the propagation distance of a wave from a coherent source to the point beyond which the phase is irreversibly destroyed (for example, by a coupling with noisy environment).

• Dephasing length

Intense laser pulses traveling through plasma can generate, for example, a wake (the region of turbulence around a solid body moving relative to a liquid, caused by its flow around the body) or X-rays. The **dephasing length** is the distance after which the electrons outrun the wake, or (for a given mismatch in speed of pulses and X-rays) laser and X-rays slip out of phase.

• Metric theory of gravity

A metric theory of gravity assumes the existence of a symmetric metric (seen as a property of space—time itself) to which matter and non-gravitational fields respond. Such theories differ by the types of additional gravitational fields, say, by dependency or not on the location and/or velocity of the local systems. General Relativity is one such theory; it contains only one gravitational field, the space—time metric itself, and it is governed by Einstein's partial differential equations. It has been found empirically that, besides Nordstrom's 1913 conformally-flat scalar theory, every other metric theory of gravity introduces auxiliary gravitational fields.

A bimetric theory of gravity is a metric theory of gravity in which two, instead of one, metric tensors are used for, say, effective Riemannian and background Minkowski space—times.

Østvang (2001) proposed a quasi-metric framework for relativistic gravity.

• Gravitational radius

The **gravitational radius** (or *event horizon*) is the radius that a spherical mass must be compressed to in order to transform it into a black hole. The *Schwarzschild radius* is the gravitational radius $\frac{2Gm}{c^2}$ of a Schwarzschild black hole with mass m.

A "typical" black hole has mass $\approx 6\,M_{Sun}$, diameter $\approx 18\,\mathrm{km}$, temperature $\approx 10^{-8}\,\mathrm{K}$ and lifetime $\approx 2\times 10^{68}\,\mathrm{years}$. The central black hole of the galaxy M87 (in the center of our Virgo Supercluster) has mass 3 billions suns and diameter at least one light-day. Usually central black hole weighs $\approx 0.1\%$ of the surrounding galactic budge.

On the other hand, a hypothetical quantum mechanical black hole has mass $10^3 M_{proton}$, size 10^{-18} m, temperature 10^{16} K and lifetime 10^{-27} s. The smallest black hole is a hypothetical *Planck particle*, i.e., one whose Schwarzschild radius and Compton wavelength are equal to the Planck length $(10^{-20}$ times the proton's radius). Its mass is the Planck mass $(13 \times 10^{18} M_{proton})$; its lifetime is 0.26 times the Planck time.

• Binding energy

The **binding energy** of a system is the mechanical energy required to separate its parts so that their relative distances become infinite. For example, the binding energy of an electron or proton is the energy needed to remove it from the atom or the nucleus, respectively, to an infinite distance.

In Astrophysics, gravitational binding energy of a celestial body is the energy required to disassemble it into dust and gas, while the lower gravitational potential energy is needed to separate two bodies to infinite distance, keeping each intact.

• Acoustic metric

In Acoustics, the **acoustic metric** (or **sonic metric**) is a characteristic of sound-carrying properties of a given medium: air, water, etc.

In General Relativity and Quantum Gravity, it is a characteristic of signal-carrying properties in a given *analog model* (with respect to Condensed Matter Physics) where, for example, the propagation of scalar fields in curved *space-time* is modeled (see, for example, a survey [BLV05]) as the propagation of sound in a moving fluid, or slow light in a moving fluid dielectric, or *superfluid* (quasi-particles in quantum fluid).

The passage of a signal through an acoustic metric modifies the metric; for example, the motion of sound in air moves air and modifies the local speed of the sound. Such "effective" (i.e., recognized by its "effects") **Lorentzian metric** (cf. Chap. 7) governs, instead of the background metric, the propagation of fluctuations: the particles associated to the perturbations follow geodesics of that metric.

In fact, if a fluid is barotropic and inviscid, and the flow is irrotational, then the propagation of sound is described by an **acoustic metric** which depends on the density ρ of flow, velocity ${\bf v}$ of flow and local speed s of sound in the fluid. It can be given by the *acoustic tensor*

$$g = g(t, \mathbf{x}) = \frac{\rho}{s} \begin{pmatrix} -(s^2 - v^2) \vdots -\mathbf{v}^T \\ \cdots & \cdots \\ -\mathbf{v} & \vdots & 1_3 \end{pmatrix},$$

where 1_3 is the 3×3 identity matrix, and $v = ||\mathbf{v}||$. The acoustic line element can be written as

$$ds^2 = \frac{\rho}{s} \left(-(s^2-v^2)dt^2 - 2\mathbf{v}d\mathbf{x}dt + (d\mathbf{x})^2 \right) = \frac{\rho}{s} \left(-s^2dt^2 + (d\mathbf{x} - \mathbf{v}dt)^2 \right).$$

The signature of this metric is (3,1), i.e., it is a **Lorentzian metric**. If the speed of the fluid becomes supersonic, then the sound waves will be unable to come back, i.e., there exists a *mute hole*, the acoustic analog of a *black hole*.

The **optical metrics** are also used in analog gravity and effective metric techniques; they correspond to the representation of a gravitational field by an equivalent optical medium with magnetic permitivity equal to electric one.

• Aichelburg-Sexl metric

In Quantum Gravity, the **Aichelburg–Sexl metric** (Aichelburg and Sexl 1971) is a four-dimensional metric created by a relativistic particle (having an energy of the order of the Planck mass) of momentum p along the x-axis, described by its $line\ element$

$$ds^{2} = dudv - d\rho^{2} - \rho^{2}d\phi^{2} + 8p\ln\frac{\rho}{\rho_{0}}\delta(u)du^{2},$$

where u = t - x, v = t + x are null coordinates, ρ and ϕ are standard polar coordinates, $\rho = \sqrt{y^2 + z^2}$, and ρ_0 is an arbitrary scale constant.

This metric admits an *n*-dimensional generalization (de Vega and Sánchez 1989), given by the *line element*

$$ds^{2} = dudv - (dX^{i})^{2} + f_{n}(\rho)\delta(u)du^{2},$$

where u and v are the above null coordinates, X^i are the traverse coordinates, $\rho = \sqrt{\sum_{1 \leq i \leq n-2} (X^i)^2}$, $f_n(\rho) = K(\frac{\rho}{\rho_0})^{4-n}$, $k = \frac{8\pi^{2-0.5n}}{n-4}\Gamma(0.5n-1)$ GP, n > 4, $f_4 = 8GP \ln \frac{\rho}{\rho_0}$, G is the gravitational constant and P is the momentum of the considered particle.

This metric describes the gravitational field created, according to General Relativity, during the interaction of spinless neutral particles with rest mass much smaller than the Planck mass m_P and only one of them having an energy of the order m_P .

• Quantum metrics

A quantum metric is a general term used for a metric expected to describe the space-time at quantum scales (of order Planck length $l_P \approx 1.6162 \times 10^{-35}$ m). Extrapolating predictions of Quantum Mechanics and General Relativity, the metric structure of this space-time is determined by vacuum fluctuations of very high energy (10^{19} GeV corresponding to the Planck mass $m_P \approx 2.176 \times 10^{-8}$ kg) creating black holes with radii of order l_P . The space-time becomes "quantum foam:" violent warping and turbulence. It loses the smooth continuous structure (apparent macroscopically) of a Riemannian manifold, to become discrete, fractal, non-differentiable: breakdown at l_P of the functional integral in the classical field equations.

Examples of quantum metric spaces are: Rieffel's **compact quantum metric space**, **Fubini–Study metric** on quantum states, statistical geometry of fuzzy lumps [ReRo01] and quantization of the **metric cone** (cf. Chap. 1) in [IsKuPe90].

• Quantal distances

A quantal distance is a distance between quantum states.

The pure states correspond to the rays in the Hilbert space of wave functions.

The Wootters distance between pure states ω_1 and ω_2 is $\cos^{-1}|\langle \omega_1, \omega_2 \rangle|$. The Fubini–Study distance between pure states ω_1 and ω_2 is $\sqrt{2}(1-|\langle \omega_1, \omega_2 \rangle|^2)^{\frac{1}{2}}$; cf. Fubini–Study metric in Chap. 7.

The mixed quantum states are represented by density operators (i.e., positive operators of unit trace) in the complex projective space over the infinite-dimensional Hilbert space. The m-dimensional version corresponds to the m-qubit quantum states represented by $2^m \times 2^m$ density matrices.

Let X denote the set of all density operators in this Hilbert space. For two given quantum states, represented by density operators $x, y \in X$, we mention the following distances on X.

The **Hilbert–Schmidt norm metric** (cf. Chap. 13) is $||x-y||_{HS}$ with $||A||_{HS} = \sqrt{Tr(A^TA)}$ is the *Hilbert–Schmidt norm* of an operator A.

The **trace norm metric** (cf. Chap. 12) is $||x - y||_{tr}$, where $||A||_{tr} = Tr\sqrt{(A^TA)}$ is the *trace norm* of an operator A. The maximum probability that a quantum measurement will distinguish x from y is $\frac{1}{2}||x - y||_{tr}$.

The Bures–Uhlmann distance is $\sqrt{2(1-Tr((\sqrt{x}y\sqrt{x})^2)}$ (cf. Bures metric in Chap. 7). The Fidelity similarity is $Tr((\sqrt{x}y\sqrt{x})^2)$.

The **Gudder distance** is $\inf_{\lambda \in [0,1]} : (1-\lambda)x + \lambda x' = (1-\lambda)y + \lambda y'; x', y' \in X$. In fact, X is convex, i.e., $\lambda x + (1-\lambda)y \in X$ whenever $x, y \in X$ and $\lambda \in (0,1)$.

Examples of other distances used in the quantal setting are the **Frobenius norm metric** (cf. Chap. 12), **Sobolev metric** (cf. Chap. 13), **Monge–Kantorovich metric** (cf. Chap. 21).

• Action at a distance (in Physics)

An action at a distance is the interaction, without known mediator, of two objects separated in space. Einstein used the term spooky action

at a distance for quantum mechanical interaction (like entanglement and quantum non-locality) which is instantaneous, regardless of distance. His **principle of locality** is: distant objects cannot have direct influence on one another, an object is influenced directly only by its immediate surroundings.

Alice—Bob distance is the distance between two entangled particles, "Alice" and "Bob." In 2004, Zellinger et al. teleported (across a distance 600 m) some quantum information – the polarization property of a photon – to its mate in an entangled pair of photons. In 2006, Ursin et al. transmitted an entangled photon over a distance of 144 km (between the Canary Islands of La Palma and Tenerife) via an optical free-space link. Quantum Theory predicts that the correlations based on quantum entanglement should be maintained over arbitrary Alice—Bob distances. But a strong non-locality, i.e., a measurable action at a distance (a superluminal propagation of real, physical information) never was observed and is generally not expected.

Already controversial (since the speed of light is maximal) long range non-quantum interaction becomes marginal for "mental action at a distance" (telepathy, precognition, psychokinesis, etc.). But if Penrose's intuition that the human brain utilizes quantum mechanical processes is right, then such "psychic non-local" communication looks possible.

The term *short range interaction* is also used for the transmission of action at a distance by a material medium from point to point with a certain velocity dependent on properties of this medium. Also, in Information Storage, the term *near-field interaction* is used for very short distance interaction using scanning probe techniques. *Near-field communication* is a standards-based technology enabling convenient short-range wireless communication between electronic devices.

• Entanglement distance

The entanglement distance is the maximal distance between two entangled electrons in a degenerate electron gas beyond which all entanglement is observed to vanish. Degenerate matter (for example, in White Dwarf stars) is matter having so high density that the dominant contribution to its pressure arises from the Pauli exclusion principle: no two identical fermions may occupy the same quantum state simultaneously.

24.2 Distances in Chemistry and Crystallography

Main chemical substances are ionic (held together by ionic bonds), metallic (giant close packed structures held together by metallic bonds), giant covalent (as diamond and graphite), or molecular (small covalent). Molecules are made of a fixed number of atoms joined together by covalent bonds; they range from small (single-atom molecules in the noble gases) to very

large ones (as in polymers, proteins or DNA). The **interatomic distance** of two atoms is the distance (in angstroms or picometers) between their nuclei.

• Atomic radius

Quantum Mechanics implies that an atom is not a ball having an exactly defined boundary. Hence, **atomic radius** is defined as the distance from the atomic nucleus to the outermost stable electron orbital in a atom that is at equilibrium. Atomic radii represent the sizes of isolated, electrically neutral atoms, unaffected by bonding.

Atomic radii are estimated from **bond distances** if the atoms of the element form bonds; otherwise (like the noble gases), only **Van-der-Waals** radii are used.

The atomic radii of elements increase as one moves down the column (or to the left) in the Periodic Table of Elements.

• Bond distance

The **bond distance** (or *bond length*) is the distance between the nuclei of two bonded atoms. For example, typical bond distances for carbon-carbon bonds in an organic molecule are 1.53, 1.34 and 1.20 Å for single, double and triple bonds, respectively. The atomic nuclei repel each other; the **equilibrium distance** between two atoms in a molecule is the internuclear distance at the minimum of the electronic (or potential) energy surface.

Depending on the type of bonding of the element, its atomic radius is called *covalent* or *metallic*. The *metallic radius* is one half of the **metallic distance**, i.e., the closest internuclear distance in a *metallic crystal* (a closely packed crystal lattice of metallic element).

Covalent radii of atoms (of elements that form covalent bonds) are inferred from bond distances between pairs of covalently-bonded atoms: they are equal to the sum of the covalent radii of two atoms. If the two atoms are of the same kind, then their covalent radius is one half of their bond distance. Covalent radii for elements whose atoms cannot bond to each another is inferred by combining the radii of those that bond with bond distances between pairs of atoms of different kind.

• Van-der-Waals contact distance

Intermolecular distance data are interpreted by viewing atoms as hard spheres. The spheres of two neighboring non-bonded atoms (in touching molecules or atoms) are supposed to just touch. So, their interatomic distance, called the **Van-der-Waals contact distance**, is the sum of radii, called **Van-der-Waals radii**, of their hard spheres. The Van-der-Waals radius of carbon is 1.7 Å, while its covalent radius is 0.76. The Van-der-Waals contact distance corresponds to a "weak bond," when repulsion forces of electronic shells exceed London (attractive electrostatic) forces.

• Interionic distance

An *ion* is an atom that has a positive or negative electrical charge. The **interionic distance** is the distance between the centers of two adjacent (bonded) ions. **Ionic radii** (Goldschmitt and Pauling, independently, in 1920's) are inferred from ionic bond distances in real molecules and crystals.

The ion radii of *cations* (positive ions, for example, sodium Na⁺) are smaller than the atomic radii of the atoms they come from, while *anions* (negative ions, for example, chlorine Cl⁻) are larger than their atoms.

• Intermicellar distance

Micelle is an electrically charged particle built up from polymeric molecules or ions and occurring in certain colloidal electrolytic solutions like soaps and detergents. This term is also used for a submicroscopic aggregation of molecules, such as a droplet in a colloidal system, and for a coherent strand or structure in a fiber.

The intermicellar distance is the average distance between micelles.

• Range of molecular forces

Molecular forces (or interactions) are the following electromagnetic forces: ionic bonds (charges), hydrogen bonds (dipolar), dipole–dipole interactions, London forces (the attraction part of Van-der-Waals forces) and steric repulsion (the repulsion part of Van-der-Waals forces). If the distance (between two molecules or atoms) is d, then (experimental observation) the potential energy function P relates inversely to d^n with n=1,3,3,6,12 for the above five forces, respectively.

The range (or the radius) of an interaction is considered short if P approaches 0 rapidly as d increases. It is also called short if it is at most 3 Å; so, only the range of steric repulsion is short (cf. range of fundamental forces).

An example: for polyelectrolyte solutions, the long range ionic solventwater force competes with the shorter range water-water (hydrogen bonding) force.

In protein molecule, the range of London–Van-der-Waals force is $\approx 5 \, \text{Å}$, and the range of hydrophobic effect is up to $12 \, \text{Å}$, while the length of hydrogen bond is $\approx 3 \, \text{Å}$, and the length of peptide bond (when the carboxyl group of one molecule reacts with the amino group of the other molecule, thereby releasing a molecule of water) is $\approx 1.5 \, \text{Å}$.

• Chemical distance

Various chemical systems (single molecules, their fragments, crystals, polymers, clusters) are well represented by graphs where vertices (say, atoms, molecules acting as monomers, molecular fragments) are linked by, say, chemical bonding, Van-der-Waals interactions, hydrogen bonding, reactions path.

In Organic Chemistry, a molecular graph G(x) = (V(x), E(x)) is a graph representing a molecule x, so that the vertices $v \in V(x)$ are atoms and the

edges $e \in E(x)$ correspond to electron pair bonds. The usual and complementary reciprocal **Wiener numbers** of G(x) are $\frac{1}{2} \sum_{a,b \in V(x)} d(a,b)$ and $\frac{1}{2} \sum_{a,b \in V(x)} (1 + D - d(a,b))^{-1}$, where D is the diameter of G(x).

The (bonds and electrons) *BE-matrix* of a molecule x is the $|V(x)| \times |V(x)|$ matrix $((e_{ij}(x)))$, where $e_{ii}(x)$ is the number of free unshared valence electrons of the atom A_i and, for $i \neq j$, $e_{ij}(x) = e_{ji}(x) = 1$ if there is a bond between atoms A_i and A_j , and A_j and A_j and A_j are A_j and A_j and A_j are A_j and A_j and A_j are A_j and A_j and A_j and A_j are A_j and A_j and A_j and A_j are A_j and A_j are A_j and A_j are A_j and A_j are A_j and A_j and A_j are A_j are A_j and A_j and A_j are A_j and A_j are A_j and A_j are A_j are A_j are A_j are A_j and A_j are A_j and A_j are A_j are A_j are A_j are A_j and A_j are A_j are A_j are A_j and A_j are A_j are A_j are A_j and A_j are A_j

Given two *stoichiometric* (i.e., with the same number of atoms) molecules x and y, their **Dugundji–Ugi chemical distance** is the **Hamming metric**

$$\sum_{1 < i, j < |V|} |e_{ij}(x) - e_{ij}(y)|,$$

and their Pospichal-Kvasnička chemical distance is

$$\min_{\pi} \sum_{1 \le i, j \le |V|} |e_{ij}(x) - e_{\pi(i)\pi(j)}(y)|,$$

where π is any permutation of the atoms.

The above distance is equal to |E(x)|+|E(y)|-2|E(x,y)|, where E(x,y) is the edge-set of the maximum common subgraph (not induced, in general) of the molecular graphs G(x) and G(y). (Cf. **Zelinka distance** in Chap. 15 and **Mahalanobis distance** in Chap. 17.)

The **Pospichal–Kvasnička reaction distance**, assigned to a molecular transformation $x \to y$, is the minimum number of *elementary* transformations needed to transform G(x) onto G(y).

• Molecular RMS radius

The **molecular RMS radius** (or *radius of gyration*) is the root-mean-square distance of atoms in a molecule from their common center of gravity; it is

$$\sqrt{\frac{\sum_{1 \le i \le n} d_{0i}^2}{n+1}} = \sqrt{\frac{\sum_{i} \sum_{j} d_{ij}^2}{(n+1)^2}},$$

where n is the number of atoms, d_{0i} is the Euclidean distance of the i-th atom from the center of gravity of the molecule (in a specified conformation), and d_{ij} is the Euclidean distance between the i-th and j-th atoms.

• Mean molecular radius

The **mean molecular radius** is the number $\frac{r_i}{n}$, where n is the number of atoms in the molecule, and r_i is the Euclidean distance of the i-th atom from the geometric center $\frac{\sum_j x_{ij}}{n}$ of the molecule (here x_{ij} is the i-th Cartesian coordinate of the j-th atom).

• Molecular similarities

Given two three-dimensional molecules x and y characterized by some structural (shape or electronic) property P, their similarities are called

molecular similarities. The main electronic similarities correspond to some correlation similarities from Sect. 17.4 for example, the **Spearman rank correlation** and the two that follow now.

The Carbo similarity (Carbo, Leyda and Arnau 1980) is the cosine similarity (or normalized dot product, cf. Chap. 17) defined by

$$\frac{\langle f(x), f(y) \rangle}{||f(x)||_2 \cdot ||f(y)||_2},$$

where the electron density function f(z) of a molecule z is the volumic integral $\int P(z)dv$ over the whole space.

The *Hodgkin-Richards similarity* (1991) is defined (cf. the **Morisita-Horn similarity** in Chap. 17) by

$$\frac{2\langle f(x), f(y) \rangle}{||f(x)||_2^2 + ||f(y)||_2^2},$$

where f(z) is the electrostatic potential or electrostatic field of a molecule z.

Petitjean (1995) proposed to use the distance $V(x \cup y) - V(x \cap y)$, where the *volume* V(z) of a molecule z is the union of *Van-der-Waals spheres* of its atoms. Cf. **Van-der-Waals contact distance** and, in Chap. 9, **Nikodym metric** $V(x\Delta y)$.

• Persistence length

A *polymer* is a large macromolecule composed of repeating structural units connected by covalent chemical bonds. The **persistence length** of a polymer chain is the length over which correlations in the direction of the tangent are lost.

The molecule behaves as a flexible elastic rod for shorter segments, while for much longer ones it can only be described statistically. Cf. **correlation length**.

Twice the persistence length is the *Kuhn length*, i.e., the length of hypothetical segments which can be thought of as if they are freely jointed with each other in order to form given polymer chain.

• Repeat distance

Given a periodic layered structure, its **repeat distance** is the period, i.e., the **spacing distance** between layers (say, lattice planes, bilayers in a liquid-crystal system, or graphite sheets along the unit cell's hexagonal axis).

A crystal lattice, the unit cell in it and the cell spacing are called also a repeat pattern, the basic repeat unit and the cell repeat distance (or lattice spacing, interplaner distance), respectively.

The repeat distance in a polymer is the ratio of the unit cell length that is parallel to the axis of propagation of the polymer to the number of monomeric units this length covers.

• Metric symmetry

The full crystal symmetry is given by its *space group*.

The **metric symmetry** of the crystal lattice is its symmetry without taking into account the arrangement of the atoms in the unit cell.

In between lies the *Laue group* giving equivalence of different reflexions, i.e., the symmetry of the crystal diffraction pattern. In other words, it is the symmetry in the *reciprocal space* (taking into account the reflex intensities).

The Laue symmetry can be lower than the metric symmetry (for example, an orthorhombic unit cell with a=b is metrically tetragonal) but never higher.

There are seven *crystal systems* (triclinic, monoclinic, orthorhombic, tetragonal, trigonal, hexagonal, and cubic); taken together with possible *lattice centerings*, there are 14 *Bravais lattices*.

• Dislocation distances

In Crystallography, a *dislocation* is a defect extending through a crystal for some distance (**dislocation path length**) along a *dislocation line*. It either forms a complete loop within the crystal or ends at a surface or other dislocation.

The **mean free path** of a dislocation is (Gao, Chen, Kysar, Lee and Gan 2007), in 2D, the average distance between its origin and the nearest particle or, in 3D, the maximum radius of a dislocation loop before it reaches a particle in the slip plane.

The Burgers vector of a dislocation is a crystal vector denoting the direction and magnitude of the atomic displacement that occurs within a crystal when a dislocation moves through the lattice. A dislocation is called edge, screw or mixed if the angle between its line vector and the Burgers vector is 90°, 0° or otherwise, respectively. The edge dislocation width is the distance over which the magnitude of the displacement of the atoms from their perfect crystal position is greater than $\frac{1}{4}$ of the magnitude of the Burgers vector.

The dislocation density ρ is the total length of dislocation lines per unit volume; typically, it is $10\,\mathrm{km\,cm^{-3}}$ but can reach $10^6\,\mathrm{km\,cm^{-3}}$ in a heavily deformed metal. The **average distance** between dislocations depends on their arrangement; it is $\rho^{-\frac{1}{2}}$ for a quadratic array of parallel dislocations. If the average distance decreases, dislocations start to cancel each other's motion.

The **spacing dislocation distance** is the minimum distance between two dislocations which can coexist on separate planes without recombining spontaneously.

• Dynamical diffraction distances

Diffraction is the apparent bending of propagating waves around obstacles of about the wavelength size. Diffraction from a 3D periodic structure such as an atomic crystal is called *Bragg diffraction*. It is a convolution of

the simultaneous scattering of the probe beam (light as X-rays, or matter waves such as electrons or neutrons) by the sample and interference (superposition of waves reflecting from different crystal planes).

The Bragg Law, modeling diffraction as reflexion from crystal planes of atoms, states that waves (with wavelength λ scattered under angle θ from planes at spacing distance d) interfere only if they remain in phase, i.e., $\frac{2d\sin\theta}{\lambda}$ is an integer.

The decay of intensity with depth traversed in the crystal occurs by dynamical extinction, redistributing energy within the wave field, and by photoelectric absorption (a loss of energy from the wave field to the atoms of the crystal).

The *dynamical* (multiple diffraction) theory is used to model the *perfect* (no disruptions in the periodicity) crystals. It considers the incident and diffracted wave fronts as coupled parts of a wave field that interact with each other and the periodically varying electrical susceptibility of the medium so as to satisfy the Maxwell equations.

The former kinematic theory works for imperfect crystals and estimates absorption.

Dynamical theory distinguishes two cases: Laue (or transmission) and Bragg (or reflexion) case, when the reflected wave is directed toward the inside and, respectively, outside of the crystal. The wave field is represented visually by its dispersion surface. The inverse of the diameter of this surface is called (Autier 2001) the **Pendellösung distance** Λ_L in the Laue case and the **extinction distance** Λ_B in the Bragg case.

At the exit face of the crystal, the wave splits into two single waves with different directions: incident 0-beam and diffracted H-beam. With increasing thickness of the crystal, the wave leaving it will first appear mainly in the 0-beam, then entirely in the H-beam at thickness $\frac{\Lambda_L}{2}$, and subsequently it will oscillate between these beams with a period Λ_L , called the **Pendellösung length**; cf. similar **coupling length**.

The wave amplitude (and the intensity of the diffracted beam) is transferred back-and-forth once, i.e., the physical distance acquires a phase change of 2π . Pendellösung oscillations happen also in Bragg case, but with very rapidly decaying amplitudes, and *Pendellösung fringes* are visible only for θ close to 0^0 or 45^0 .

Diffraction that involves multiple scattering events is called *extinction* since it reduces the observed integrated diffracted intensity. Extinction is very significant for perfect crystals and is then called *primary extinction*. In the Bragg case, the **primary extinction length** (James 1964) is the inverse of the *extinction factor* (maximum extinction coefficient for the middle of the range of total reflection):

$$\frac{\pi V \cos \theta}{\lambda r_e |F|C},$$

where F, C (valued 1 or $\cos 2\theta$) are the structure and polarization factors, V is the volume of unit cell, $r_e \approx 2.81794 \times 10^{-15} \mathrm{m}$ is the classical electron radius and λ is the X-ray beam wavelength. The diffracted intensity with sufficiently large thickness no longer increases significantly with increased thickness.

The **extinction length** of an electron or neutron diffraction is $\frac{\pi V \cos \theta}{\lambda |F|}$. Half of it gives the number of atom planes needed to reduce the incident beam to zero intensity.

The **X-ray penetration depth** (or attenuation length, mean free path, extinction distance) is (Wolfstieg 1976) the depth into the material where the intensity of the diffracted beam has decreased e-fold. Cf. **penetration depth**.

In Gullity (1956) X-ray penetration depth is the depth z such that $\frac{I_z}{I_{\infty}} = 1 - \frac{1}{e}$, where I_{∞} , I_z are the total diffraction intensities given from the whole specimen and, respectively, the range between the surface and the depth, z, from it.

• X-ray absorption length

The absorption edge is a sharp discontinuity in the absorption spectrum of X-rays by an element that occurs when the energy of the photon is just above the **binding energy** of an electron in a specific shell of the atom. The **X-ray absorption length** of a crystal is the thickness s of the sample such that the intensity of the X-rays incident upon it at an energy 50 eV above the absorption edge is attenuated e-fold.

For an X-ray laser, the *extinction length* is the thickness needed to fully reflect the beam; usually, it is a few micrometers while the absorption length is much larger.

In Segmüller (1968) the absorption length is $\frac{\sin \theta}{\mu}$, where μ is the linear absorption coefficient, and the beam enters the crystal at an angle θ .

Chapter 25 Distances in Geography, Geophysics, and Astronomy

25.1 Distances in Geography and Geophysics

• Great circle distance

The great circle distance (or spherical distance, orthodromic distance) is the shortest distance between points x and y on the surface of the Earth measured along a path on the Earth's surface. It is the length of the great circle arc, passing through x and y, in the spherical model of the planet.

Let δ_1 and ϕ_1 be, respectively, the latitude and the longitude of x, and δ_2 and ϕ_2 those of y; let r be the Earth's radius. Then the great circle distance is equal to

$$r \arccos(\sin \delta_1 \sin \delta_2 + \cos \delta_1 \cos \delta_2 \cos(\phi_1 - \phi_2)).$$

In the spherical coordinates (θ, ϕ) , where ϕ is the azimuthal angle and θ is the colatitude, the great circle distance between $x = (\theta_1, \phi_1)$ and $y = (\theta_2, \phi_2)$ is equal to

$$r \arccos (\cos \theta_1 \cos \theta_2 + \sin \theta_1 \sin \theta_2 \cos(\phi_1 - \phi_2))$$
.

For $\phi_1 = \phi_2$, the formula above reduces to $r|\theta_1 - \theta_2|$.

The **spheroidal distance** is the distance between two points on the Earth's surface in the spheroidal model of the planet. The shape of the Earth more closely resembles a flattened spheroid with extreme values for the radius of curvature of 6,336 km at the equator and 6,399 km at the poles.

Loxodromic distance

A rhumb line (or loxodromic curve) is a curve on the Earth's surface that crosses each meridian at the same angle. It is the path taken by a ship or plane that maintains a constant compass direction.

The **loxodromic distance** is a distance between two points on the Earth's surface on the rhumb line joining them. It is never shorter than the great circle distance.

The **nautical distance** is the length in nautical miles of the rhumb line joining any two places on the Earth's surface. One nautical mile is equal to 1,852 m (roughly, it is 1' of latitude; cf. **nautical length units** in Chap. 27).

• Continental shelf distance

Article 76 of the United Nations Convention on the Law of the Sea (1999), defined the continental shelf of a coastal state (its sovereignty domain) as the seabed and subsoil of the submarine areas that extend beyond its territorial sea throughout the natural prolongation of its land territory to the outer edge of the continental margin. It postulated that the **continental shelf distance**, i.e., the **range distance** from the baselines from which the breadth of the territorial sea is measured to above the other edge, should be within 200–350 nautical miles, and gave rules of its (almost) exact determination.

Article 47 of the same convention postulated that, for an archipelagic state, the ratio of the area of its archipelagic waters (sovereignty domain) to the area of its land, including atolls, should be between 1–1 and 9–1, and elaborated case-by-case rules.

• Radio distances

The main modes of electromagnetic wave (radio, light, X-rays, etc.) propagation are *direct wave* (line-of-sight), *surface wave* (interacting with the Earth's surface and following its curvature) and *skywave* (relying on refraction in the ionosphere).

The **line-of-sight distance** is the distance which radio signals travel, from one antenna to another, by a *line of sight path*, where both antennas are visible to one another, and there are no metallic obstructions.

The **horizon distance** is the distance on the Earth's surface reached by a direct wave; due to ionospheric refraction or tropospheric events, it is sometimes greater than the distance to the visible horizon.

In Television, the **horizon distance** is the distance of the farthest point on the Earth's surface visible from a transmitting antenna.

The **skip distance** is the shortest distance that permits a radio signal (of given frequency) to travel as a skywave from the transmitter to the receiver by reflection (hop) in the ionosphere.

If two radio frequencies are used (for instance, 12.5 and 25 kHz in maritime communication), the **interoperability distance** and **adjacent channel separation distance** are the range within which all receivers work with all transmitters and, respectively, the minimal distance which should separate adjacent tunes for narrow-band transmitters and wideband receivers, in order to avoid interference.

DX is amateur radio slang (and Morse code) for distance; **DXing** is a distant radio exchange (amplifiers required).

• Tolerance distance

In GIS (computer-based Geographic Information System), the **tolerance** distance is the maximal distance between points which must be established so that gaps and overshoots can be corrected (lines snapped together) as long as they fall within the tolerance distance.

• Map's distance

The **map's distance** is the distance between two points on the map (not to be confused with **map distance** between two loci on a genetic map from Chap. 25).

The **horizontal distance** is determined by multiplying the map's distance by the numerical scale of the map.

Map resolution is the size of the smallest feature that can be represented on a surface; more generally, it is the accuracy at which the location and shape of map features can be depicted for a given map scale.

• Equidistant map

An **equidistant map** is a map projection of Earth having a well-defined non-trivial set of *standard lines*, i.e., lines (straight or not) with constant scale and length proportional to corresponding lines on Earth. Some examples are:

Sanson-Flamsteed equatorial map: all parallels are straight lines;

Cylindric equidistant map: the vertical lines and equator are straight lines;

Azimuthal equidistant map preserves distances along any line through the central point; Werner cordiform map preserves, moreover, distances along any arc centered at that point.

• Horizontal distance

The horizontal distance (or ground distance) is the distance on a true level plane between two points, as scaled off of the map (it does not take into account the relief between two points). There are two types of horizontal distance: straight-line distance (the length of the straight-line segment between two points as scaled off of the map), and distance of travel (the length of the shortest path between two points as scaled off of the map, in the presence of roads, rivers, etc.).

• Slope distance

The **slope distance** (or **slant distance**) is the inclined distance (as opposed to true horizontal or vertical distance) between two points.

Walking uphill, humans and animals minimize metabolic energy expenditure; so, at critical slopes, they shift to zigzag walking. Langmuir's mountain hiking handbook advises to do it at 25°. Llobera and Sluckin (2007) explain switchbacks in hill trails by the need, for walkers, to zigzag in order to maintain the critical slope, $\approx 16^{\circ}$ uphill and $\approx 12.4^{\circ}$ downhill. Skiing and sailing against the wind also require zigzagging.

• Setback distance

In land use, a **setback distance** (or *setback*, *buffer distance*) is the minimum horizontal distance at which a building or other structure must

legally be from property lines, or street, watercourse, or any other place which needs protection. Setbacks may also allow for public utilities to access the buildings, and for access to utility meters. Cf. also **buffer distance** and **clearance distance** in Chap. 29.

• Distance-based numbering

The distance-based exit number is a number assigned to a road junction, usually an exit from a freeway, expressing in miles (or kilometers) the distance from the beginning of the highway to the exit. A *milestone* (or *kilometer sign*) is one of a series of numbered markers placed along a road at regular intervals. Zero Milestone in Washington, DC is designated as the reference point for all road distances in US.

Distance-based house addressing is the system (especially, in US) when buildings and blocks are numbered according to distance from a given baseline. For example, in Florida Keys, house number 67,430 is 67.4 miles from Mile Marker 0 in Key West; in Naperville, house number 67W430 is 67 miles west of downtown Chicago. The GIS-inspired guideline in US state Georgia is to use the address $n = \frac{d}{10} + 100$, where d is the distance in feet of the house from the reference point; roughly, this distance in miles is $\frac{n}{500}$.

Metes and bounds is a traditional system of land description (in real estate and town boundaries determination) by courses and distances. Metes is a boundary defined by the measurement of each straight run specified as displacement, i.e., by the distance and direction. Bounds refers to a general boundary description in terms of local geography (along some watercourse, public road, wall etc.). The boundaries are described in a running prose style, all the way around the parcel of land in sequence. Surveying is the technique and science of determining the terrestrial and spatial position of points and the distances and angles between them; cf., for example, Surveyor's Chain measures among Imperial length measures in Chap. 27.

• Road travel distance

The **road travel** (or *road*, *driving*, *wheel*, *actual*) **distance** between two locations (say, cities) of a region is the length of the shortest road connecting them.

Some GISs (Geographic Information Systems) approximate road distances as the l_p -metric with $p \approx 1.7$ or as a linear function of **great circle distances**; in US the multiplier is ≈ 1.15 in an east-west direction and ≈ 1.21 in the north-south direction. Several relevant notions of distance follow.

The **GPS** navigation distance: the distance directed by GPS (Global Positioning System, cf. radio distance measurement in Chap. 29) navigation devices. But this shortest route, from the GPS system point of view, is not always the best, for instance, when it directs a large truck to drive through a tiny village. Cf. also the little boy's paradox among quotes on "near-far" distances in Chap. 28.

The **official distance**: the recognized driving distance between two locations that will be used for travel or payment of billing.

The distance between zip codes (in general, postal or telephone area codes) is the estimated driving distance (or driving time) between two corresponding locations.

The **journey length** is a general notion of distance (usually in kilometers) used as a reference in transport studies. It can refer to, for example, the average distance traveled per person by some mode of transport (walk, cycle, car, bus, rail, taxi) or a statutory vehicle distance as in the evaluation of aircraft fuel consumption.

The world's remotest place is on the Tibetan plateau (34.7°N, 85.7°E): 1 day by car and 20 by foot to Lhasa.

A *trip meter* is a device used for recording the distance traveled by an automobile in any particular journey.

• Road sight distances

In Transportation Engineering, the *normal visual acuity* is the ability of a person to recognize a letter (or an object) of size ≈ 8.5 mm from a distance ≈ 6 m.

The **sight distance** (or *clear sight distance*) is the length of highway visible to a driver. The **safe sight distance** is the necessary sight distance needed by the driver in order to accomplish a fixed task; the main safe distances, used in road design, are:

The **stopping sight distance** – to stop the vehicle before reaching an unexpected obstacle

The $maneuver\ sight\ distance$ – to drive around an unexpected small obstacle

The road view sight distance – to anticipate the alignment (eventually curved and horizontal/vertical) of the road (for instance, choosing a speed)

The passing sight distance – to overtake safely (the distance the opposing vehicle travels during the overtaking manoeuvre)

The safe overtaking distance is the sum of four distances: the passing sight distance, the perception-reaction distance (between decision and action), the distance physically needed for overtaking and the buffer safety distance.

Also, adequate sight distances are required locally: at intersections and in order to process information on traffic signs.

• Atmospheric visibility distances

Atmospheric extinction (or attenuation) is a decrease in the amount of light going in the initial direction due to absorption (stopping) and scattering (direction change) by particles (solid or liquid, with diameter $0.002-100\,\mu\text{m}$) or gaze molecules. The dominant processes responsible for it are Rayleigh scattering (by particles smaller than the wavelength of the incident light) and absorption by dust, ozone O_3 and water.

In extremely clean air in the Arctic or mountainous areas, the visibility can reach 70–100 km. But it is often reduced by air pollution and high humidity: haze (in dry air) or mist (moist air). *Haze* is an atmospheric

condition where dust, smoke and other dry particles (from farming, traffic, industry, fires, etc.) obscure the sky.

The World Meteorological Organization classifies the horizontal obscuration into the categories of fog (a cloud in contact with the ground), ice fog, steam fog, mist, haze, smoke, volcanic ash, dust, sand and snow. Fog and mist are composed mainly of water droplets, haze and smoke can be of smaller particle size.

Visibility of less than $100 \,\mathrm{m}$ is usually reported as zero. The international definition of fog, mist and haze is a visibility of <1, 1-2, and $2-5 \,\mathrm{km}$, respectively. Visibility is especially useful for safety reasons in traffic (roads, sailing and aviation).

In air pollution literature, **visibility** (or *daylight visual range*) is the distance at which the contrast of a visual target against the background (usually, the sky) is equal to the threshold contrast value for the human eye, necessary for object identification, while **visual range** is the distance at which the target is just visible. Visibility can be smaller than the visual range since it requires recognition of the object.

Visibility is usually characterized by either visual range or by the *extinction coefficient* (attenuation of light per unit distance due to four components: scattering and absorption by gases and particles in the atmosphere). It has units of inverse length and, under certain conditions, is inversely related to the visual range.

Meteorological range (or standard visibility, standard visual range) is an instrumental daytime measurement of the (daytime sensory) visual range of a target. It is the furthest distance at which a black object silhouetted against a sky would be visible assuming a 2% threshold value for an object to be distinguished from the background. Numerically, it is ln 50 divided by the extinction coefficient.

In Meteorology, **visibility** is the distance at which an object or light can be clearly discerned with the unaided eye under any particular circumstances. It is the same in darkness as in daylight for the same air. **Visual range** is defined as the greatest distance in a given direction at which it is just possible to see and identify with the unaided eye in the daytime, a prominent dark object against the sky at the horizon, and at night, a known, preferably unfocused, moderately intense light source.

The International Civil Aviation Organization defines the **nighttime visual range** (or *transmission range*) as the greatest distance at which lights of 1,000 cd can be seen and identified against an unlit background. Daytime and nighttime ranges measure the atmospheric attenuation of contrast and flux density, respectively.

In Aviation Meteorology, the **runway visual range** is the maximum distance along a runway at which the runway markings are visible to a pilot after touchdown. It is measured assuming constant contrast and illuminance thresholds.

Oblique visual range (or *slant visibility*) is the greatest distance at which a target can be perceived when viewed along a line of sight inclined to the horizontal.

• Atmosphere distances

The atmosphere distances are the altitudes above Earth's surface (mean sea level) which indicate approximately the following specific (in terms of temperature, electromagnetism, etc.) layers of its atmosphere.

Below 1–2 km: planetary boundary layer where the effects of friction (diurnal heat, moisture or momentum transfer to or from the surface) are significant. A thin (roughly, below 10 m) surface boundary layer where friction effects are more or less constant throughout (as opposed to decreasing with height, as they do above it) and, like friction, the effects of insulation and radiational cooling are strongest.

Below 8–16 km (over the poles and equator, respectively): *troposphere*, temperature decreases with height (the weather and clouds occur here).

From 8 km: death zone for human climbers (lack of oxygen); 8.848 km: summit of Mount Everest.

From 7-17 to $50 \,\mathrm{km}$: *stratosphere*, temperature increases with height (the ozone layer is at $19-48 \,\mathrm{km}$).

From 50 to 80–85 km: mesosphere, temperature decreases with height.

From 80–85 to 640–690 km: *thermosphere*, temperature increases with height (the altitude of International Space Station is 278–460 km).

100 km: *Kármán line* prescribed by FAI (Fédération Aéronautique International) as the boundary separating aeronautics and astronautics, i.e., the beginning of *outer space*. Above this altitude, a vehicle should travel faster than orbital velocity in order to get aerodynamic lift from the atmosphere needed to support itself.

From $\approx 500 \,\mathrm{km}$ upwards: exosphere (or outer atmosphere), where atoms rarely undergo collisions and so can escape into space. The altitudes of the Hubble Space Telescope, Landsat and GPS (Global Positioning Satellites) are 595, 705 and 20,200 km, respectively.

From 50–80 to 2,000 km: *ionosphere*, electrically conducting region, while *neutrosphere* is the region from the Earth's surface upward in which the atmospheric constituents are mainly unionized; the region of transition between the neutrosphere and the ionosphere is 70–90 km depending on latitude and season.

Up to 6–10 Earth's radii on sunward side: magnetosphere, where Earth's magnetic field still dominates that of the solar wind. Geospace is the region of space that stretches from the beginning of Earth's ionosphere to the end of its magnetosphere.

From about $90,000 \,\mathrm{km}$: the $100-1,000 \,\mathrm{km}$ thick Earth's bow shock (boundary between the magnetosphere and an ambient medium).

From 650 to $65,000\,\mathrm{km}$: Van Allen radiation belt of intense ionizing radiation.

 $80-100 \,\mathrm{km}$: the upper limit of homosphere, where the Earth's atmosphere has relatively uniform gaseous composition; the limit of the atmosphere is the level at which the atmospheric density becomes about one particle per cubic centimeter, i.e., the same as the density of interplanetary space (also, the altitude where a molecule of air ceases to be held in free paths which are segments of Earth orbits).

From 320,000 km: lunar gravity exceeds Earth's.

• Distances in Oceanography

The average and maximal depths of the ocean are $\approx 3,800$ and $11,524\,\mathrm{m}$ (Mindanao Deep), while the average and maximal land heights are ≈ 840 and $8,840\,\mathrm{m}$ (Mt. Everest).

Decay distance: the distance through which ocean waves travel as swell after leaving the generating area.

Deep water (or *short*, *Stokesian*) **wave:** a surface ocean wave that is traveling in water depth greater than one-quarter of its wavelength; the velocity of deep-water waves is independent of the depth. **Shallow water** (or *long*, *Lagrangian*) **wave**: a surface ocean wave of length 25 or more times larger than water depth.

Littoral (or *intertidal*): the zone is between high and low water marks. Sometimes, *littoral* refers to the zone between the shore and water depths of $\approx 200 \,\mathrm{m}$.

Oceanographic equator (or thermal equator): the zone of maximum sea surface temperature located near the geographic equator. Sometimes, it is defined more specifically as the zone within which the sea surface temperature exceeds 28°C.

Standard depth: a depth below the sea surface at which water properties should be measured and reported (according to the proposal by the International Association of Physical Oceanography in 1936), namely (in meters): 0, 10, 20, 30, 50, 75, 100, 150, 200, 250, 300, 400, 500, 600, 800, 1,000, 1,200, 1,500, 2,000, 2,500, 3,000, 4,000, 5,000, 6,000, 7,000, 8,000, 9,000, 10,000.

Charted depth: the recorded vertical distance from the tidal datum to the sea-bed.

Depth of no motion: a reference depth in a body of water at which it is assumed that the horizontal velocities are practically zero.

The **thermocline** and **pycnocline**: the layers where the water temperature and density, respectively, change rapidly with depth.

Depth of compensation: the depth at which illuminance has diminished to the extent that oxygen production through photosynthesis and oxygen consumption through respiration by plants are equal. The maximum depth for photosynthesis depends on plants and weather.

Depth of the effective sunlight penetration: the depth at which $\approx 1\%$ of solar energy penetrates; in general, it does not exceed 100 m. The ocean is opaque to electromagnetic radiation with a small window in the visible spectrum, but it is transparent to acoustic transmission.

Sound energy is absorbed (due to viscosity) and scattered (due to inhomogeneities such as temperature, bubbles, plankton) in the ocean, together called *attenuation* (or *extinction*). About 6% of the energy is back-scattered, but absorption is generally the largest of the two terms. The **extinction distance** is the distance over which the sound intensity falls to $\frac{1}{e}$ of its value. For wavelength $\lambda = 10^a \times 15\,\mathrm{cm}$, the extinction distance is $\approx 437 \times 100^a\,\mathrm{km}$.

A SOFAR channel (SOund Fixing And Ranging): a layer of water deep in the ocean where the speed of sound is at a minimum, because water pressure, temperature and salinity cause a minimum of water density. Sound waves can get caught and bent in this channel and travel hundreds of kilometers. In low and middle latitudes, the SOFAR channel axis lies 600–1,200 m below the sea surface; it is deepest in the subtropics and comes to the surface in high latitudes.

The **pelagic zone** (or open-ocean zone) consists of all the sea other than that near the coast or the sea floor. Within the epipelagic zone (above $\approx 200\,\mathrm{m}$) there is enough light for photosynthesis, and thus plants and animals are largely concentrated here; the mesopelagic zone ($\approx 200-1,000\,\mathrm{m}$) is the twilight zone. Below the epipelagic zone lies the aphotic zone which is not exposed to sunlight; the transition to it happens roughly at the depth of compensation or depth of effective sunlight penetration.

Depth of frictional resistance: the depth at which the wind-induced current direction is 180° from that of the true wind.

Mixing length: the distance which an eddy (a circular movement of water) maintains its identity until it mixes again; analogous to the mean free path of a molecule.

Mixed layer depth: the depth of the bottom of the *mixed layer*, i.e., a nearly isothermal surface layer of 40–150 m depth where water is mixed through wave action or thermohaline convection.

Depth of exponential mixing or depth of homogeneous mixing refers to a surface turbulent mixing layer in which the distribution of a constituent decreases exponentially, or is constant, respectively, with height.

Monin-Obukhov length: a rough measure of the height over the ground, where mechanically produced (by vertical wind shear) turbulence becomes smaller than the buoyant production of turbulent energy (dissipative effect of negative buoyancy). In the daytime over land, it is usually 1–50 m.

• Moho distance

The **Moho** distance is the distance from a point on the Earth's surface to the *Moho* interface (or *Mohorovicic seismic discontinuity*) beneath it. The *Moho* interface is the boundary between the Earth's brittle outer crust and the hotter softer mantle; the Moho distance ranges between 5 and 10 km beneath the ocean floor to 35–65 km beneath the continents.

Cf. the world deepest cave (Krubera, Caucasus: 2.1 km), deepest mine (Western Deep Levels gold mine, South Africa: about 4 km) and deepest drill (Kola Superdeep Borehole: 12.3 km). The temperature rises usually by 1° every 33 m. The Japanese research vessel *Chikyu* started to drill (from September 2007, 200 km off Nagoya cost) to the Moho interface. The ice drills by EPICA (European Program for Ice Coring in Antarctica) went 3.2 km deep and, in terms of climate data, 900,000 years back.

The Earth's mantle extends from the Moho discontinuity to the mantle-core boundary at a depth of approximately $2,890\,\mathrm{km}$. Liquid outer core of radius $3,400\,\mathrm{km}$ contains solid inner core of radius $1,220\,\mathrm{km}$. The mantle is divided into the upper and the lower mantle by a discontinuity at about $660\,\mathrm{km}$. Other seismic discontinuities are at about $60-90\,\mathrm{km}$ (Hales discontinuity), $50-150\,\mathrm{km}$ (Gutenberg discontinuity), $220\,\mathrm{km}$ (Lehmann discontinuity), $410\,\mathrm{km}$, $520\,\mathrm{km}$, and $710\,\mathrm{km}$.

The tectonic plates are large parts of the lithosphere (the solid layer of the crust and the upper mantle), up to 60 km deep. They float on the more plastic part of the mantle, the asthenosphere, 100–200 km deep.

• Distances in Seismology

The Earth's crust is broken into tectonic plates that move around (at some centimeters per year) driven by the thermal convection of the deeper mantle and by gravity. At their boundaries, plates stick most of the time and then slip suddenly. An earthquake, i.e., a sudden (several seconds) motion or trembling in the Earth, caused by the abrupt release of slowly accumulated strain, was, from 1906, seen mainly as a rupture (sudden appearance, nucleation and propagation of new crack or fault) due to elastic rebound. However, from 1966, it is seen within the framework of slippage along a pre-existing fault or plate interface, as the result of stick-slip frictional instability. So, an earthquake happens when dynamic friction becomes less than static friction. The advancing boundary of the slip region is called the rupture front. The standard approach assumes that the fault is a definite surface of tangential displacement discontinuity, embedded in a liner elastic crust.

Ninety percent of earthquakes are of tectonic origin, but they can be caused also by volcanic eruption, nuclear explosion and work in a large dam, well or mine. Earthquakes can be measured by **focal depth**, speed of slip, intensity (modified Mercali scale of earthquake effects), magnitude, acceleration (main destruction factor), etc. The Richter logarithmic scale of magnitude is computed from the amplitude and frequency of shock waves received by a seismograph, adjusted to account for **epicentral distance**. An increase of 1.0 of the Richter magnitude corresponds to an increase of 10 times in amplitude of the waves and ≈ 31 times in energy; the largest recorded value is 9.5 (Chile 1960).

An earthquake first releases energy in the form of shock pressure waves that move quickly through the ground with an up-and-down motion. Next come shear waves, which move along the surface, causing much damage: Love waves in a side-to-side fashion, followed by Rayleigh waves, which have a rolling motion.

Distance attenuation models (cf. **distance decay** in Chap. 29), used in earthquake engineering for buildings and bridges, postulate usually acceleration decay with increase of some **site-source distance**, i.e., the distance between seismological stations and the crucial (for the given model) "central" point of the earthquake.

The simplest model is the *hypocenter* (or focus), i.e., the point inside the Earth from which an earthquake originates (the waves first emanate, the seismic rupture or slip begins). The *epicenter* is the point of the Earth's surface directly above the hypocenter. This terminology is also used for other catastrophes, such as an impact or explosion of a nuclear weapon, meteorite or comet but, for an explosion in the air, the term *hypocenter* refers to the point on the Earth's surface directly below the burst. A list of the main Seismology distances follows.

The **focal depth:** the distance between the hypocenter and epicenter; the average focal depth is 100-300 km.

The **hypocentral distance**: the distance from the station to the hypocenter.

The epicentral distance (or earthquake distance:) the great circle distance from the station to the epicenter.

The **Joyner–Boore distance**: the distance from the station to the closest point of the Earth's surface, located over the *rupture surface*, i.e., the rupturing portion of the fault plane.

The **rupture distance:** the distance from the station to the closest point on the rupture surface.

The **seismogenic depth distance:** the distance from the station to the closest point of the rupture surface within the *seismogenic zone*, i.e., the depth range where the earthquake may occur; usually at depth 8–12 km.

The **cross-over distance**: the distance on a seismic refraction survey time-distance chart at which the travel times of the direct and refracted waves are the same.

Also used are the distances from the station to:

- The center of static energy release and the center of static deformation of the fault plane
- The surface point of maximal macroseismic intensity, i.e., of maximal ground acceleration (it can be different from the epicenter)
- The epicenter such that the reflection of body waves from the Moho interface (the crust-mantle boundary) contribute more to ground motion than directly arriving shear waves (it called the critical Moho distance)
- The sources of noise and disturbances: oceans, lakes, rivers, railroads, buildings

The **space–time link distance** between two earthquakes x and y is defined by

$$\sqrt{d^2(x,y) + C|t_x - t_y|^2},$$

where d(x, y) is the distance between their epicenters or hypocenters, $|t_x - t_y|$ is the time lag, and C is a scaling constant needed to connect distance d(x, y) and time.

The *earthquake distance effect*: at greater distances from its center, the perception of an earthquake becomes weaker and the lower frequency shaking dominates it.

Another space—time measure for catastrophic events is **distance between landfalls** for hurricanes hitting a given US state. It is (Landreneau 2003) the length of this state's coastline divided by the number of hurricanes which have affected this state from 1899.

25.2 Distances in Astronomy

A celestial object (or celestial body) is a term describing astronomical objects such as stars and planets. The celestial sphere is the projection of celestial objects into their apparent positions in the sky as viewed from the Earth. The celestial equator is the projection of the Earth's equator onto the celestial sphere. The celestial poles are the projections of Earth's North and South Poles onto the celestial sphere. The hour circle of a celestial object is the great circle of the celestial sphere, passing through the object and the celestial poles. The ecliptic is the intersection of the plane that contains the orbit of the Earth with the celestial sphere: seen from the Earth, it is the path that the Sun appears to follow across the sky over the course of a year. The vernal equinox point (or the First point in Aries) is one of the two points on the celestial sphere, where the celestial equator intersects the ecliptic: it is the position of the Sun on the celestial sphere at the time of the vernal equinox.

The *horizon* is the line that separates Earth from sky. It divides the sky into the upper hemisphere that the observer can see, and the lower hemisphere that he can not. The pole of the upper hemisphere (the point of the sky directly overhead) is called the *zenith*, the pole of the lower hemisphere is called the *nadir*.

In general, an **astronomical distance** is a distance from one celestial body to another (measured in light-years, parsecs, or Astronomical Units). The average distance between stars (in a galaxy like our own) is several light-years. The average distance between galaxies (in a cluster) is only about 20 times their diameter, i.e., several megaparsecs.

• Latitude

In spherical coordinates (r, θ, ϕ) , the **latitude** is the **angular distance** δ from the xy-plane (fundamental plane) to a point, measured from the origin; $\delta = 90^{\circ} - \theta$, where θ is the **colatitude**.

In a geographic coordinate system (or earth-mapping coordinate system), the **latitude** is the angular distance from the Earth's equator to an object,

measured from the center of the Earth. Latitude is measured in degrees, from -90° (South pole) to $+90^{\circ}$ (North pole). Parallels are the lines of constant latitude.

In Astronomy, the **celestial latitude** is the latitude of a celestial object on the celestial sphere from the intersection of the fundamental plane with the celestial sphere in a given *celestial coordinate system*. In the *equatorial coordinate system* the fundamental plane is the plane of the Earth's equator; in the *ecliptic coordinate system* the fundamental plane is the plane of the ecliptic; in the *galactic coordinate system* the fundamental plane is the plane of the Milky Way; in the *horizontal coordinate system* the fundamental plane is the observer's horizon. Celestial latitude is measured in degrees.

• Longitude

In spherical coordinates (r, θ, ϕ) , the **longitude** is the **angular distance** ϕ in the xy-plane from the x-axis to the intersection of a great circle, that passes through the point, with the xy-plane.

In a geographic coordinate system (or earth-mapping coordinate system), the **longitude** is the angular distance measured eastward along the Earth's equator from the *Greenwich meridian* (or *Prime meridian*) to the intersection of the meridian that passes through the object. Longitude is measured in degrees, from 0° to 360° . A *meridian* is a great circle, passing through Earth's North and South Poles; the meridians are the lines of constant longitude.

In Astronomy, the **celestial longitude** is the longitude of a celestial object on the celestial sphere measured eastward, along the intersection of the fundamental plane with the celestial sphere in a given *celestial coordinate system*, from the chosen point. In the *equatorial coordinate system* the fundamental plane is the plane of the Earth's equator; in the *ecliptic coordinate system* – the plane of the ecliptic; in the *galactic coordinate system* – the plane of the Milky Way; in the *horizontal coordinate system* – the observer's horizon. Celestial longitude is measured in units of time.

Colatitude

In spherical coordinates (r, θ, ϕ) , the **colatitude** is the **angular distance** θ from the z-axis to the point, measured from the origin; $\theta = 90^{\circ} - \delta$, where δ is the **latitude**.

In a geographic coordinate system (or earth-mapping coordinate system), the **colatitude** is the angular distance from the Earth's North Pole to an object, measured from the center of the Earth. Colatitude is measured in degrees.

Declination

In the equatorial coordinate system (or geocentric coordinate system), the declination δ is the celestial latitude of a celestial object on the celestial sphere, measured from the celestial equator. Declination is measured in degrees, from -90° to $+90^{\circ}$.

• Right ascension

In the equatorial coordinate system (or geocentric coordinate system), fixed to the stars, the **right ascension** RA is the **celestial longitude** of a celestial object on the celestial sphere, measured eastward along the celestial equator from the First point in Aries to the intersection of the hour circle of the celestial object. Right ascension is measured in units of time (hours, minutes and seconds) with 1 h of time approximately equal to 15 $^{\circ}$.

The time needed for one complete cycle of the precession of the equinoxes is called the *Platonic year* (or *Great year*); it is about 257 centuries and slightly decreases. This cycle is important in the Maya calendar and Astrology. The time it takes Earth's solar system to revolve once around the Milky Way center is called the *Galactic year*. It is estimated to be within 220–250 million Earth years.

• Hour angle

In the equatorial coordinate system (or geocentric coordinate system), fixed to the Earth, the hour angle is the celestial longitude of a celestial object on the celestial sphere, measured along the celestial equator from the observer's meridian to the intersection of the hour circle of the celestial object. The hour angle is measured in units of time (hours, minutes and seconds). It gives the time elapsed since the celestial object's last transit at the observer's meridian (for a positive hour angle), or the time until the next transit (for a negative hour angle).

• Polar distance

In the equatorial coordinate system (or geocentric coordinate system), the **polar distance** (or codeclination) PD is the **colatitude** of a celestial object, i.e., the **angular distance** from the celestial pole to a celestial object on the celestial sphere, similarly as the **declination** δ is measured from the celestial equator: $PD = 90^{\circ} \pm \delta$. Polar distance is expressed in degrees, and cannot exceed 90° in magnitude. An object on the celestial equator has $PD = 90^{\circ}$.

• Ecliptic latitude

In the *ecliptic coordinate system*, the **ecliptic latitude** is the **celestial latitude** of a celestial object on the celestial sphere from the ecliptic. Ecliptic latitude is measured in degrees.

• Ecliptic longitude

In the *ecliptic coordinate system*, the **ecliptic longitude** is the **celestial longitude** of a celestial object on the celestial sphere measured eastward along the ecliptic from the First point in Aries. Ecliptic longitude is measured in units of time.

• Altitude

In the horizontal coordinate system (or Alt/Az coordinate system), the altitude ALT is the celestial latitude of an object from the horizon. It is the complement of the zenith distance ZA: $ALT = 90^{\circ} - ZA$. Altitude is measured in degrees.

Azimuth

In the horizontal coordinate system (or Alt/Az coordinate system), the azimuth is the celestial longitude of an object, measured eastward along the horizon from the North point. Azimuth is measured in degrees, from 0° to 360° .

• Zenith distance

In the horizontal coordinate system (or Alt/Az coordinate system), the **zenith distance** (or North polar distance, zenith angle) ZA is the **colatitude** of an object, measured from the zenith. It is the complement the **altitude** ALT: $ZA = 90^{\circ} - ALT$.

• Lunar distance

The lunar distance is the angular distance between the Moon and another celestial object.

• Elliptic orbit distance

The elliptic orbit distance is the distance from a mass M which a satellite body has in an elliptic orbit about the mass M at the focus. This distance is given by

$$\frac{a(1-e^2)}{1+e\cos\theta},$$

where a is the semi-major axis, e is the eccentricity, and θ is the orbital angle.

The semi-major axis a of an ellipse (or an elliptic orbit) is half of its major diameter; it is the average (over the eccentric anomaly) elliptic orbit distance. Such average distance over the **true anomaly** is the semi-minor axis, i.e., half of its minor diameter. The eccentricity e of an ellipse (or an elliptic orbit) is the ratio of half the distance between the foci e and the semi-major axis $e = \frac{e}{a}$. For an elliptic orbit, $e = \frac{r_+ - r_-}{r_+ + r_-}$, where e is the apoapsis distance, and e is the periapsis distance.

• Periapsis distance

The **periapsis distance** is the closest distance r_{-} a body reaches in an elliptic orbit about a mass M: $r_{-} = a(1 - e)$, where a is the *semi-major axis* and e is the *eccentricity*.

The **perigee** is the periapsis of an elliptical orbit around the Earth. The **perihelion** is the periapsis of an elliptical orbit around the Sun. The **periastron** is the point in the orbit of a double star where the smaller star is closest to its primary.

• Apoapsis distance

The **apoapsis distance** is the farthest distance r_+ a body reaches in an elliptic orbit about a mass M: $r_+ = a(1 + e)$, where a is the *semi-major axis*, and e is the *eccentricity*.

The **apogee** is the apoapsis of an elliptical orbit around the Earth. The **aphelion** is the apoapsis of an elliptical orbit around the Sun. The **apastron** is the point in the orbit of a double star where the smaller star is farthest from its primary.

• True anomaly

The **true anomaly** is the **angular distance** of a point in an orbit past the point of **periapsis** measured in degrees.

• Titius-Bode law

Titius–Bode law is an empirical (not explained well yet) law approximating the mean planetary distance from the Sun (i.e., its orbital *semi-major axis*) by $\frac{3k+4}{10}$ AU. Here 1 AU denotes such mean distance for Earth (i.e., about 1.5×10^8 km ≈ 8.3 light-minutes) and $k = 0, 2^0, 2^1, 2^2, 2^3, 2^4, 2^5, 2^6, 2^7$ for Mercury, Venus, Earth, Mars, Ceres (the largest one in the Asteroid Belt, $\approx \frac{1}{3}$ of its mass), Jupiter, Saturn, Uranus, Pluto. However, Neptune does not fit in the law while Pluto fits Neptune's spot $k = 2^7$.

• Primary-satellite distances

Consider two celestial bodies: a primary M and a smaller one m (a satellite, orbiting around M, or a secondary star, or a comet passing by).

The **mean distance** is the arithmetic mean of the maximum and minimum distances of a body m from its primary M.

Let ρ_M , ρ_m and R_M , R_m denote the densities and radii of M and m. Then the **Roche limit** of the pair (M,m) is the maximal distance between them within which m will disintegrate due to the tidal forces of M exceeding the gravitational self-attraction of m. This distance is $R_M \sqrt[3]{2\frac{\rho_M}{\rho_m}} \approx 1.26R_M \sqrt[3]{\frac{\rho_M}{\rho_m}}$ if m is a rigid spherical body, and it is about $2.423R_M \sqrt[3]{\frac{\rho_M}{\rho_m}}$ if body m is fluid. The Roche limit is relevant only if it exceeds R_M . It is $0.80R_M$, $1.49R_M$ and $2.80R_M$ for the pairs (the Sun, the Earth), (the Earth, the Moon) and (the Earth, a comet), respectively. A possible origin of the rings of Saturn is a moon which came closer to Saturn than its Roche limit.

Let d(m, M) denote the distance between m and M; let S_m and S_M denote the masses of m and M. Then the **Hill sphere of** m **in presence of** M is an approximation to the gravitational sphere of influence of m in the face of perturbation from M. Its radius is about $d(m, M) \sqrt[3]{\frac{S_m}{3S_M}}$. For example, the radius of Hill sphere of the Earth is 0.01 AU; the Moon, at distance 0.0025 AU, is within the Hill sphere of the Earth.

The pair (M, m) can be characterized by five **Lagrange points** L_i , $1 \le i \le 5$, where a third, much smaller body (say, a spacecraft) will be relatively stable because its centrifugal force is equal to the combined gravitational attraction of M and m. These points are:

 L_1 , L_2 , L_3 lying on the line through the centers of M and m so that $d(L_3,m)=2d(M,m)$, $d(M,L_2)=d(M,L_1)+d(L_1,m)+d(m,L_2)$, $d(L_1,m)=d(m,L_2)$, respectively; The satellite SOHO (Solar and Heliospheric Observatory) is at the semi-stable point L_1 of the Sun-Earth gravitational system, where the view of the Sun is uninterrupted. The satellite WMAP (Wilkinson Microwave Anisotropy Probe) is at L_2 ; the Planck Surveyor will be there in 2008. In 2013 NASA will put at L_2 the NGST

(Next Generation Space Telescope) since the cold and stable temperature of L_2 enhances infrared observations of faint and very distant objects.

 L_4 and L_5 lying on the orbit of m around M and forming equilateral triangles with the centers of M and m. These two points are more stable; each of them forms with M and m a partial solution of the (unsolved) gravitational three-body problem. Objects orbiting at the L_4 and L_5 points are called Trojans. The Moon was created 4.5 billion years ago by a sidewise impact on the Earth Big Splash by a Mars-sized Trojan planetoid slowly approaching from the L_4 Lagrange point of the Sun–Earth system.

The most tenuously linked binary in the solar system is 2001 QW322: two icy bodies in the Kuiper belt, at mean distance $\geq 10^4$ km, orbiting each other at 3 km h^{-1} .

• Solar distances

The mean distance of Sun from Earth is $1 \text{ AU} \approx 1.496 \times 10^{11} \text{ m}$.

The mean distance of Sun from the Milky Way core is $\approx 2.5 \times 10^{20}$ m (26,000 light-years).

The Sun's radius is 6.955×10^8 m; it is measured from its center to the edge of the *photosphere* (≈ 500 km thick layer below which the Sun is opaque to visible light).

The Sun does not have a definite boundary, but it has a well-defined interior structure: the *core* extending from the center to ≈ 0.2 solar radii, the *radiative zone* at ≈ 0.2 –0.8 solar radii, where thermal radiation is sufficient to transfer the intense heat of the core outward, the *tachocline* (transition layer) and the *convection zone*, where thermal columns carry hot material to the surface (photosphere) of the Sun.

The principal zones of the *solar atmosphere* (parts above the photosphere) are: temperature minimum, chromosphere, transition region, corona, and heliosphere.

The *chromosphere*, $\approx 3,000$ km deep layer, is more visually transparent. The *corona* is a highly rarefied tenuous region continually varying in size and shape; it is visible only during a total solar eclipse. The chromosphere-corona region is much hotter than the Sun's surface. As the corona extends further, it becomes the *solar wind*, a very thin gas of charged particles that travels through the solar system.

The *heliosphere* is the teardrop-shaped region around the Sun created by the solar wind and filled with solar magnetic fields and outward-moving gas. It extends from ≈ 20 solar radii (0.1 AU) outward 86–100 AU past the orbit of Pluto to the *heliopause*, where the interstellar medium and solar wind pressures balance.

The interstellar medium and the solar wind are moving supersonically in opposite directions, towards and away from the Sun. The point, $\approx 80 \,\mathrm{AU}$ from the Sun, where the solar wind becomes subsonic is the *termination shock*. The point, $\approx 230 \,\mathrm{AU}$ from the Sun, where the interstellar medium becomes subsonic is the *bow shock*.

• Habitable zone radii

The habitable zone radii of a star are the minimal and maximal orbital radii such that liquid water may exist on a terrestrial (i.e., primarily composed of silicate or, possibly, carbon rocks) planet orbiting within this range, so that life could develop there in a similar way as on the early Earth. For our Sun, such radii are around 0.95 and 1.37 AU and include only Earth. Maximally Earth-like mean temperature is expected at the distance $\sqrt{\frac{L_{star}}{L_{sun}}}$ AU from a star, where L is total radiant energy (Earth's average temperature is 14.4°C). The Sun is becoming hotter (at least 30% more since the formation of Earth and another 10% over the next 1.1 billion years); so, it will be too hot, even for microbial life, in 500–900 million years.

The only known extrasolar planets orbiting near the habitable zone of their star are *super-earths* (terrestrial planets of 1–10 Earth's mass) Gl 581c and Gl 581d orbiting near, respectively, the inner and outer edges of the habitable zone of red dwarf Gliese 581 in the constellation Libra of our galaxy, 20.4 light-years from Sun. Gl 581c has temperature within [-3°C, 40°C], while Earth's life exists within [-15°C, 121°C].

According to Lineweaver, Fenner and Gibson (2004) the galactic habitable zone of our Milky Way is a slowly expanding annular region between 7 and 9 kpc of **galactocentric distance**; so, the minimal and maximal radii of this zone are 22,000 and 28,000 light-years. They used four prerequisites for complex life: the presence of a host star, enough heavy elements to form terrestrial planets, sufficient time (4 \pm 1 billion years) for biological evolution and an environment free of supernovae.

The **Dyson radius** of a star is the radius of a hypothetical *Dyson sphere* around it, i.e., a megastructure (say, a system of orbiting star-powered satellites) meant to completely encompass a star and capture a large part of its energy output. The solar energy, available at distance d (measured in AU) from the Sun, is $\frac{1,366}{d^2}$ W m⁻². The inner surface of the sphere is intended to be used as habitat. An example of such speculations: at Dyson radius 300×10^6 km from Sun a continuous structure with thermal ambient 20° C (on the inner surface) and efficiency 3% of power generation (by a heat flux to -3° C on the outer surface) is conceivable.

• SETI detection ranges

SETI (Search for Extra Terrestrial Intelligence) activity involves using sensitive radio telescopes to search for a possible alien radio transmission. The recorded signals are mostly random noise but in 1977 a very strong signal (called WOW!) was received at $\leq\!50\,\mathrm{kHz}$ of the frequency 1420.406 MHz of hydrogene line.

There are **SETI detection ranges**, i.e., the maximal distances over which detection is still possible using given assumptions about frequency, antenna dish size, receiver bandwidth, etc. They are low for broadband signals from Earth (from 0.007 AU for AM radio up to 5.4 AU for EM radio) but reach 720 light-years for the S-Band of the world's largest radio telescope Arecibo.

Positive SETI consists of sending signals into space in the hope that they will be picked up by an alien intelligence. The first radio signals from Earth to reach space were produced around 1940 but television and radio signals actually decompose into static within 1-2 light years. In 1974 Arecibo sent a very elaborate radio signal aimed at the globular star cluster M13, 25,000 light-years away.

Chapter 26 Distances in Cosmology and Theory of Relativity

26.1 Distances in Cosmology

The *Universe* is defined as the whole space—time continuum in which we exist, together with all the energy and matter within it.

Cosmology is the study of the large-scale structure of the Universe. Specific cosmological questions of interest include the *isotropy* of the Universe (on the largest scales, the Universe looks the same in all directions, i.e., is invariant to rotations), the *homogeneousness* of the Universe (any measurable property of the Universe is the same everywhere, i.e., it is invariant to translations), the density of the Universe, the equality of matter and anti-matter, and the origin of density fluctuations in galaxies.

In 1929, Hubble discovered that all galaxies have a positive redshift, i.e., all galaxies, except for a few nearby galaxies like Andromeda, are receding from the Milky Way. By the Copernican principle (that we are not at a special place in the Universe), we deduce that all galaxies are receding from each other, i.e., we live in a dynamic, expanding Universe, and the further a galaxy is away from us, the faster it is moving away (this is now called the Hubble law). The Hubble flow is the general outward movement of galaxies and clusters of galaxies resulting from the expansion of the Universe. It occurs radially away from the observer, and obeys the Hubble law. Galaxies can overcome this expansion on scales smaller than that of clusters of galaxies; the clusters, however, are being forever driven apart by the Hubble flow.

In Cosmology, the prevailing scientific theory about the early development and shape of the Universe is the *Big Bang Theory*. The observation that galaxies appear to be receding from each other can be combined with the General Theory of Relativity to extrapolate the condition of the Universe back in time. This leads to the construction that, as one goes back in time, the Universe becomes increasingly hot and dense, then leads to a gravitational singularity, at which all distances become zero, and temperatures and pressures become infinite. The term *Big Bang* is used to refer to a hypothesized point in time when the observed expansion of the Universe began.

Based on measurements of the expansion of the Universe, it is currently believed that the Universe has an age of 13.7 ± 0.2 billion years. It should be longer if the expansion accelerates, as was supposed recently. Basing on the abundance ratio of uranium/thorium chondritic meteorites, [Dau05] estimated this age as 14.5 ± 2 billion years.

In Cosmology (or, more exactly, *Cosmography*, the measurement of the Universe) there are many ways to specify the distance between two points, because in the expanding Universe, the distances between comoving objects are constantly changing, and Earth-bound observers look back in time as they look out in distance. The unifying aspect is that all distance measures somehow measure the separation between events on *radial null trajectories*, i.e., trajectories of photons which terminate at the observer. In general, the **cosmological distance** is a distance far beyond the boundaries of our Galaxy.

The geometry of the Universe is determined by several cosmological parameters: the expansion parameter (or the scale factor) a, the Hubble constant H, the density ρ and the critical density ρ_{crit} (the density required for the Universe to stop expansion and, eventually, collapse back onto itself), the cosmological constant Λ , the curvature k of the Universe. Many of these quantities are related under the assumptions of a given cosmological model. The most common cosmological models are the closed and open Friedmann–Lemaître cosmological models and the Einstein–de Sitter cosmological model.

The Einstein–de Sitter cosmological model assumes a homogeneous, isotropic, constant curvature Universe with zero cosmological constant Λ and pressure P. For constant mass M of the Universe, $H^2 = \frac{8}{3}\pi G\rho$, $t = \frac{2}{3}H^{-1}$, $a = \frac{1}{R_C}(\frac{9GM}{2})^{1/3}t^{2/3}$, where $G = 6.67 \times 10^{-11}m^3kg^{-1}s^{-2}$ is the gravitational constant, $R_C = |k|^{-\frac{1}{2}}$ is the radius of curvature, and t is the age of the Universe.

The expansion parameter a=a(t) is a scale factor, relating the size of the Universe R=R(t) at time t to the size of the Universe $R_0=R(t_0)$ at time t_0 by $R=aR_0$. Most commonly in modern usage it is chosen to be dimensionless, with $a(t_{obser})=1$, where t_{obser} is the present age of the Universe.

The Hubble constant H is the constant of proportionality between the speed of expansion v and the size of the Universe R, i.e., v = HR. This equality is just the Hubble law with the Hubble constant $H = \frac{a'(t)}{a(t)}$. This is a linear redshift-distance relationship, where redshift is interpreted as recessional velocity v, typically expressed in km/s.

The current value of the Hubble constant is $H_0 = 71 \pm 4 \,\mathrm{km} \,\mathrm{s}^{-1} \mathrm{Mpc}^{-1}$, where the subscript 0 refers to the present epoch because H changes with time. The Hubble time and the **Hubble distance** are defined by $t_H = \frac{1}{H_0} \approx 4.35 \times 10^{17} \,\mathrm{s}$ and $D_H = \frac{c}{H_0}$ (here c is the speed of light), respectively. The Hubble volume is the volume of universe with a comoving size of $\frac{c}{H_0}$ (a sphere with radius $\approx 14,000 \,\mathrm{Mpc}$, mass $\approx 10^{60} \,\mathrm{kg}$ and $\approx 10^{80} \,\mathrm{atoms}$).

The mass density $\rho = \rho_0$ in the present epoch and the value of the cosmological constant Λ are dynamical properties of the Universe. They can be made into dimensionless density parameters Ω_M and Ω_{Λ} by $\Omega_M = \frac{8\pi G \rho_0}{3H_0^3}$, $\Omega_{\Lambda} = \frac{\Lambda}{3H_0^3}$. A third density parameter Ω_R measures the "curvature of space," and can be defined by the relation $\Omega_M + \Omega_{\Lambda} + \Omega_R = 1$.

These parameters totally determine the geometry of the Universe if it is homogeneous, isotropic, and matter-dominated.

The velocity of a galaxy is measured by the *Doppler effect*, i.e., the fact that light emitted from a source is shifted in wavelength by the motion of the source. (The Doppler shift is reversed in some metamateials: a light source moving toward an observer appears to reduce its frequency.) A relativistic form of the Doppler shift exists for objects traveling very quickly, and is given by $\frac{\lambda_{obser}}{\lambda_{emit}} = \sqrt{\frac{c+v}{c-v}}$, where λ_{emit} is the emitted wavelength, and λ_{obser} is the shifted (observed) wavelength. The change in wavelength with respect to the source at rest is called the redshift (if moving away), and is denoted by the letter z. The relativistic redshift z for a particle is given by $z = \frac{\Delta \lambda_{obser}}{\lambda_{emit}} =$

$$\frac{\lambda_{obser}}{\lambda_{emit}} - 1 = \sqrt{\frac{c+v}{c-v}} - 1.$$

The cosmological redshift is directly related to the scale factor a = a(t): $z+1=\frac{a(t_{obser})}{a(t_{emit})}$. Here $a(t_{obser})$ is the value of the scale factor at the time the light from the object is observed, and $a(t_{emit})$ is the value of the scale factor at the time it was emitted.

• Hubble distance

The **Hubble distance** is a constant

$$D_H = \frac{c}{H_0} \approx 4220\,\mathrm{Mpc} \approx 1.3\times 10^{26}\,\mathrm{m} \approx 13.7\times 10^9 \mathrm{\ light-years},$$

where c is the speed of light, and $H_0 = 71 \pm 4 \,\mathrm{km}\,\mathrm{s}^{-1}\,\mathrm{Mpc}^{-1}$ is the Hubbleconstant.

It is the distance from us to the cosmic light horizon which marks the edge of the visible Universe, i.e., the radius of a sphere, centered upon the Earth, which is approximately 13.7 billion light-years. It is often referred to as the **lookback distance** because astronomers, who view distant objects, are "looking back" into the history of the Universe.

For small v/c or small distance d in the expanding Universe, the velocity is proportional to the distance, and all distance measures, for example, angular diameter distance, luminosity distance, etc., converge. Taking the linear approximation, this reduces to $d \approx zD_H$, where z is the redshift. But this is true only for small redshifts.

Comoving distance

The standard Big Bang model uses comoving coordinates, where the spatial reference frame is attached to the average positions of galaxies. With this set of coordinates, both the time and expansion of the Universe can be ignored and the shape of space is seen as a spatial hypersurface at constant cosmological time.

The **comoving distance** (or coordinate distance, cosmological distance, χ) is a distance in comoving coordinates between two points in space at a single cosmological time, i.e., the distance between two nearby objects in the Universe which remains constant with epoch if the two objects are moving with the Hubble flow. It is the distance between them which would be measured with rulers at the time they are being observed (the **proper distance**) divided by the ratio of the scale factor of the Universe then to now. In other words, it is the proper distance multiplied by (1+z), where z is the redshift:

$$d_{comov}(x,y) = d_{proper}(x,y) \cdot \frac{a(t_{obser})}{a(t_{emit})} = d_{proper}(x,y) \cdot (1+z).$$

At the time t_{obser} , i.e., in the present epoch, $a=a(t_{obser})=1$, and $d_{comov}=d_{proper}$, i.e., the comoving distance between two nearby events (close in redshift or distance) is the proper distance between them. In general, for a cosmological time t, $d_{comov}=\frac{d_{proper}}{a(t)}$.

The total line-of sight comoving distance D_C from us to a distant object is computed by integrating the infinitesimal $d_{comov}(x, y)$ contributions between nearby events along the time ray from the time t_{emit} , when the light from the object was emitted, to the time t_{obser} , when the object is observed:

$$D_C = \int_{t_{emit}}^{t_{obser}} \frac{cdt}{a(t)}.$$

In terms of redshift, D_C from us to a distant object is computed by integrating the infinitesimal $d_{comov}(x,y)$ contributions between nearby events along the radial ray from z=0 to the object: $D_C=D_H\int_0^z \frac{dz}{E(z)}$, where D_H is the **Hubble distance**, and $E(z)=(\Omega_M(1+z)^3+\Omega_R(1+z)^2+\Omega_\Lambda)^{\frac{1}{2}}$.

In a sense, the comoving distance is the fundamental distance measure in Cosmology since all other distances can simply be derived in terms of it.

• Proper distance

The **proper distance** (or **physical distance**, ordinary distance) is a distance between two nearby events in the frame in which they occur at the same time. It is the distance measured by a ruler at the time of observation. So, for a cosmological time t,

$$d_{proper}(x, y) = d_{comov} \cdot a(t),$$

where d_{comov} is the **comoving distance**, and a(t) is the *scale factor*.

In the present epoch (i.e., at the time t_{obser}) $a = a(t_{obser}) = 1$, and $d_{proper} = d_{comov}$. So, the proper distance between two nearby events (i.e.,

close in redshift or distance) is the distance which we would measure locally between the events today if those two points were locked into the Hubble flow.

• Proper motion distance

The proper motion distance (or transverse comoving distance, contemporary angular diameter distance) D_M is a distance from us to a distant object, defined as the ratio of the actual transverse velocity (in distance over time) of the object to its proper motion (in radians per unit time). It is given by

$$D_{M} = \begin{cases} D_{H} \frac{1}{\sqrt{\Omega_{R}}} \sinh(\sqrt{\Omega_{R}} D_{C}/D_{H}), & \text{for } \Omega_{R} > 0, \\ D_{C}, & \text{for } \Omega_{R} = 0, \\ D_{H} \frac{1}{\sqrt{|\Omega_{R}|}} \sin(\sqrt{|\Omega_{R}|} D_{C}/D_{H}), & \text{for } \Omega_{R} < 0, \end{cases}$$

where D_H is the **Hubble distance**, and D_C is the **line-of-sight** comoving distance. For $\Omega_{\Lambda} = 0$, there is an analytic solution (z is the redshift):

$$D_{M} = D_{H} \frac{2(2 - \Omega_{M}(1 - z) - (2 - \Omega_{M})\sqrt{1 + \Omega_{M}z})}{\Omega_{M}^{2}(1 + z)}.$$

The proper motion distance D_M coincides with the line-of-sight comoving distance D_C if and only if the curvature of the Universe is equal to zero. The **comoving distance** between two events at the same redshift or distance but separated in the sky by some angle $\delta\theta$ is equal to $D_M\delta\theta$.

The distance D_M is related to the **luminosity distance** D_L by $D_M = \frac{D_L}{1+z}$, and to the **angular diameter distance** D_A by $D_M = (1+z)D_A$.

• Luminosity distance

The luminosity distance D_L is a distance from us to a distant object, defined by the relationship between the observed flux S and emitted luminosity L:

$$D_L = \sqrt{\frac{L}{4\pi S}}.$$

This distance is related to the **proper motion distance** D_M by $D_L = (1+z)D_M$, and to the **angular diameter distance** D_A by $D_L = (1+z)^2D_A$, where z is the redshift.

The luminosity distance does take into account the fact that the observed luminosity is attenuated by two factors, the relativistic redshift and the Doppler shift of emission, each of which contributes an (1+z) attenuation: $L_{obser} = \frac{L_{emiss}}{(1+z)^2}$.

The corrected luminosity distance D'_{L} is defined by $D'_{L} = \frac{D_{L}}{1+z}$.

• Distance modulus

The distance modulus DM is defined by $DM = 5 \ln(\frac{D_L}{10pc})$, where D_L is the luminosity distance. The distance modulus is the difference

between the absolute magnitude and apparent magnitude of an astronomical object. Distance moduli are most commonly used when expressing the distances to other galaxies. For example, the Large Magellanic Cloud is at a distance modulus 18.5, the Andromeda Galaxy's distance modulus is 24.5, and the Virgo Cluster has the DM equal to 31.7.

• Angular diameter distance

The angular diameter distance (or angular size distance) D_A is a distance from us to a distant object, defined as the ratio of an object's physical transverse size to its angular size (in radians). It is used to convert angular separations in telescope images into proper separations at the source. It is special for not increasing indefinitely as $z \to \infty$; it turns over at $z \sim 1$, and therefore more distant objects actually appear larger in angular size. Angular diameter distance is related to the **proper motion distance** D_M by $D_A = \frac{D_M}{1+z}$, and to the **luminosity distance** D_L by $D_A = \frac{D_L}{(1+z)^2}$, where z is the redshift.

The distance duality $\frac{D_L(z)}{D_A(z)} = (1+z)^2$ links D_L , based on the apparent luminosity of standard candles (for example, supernovae) and D_A , based on the apparent size of standard rulers (for example, baryon oscillations). It holds for any general **metric theory of gravity** (see Chap. 24) in any background in which photons travel on unique null geodesics.

If the angular diameter distance is based on the representation of object diameter as angle \times distance, the **area distance** is defined similarly according the representation of object area as solid angle \times distance².

• Einstein radius

General Relativity predicts $gravitational\ lensing$, i.e., deformation of the light from a source (a galaxy or star) in the presence of a $gravitational\ lens$, i.e., a body of large mass M (another galaxy, or a black hole) bending it.

If the source S, lens L and observer O are all aligned, the gravitational deflection is symmetric around the lens. The **Einstein radius** is the radius of the resulting *Einstein* (or *Chwolson*) ring. In radians (for the gravitational constant G and the speed C of light) it is

$$\sqrt{M\frac{4G}{c^2}\frac{D(L,S)}{D(O,L)D(O,S)}},$$

where D(O, L) and D(O, S) are the **angular diameter distances** of the lens and source, while D(L, S) is the angular diameter distance between them.

• Light-travel distance

The **light-travel distance** (or *light-travel time distance*) D_{lt} is a distance from us to a distant object, defined by $D_{lt} = c(t_{obser} - t_{emit})$, where t_{obser} is the time when the object was observed, and t_{emit} is the time when the

light from the object was emitted. It is not a very useful distance, because it is very hard to determine t_{emit} , the age of the Universe at the time of emission of the light which we see.

• Parallax distance

The **parallax distance** D_P is a distance from us to a distant object, defined from measuring of *parallaxes*, i.e., its apparent changes of position in the sky caused by the motion of the observer on the Earth around the Sun.

The cosmological parallax is measured as the difference in the angles of line of sight to the object from two endpoints of the diameter of the orbit of the Earth which is used as a baseline. Given a baseline, the parallax $\alpha-\beta$ depends on the distance, and knowing this and the length of the baseline (two astronomical units AU, where $AU \approx 150$ million kilometers is the distance from the Earth to the Sun) one can compute the distance to the star by the formula

$$D_P = \frac{2}{\alpha - \beta},$$

where D_P is in parsecs, α and β are in arc-seconds.

In Astronomy, "parallax" usually means the annual parallax p which is the difference in the angles of a star seen from the Earth and from the Sun. Therefore the distance of a star (in parsecs) is given by $D_P = \frac{1}{n}$.

• Kinematic distance

The **kinematic distance** is the distance to a galactic source, which is determined from differential rotation of the galaxy: the *radial velocity* of a source directly corresponds to its Galactocentric radius. But the *kinematic distance ambiguity* arises since, in our inner galaxy, any given Galactocentric radius corresponds to two distances along the line of sight, *near* and *far* kinematic distances.

This problem is solved, for some galactic regions, by measurement of their absorption spectra, if there is an interstellar cloud between the region and observer.

• Radar distance

The radar distance D_R is a distance from us to a distant object, measured by a radar. Radar typically consists of a high frequency radio pulse sent out for a short interval of time. When it encounters a conducting object, sufficient energy is reflected back to allow the radar system to detect it. Since radio waves travel in air at close to their speed in vacuum, one can calculate the distance D_R of the detected object from the round-trip time t between the transmitted and received pulses as

$$D_R = \frac{1}{2}ct,$$

where c is the speed of light.

• Cosmological distance ladder

For measuring distances to astronomical objects, one uses a kind of "ladder" of different methods; each method applies only for a limited distance, and each method which applies for a larger distance builds on the data of the preceding methods.

The starting point is knowing the distance from the Earth to the Sun; this distance is called one $astronomical\ unit\ (AU)$, and is roughly 150 million kilometers. Copernicus made the first, roughly accurate, solar system model, using data taken in ancient times, in his famous $De\ Revolutionibus\ (1543)$. Distances in the inner solar system are measured by bouncing radar signals off planets or asteroids, and measuring the time until the echo is received. Modern models are very accurate.

The next step in the ladder consists of simple geometrical methods; with them, one can go to a few hundred light-years. The distance to nearby stars can be determined by their parallaxes: using Earth's orbit as a baseline, the distances to stars are measured by triangulation. This is accurate to about 1% at 50 light-years, 10% at 500 light-years.

Using data acquired by the geometrical methods, and adding photometry (i.e., measurements of the brightness) and spectroscopy, one gets the next step in the ladder for stars so far away that their parallaxes are not measurable yet. As the brightness decreases proportionally to the square of the distance, if we know the absolute brightness of a star (i.e., its brightness at the standard reference distance 10 pc), and its apparent brightness (i.e., the actual brightness which we observe on the Earth) we can say how far away the star is. To define the absolute brightness, one can use the Hertzsprung–Russel diagram: stars of similar type have similar brightnesses; thus, if we know a star's type (from its color and/or spectrum), we can find its distance by comparing its apparent and absolute magnitudes; the latter derived from geometric parallaxes to nearby stars.

For even larger distances in the Universe, one needs an additional element: standard candles, i.e., several types of cosmological objects, for which one can determine their absolute brightness without knowing their distances. Primary standard candles are the Cepheid variable stars. They periodically change their size and temperature. There is a relationship between the brightness of these pulsating stars and the period of their oscillations, and this relationship can be used to determine their absolute brightness. Cepheids can be identified as far as in the Virgo Cluster (60 million light-years). Another type of standard candle (secondary standard candles) which is brighter than the Cepheids and, hence can be used to determine the distances to galaxies even hundreds of millions of light-years away, are supernovae and entire galaxies. However, Howell et al. (2007) suggested that the brightness of supernova explosions diminish over time on average, casting doubt on their use as accurate distance gauges (in particular, for dark energy measurements).

For really large distances (several hundreds of millions of light-years or even several billions of light-years), the cosmological redshift and the Hubble law are used. A complication is that it is not clear what is meant by "distance" here, and there are several types of distances used in Cosmology (luminosity distance, proper motion distance, angular diameter distance, etc.).

Depending on the situation, there is a large variety of special techniques to measure distances in Cosmology, such as **light echo distance**, **Bondi radar distance**, **RR Lyrae distance** and *secular*, *statistical*, *expansion*, *spectroscopic* parallax distances. For example, since the year 2000 NASA's Chandra X-ray Observatory measures the distance to a distant source via the delay of the halo of scattering material (interstellar dust grains) between it and the Earth.

26.2 Distances in Theory of Relativity

The Minkowski space-time (or Minkowski space, Lorentz space-time, flat space-time) is the usual geometric setting for the Einstein Special Theory of Relativity. In this setting the three ordinary dimensions of space are combined with a single dimension of time to form a four-dimensional space-time $\mathbb{R}^{1,3}$ in the absence of gravity.

Vectors in $\mathbb{R}^{1,3}$ are called four-vectors (or events). They can be written as (ct,x,y,z), where the first component is called the time-like component (c) is the speed of light, and (c) is the time) while the other three components are called spatial components. In spherical coordinates, they can be written as (ct,r,θ,ϕ) . In the Theory of Relativity, the spherical coordinates are a system of curvilinear coordinates (ct,r,θ,ϕ) , where (c) is the speed of light, (c) is the time, (c) is the radius from a point to the origin with (c) is the azimuthal angle in the (c) plane from the (c) axis with (c) is the polar angle from the (c) axis with (c)

$$||v||^2 = \langle v, v \rangle = c^2 t^2 - x^2 - y^2 - z^2.$$

They are said to be *time-like*, *space-like*, and *light-like* (*isotropic*) if their squared norms are positive, negative, or equal to zero, respectively.

The set of all light-like vectors forms the *light cone*. If the coordinate origin is singled out, the space can be broken up into three domains: domains of *absolute future* and *absolute past*, falling within the light cone, whose points are joined to the origin by time-like vectors with positive or negative value of time coordinate, respectively, and the domain of *absolute elsewhere*, falling outside of the light cone, whose points are joined to the origin by space-like vectors.

A world line of an object is the sequence of events that marks the time history of the object. A world line traces out the path of a single point in the Minkowski space. It is a one-dimensional curve, represented by the coordinates as a function of one parameter. A world line is a time-like curve in space—time, i.e., at any point its tangent vector is a time-like four-vector. All world lines fall within the light cone, formed by light-like curves, i.e., the curves whose tangent vectors are light-like four-vectors correspond to the motion of light and other particles of zero rest mass.

World lines of particles at constant speed (equivalently, of free falling particles) are called *geodesics*. In Minkowski space they are straight lines.

A geodesic in Minkowski space, which joins two given events x and y, is the longest curve among all world lines which join these two events. This follows from the *Einstein time* (triangle) *inequality* (cf. **inverse triangle inequality** and, in Chap. 5, **reverse triangle inequality**):

$$||x + y|| \ge ||x|| + ||y||,$$

according to which a time-like broken line joining two events is shorter than the single time-like geodesic joining them, i.e., the proper time of the particle moving freely from x to y is greater than the proper time of any other particle whose world line joins these events. This fact is usually called $twin\ paradox$.

The *space-time* is a four-dimensional *manifold* which is the usual mathematical setting for the Einstein General Theory of Relativity. Here the three spatial components with a single time-like component form a four-dimensional space-time in the presence of gravity. Gravity is equivalent to the geometric properties of space-time, and in the presence of gravity the geometry of space-time is curved. Thus, the space-time is a four-dimensional curved manifold for which the tangent space to any point is the Minkowski space, i.e., it is a *pseudo-Riemannian manifold* – a manifold, equipped with a non-degenerate indefinite metric, called **pseudo-Riemannian metric** – of signature (1,3).

In the General Theory of Relativity, gravity is described by the properties of the local geometry of space—time. In particular, the gravitational field can be built out of a **metric tensor**, a quantity describing geometrical properties space—time such as distance, area, and angle. Matter is described by its *stressenergy tensor*, a quantity which contains the density and pressure of matter. The strength of coupling between matter and gravity is determined by the *gravitational constant*.

The Einstein field equation is an equation in the General Theory of Relativity, that describes how matter creates gravity and, conversely, how gravity affects matter. A solution of the Einstein field equation is a certain **Einstein metric** appropriated for the given mass and pressure distribution of the matter.

A *black hole* is a massive astrophysical object that is theorized to be created from the collapse of a neutron or "quark" star. The gravitational forces are so

strong in a black hole that they overcome neutron degeneracy pressure and, roughly, collapse to a point (called a *singularity*). Even light cannot escape the gravitational pull of a black hole within the black hole's **gravitational radius** (or *event horizon*). Uncharged, zero angular momentum black holes are called *Schwarzschild black holes*. Uncharged non-zero angular momentum black holes are called *Kerr black holes*. Non-spinning charged black holes are called *Reissner-Nordström black holes*. Charged, spinning black holes are called *Kerr-Newman black holes*. Corresponding metrics describe how spacetime is curved by matter in the presence of these black holes.

For an additional information see, for example, [Wein72].

• Minkowski metric

The Minkowski metric is a pseudo-Riemannian metric, defined on the *Minkowski space* $\mathbb{R}^{1,3}$, i.e., a four-dimensional real vector space which is considered as the *pseudo-Euclidean space* of signature (1,3). It is defined by its metric tensor

$$((g_{ij})) = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}.$$

The line element ds^2 , and the space-time interval element ds of this metric are given by

$$ds^2 = c^2 dt^2 - dx^2 - dy^2 - dz^2.$$

In spherical coordinates (ct, r, θ, ϕ) , one has $ds^2 = c^2 dt^2 - dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$.

The pseudo-Euclidean space $\mathbb{R}^{3,1}$ of signature (3,1) with the line element

$$ds^2 = -c^2 dt^2 + dx^2 + dy^2 + dz^2$$

can also be used as a space—time model of the Einstein Special Theory of Relativity. Usually, signature (1,3) is used in Particle Physics, whereas signature (3,1) is used in Relativity Theory.

• Affine space—time distance

Given a space-time (M^4,g) , there is a unique affine parametrization $s \to \gamma(s)$ for each light ray (i.e., light-like geodesic) through the observation event p_{obser} , such that $\gamma(0) = p_{obser}$ and $g(\frac{d\gamma}{dt}, U_{obser}) = 1$, where U_{obser} is the 4-velocity of the observer at p_{obser} (i.e., a vector with $g(U_{obser}, U_{obser}) = -1$).

In this case, the **affine space—time distance** is the *affine parameter* s, viewed as a distance measure.

The affine space—time distance is monotone increasing along each ray, and it coincides, in an infinitesimal neighborhood of p_{obser} , with the Euclidean distance in the rest system of U_{obser} .

• Lorentz metric

A Lorentz metric (or Lorentzian metric) is a pseudo-Riemannian metric (i.e., non-degenerate indefinite metric) of signature (1, p).

The curved space—time of the General Theory of Relativity can be modeled as a *Lorentzian manifold* (a manifold equipped with a Lorentz metric) of signature (1,3). The *Minkowski space* $\mathbb{R}^{1,3}$ with the flat **Minkowski metric** is a model of a flat Lorentzian manifold.

In Lorentzian Geometry the following definition of distance is commonly used. Given a rectifiable non-space-like curve $\gamma:[0,1]\to M$ in the spacetime M, the length of the curve is defined as $l(\gamma)=\int_0^1\sqrt{-\langle \frac{d\gamma}{dt},\frac{d\gamma}{dt}\rangle}dt$. For a space-like curve we set $l(\gamma)=0$. Then the **canonic Lorentz distance** between two points $p,q\in M$ is defined as

$$\sup_{\gamma \in \Gamma} l(\gamma)$$

if $p \prec q$, i.e., if the set Γ of future directed non-space-like curves from p to q is non-empty; otherwise, this distance is 0.

For two points x, y of space—time at geodesic distance d(x, y), their world function is $\pm \frac{1}{2}d^2(x, y)$, where the sign depends on whether x and y are or are not, respectively, causally related, i.e., can be joined by a time-like or null path.

• Kinematic metric

Given a set X, a kinematic metric (or time-like metric, abstract Lorenzian distance) is a function $\tau: X \times X \to \mathbb{R}_{\geq 0} \cup \{\infty\}$ such that, for all $x, y, z \in X$:

- 1. $\tau(x,x) = 0$
- 2. $\tau(x,y) > 0$ implies $\tau(y,x) = 0$ (anti-symmetry)
- 3. $\tau(x,y), \tau(y,z) > 0$ implies $\tau(x,z) > \tau(x,y) + \tau(y,z)$ (inverse triangle inequality)

The space-time set X consists of events $x=(x_0,x_1)$ where, usually, $x_0 \in \mathbb{R}$ is the time and $x_1 \in \mathbb{R}^3$ is the spatial location of the event x. The inequality $\tau(x,y) > 0$ means causality, i.e., x can influence y; usually, it is equivalent to $y_0 > x_0$ and the value $\tau(x,y) > 0$ can be seen as the largest (since it depends on the speed) proper (i.e., subjective) time of moving from x to y.

If the gravity is negligible, then $\tau(x,y) > 0$ implies $y_0 - x_0 \ge ||y_1 - x_1||_2$, and $\tau_p(x,y) = ((y_0 - x_0)^p - ||y_1 - x_1||_2^p))^{\frac{1}{p}}$ (as defined by Busemann in 1967) is a real number. For $p \approx 2$ it is consistent with Special Relativity observations.

A kinematic metric is not our usual distance metric; also it is not related to the **kinematic distance** in Astronomy.

• Lorentz-Minkowski distance

The **Lorentz–Minkowski distance** is a distance on \mathbb{R}^n (or on \mathbb{C}^n), defined by

$$\sqrt{|x_1 - y_1|^2 - \sum_{i=2}^{n} |x_i - y_i|^2}.$$

• Galilean distance

The **Galilean distance** is a distance on \mathbb{R}^n , defined by

$$|x_1 - y_1|$$

if $x_1 \neq y_1$, and by

$$\sqrt{(x_2-y_2)^2+\cdots+(x_n-y_n)^2}$$

if $x_1 = y_1$. The space \mathbb{R}^n equipped with the Galilean distance is called Galilean space. For n=4, it is a mathematical setting for the space–time of classical mechanics according to Galilei–Newton in which the distance between two events taking place at the points p and q at the moments of time t_1 and t_2 is defined as the time interval $|t_1 - t_2|$, while if these events take place at the same time, it is defined as the distance between the points p and q.

• Einstein metric

In the General Theory of Relativity, describing how space—time is curved by matter, the **Einstein metric** is a solution to the *Einstein field equation*

$$R_{ij} - \frac{g_{ij}R}{2} + \Lambda g_{ij} = \frac{8\pi G}{c^4} T_{ij},$$

i.e., a **metric tensor** $((g_{ij}))$ of signature (1,3), appropriated for the given mass and pressure distribution of the matter. Here $E_{ij} = R_{ij} - \frac{g_{ij}R}{2} + \Lambda g_{ij}$ is the *Einstein curvature tensor*, R_{ij} is the *Ricci curvature tensor*, R is the *Ricci scalar*, Λ is the *cosmological constant*, G is the *gravitational constant*, and T_{ij} is a *stress-energy tensor*. Empty space (vacuum) corresponds to the case of zero Ricci tensor: $R_{ij} = 0$.

The static Einstein metric for a homogeneous and isotropic Universe is given by the *line element*

$$ds^{2} = -dt^{2} + \frac{dr^{2}}{(1 - kr^{2})} + r^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2}),$$

where k is the curvature of the space–time, and the *scale factor* is equal to 1.

• de Sitter metric

The **de Sitter metric** is a maximally symmetric vacuum solution to the *Einstein field equation* with a positive cosmological constant Λ , given by the *line element*

$$ds^{2} = dt^{2} + e^{2\sqrt{\frac{\Lambda}{3}}t}(dr^{2} + r^{2}d\theta^{2} + r^{2}\sin^{2}\theta d\phi^{2}).$$

Without a cosmological constant (i.e., with $\Lambda = 0$), the most symmetric solution to the Einstein field equation in a vacuum is the flat **Minkowski** metric.

The anti-de Sitter metric corresponds to a negative value of Λ .

• Schwarzschild metric

The **Schwarzschild metric** is a solution to the *Einstein field equation* for empty space (vacuum) around a spherically symmetric mass distribution; this metric gives a representation of the Universe around a black hole of a given mass, from which no energy can be extracted. It was found by Schwarzschild in 1916, only a few months after the publication of the Einstein field equation, and was the first exact solution of this equation.

The *line element* of this metric is given by

$$ds^{2} = \left(1 - \frac{r_{g}}{r}\right)c^{2}dt^{2} - \frac{1}{\left(1 - \frac{r_{g}}{r}\right)}dr^{2} - r^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2}),$$

where $r_g = \frac{2Gm}{c^2}$ is the Schwarzschild radius, m is the mass of the black hole, and G is the gravitational constant.

This solution is only valid for radii larger than r_g , as at $r=r_g$ there is a coordinate singularity. This problem can be removed by a transformation to a different choice of space–time coordinates, called Kruskal–Szekeres coordinates. As $r \to +\infty$, the Schwarzschild metric approaches the **Minkowski metric**.

• Kruskal-Szekeres metric

The **Kruskal–Szekeres metric** is a solution to the *Einstein field equation* for empty space (vacuum) around a static spherically symmetric mass distribution, given by the *line element*

$$ds^{2} = 4\frac{r_{g}}{r} \left(\frac{r_{g}}{R}\right)^{2} e^{-\frac{r}{r_{g}}} (c^{2}dt^{'2} - dr^{'2}) - r^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2}),$$

where $r_g = \frac{2Gm}{c^2}$ is the Schwarzschild radius, m is the mass of the black hole, G is the gravitational constant, R is a constant, and the Kruskal–Szekeres coordinates $(t^{'}, r^{'}, \theta, \phi)$ are obtained from the spherical coordinates (ct, r, θ, ϕ) by the Kruskal–Szekeres transformation $r^{'2} - ct^{'2} = R^2 \left(\frac{r}{r_g} - 1\right) e^{\frac{r}{r_g}}$, $\frac{ct^{'}}{r^{'}} = \tanh\left(\frac{ct}{2r_g}\right)$.

In fact, the Kruskal–Szekeres metric is the **Schwarzschild metric**, written in Kruskal–Szekeres coordinates. It shows that the singularity of the space–time in the Schwarzschild metric at the Schwarzschild radius r_g is not a real physical singularity.

• Kottler metric

The **Kottler metric** is the unique spherically symmetric vacuum solution to the *Einstein field equation* with a cosmological constant Λ . It is given by the *line element*

$$ds^{2} = -\left(1 - \frac{2m}{r} - \frac{\Lambda r^{2}}{3}\right)dt^{2} + \left(1 - \frac{2m}{r} - \frac{\Lambda r^{2}}{3}\right)^{-1}dr^{2} + r^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2}).$$

It is called also the Schwarzschild–de Sitter metric for $\Lambda > 0$, and the Schwarzschild–anti-de Sitter metric for $\Lambda < 0$.

• Reissner-Nordström metric

The **Reissner–Nordström metric** is a solution to the *Einstein field equation* for empty space (vacuum) around a spherically symmetric mass distribution in the presence of a charge; this metric gives a representation of the Universe around a charged black hole.

The *line element* of this metric is given by

$$ds^{2} = \left(1 - \frac{2m}{r} + \frac{e^{2}}{r^{2}}\right)dt^{2} - \left(1 - \frac{2m}{r} + \frac{e^{2}}{r^{2}}\right)^{-1}dr^{2} - r^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2}),$$

where m is the mass of the hole, e is the charge (e < m), and we have used units with the speed of light c and the gravitational constant G equal to one.

• Kerr metric

The **Kerr metric** (or **Kerr–Schild metric**) is an exact solution to the *Einstein field equation* for empty space (vacuum) around an axially symmetric, rotating mass distribution; this metric gives a representation of the Universe around a rotating black hole.

Its line element is given (in Boyer-Lindquist form) by

$$ds^{2} = \rho^{2} \left(\frac{dr^{2}}{\Delta} + d\theta^{2} \right) + (r^{2} + a^{2}) \sin^{2} \theta d\phi^{2} - dt^{2} + \frac{2mr}{\rho^{2}} (a \sin^{2} \theta d\phi - dt)^{2},$$

where $\rho^2 = r^2 + a^2 \cos^2 \theta$ and $\Delta = r^2 - 2mr + a^2$. Here m is the mass of the black hole and a is the angular velocity as measured by a distant observer.

The generalization of the Kerr metric for a charged black hole is known as the **Kerr–Newman metric**. When a=0, the Kerr metric becomes the **Schwarzschild metric**. A black hole can be diagnosed as rotating if radiation processes are observed inside its Scwarzchild radius but outside its Kerr radius.

• Kerr-Newman metric

The **Kerr–Newman metric** is an exact, unique and complete solution to the *Einstein field equation* for empty space (vacuum) around an axially symmetric, rotating mass distribution in the presence of a charge; this metric gives a representation of the Universe around a rotating charged black hole.

The *line element* of the exterior metric is given by

$$ds^{2} = -\frac{\Delta}{\rho^{2}}(dt - a\sin^{2}\theta d\phi)^{2} + \frac{\sin^{2}\theta}{\rho^{2}}((r^{2} + a^{2})d\phi - adt)^{2} + \frac{\rho^{2}}{\Delta}dr^{2} + \rho^{2}d\theta^{2},$$

where $\rho^2 = r^2 + a^2 \cos^2 \theta$ and $\Delta = r^2 - 2mr + a^2 + e^2$. Here m is the mass of the black hole, e is the charge, and a is the angular velocity. When e = 0, the Kerr–Newman metric becomes the **Kerr metric**.

• Static isotropic metric

The **static isotropic metric** is the most general solution to the *Einstein field equation* for empty space (vacuum); this metric can represent a static isotropic gravitational field. The *line element* of this metric is given by

$$ds^{2} = B(r)dt^{2} - A(r)dr^{2} - r^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2}),$$

where B(r) and A(r) are arbitrary functions.

• Eddington-Robertson metric

The **Eddington–Robertson metric** is a generalization of the **Schwarzschild metric** which allows that the mass m, the gravitational constant G, and the density ρ are altered by unknown dimensionless parameters α , β , and γ (all equal to 1 in the Einstein field equation).

The *line element* of this metric is given by

$$ds^{2} = \left(1 - 2\alpha \frac{mG}{r} + 2(\beta - \alpha\gamma) \left(\frac{mG}{r}\right)^{2} + \dots\right) dt^{2}$$
$$-\left(1 + 2\gamma \frac{mG}{r} + \dots\right) dr^{2} - r^{2} (d\theta^{2} + \sin^{2}\theta d\phi^{2}).$$

• Janis-Newman-Wincour metric

The Janis-Newman-Wincour metric is the most general spherically symmetric static and asymptotically flat solution to the *Einstein field equation* coupled to a massless scalar field. It is given by the *line element*

$$ds^{2} = -\left(1 - \frac{2m}{\gamma r}\right)^{\gamma} dt^{2} + \left(1 - \frac{2m}{\gamma r}\right)^{-\gamma} dr^{2} + \left(1 - \frac{2m}{\gamma r}\right)^{1-\gamma} r^{2} (d\theta^{2} + \sin^{2}\theta d\phi^{2}),$$

where m and γ are constants. For $\gamma = 1$ one obtains the **Schwarzschild** metric. In this case the scalar field vanishes.

• FLRW metric

The **FLRW** metric (or **Friedmann–Lemaître–Robertson–Walker** metric) is a exact solution to the *Einstein field equation* for a simply connected, homogeneous, isotropic expanding (or contracting) Universe filled with a constant density and negligible pressure. This metric gives a representation of a matter-dominated Universe filled with a pressureless dust. The FLRW metric models the metric expansion of space, i.e., the averaged increase of measured distance (an intrinsic expansion) between objects in the Universe with time.

The *line element* of this metric is usually written in the *spherical coordinates* (ct, r, θ, ϕ) :

$$ds^{2} = c^{2}dt^{2} - a(t)^{2} \cdot \left(\frac{dr^{2}}{1 - kr^{2}} + r^{2} \cdot (d\theta^{2} + \sin^{2}\theta d\phi^{2})\right),$$

where a(t) is the *scale factor* and k is the *curvature* of the space–time. There is also another form for the *line element*:

$$ds^{2} = c^{2}dt^{2} - a(t)^{2} \cdot (dr'^{2} + \tilde{r}^{2} \cdot (d\theta^{2} + \sin^{2}\theta d\phi^{2})),$$

where r' gives the **comoving distance** from the observer, and \tilde{r} gives the **proper motion distance**, i.e., $\tilde{r} = R_C \sinh(r'/R_C)$, or r', or $R_C \sin(r'/R_C)$ for negative, zero or positive curvature, respectively, where $R_C = 1/\sqrt{|k|}$ is the absolute value of the radius of curvature.

• Bianchi metrics

The **Bianchi metrics** are solutions to the Einstein field equation for cosmological models that have spatially homogeneous sections, invariant under the action of a three-dimensional Lie group, i.e., they are real four-dimensional metrics with a three-dimensional isometry group, transitive on 3-surfaces. Using the Bianchi classification of three-dimensional Lie algebras over Killing vector fields, we obtain the nine types of Bianchi metrics.

Each Bianchi model B defines a transitive group G_B on some threedimensional simply connected manifold M; so, the pair (M, G) (where G is the maximal group acting on X and containing G_B) is one of eight Thurston model geometries if M/G' is compact for a discrete subgroup G' of G. In particular, Bianchi type IX corresponds to the model geometry S^3 .

The Bianchi type I metric is a solution to the Einstein field equation for an anisotropic homogeneous Universe, given by the *line element*

$$ds^2 = -dt^2 + a(t)^2 dx^2 + b(t)^2 dy^2 + c(t)^2 dz^2,$$

where the functions a(t), b(t), and c(t) are determined by the Einstein equation. It corresponds to flat spatial sections, i.e., is a generalization of the **FLRW metric**.

The Bianchi type IX metric, or **Mixmaster metric**, exhibits complicated dynamic behavior near its curvature singularities.

• Kasner metric

The **Kasner metric** is one of the Bianchi type I metrics, which is a vacuum solution to the Einstein field equation for an anisotropic homogeneous Universe, given by the *line element*

$$ds^{2} = -dt^{2} + t^{2p_{1}}dx^{2} + t^{2p_{2}}dy^{2} + t^{2p_{3}}dz^{2},$$

where $p_1 + p_2 + p_3 = p_1^2 + p_2^2 + p_3^2 = 1$.

The Kasner metric can also be written as

$$ds^{2} = -dt^{2} + t^{\frac{2}{3}} \left(t^{\frac{4}{3}\cos(\phi + \frac{\pi}{3})} dx^{2} + t^{\frac{4}{3}\cos(\phi - \frac{\pi}{3})} dy^{2} + t^{-\frac{4}{3}\cos\phi} dz^{2} \right),$$

and is called in this case the Kasner circle.

One of the Kasner metrics, the *Kasner-like metric*, is given by the *line* element

$$ds^{2} = -dt^{2} + t^{2q}(dx^{2} + dy^{2}) + t^{2-4q}dz^{2}.$$

The Kasner axisymmetric metric is given by the line element

$$ds^{2} = -\frac{dt^{2}}{\sqrt{t}} + \frac{dx^{2}}{\sqrt{t}} + tdy^{2} + tdz^{2}.$$

• Kantowski-Sashs metric

The **Kantowski–Sashs metric** is a solution to the Einstein field equation, given by the *line element*

$$ds^{2} = -dt^{2} + a(t)^{2}dz^{2} + b(t)^{2}(d\theta^{2} + \sin\theta d\phi^{2}),$$

where the functions a(t) and b(t) are determined by the Einstein equation. It is the only homogeneous model without a three-dimensional transitive subgroup.

In particular, the Kantowski–Sashs metric with the line element

$$ds^{2} = -dt^{2} + e^{2\sqrt{\Lambda}t}dz^{2} + \frac{1}{\Lambda}(d\theta^{2} + \sin^{2}\theta d\phi^{2})$$

describes a Universe with two spherical dimensions having a fixed size during the cosmic evolution, and the third dimension is expanding exponentially.

• GCSS metric

A GCSS (i.e., general cylindrically symmetric stationary) metric is a solution to the *Einstein field equation*, given by the *line element*

$$ds^{2} = -fdt^{2} + 2kdtd\phi + e^{\mu}(dr^{2} + dz^{2}) + ld\phi^{2},$$

where the space-time is divided into two regions: the interior, with $0 \le r \le R$, to a cylindrical surface of radius R centered along z, and the exterior, with $R \le r < \infty$. Here f, k, μ and l are functions only of r, and $-\infty < t, z < \infty, \ 0 \le \phi \le 2\pi$; the hypersurfaces $\phi = 0$ and $\phi = 2\pi$ are identical.

• Lewis metric

The Lewis metric is a cylindrically symmetric stationary metric which is a solution to the *Einstein field equation* for empty space (vacuum) in the exterior of a cylindrical surface. The *line element* of this metric has the form

$$ds^{2} = -fdt^{2} + 2kdtd\phi - e^{\mu}(dr^{2} + dz^{2}) + ld\phi^{2},$$

where $f=ar^{-n+1}-\frac{c^2}{n^2a}r^{n+1}$, k=-Af, $l=\frac{r^2}{f}-A^2f$, $e^{\mu}=r^{\frac{1}{2}(n^2-1)}$ with $A=\frac{cr^{n+1}}{naf}+b$. The constants n,a,b, and c can be either real or complex, the corresponding solutions belong to the Weyl class or Lewis class, respectively. In the last case, the metric coefficients become $f=r(a_1^2-b_1^2)\cos(m\ln r)+2ra_1b_1\sin(m\ln r)$, $k=-r(a_1a_2-b_1b_2)\cos(m\ln r)-r(a_1b_2+a_2b_1)\sin(m\ln r)$, $l=-r(a_2^2-b_2^2)\cos(m\ln r)-2ra_2b_2\sin(m\ln r)$, $e^{\mu}=r^{-\frac{1}{2}(m^2+1)}$, where m,a_1,a_2,b_1 , and b_2 are real constants with $a_1b_2-a_2b_1=1$. Such metrics form a subclass of the Kasner type metrics.

• van Stockum metric

The **van Stockum metric** is a stationary cylindrically symmetric solution to the *Einstein field equation* for empty space (vacuum) with a rigidly rotating infinitely long dust cylinder. The *line element* of this metric for the interior of the cylinder is given (in comoving, i.e., corotating coordinates) by

$$ds^{2} = -dt^{2} + 2ar^{2}dtd\phi + e^{-a^{2}r^{2}}(dr^{2} + dz^{2}) + r^{2}(1 - a^{2}r^{2})d\phi^{2},$$

where $0 \le r \le R$, R is the radius of the cylinder, and a is the angular velocity of the dust particles. There are three vacuum exterior solutions (i.e., **Lewis metrics**) that can be matched to the interior solution, depending on the mass per unit length of the interior (the *low mass case*, the *null case*, and the *ultrarelativistic case*). Under some conditions (for example, if ar > 1), the existence of *closed time-like curves* (and, hence, time-travel) is allowed.

• Levi-Civita metric

The **Levi–Civita metric** is a static cylindrically symmetric vacuum solution to the *Einstein field equation*, with the *line element*, given (in the Weyl form) by

$$ds^{2} = -r^{4\sigma}dt^{2} + r^{4\sigma(2\sigma-1)}(dr^{2} + dz^{2}) + C^{-2}r^{2-4\sigma}d\phi,$$

where the constant C refers to the deficit angle, and the parameter σ is mostly understood in accordance with the Newtonian analogy of the Levi–Civita solution – the gravitational field of an infinite uniform line-mass (infinite wire) with the linear mass density σ . In the case $\sigma = -\frac{1}{2}$, C = 1 this metric can be transformed either into the Taub's plane symmetric metric, or into the Robinson-Trautman metric.

• Weyl-Papapetrou metric

The Weyl-Papapetrou metric is a stationary axially symmetric solution to the *Einstein field equation*, given by the *line element*

$$ds^{2} = Fdt^{2} - e^{\mu}(dz^{2} + dr^{2}) - Ld\phi^{2} - 2Kd\phi dt,$$

where F, K, L and μ are functions only of r and $z, LF + K^2 = r^2, \infty < t, z < \infty, 0 \le r < \infty$, and $0 \le \phi \le 2\pi$; the hypersurfaces $\phi = 0$ and $\phi - 2\pi$ are identical.

• Bonnor dust metric

The **Bonnor dust metric** is a solution to the *Einstein field equation*, which is an axially symmetric metric describing a cloud of rigidly rotating dust particles moving along circular geodesics about the z-axis in hypersurfaces of z = constant. The *line element* of this metric is given by

$$ds^{2} = dt^{2} + (r^{2} - n^{2})d\phi^{2} + 2ndtd\phi + e^{\mu}(dr^{2} + dz^{2}),$$

where, in Bonnor comoving (i.e., corotating) coordinates, $n=\frac{2hr^2}{R^3}$, $\mu=\frac{h^2r^2(r^2-8z^2)}{2R^8}$, $R^2=r^2+z^2$, and h is a rotation parameter. As $R\to\infty$, the metric coefficients tend to Minkowski values.

• Weyl metric

The **Weyl metric** is a general static axially symmetric vacuum solution to the *Einstein field equation* given, in Weyl canonical coordinates, by the *line element*

$$ds^{2} = e^{2\lambda}dt^{2} - e^{-2\lambda}\left(e^{2\mu}(dr^{2} + dz^{2}) + r^{2}d\phi^{2}\right),$$

where λ and μ are functions only of r and z such that $\frac{\partial^2 \lambda}{\partial r^2} + \frac{1}{r} \cdot \frac{\partial \lambda}{\partial r} + \frac{\partial^2 \lambda}{\partial z^2} = 0$, $\frac{\partial \mu}{\partial r} = r(\frac{\partial \lambda^2}{\partial r} - \frac{\partial \lambda^2}{\partial z})$, and $\frac{\partial \mu}{\partial z} = 2r\frac{\partial \lambda}{\partial r}\frac{\partial \lambda}{\partial z}$.

• Zipoy-Voorhees metric

The Zipoy–Voorhees metric (or γ -metric) is a Weyl metric, obtained for $e^{2\lambda}=\left(\frac{R_1+R_2-2m}{R_1+R_2+2m}\right)^{\gamma}, e^{2\mu}=\left(\frac{(R_1+R_2+2m)(R_1+R_2-2m)}{4R_1R_2}\right)^{\gamma^2}$, where $R_1^2=r^2+(z-m)^2, R_2^2=r^2+(z+m)^2$. Here λ corresponds to the Newtonian potential of a line segment of mass density $\gamma/2$ and length 2m, symmetrically distributed along the z-axis.

The case $\gamma = 1$ corresponds to the **Schwartzschild metric**, the cases $\gamma > 1$ ($\gamma < 1$) correspond to an oblate (prolate) spheroid, and for $\gamma = 0$ one obtains the flat Minkowski space—time.

• Straight spinning string metric

The straight spinning string metric is given by the line element

$$ds^{2} = -(dt - ad\phi)^{2} + dz^{2} + dr^{2} + k^{2}r^{2}d\phi^{2},$$

where a and k>0 are constants. It describes the space–time around a straight spinning string. The constant k is related to the string's mass-per-length μ by k=1–4 μ , and the constant a is a measure of the string's spin. For a=0 and k=1, one obtains the **Minkowski metric** in cylindrical coordinates.

• Tomimatsu-Sato metric

A **Tomimatsu–Sato metric** [ToSa73] is one of the metrics from an infinite family of spinning mass solutions to the *Einstein field equation*, each of which has the form $\xi = U/W$, where U and W are some polynomials. The simplest solution has $U = p^2(x^4 - 1) + q^2(y^4 - 1) - 2ipqxy(x^2 - y^2)$, $W = 2px(x^2 - 1) - 2iqy(1 - y^2)$, where $p^2 + q^2 = 1$. The *line element* for this solution is given by

$$ds^{2} = \Sigma^{-1} \left((\alpha dt + \beta d\phi)^{2} - r^{2} (\gamma dt + \delta d\phi)^{2} \right) - \frac{\Sigma}{p^{4} (x^{2} - y^{2})^{4}} (dz^{2} + dr^{2}),$$

where
$$\alpha = p^2(x^2 - 1)^2 + q^2(1 - y^2)^2$$
, $\beta = -\frac{2q}{p}W(p^2(x^2 - 1)(x^2 - y^2) + 2(px + 1)W)$, $\gamma = -2pq(x^2 - y^2)$, $\delta = \alpha + 4((x^2 - 1) + (x^2 + 1)(px + 1))$, $\Sigma = \alpha\delta - \beta\gamma = |U + W|^2$.

• Gödel metric

The **Gödel metric** is an exact solution to the *Einstein field equation* with cosmological constant for a rotating Universe, given by the *line element*

$$ds^{2} = -(dt^{2} + C(r)d\phi)^{2} + D^{2}(r)d\phi^{2} + dr^{2} + dz^{2},$$

where (t,r,ϕ,z) are the usual cylindrical coordinates. The Gödel Universe is homogeneous if $C(r)=\frac{4\Omega}{m^2}\sinh^2\left(\frac{mr}{2}\right),\ D(r)=\frac{1}{m}\sinh(mr)$, where m and Ω are constants. The Gödel Universe is singularity-free. There are closed time-like curves through every event, and hence time-travel is possible here. The condition required to avoid such curves is $m^2>4\Omega^2$.

• Conformally stationary metric

The **conformally stationary metrics** are models for gravitational fields that are time-independent up to an overall conformal factor. If some global regularity conditions are satisfied, the space–time must be a product $\mathbb{R} \times M^3$ with a (Hausdorff and paracompact) 3-manifold M^3 , and the line element of the metric is given by

$$ds^{2} = e^{2f(t,x)} \left(-(dt + \sum_{\mu} \phi_{\mu}(x)dx_{\mu})^{2} + \sum_{\mu,\nu} g_{\mu\nu}(x)dx_{\mu}dx_{\nu} \right),$$

where $\mu, \nu = 1, 2, 3$. The conformal factor e^{2f} does not affect the light-like geodesics apart from their parametrization, i.e., the paths of light rays are completely determined by the Riemannian metric $g = \sum_{\mu,\nu} g_{\mu\nu}(x) dx_{\mu} dx_{\nu}$ and the one-form $\phi = \sum_{\mu} \phi_{\mu}(x) dx_{\mu}$ which both live on M^3 .

In this case, the function f is called the *redshift potential*, the metric g is called the **Fermat metric**, and the one-form ϕ is called the *Fermat one-form*.

For a static space–time, the geodesics in the Fermat metric are the projections of the null geodesics of space–time.

In particular, the **spherically symmetric and static metrics**, including models for non-rotating stars and black holes, wormholes, monopoles, naked singularities, and (boson or fermion) stars, are given by the *line element*

$$ds^2 = e^{2f(r)} (-dt^2 + S(r)^2 dr^2 + R(r)^2 (d\theta^2 + \sin^2 \theta d\phi^2)).$$

Here, the one-form ϕ vanishes, and the Fermat metric g has the special form

$$g = S(r)^2 dr^2 + R(r)^2 (d\theta^2 + \sin^2 \theta d\phi^2).$$

For example, the conformal factor $e^{2f(r)}$ of the **Schwartzschild metric** is equal to $1 - \frac{2m}{r}$, and the corresponding Fermat metric has the form

$$g = (1 - \frac{2m}{r})^{-2} (1 - \frac{2m}{r})^{-1} r^2 (d\theta^2 + \sin\theta d\phi^2).$$

pp-wave metric

The **pp-wave metric** is an exact solution to the Einstein field equation, in which radiation moves at the speed of light. The *line element* of this metric is given (in Brinkmann coordinates) by

$$ds^{2} = H(u, x, y)du^{2} + 2dudv + dx^{2} + dy^{2},$$

where H is any smooth function. The term "pp" stands for *plane-fronted* waves with parallel propagation introduced by Ehlers and Kundt in 1962.

The most important class of particularly symmetric pp-waves are the **plane wave metrics**, in which H is quadratic. The wave of death, for example, is a gravitational (i.e., the space-time curvature fluctuates) plane wave exhibiting a strong nonscalar null curvature singularity, which propagates through an initially flat space-time, progressively destroying the Universe.

Examples of axisymmetric pp-waves include the *Aichelburg–Sexl ultra-boost* which models the physical experience of an observer moving past a spherically symmetric gravitating object at nearly the speed of light, and the *Bonnor beam* which models the gravitational field of an infinitely long beam of incoherent electromagnetic radiation. The Aichelburg–Sexl wave is obtained by boosting the Schwarzschild solution to the speed of light at fixed energy, i.e., it describes a Schwarzschild black hole moving at the speed of light. Cf. **Aichelburg–Sexl metric** in Chap. 24.

• Bonnor beam metric

The **Bonnor beam metric** is an exact solution to the Einstein field equation, which models an infinitely long, straight beam of light. It is an example of a **pp-wave metric**.

The interior part of the solution (in the uniform plane wave interior region, which is shaped like the world tube of a solid cylinder) is defined by the *line element*

$$ds^{2} = -8\pi mr^{2}du^{2} - 2dudv + dr^{2} + r^{2}d\theta^{2}.$$

where $-\infty < u, v < \infty$, $0 < r < r_0$, and $-\pi < \theta < \pi$. This is a null dust solution and can be interpreted as incoherent electromagnetic radiation.

The exterior part of the solution is defined by

$$ds^{2} = -8\pi m r_{0}^{2} (1 + 2\log(r/r_{0}))du^{2} - 2dudv + dr^{2} + r^{2}d\theta^{2},$$

where $-\infty < u, v < \infty, r_0 < r < \infty, \text{ and } -\pi < \theta < \pi.$

The Bonnor beam can be generalized to several parallel beams traveling in the same direction.

• Plane wave metric

The **plane wave metric** is a vacuum solution to the *Einstein field equation*, given by the *line element*

$$ds^{2} = 2dwdu + 2f(u)(x^{2} + y^{2})du^{2} - dx^{2} - dy^{2}.$$

It is conformally flat, and describes a pure radiation field. The space—time with this metric is called the *plane gravitational wave*. It is an example of a **pp-wave metric**.

• Wils metric

The **Wils metric** is a solution to the *Einstein field equation*, given by the *line element*

$$ds^{2} = 2xdwdu - 2wdudx + (2f(u)x(x^{2} + y^{2}) - w^{2})du^{2} - dx^{2} - dy^{2}.$$

It is conformally flat, and describes a pure radiation field which is not a plane wave.

• Koutras-McIntosh metric

The **Koutras–McIntosh metric** is a solution to the *Einstein field equation*, given by the *line element*

$$ds^{2} = 2(ax+b)dwdu - 2awdudx + (2f(u)(ax+b)(x^{2}+y^{2}) - a^{2}w^{2})$$
$$du^{2} - dx^{2} - dy^{2}.$$

It is conformally flat and describes a pure radiation field which, in general, is not a *plane wave*. It gives the **plane wave metric** for a = 0, b = 1, and the **Wils metric** for a = 1, b = 0.

• Edgar-Ludwig metric

The **Edgar–Ludwig metric** is a solution to the *Einstein field equation*, given by the *line element*

$$ds^2 = 2(ax+b)dwdu - 2awdudx +$$

+
$$(2f(u)(ax + b)(g(u)y + h(u) + x^2 + y^2) - a^2w^2) du^2 - dx^2 - dy^2$$
.

This metric is a generalization of the **Koutras–McIntosh metric**. It is the most general metric which describes a conformally flat pure radiation (or null fluid) field which, in general, is not a *plane wave*. If plane waves are excluded, it has the form

$$ds^{2} = 2xdwdu - 2wdudx + (2f(u)x(g(u)y + h(u) + x^{2} + y^{2}) - w^{2})$$
$$du^{2} - dx^{2} - du^{2}.$$

• Bondi radiating metric

The **Bondi radiating metric** describes the asymptotic form of a radiating solution to the *Einstein field equation*, given by the *line element*

$$ds^{2} = -\left(\frac{V}{r}e^{2\beta} - U^{2}r^{2}e^{2\gamma}\right)du^{2} - 2e^{2\beta}dudr - 2Ur^{2}e^{2\gamma}dud\theta + r^{2}(e^{2\gamma}d\theta^{2} + e^{-2\gamma}\sin^{2}\theta d\phi^{2}),$$

where u is the retarded time, r is the **luminosity distance**, $0 \le \theta \le \pi$, $0 \le \phi \le 2\pi$, and U, V, β, γ are functions of u, r, and θ .

• Taub-NUT de Sitter metric

The **Taub-NUT de Sitter metric** (cf. **de Sitter metric**) is a positive-definite (i.e., Riemannian) solution to the *Einstein field equation* with a cosmological constant Λ , given by the *line element*

$$ds^{2} = \frac{r^{2} - L^{2}}{4\Delta}dr^{2} + \frac{L^{2}\Delta}{r^{2} - L^{2}}(d\psi + \cos\theta d\phi)^{2} + \frac{r^{2} - L^{2}}{4}(d\theta^{2} + \sin^{2}\theta d\phi^{2}),$$

where $\Delta = r^2 - 2Mr + L^2 + \frac{\Lambda}{4}(L^4 + 2L^2r^2 - \frac{1}{3}r^4)$, L and M are parameters, and θ , ϕ , ψ are the *Euler angles*. If $\Lambda = 0$, one obtains the **Taub-NUT metric**, using some regularity conditions. NUT manifold was discovered in Ehlers (1957) and rediscovered in Newman, Tamburino, Unti (1963); it is closely related to the metric in Taub (1951).

• Eguchi–Hanson de Sitter metric

The Eguchi–Hanson de Sitter metric (cf. de Sitter metric) is a positive-definite (i.e., Riemannian) solution to the *Einstein field equation* with a cosmological constant Λ , given by the *line element*

$$ds^{2} = \left(1 - \frac{a^{4}}{r^{4}} - \frac{\Lambda r^{2}}{6}\right)^{-1} dr^{2} + \frac{r^{2}}{4} \left(1 - \frac{a^{4}}{r^{4}} - \frac{\Lambda r^{2}}{6}\right) (d\psi + \cos\theta d\phi)^{2} + \frac{r^{2}}{4} (d\theta^{2} + \sin^{2}\theta d\phi^{2}),$$

where a is a parameter, and θ , ϕ , ψ are the Euler angles. If $\Lambda = 0$, one obtains the **Eguchi–Hanson metric**.

• Barriola-Vilenkin monopole metric

The Barriola-Vilenkin monopole metric is given by the line element

$$ds^{2} = -dt^{2} + dr^{2} + k^{2}r^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2}),$$

with a constant k < 1. There is a deficit solid angle and a singularity at r = 0; the plane t = constant, $\theta = \frac{\pi}{2}$ has the geometry of a cone. This metric is an example of a conical singularity; it can be used as a model for monopoles that might exist in the Universe.

A magnetic monopole is a hypothetical isolated magnetic pole, "a magnet with only one pole." It has been theorized that such things might exist in the form of tiny particles similar to electrons or protons, formed from topological defects in a similar manner to cosmic strings, but no such particle has ever been found. Cf. **Gibbons–Manton metric** in Chap. 7.

• Bertotti-Robinson metric

The **Bertotti–Robinson metric** is a solution to the *Einstein field equation* in a Universe with a uniform magnetic field. The *line element* of this metric is

$$ds^2 = Q^2(-dt^2 + \sin^2t dw^2 + d\theta^2 + \sin^2\theta d\phi^2),$$

where Q is a constant, $t \in [0, \pi], w \in (-\infty, +\infty), \theta \in [0, \pi], \text{ and } \phi \in [0, 2\pi].$

• Morris-Thorne metric

The Morris-Thorne metric is a wormhole solution to the Einstein field equation with the line element

$$ds^{2} = e^{\frac{2\Phi(w)}{c^{2}}}c^{2}dt^{2} - dw^{2} - r(w)^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2}),$$

where $w \in [-\infty, +\infty]$, r is a function of w that reaches some minimal value above zero at some finite value of w, and $\Phi(w)$ is a gravitational potential allowed by the space–time geometry.

A wormhole is a hypothetical "tube" in space connecting widely separated positions in a Universe. All wormholes require exotic material with negative energy density in order to hold them open.

• Misner metric

The **Misner metric** is a metric, representing two black holes. Misner (1960) provided a prescription for writing a metric connecting a pair of black holes, instantaneously at rest, whose throats are connected by a wormhole. The *line element* of this metric has the form

$$ds^{2} = -dt^{2} + \psi^{4}(dx^{2} + dy^{2} + dz^{2}),$$

where the conformal factor ψ is given by

$$\psi = \sum_{n=-N}^{N} \frac{1}{\sinh(\mu_0 n)} \frac{1}{\sqrt{x^2 + y^2 + (z + \coth(\mu_0 n))^2}}.$$

The parameter μ_0 is a measure of the ratio of mass to separation of the throats (equivalently, a measure of the distance of a loop in the surface, passing through one throat and out of the other). The summation limit N tends to infinity.

The topology of the *Misner space-time* is that of a pair of asymptotically flat sheets connected by a number of Einstein–Rosen bridges. In the simplest case, the Misner space can be considered as a two-dimensional space with topology $\mathbb{R} \times S^1$ in which light progressively tilts as one moves forward in time, and has *closed time-like curves* after a certain point.

• Alcubierre metric

The Alcubierre metric (Alcubierre 1994) is a solution to the Einstein field equation, representing warp drive space-time in which the existence of closed time-like curves is allowed. What is violated in this case is only the relativistic principle that a space-traveler may move with any velocity up to, but not including or exceeding, the speed of light. The Alcubierre construction corresponds to a warp (i.e., faster than light) drive in that it causes space-time to contract in front of a spaceship bubble and expand behind, thus providing the spaceship with a velocity that can be much greater than the speed of light relative to distant objects, while the spaceship never locally travels faster than light.

The line element of this metric has the form

$$ds^{2} = -dt^{2} + (dx - vf(r)dt)^{2} + dy^{2} + dz^{2},$$

with $v = \frac{dx_s(t)}{dt}$ as the apparent velocity of the warp drive spaceship, and $x_s(t)$ as the trajectory of the spaceship along the coordinate x, the radial coordinate being defined by $r = ((x - x_s(t))^2 + y^2 + z^2)^{\frac{1}{2}}$, and f(r) an arbitrary function subject to the boundary conditions that f = 1 at r = 0 (the location of the spaceship), and f = 0 at infinity.

• Rotating C-metric

The **rotating** C-metric is a solution to the Einstein-Maxwell equations, describing two oppositely charged black holes, uniformly accelerating in opposite directions. The line element of this metric has the form

$$ds^{2} = A^{-2}(x+y)^{-2} \left(\frac{dy^{2}}{F(y)} + \frac{dx^{2}}{G(x)} + k^{-2}G(X)d\phi^{2} - k^{2}A^{2}F(y)dt^{2} \right),$$

where $F(y) = -1 + y^2 - 2mAy^3 + e^2A^2y^4$, $G(x) = 1 - x^2 - 2mAx^3 - e^2A^2x^4$, m, e, and A are parameters related to the mass, charge and acceleration of the black holes, and k is a constant fixed by regularity conditions.

This metric should not be confused with the C-metric from Chap. 11.

• Myers-Perry metric

The **Myers–Perry metric** describes a five-dimensional rotating black hole. Its *line element* is given by

$$ds^{2} = -dt^{2} + \frac{2m}{\rho^{2}}(dt - a\sin^{2}\theta d\phi - b\cos^{2}\theta d\psi)^{2} +$$
$$+\frac{\rho^{2}}{R^{2}}dr^{2} + \rho^{2}d\theta^{2} + (r^{2} + a^{2})\sin^{2}\theta d\phi^{2} + (r^{2} + b^{2})\cos^{2}\theta d\psi^{2},$$

where
$$\rho^2 = r^2 + a^2 \cos^2 \theta + b^2 \sin^2 \theta$$
, and $R^2 = \frac{(r^2 + a^2)(r^2 + b^2) - 2mr^2}{r^2}$.

• Kaluza-Klein metric

The Kaluza–Klein metric is a metric in the Kaluza–Klein model of five-dimensional (in general, multidimensional) space–time which seeks to unify classical gravity and electromagnetism.

Kaluza (1919) found that, if the Einstein theory of pure gravitation is extended to a five-dimensional space—time, the *Einstein field equations* can be split into an ordinary four-dimensional gravitation tensor field, plus an extra vector field which is equivalent to the Maxwell equation for the electromagnetic field, plus an extra scalar field (known as the "dilation") which is equivalent to the massless Klein—Gordon equation.

Klein (1926) assumed the fifth dimension to have circular topology, so that the fifth coordinate is periodic, and the extra dimension is curled up to an unobservable size. An alternative proposal is that the extra dimension is (extra dimensions are) extended, and the matter is trapped in a four-dimensional submanifold. This approach has properties similar to the four-dimensional – all dimensions are extended and equal at the beginning, and the signature has the form (p, 1).

In a model of a large extra dimension, the fifth-dimensional metric of a Universe can be written in Gaussian normal coordinates in the form

$$ds^{2} = -(dx_{5})^{2} + \lambda^{2}(x_{5}) \sum_{\alpha,\beta} \eta_{\alpha\beta} dx_{\alpha} dx_{\beta},$$

where $\eta_{\alpha\beta}$ is the four-dimensional **metric tensor**, and $\lambda^2(x_5)$ is an arbitrary function of the fifth coordinate.

• Prasad metric

A de Sitter Universe can be described as the sum of the external space and the internal space.

The internal space has a negative constant curvature $-\frac{1}{r^2}$ and can be characterized by the symmetry group $SO_{3,2}$. The **Prasad metric** of this space is given, in hyperspherical coordinates, by the *line element*

$$ds^{2} = r^{2} \cos^{2} t (d\chi^{2} + \sinh^{2} \chi (d\theta^{2} + \sin^{2} \theta d\phi^{2})) - r^{2} dt^{2}.$$

The value $\sin \chi$ is the so-called *a dimensional normalized radius* of the de Sitter Universe.

The external space has constant curvature $\frac{1}{R^2}$ and can be characterized by the symmetry group $SO_{4,1}$. Its metric has the *line element* of the form

$$ds^{2} = R^{2} \cosh^{2} t (d\chi^{2} + \sin^{2} \chi (d\theta^{2} + \sin^{2} \theta d\phi^{2})) - R^{2} dt^{2}.$$

• Ponce de León metric

The **Ponce de León metric** is a five-dimensional background metric, given by the *line element*

$$ds^{2} = l^{2}dt^{2} - (t/t_{0})^{2}pl^{\frac{2p}{p-1}}(dx^{2} + dy^{2} + dz^{2}) - \frac{t^{2}}{(p-1)^{2}}dl^{2},$$

where l is the fifth (space-like) coordinate. This metric represents a five-dimensional apparent vacuum, but it is not flat.

Part VII Real-world Distances

Chapter 27 Length Measures and Scales

The term *length* has many meanings: distance lengthwise, extent lengthwise, linear measure, span, reach, end, limit, etc.; for example, length of a train, a meeting, a book, a travel, a shirt, a vowel, a proof. The *length* of an object is the distance between its ends, its linear extent, while the *height* is the vertical extent, and *width* (or *breadth*) is the distance from one side to the other at right angles to the length. The *depth* is distance downward, distance inward, deepness, vertical extent, drop.

The ancient Greek mathematicians regarded all numbers as lengths (of straight line segments), areas or volumes.

In Engineering and Physics, "length" usually means "distance." **Unit distance** is a distance taken as a convenient unit of length in a given context.

In Mathematics, **length function** is a function $l: G \to \mathbb{R}_{\geq 0}$ on a group (G, +, 0) such that l(0) = 0 and l(g) = l(-g), $l(g + g') \leq l(g) + l(g')$ for all $g, g' \in G$.

But in this chapter we consider length only as a measure of physical distance. We give selected information on the most important length units and present, in length terms, a list of interesting physical objects.

27.1 Length scales

The main length measure systems are: Metric, Imperial (British and American), Japanese, Thai, Chinese Imperial, Old Russian, Ancient Roman, Ancient Greek, Biblical, Astronomical, Nautical, and Typographical.

There are many other specialized length scales; for example, to measure cloth, shoe size, gauges (such as interior diameters of shotguns, wires, jewelry rings), sizes for abrasive grit, sheet metal thickness, etc. Also, many units express relative or reciprocal distances. For example, the reciprocal of distance (say, the focal length of a lens, radius of curvature and the convergence of an optical beam) are measured in *diopters*, i.e., reciprocal meters m^{-1} . Also, the *hertz* Hz is the SI unit of frequency (inverse second s^{-1}).

• International Metric System

The International Metric System (or SI, short for Système International), also known as MKSA (meter-kilogram-second-ampere), is a modernized version of the metric system of units, established by an international treaty (the Treaty of the Meter from 20 May 1875), which provides a logical and interconnected framework for all measurements in science, industry and commerce. The system is built on a foundation consisting of the following seven SI base units, assumed to be mutually independent:

(1) Length: **meter** (m); it is equal to the distance traveled by light in a vacuum in 1/299,792,458 of a second; (2) time: *second* (s); (3) mass: *kilogram* (kg); (4) temperature: *kelvin* (K); (5) electric current: *ampere* (A); (6) luminous intensity: *candela* (cd); (7) amount of substance: *mole* (mol).

Originally, on 26 March 1791, the $m\grave{e}tre$ (French for meter) was defined as $\frac{1}{10,000,000}$ of the distance from the North Pole to the equator along the meridian that passes through Dunkirk in France and Barcelona in Spain. The name $m\grave{e}tre$ was derived from the Greek metron (measure). In 1799 the standard of $m\grave{e}tre$ became a meter-long platinum-iridium bar kept in Sèvres, a town outside Paris, for people to come and compare their rulers with. (The metric system, introduced in 1793, was so unpopular that Napoleon was forced to abandon it and France returned to the $m\grave{e}tre$ only in 1837.) In 1960, the **meter** was officially defined in terms of wavelength.

The initial metric unit of mass, the *gram*, was defined as the mass of one cubic centimeter of water at its temperature of maximum density. For capacity, the *litre* (liter) was defined as the volume of a cubic decimeter.

A metric meterstick is a rough rule of thumb for comprehending a metric unit in terms of everyday life; for example, 5 cm is the side of matchbox, and 1 km is about 10-min walk.

• Metrication

The **metrication** is an ongoing (especially, in US, UK and Caribbean countries) process of conversion to the **International Metric System** SI. Officially, only US, Liberia and Myanmar have not switched to SI. For example, US uses only miles for road distance signs (milestones). Altitudes in aviation are usually described in feet; in shipping, nautical miles and knots are used. Resolutions of output devices are frequently specified in *dpi* (dots per inch).

Hard metric means designing in the metric measures from the start and conformation, where appropriate, to internationally recognized sizes and designs.

Soft metric means multiplying an inch-pound number by a metric conversion factor and rounding it to an appropriate level of precision; so, the soft converted products do not change size. The *American Metric System* consists of converting traditional units to embrace the uniform base 10 method that the Metric System uses. Such SI-Imperial hybrid units, used in soft metrication, are, for example, *kiloyard* (914.4 m), *kilofoot* (304.8 m), *mil* or *milli-inch* (24.5 µm), and *min* or *microinch* (25.4 nm).

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Metric inch $(2.5 \,\mathrm{cm})$ approximating the inch $(2.54 \,\mathrm{cm})$ was used in some Soviet computers when building from American blueprints.

In athletics, races of 1,500 or 1,600 m are often called *metric miles*.

• Meter-related terms

We present this large family of non-mathematical terms by the following examples (besides the unit of length).

Meter, in Poetry, (or cadence) is a measure of rhythmic quality, the regular linguistic sound patterns of verse; hypermeter is a part of verse with an extra syllable. Metromania, in Psychiatry, is a mania for writing verses.

Meter, in Music, (or metre) is the regular rhythmic patterns of musical line, the division of a composition into parts of equal time, and the subdivision of them. Isometre is the use of pulse (unbroken series of periodically occurring short stimuli) without regular meter, polymetre is the use of two or more different meters simultaneously whereas multimetre is using them in succession.

Metrometer: in Medicine, an instrument measuring the size of the womb. The names of various measuring instruments contain *meter* at the end.

Metra, in Medicine, is a synonym of uterus. So, metropathy is any disease of the uterus (say, metritis (inflammation), metratonia (atony), metratrophia, metrofibroma), and metrocyte is the mother cell.

Metrography (or hysterography) has two meanings: (1) radiographic examination of the uterine cavity filled with a contrasting medium and (2) a graphic procedure used to record uterine contractions. But also, metrograph is an instrument attached to a locomotive for recording its speed and the number and duration of its stops.

Metering: an equivalent term for a measuring; *micrometry*: measurement under the microscope.

Metric, as a non-mathematical term, is a standard unit of measure (for example, font metrics refer to numeric values relating to size and space in the font) or, more generally, part of a system of parameters; cf. quality metrics in Chap. 29.

Metrology: the science of, or a system of, weights and measures.

Metronomy: measurement of time by an instrument.

Metrosophy: a cosmology based on a strict number correspondences.

Telemetry: technology that allows remote measurement; archeometry: the science of exact measuring referring to the remote past; psychometry: alleged psychic power enabling one to divine facts by handling objects, and so on.

Psychometrics: the study concerned with the theory and technique of psychological measurement; psychrometrics: the field of engineering concerned with the determination of physical and thermodynamic properties of gas—vapor mixtures; biometrics: the study of automated methods for uniquely recognizing humans based upon one or more intrinsic physical or behavioral traits, and so on.

Isometropia: equality of refraction in both eyes.

Isometric exercise: a muscle strength training exercise in which a force is applied to a resistant object (cf. **isometric muscle action** in Chap. 29).

Isometric particle: a virus which (at the stage of virion capsid) has icosahedral symmetry.

Isometric process: a thermodynamic process at constant volume.

Isometric projection (or isometric view): the representation of 3D objects in 2D in which the angles between three projection axes are the same, or $\frac{2\pi}{3}$.

Isometric action RPG: a role-playing computer game played from a " $\frac{3}{4}$ -view," i.e., third-person isometric view when the camera is set above and away from the main characters, to view the action (usually real-time combat) from an angled, top-down perspective.

Isometric crystal system: the cubic crystal lattice; cf. crystal metric symmetry in Chap. 24. Metrohedry: overlap in 3D of the lattices of twin domains in a crystal.

Metrohedry: overlap in 3D of the lattices of twin domains in a crystal.

Metrio: the Greek coffee with one teaspoon of sugar (medium sweet). In Anthropology, *metriocranic* means having a skull that is moderately high compared with its width, with a breadth-height index 92–98.

Metroid is the name of a series of video games produced by Nintendo (10 games released in 1986–2007); metroids are a fictional species of parasitic alien creatures from those games. Metric is a Canadian New Wave rock band.

Examples of companies with a meter-related term in their name are: Metron, Metric Inc., Metric Engineering, World Wide Metric.

• Metric length measures

```
\begin{array}{l} \textit{kilometer} \; (\text{km}) \; = 1,\!000 \, \text{meters} \; = 10^3 \, \text{m}; \\ \textit{meter} \; (\text{m}) \; = 10 \, \text{decimeters} \; = 10^0 \, \text{m}; \\ \textit{decimeter} \; (\text{dm}) \; = 10 \, \text{centimeters} \; = 10^{-1} \, \text{m}; \\ \textit{centimeter} \; (\text{cm}) \; = 10 \, \text{millimeters} \; = 10^{-2} \, \text{m}; \\ \textit{millimeter} \; (\text{mm}) \; = 1,\!000 \, \text{micrometers} \; = 10^{-3} \, \text{m}; \\ \textit{micrometer} \; (\text{or micron}, \; \mu) \; = 1,\!000 \, \text{nanometers} = 10^{-6} \, \text{m}; \\ \textit{nanometer} \; (\text{nm}) \; = 10 \, \text{angstr\"oms} = 10^{-9} \, \text{m}. \end{array}
```

The lengths 10^{3t} m, $t=-8,-7,\ldots,-1,1,\ldots,7,8$, are given by *metric prefixes*: yocto-, zepto-, atto-, fempto-, pico-, nano-, micro-, milli-, kilo-, mega-, giga-, tera-, peta-, exa-, zetta-, yotta-, respectively. The lengths 10^t m, t=-2,-1,1,2, are given by: centi-, deci-, deca-, hecto-, respectively.

In computers, a bit is 1 or 0, a byte is 8 bits and 10^{3t} bytes for $t=1,\ldots,7$ are kilo-, mega-, giga-, tera-, peta-, exa-, zettabyte, respectively.

• Imperial length measures

The Imperial length measures (as slightly adjusted by international agreement on 1 July 1959) are:

```
league = 3 \text{ miles};
(US survey) mile = 5,280 \text{ feet } \approx 1,609.347 \text{ m};
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```
international\ mile = 1,609.344 \,\mathrm{m};
   yard = 3 \text{ feet } = 0.9144 \text{ m};
   foot = 12 \text{ inches} = 0.3048 \text{ m};
   inch = 2.54 \,\mathrm{cm} (for firearms, caliber);
   line = \frac{1}{12} inch;
   agate line = \frac{1}{14} inch;

mickey = \frac{1}{200} inch;
   mil (British thou) = \frac{1}{1,000} inch (mil is also an angle measure
\frac{\pi}{3.200} \approx 0.001 \, \text{rad}).
```

In addition, Surveyor's Chain measures are: furlong = 10 chains $=\frac{1}{8}$ mile; chain = 100 links =66 feet; rope = 20 feet; rod (or pole) =16.5feet; link = 7.92 inches. Mile, furlong and fathom (6 feet) come from the slightly shorter Greco-Roman miles (milliare), stadion and orguia, mentioned in the New Testament.

For measuring cloth, old measures are used: bolt = 40 yards; $ell = \frac{5}{4}$

yard; $goad = \frac{3}{2}$ yard; $quarter = \frac{1}{4}$ yard; $finger = \frac{1}{8}$ yard; $nail = \frac{1}{16}$ yard. The following are also old English units of length (cf. **cubit**): barl $eycorn = \frac{1}{3}$ inch; $digit = \frac{3}{4}$ inches; palm = 3 inches; hand = 4 inches; shaftment = 6 inches; span = 9 inches; cubit = 18 inches.

Cubit

The **cubit**, originally the distance from the elbow to the tip of the fingers of an average person, is the ordinary unit of length in the ancient Near East which varied among the cultures and with time. It is the oldest recorded measure of length; the cubit was used in the temples of Ancient Egypt from at least 2700 BC with the following proportions: 1 ordinary Egyptian cubit=6 palms=24 digits=450 mm (18 inches), and 1 royal Egyptian $cubit = 7 \ palms = 28 \ digits = 525 \ mm$. Relevant Sumerian measures were: 1 ku (ordinary Mesopotamian cubit) = 30 shusi = 25 uban = 500 mm, and 1 kus (great Mesopotamian cubit) = $36 \ shusi = 30 \ uban = 600 \ mm$.

Biblical measures of length are the *cubit* and its multiples by $4, \frac{1}{2}, \frac{1}{6}$ $\frac{1}{24}$ called fathom, span, palm, digit, respectively. But the basic length of the Biblical cubit is unknown; it is estimated now as about 44.5 cm (as Roman *cubitus*) for the common cubit, used in commerce, and 51–56 cm for the sacred one, used for building.

The Talmudic cubit is ≈ 56 cm.

The pyramid cubit is ≈ 63.5 cm; this unit, derived in Newton's Biblical studies, is supposed to be the basic one in the dimensions of the Great Pyramid and in far-reaching numeric relations on them.

• Nautical length units

The **nautical length units** (also used in aerial navigation) are:

 $sea\ league = 3\ sea\ (nautical)\ miles;$

nautical mile = 1.852 m (originally, it was defined as 1' of arc along the great circle of Earth);

 $qeographical\ mile \approx 1,855.325\,\mathrm{m}$ (the average distance on the Earth's surface, represented by 1' of arc along the Earth's equator);

```
cable = 120 \, \mathrm{fathoms} = 720 \, \mathrm{feet} = 219.456 \, \mathrm{m}; short \, cable = \frac{1}{10} \, \mathrm{nautical \ mile} \approx 608 \, \mathrm{feet}; fathom = 6 \, \mathrm{feet}.
```

• ISO paper sizes

In the widely used ISO paper size system, the height-to-width ratio of all pages is the *Lichtenberg ratio*, i.e., $\sqrt{2}$. The system consists of formats An, Bn and (used for envelopes) Cn with $0 \le n \le 10$, having widths $2^{-\frac{1}{4}-\frac{n}{2}}$, $2^{-\frac{n}{2}}$ and $2^{-\frac{1}{8}-\frac{n}{2}}$, respectively. The above measures are in meters; so, the area of An is 2^{-n} square meter. They are rounded and expressed usually in millimeters; for example, format A4 is 210×297 and format B7 (used also for EU and US passports) is 88×125 .

Typographical length units

```
point \ (PostScript) = \frac{1}{72} \text{ inch} = 100 \text{ gutenbergs} = 3.527777778 \text{ cm};

point \ (TeX) \ (\text{or } printer's \ point) = \frac{1}{72.27} \text{ inch} = 3.514598035 \text{ cm};

point \ (ATA) = 3.514598 \text{ cm};
```

kyu (Japanese) (or Q, quarter) = 2.5 cm;

point $(Didot) = \frac{1}{72}$ French royal inch = 3.761 cm, and cicero = 12 points (Didot);

pica (Postscript, TeX or ATA) = 12 points in the corresponding system; $twip = \frac{1}{20}$ of a point in the corresponding system.

• Astronomical length units

The **Hubble distance** (cf. Chap. 26) or *Hubble radius*, *Hubble length* is $D_H = \frac{c}{H_0} \approx 4.228 \,\mathrm{Gpc} \approx 13.7$ billion light-years (used to measure distances $d > \frac{1}{2} \,\mathrm{Mpc}$ in terms of redshift z: $d = zD_H$ if $z \leq 1$, and $d = \frac{(z+1)^2-1}{(z+1)^2+1} D_H$, otherwise).

 $gigaparsec = 10^3 \text{ megaparsec};$

hubble (or light-gigayear, light-Gyr, light-Ga) = 10^9 (billion) light-years ≈306.595 megaparsec;

 $megaparsec = 10^3 \text{ kiloparsec} \approx 3.262 \text{ MLY};$

 $\overrightarrow{MLY} = 10^6$ (million) light-years;

 $kiloparsec = 10^3 \text{ parsecs};$

 $siriometer = 10^6 \, \mathrm{AU} \approx 15.813 \, \mathrm{light\text{-}years}$ (about twice the distance Earth-Sirius);

 $parsec = \frac{648,000}{\pi} \text{ AU} \approx 3.261634 \text{ light-years} = 3.08568 \times 10^{16} \text{ m}$ (the distance from an imaginary star, when the lines drawn from it to the Earth and Sun form the maximum angle, i.e., parallax of 1s);

(Julian) light-year $\approx 9.46073 \times 10^{15} \,\mathrm{m} \approx 5.2595 \times 10^{5} \,\mathrm{light\text{-}minutes}$ $\approx \pi \times 10^{7} \,\mathrm{light\text{-}seconds}$ (the distance light travels in vacuum in a year; used to measure interstellar distances);

spat (used formerly) = 10^{12} m ≈ 6.6846 AU;

astronomical unit (AU) = $1.49597871 \times 10^{11} \,\mathrm{m} \approx 8.32 \,\mathrm{light}$ -minutes (the average distance between the Earth and the Sun; used to measure distances within the solar system);

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light-second $\approx 2.998 \times 10^8 \,\mathrm{m}$;

 $picoparsec \approx 30.86 \,\mathrm{km}$ (cf. other funny units such as $microcentury \approx 52.5 \,\mathrm{minutes}$, usual length of lectures, and $nanocentury \approx \pi \,\mathrm{seconds}$).

• Very small length units

Angström (A) = 10^{-10} m;

angström star (or Bearden unit): $A^* \approx 1.0000148$ angström (used, from 1965, to measure wavelengths of X-rays and distances between atoms in crystals):

X unit (or Siegbahn unit) $\approx 1.0021 \times 10^{-13}$ m (used formerly to measure wavelength of X-rays and gamma-rays);

Bohr radius (the atomic unit of length): α_0 , the mean radius, $\approx 5.291772 \times 10^{-11}$ m, of the orbit of the electron of a hydrogen atom (in the Bohr model);

reduced Compton wavelength of electron (i.e., $\frac{\hbar}{mc}$) for electron mass m_e : $\overline{\lambda}_C = \alpha \alpha_0 \approx 3.862 \times 10^{-13}$ m, where \hbar is the Dirac constant, c is the speed of light, and $\alpha \approx \frac{1}{137}$ is the fine-structure constant;

classical electron radius (Lorentz radius): $r_e = \alpha \overline{\lambda}_C = \alpha^2 \alpha_0 \approx 2.81794 \times 10^{-15} \,\mathrm{m};$

Compton wavelength of proton: $\approx 1.32141 \times 10^{-15}$ m; the majority of lengths, appearing in experiments on nuclear fundamental forces, are integer multiples of it.

Planck length (the smallest physical length): $l_P = \sqrt{\frac{\hbar G}{c^3}} \approx 1.6162 \times 10^{-35} \,\mathrm{m}$, where G is the Newton universal gravitational constant. It is the reduced Compton wavelength and also half of the gravitational radius for the Planck mass $m_P = \sqrt{\frac{\hbar c}{G}} \approx 2.176 \times 10^{-8} \,\mathrm{kg}$ (weight of a mite).

The remaining base Planck units are Planck time $t_p = \frac{l_P}{c} \approx 5.4 \times 10^{-44} \, \text{s}$, Planck temperature $T_P \approx 1.4 \times 10^{32} \, \text{K}$, Planck density $\rho_P \approx 5.1 \times 10^{96} \, \text{kg m}^{-3}$, and Planck charge $q_P \approx 1.9 \times 10^{-18} \, \text{C}$.

In fact, $10^{38}l_P \approx 1$ US mile, $10^{43}t_P \approx 54\,\mathrm{s}$, $10^9m_P \approx 21.76\,\mathrm{kg}$ (close to 1 talent, 26 kg of silver, a unit of mass in Ancient Greece), $\frac{1}{100}$ -th of $10^{-30}T_P$ is, roughly, a step on the kelvin scale, and $10^{20}q_P$ per minute is very close to 3 A of current. Cottrell (http://planck.com/humanscale. htm) proposed a "postmetric" human-scale adaptation of the Planck units system based on the above five units, calling them (Planck) mile, minute, talent, grade and score, respectively.

Natural units

In the **International Metric System** SI, velocity V, angular momentum W and energy E are derived from the primary quantities length L, time T and mass M by $V = \frac{L}{T}$, $W = \frac{ML^2}{T}$ and $E = \frac{ML^2}{T^2}$. Thus $L = \frac{VW}{E}$, $T = \frac{W}{E}$ and $M = \frac{E}{V^2}$. For the speed of light $c = 2.99792458 \times 10^8 \, \mathrm{m \, s^{-1}}$ and the Dirac constant $\hbar = 6.5821 \times 10^{-25} \, \mathrm{GeV}$ s, the equation $c\hbar = 0.197 \times 10^{-15} \, \mathrm{GeV}$ m holds. Gev (formerly BeV) means billion-electron-volts.

It is convenient, in high energy (i.e., short distance) Physics, to redefine all units by setting $c = \hbar = 1$. The primary **natural units** are c = 1 (for velocity V), $\hbar = 1$ (for angular momentum W) and $1 \text{ Gev} = 1.602 \times 10^{-10} \text{ J}$ (for energy).

Length L, time T and mass M are now the derived quantities $\frac{1}{E}$, $\frac{1}{E}$ and E with conversions $1 \,\text{GeV}^{-1} = 0.197 \times 10^{-15} \,\text{m}$, $1 \,\text{GeV}^{-1} = 6.58 \times 10^{-25} \,\text{s}$ and $1 \,\text{GeV} = 1.8 \times 10^{-27} \,\text{kg}$, respectively.

• Length scales in Physics

In Physics, a **length scale** is a length range (one or several orders of magnitude) within which given phenomena are consistently described by a theory. Roughly, the limit scales correspond to Cosmology and High Energy Particle Physics.

The macroscopic scale of our real world is followed by the mesoscopic scale (or nanoscale, $\sim 10^{-8}$ m) where materials and phenomena can be still described continuously and statistically, and average macroscopic properties (for example, temperature and entropy) are relevant.

In terms of their concept of elementary constituents, Chemistry (molecules, atoms), Nuclear (proton, neutron, electron, neutrino, photon), Hadronic (exited states) and Standard Model (quarks and leptons) are applicable at scales $\geq 10^{-10}$, $\geq 10^{-14}$, $\geq 10^{-15}$ and $\geq 10^{-18}$ m, respectively. In the atomic scale, $\sim 10^{-10}$ m, the individual atoms should be seen as

In the *atomic scale*, $\sim 10^{-10}$ m, the individual atoms should be seen as separated. The *QCD* (Quantum Chromodynamics) scale, $\sim 10^{-15}$ m, deals with strongly interacting particles.

The electroweak scale, $\sim 10^{-18}$ m (100–1,000 GeV, in **natural units**, i.e., in terms of energy), and the Planck scale, $\sim 10^{-35}$ m ($\sim 10^{19}$ GeV) follow.

In between, the GUT (Grand Unification Theories) scale, 10^{14} – $10^{16}\,\mathrm{GeV}$ is expected with grand unification of non-gravitational fundamental forces at the length $10^{-28}\,\mathrm{m}$. The compactification scale should give the size of compact extra-dimensions predicted by M-theory in order to derive all forces from a gravitational force acting on strings in high-dimensional space.

The electroweak scale will be probed by LHC (Large Hadron Collider) and ILC (International Linear Collider). A new proposal moves the *string scale* from 10^{-34} to 10^{-19} m (1 TeV region) and expects the corresponding spatial extra-dimensions to be compactified to a "large" radius of some fraction of a millimeter. LHC will test it; in particular, by looking for putative deviations from the Newton $\frac{1}{d^2}$ law of gravitation which was not tested in the sub-millimeter range below 6×10^{-5} m.

Both General Relativity and Quantum Mechanics (with their space—time being a continuous manifold and discrete lattice, respectively) indicate some form of minimum length where, by the Uncertainty Principle, the very notion of distance loses operational meaning. At short distances, classical geometry is replaced by "quantum geometry" described by 2D conformal field theory. In String Theory, space—time geometry is not fundamental and, perhaps, it only emerges at larger distance scales.

The T-duality is a symmetry between small and large distances. Two superstring theories are T-dual if one compactified on a space of large volume is equivalent to the other compactified on a space of small volume.

Within the Big Bang paradigm humans are roughly in the middle of nature's hierarchy of size scales. It supposes a minimal length scale and a smooth distribution (homogeneous and isotropic) at a large enough scale. But a new infinite fractal hierarchy paradigm posits a fractal universe infinite in all directions fractal universe without any boundary, limit to size scales and preferred reference frame. Parts of the Universe may be created or annihilated, while this fractal hierarchy remains unchanged without any temporal limits. Inflation (10^{50} times increase in 10^{-32} sec; cf. Eternal Inflation Theory by Guth, 2007) is, was and always will be occurring on an infinite number of size scales. Pietonero et al. (2008) claim that the Universe is fractal at large with fractal dimension \approx 2. Quantum fractal space—time is proposed also; it is seen as 2D at the Planck scale and gradually becomes 4D at larger scales.

Yet another paradigm – Loop Quantum Cosmology (Bojowald 2000) – posits that our universe emerged from a pre-existing one that had been expanding and then contracted due to gravity. At around *Planck density* ($5.1 \times 10^{96} \,\mathrm{kg}\,\mathrm{m}^{-3}$, i.e., a trillion suns compressed down to the size of a proton), the compressed space–time exerts outward force overriding gravity. The universe rebounds and keep expanding because of the bounce inertia.

27.2 Orders of magnitude for length

In this section we present a selection of orders of length magnitudes, expressed in meters.

 1.616×10^{-35} : Planck length (smallest physical length). At this scale is expected Wheeler's "quantum foam" (violent warping and turbulence of space–time, no smooth spatial geometry; the dominant structures are little multiply-connected wormholes and bubbles popping into existence and back out of it). But if the Holographic Principle holds, i.e., space-time is a grainy hologram, then the scale of grains is $\approx 10^{-16}$ m;

 10^{-34} : length of a putative *string* in M-theory, which suppose that all forces and elementary particles arise by vibration of such strings, and hopes to unify Quantum Mechanics and General Relativity;

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10^{-24} = 1 yoctometer;
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 $10^{-21} = 1$ **zeptometer**;

 $10^{-18} = 1$ attometer: weak nuclear force range, size of a quark;

 $10^{-15} = 1$ **femtometer** (formerly, *fermi*);

 1.6×10^{-15} : diameter of proton;

 $1.3\times 10^{-15};$ strong nuclear force range, medium-sized nucleus;

 $10^{-12} = 1$ **picometer** (formerly, *bicron* or *stigma*): distance between atomic nuclei in a White Dwarf star:

 10^{-11} : wavelength of hardest (shortest) X-rays and largest wavelength of gamma rays;

 5×10^{-11} : diameter of the smallest (hydrogen H) atom; 1.5×10^{-10} : diameter of the smallest (hydrogen H_2) molecule;

 $10^{-10} = 1$ angström: diameter of a typical atom, limit of resolution of the electron microscope;

 1.54×10^{-10} : length of a typical covalent bond (C-C);

 3.4×10^{-10} : distance between base pairs in DNA molecule;

 $10^{-9} = 1$ nanometer: diameter of typical molecule;

10⁻⁸: wavelengths of softest X-rays and most extreme ultraviolet;

 1.1×10^{-8} : diameter of prion (smallest self-replicating biological entity);

 4.5×10^{-8} : the smallest feature of computer chip (length of transistor gate dielectrics from high-K substitutes of silicon) in 2007;

 9×10^{-8} : human immunodeficiency virus, HIV; in general, known viruses range from 2×10^{-8} (adeno-associated virus) to 8×10^{-7} (Mimivirus);

 10^{-7} : size of chromosomes, maximum size of a particle that can fit through a surgical mask;

 2×10^{-7} : limit of resolution of the light microscope;

 $3.8-7.4\times10^{-7}$: wavelength of visible (to humans) light, i.e., the color range of purple through red (the sunlight wavelength is $560 \,\mathrm{nm}$);

 $10^{-6} = 1$ micrometer (formerly, micron);

 $10^{-6} - 10^{-5}$: diameter of a typical bacterium; in general, known (in non-dormant state) bacteria range from 1.5×10^{-7} (Micoplasma genitalium: "minimal cell") to 7×10^{-4} (Thiomargarita of Namibia);

 7×10^{-6} : diameter of the nucleus of a typical eukaryotic cell;

 8×10^{-6} : mean width of human hair (ranges from 1.8×10^{-6} to 18×10^{-6});

 10^{-5} : typical size of (a fog, mist, or cloud) water droplet;

 10^{-5} , 1.5×10^{-5} , and 2×10^{-5} : widths of cotton, silk, and wool fibers;

 2×10^{-4} : approximately, the lower limit for the human eye to discern an object;

 5×10^{-4} : diameter of a human ovum, MEMS micro-engine;

 $10^{-3} = 1$ millimeter;

 5×10^{-3} : length of average red ant; in general, insects range from 1.7×10^{-4} (Megaphragma caribea) to 3.6×10^{-1} (Pharnacia kirbyi);

 8.9×10^{-3} : gravitational radius ($\frac{2Gm}{c^2}$: the value below which mass m collapses into a black hole) of the Earth;

 $10^{-2} = 1$ centimeter;

 $10^{-1} = 1$ decimeter: wavelengths of the lowest microwave and highest UHF radio frequency, 3 GHz;

 $1~\mathrm{meter}:$ wavelength of the lowest UHF and highest VHF radio frequency, $300\,\mathrm{MHz}:$

1.435: standard gauge (the distance between the rails) of a railway track:

2.77-3.44: wavelength of the broadcast radio FM band, 108-87 MHz;

- 5.5, 33.3, and 55: height of the tallest animal (the giraffe), length of a blue whale (the largest animal), and length of the longest animal (bootlace worm *Lineus longissimus*);
- 10 = 1 **decameter**: wavelength of the lowest VHF and highest shortwave radio frequency, $30 \,\mathrm{MHz}$;
- 29: highest measured ocean wave, while 524 m and 1 km are estimated heights of mega-tsunamis on 10 July 1958 in Lituya Bay, Alaska, and 65 million years ago, in Yucatan (impact of the K-T asteroid that may have killed off the dinosaurs);
- 100 = 1 **hectometer**: wavelength of the lowest shortwave radio frequency and highest medium wave radio frequency, 3 MHz;
 - 115.5: height of the world's tallest tree, a sequoia Coast Redwood;
- 137, 300, 509, 553, and 818: heights of the Great Pyramid of Giza, Eiffel Tower, Taipei 101 Tower, CN Tower in Toronto, and, planned to open in 2009. Burj Dubai;
 - 187–555: wavelength of the broadcast radio AM band, 1,600–540 kHz;
 - 340: distance which sound travels in air in 1s;
 - $10^3 = 1$ kilometer;
- $\approx 1.1 \times 10^3$: mean diameter of asteroid 1950 DA which is predicted, with probability $\frac{1}{300}$, to collide with Earth on 16 March 2880 and may be devastating to human civilization;
 - 2.95×10^3 : gravitational radius of the Sun;
 - 3.79×10^3 : mean depth of oceans;
 - 10⁴: wavelength of the lowest medium wave radio frequency, 300 kHz;
- 8.8×10^3 and 10.9×10^3 : height of the highest mountain, Mount Everest, and depth of the Mindanao Trench;
- $5 \times 10^4 = 50$ km: the maximal distance at which the light of a match can be seen (at least 10 photons arrive on the retina during 0.1 s);
 - $8 \times 10^4 = 80 \,\mathrm{km}$: thickness of the ozone layer;
 - $1.11 \times 10^5 = 111$ km: 1° of latitude on the Earth;
- $1.5 \times 10^4 1.5 \times 10^7$: wavelength range of audible (to humans) sound (20 Hz-20 kHz);
- $1.69\times 10^5;$ length of the world's longest tunnel, Delaware Water Supply, New York;
- 2×10^5 : wavelength (the distance between the troughs at the bottom of consecutive waves) of a typical tsunami;
- 4.83×10^5 : diameter of Wilkes Land (Antarctica), the largest crater. This impact is suspected to cause the worst, Permian, mass extinction of life 250 millions years ago;
 - $10^6 = 1$ megameter, thickness of Earth's atmosphere;
- 2.4×10^6 : diameter of the plutoid Eris, the largest ($\approx 100 \,\mathrm{km}$ larger than Pluto) dwarf planet (round planetoids that have not "cleared the neighborhood" through their gravitational effects), at 67.67 AU from Sun; the smallest dwarf planet is Ceres (the largest asteroid in the Asteroid belt) of diameter 9.42×10^5 and at $2.77 \,\mathrm{AU}$;

- 3.48×10^6 : diameter of the Moon;
- 5×10^6 : diameter of LHS 4033, the smallest known White Dwarf star;
- 6.4×10^6 and 6.65×10^6 : length of the Great Wall of China and length of the Nile river;
 - 6.8×10^6 : diameter of the Earth's core;
- 1.28×10^7 and 4.01×10^7 : Earth's equatorial diameter and length of the equator;
 - 3.84×10^8 : Moon's orbital distance from the Earth;
 - $10^9 = 1$ gigameter;
 - 1.39×10^9 : diameter of the Sun:
- 1.4×10^9 : diameter of the nebulous spherical envelope of the comet 17P/Holmes on 9 November 2007, produced by dust ejected from its ice/rock nucleus. $3.6 \,\mathrm{km}$ in diameter.
 - 5.8×10^{10} : orbital distance of Mercury from the Sun;
- 1.496×10^{11} (1 astronomical unit, AU): mean distance between the Earth and the Sun (orbital distance of the Earth);
 - 5.7×10^{11} : length of longest observed comet tail (Hyakutake 1996);
 - $10^{12} = 1$ terameter (formerly, spat);
- $2.5\text{--}2.9\times10^{12}\text{:}$ diameter of the largest known supergiant star, VY Canis Majoris;
- $39.5\,\mathrm{AU} \approx 5.9 \times 10^{12}$: radius of the inner solar system (orbital distance of Pluto);
- 50 AU: distance from the Sun to the *Kuiper cliff*, the abrupt unexplained outer boundary of the Kuiper belt (the region, 30–50 AU around Sun, of trans-Neptunian objects):
- $523\,\mathrm{AU}\approx7.8\times10^{13}$: orbital distance of Sedna, the farthest known object in the solar system;
 - $10^{15} = 1$ petameter;
- $1.1\,\mathrm{light-year}\approx10^{16}\approx70,\!000\,\mathrm{AU}\approx0.337\,\mathrm{pc}$: the closest passage (in 1,360,000 years) of Giese 710, a red dwarf star expected to perturb the Oort Cloud dangerously;
- 50,000–100,000 AU: distance from the Sun of the outer boundary of the Oort cloud (supposed cloud of long-period comets) where galactic tide overtakes Sun's gravity;
- $3.99 \times 10^{16} = 266,715\,\mathrm{AU} = 4.22\,\mathrm{light\text{-}years} = 1.3\,\mathrm{pc}$: distance from the Sun to Proxima Centauri, the nearest star;
- $8.6\,\mathrm{light\text{-}years}\approx\!\!8.1\times10^{16}\mathrm{:}$ distance from the Sun to Sirius, the brightest star of our sky;
 - $10^{18} = 1$ exameter:
 - $1.57 \times 10^{18} \approx 50.9 \,\mathrm{pc}$: distance to supernova 1987A;
- $2.62 \times 10^{20} \approx 8.5 \,\mathrm{kpc}$ (2.77 × 10^4 light-years): distance from the Sun to the geometric center of our Milky Way galaxy (in giant black hole Sagittarius A^*);
- $3.98 \times 10^{20} \approx 12.9 \,\mathrm{kpc}$: distance to Canis Major Dwarf, the closest satellite galaxy to our Solar System (LMC, the largest one, is at 50 kps);

 $9.46 \times 10^{20} \approx 30.66 \,\mathrm{kpc} \approx 10^5 \,\mathrm{light\text{-}years}$: diameter of the galactic disk of our Milky Way galaxy;

 $10^{21} = 1$ zettameter:

- $2.23 \times 10^{22} = 725 \,\mathrm{kpc}$: distance to Andromeda Nebula, the closest (and approaching by about $140 \,\mathrm{km \, s^{-1}}$) large galaxy;
 - $5 \times 10^{22} = 1.6 \,\mathrm{Mpc}$: diameter of the Local Group of galaxies;
- $5.7 \times 10^{23} = 60 \,\text{MLY}$: distance to Virgo, the nearest (and approaching) major cluster; it dominates the Local Supercluster and contains extragalactic stars and a dark matter galaxy;

 $10^{24} = 1$ vottameter:

- $2 \times 10^{24} = 60 \,\mathrm{Mpc}$: diameter of the Local Supercluster;
- $2.36 \times 10^{24} = 250 \,\mathrm{MLY}$: distance to the Great Attractor (a gravitational anomaly of $\approx 5 \times 10^{16}$ solar masses, in the direction of the Shapley Supercluster, the largest known concentration of matter) where our galaxy is going;
- 1,370 MLY: length of the Sloan Great Wall of galaxies, the largest observed superstructure in the Universe (the space looks uniform on larger scales: "End of Greatness"). The diameter of the largest known void is ≈ 3.5 light-Gyr;
- $8,000\,\mathrm{MLY}\approx7.57\times10^{25}$ (redshift $z\geq1.0$): typical distance to the source of a GRB (Gamma Ray Burst), while the most distant event known, GRB 050904, has z=6.3, and the most distant event ever seen by human eyes without optical aid on record, GRB 080319B (19 March 2008), has z=0.94, i.e., $\approx7,500\,\mathrm{MLY}$;
- $12,080\,\mathrm{MLY} = 3,704\,\mathrm{Mpc}$: distance to the quasar SDSS J1148 + 5251 (redshift z=6.43, while 6.5 is supposed to be the "wall of invisibility" for visible light);
- 13,230 MLY: near-infrared observed distance to the farthest known in 2008 galaxy Abell 1835 IR1916 (redshift 10). The formation of the first stars (at the end of the $Dark\ Age$, when matter consisted of clouds of cold hydrogen) corresponds to $z\approx 20$ when the Universe was ≈ 200 million years old;
- $1.3 \times 10^{26} = 13.7 \, \text{light-Gyr} = 4.22 \, \text{Gpc}$ (redshift $z \approx 1,089$): the distance that cosmic background radiation has traveled since the Big Bang (Hubble radius $D_H = \frac{c}{H_0}$, the cosmic light horizon, the age of the Universe).
- $4.3 \times 10^{26} = 46$ light-Gyr: particle horizon (the present **comoving distance** to the edge of the observable Universe; it is larger than the Hubble radius, since the Universe is expanding). But if the topology of the Universe is non-simply connected, then it is compact and the estimated maximum length scale is only 5–15% of Hubble radius;

Projecting into the future: the scale of the Universe will be 10^{31} in 10^{14} years (last red dwarf stars die) and 10^{37} in 10^{20} years (stars have left galaxies). If protons decay, their half-life is $\geq 10^{35}$ years; the estimated number of protons in the Universe is 10^{77} ;

The Universe, in the current $Heat\ Death$ scenario, achieves beyond $10^{1,000}$ years such a low-energy state, that quantum events became ma-

jor macroscopic phenomena, and space—time looses its usual meaning again, as below the Planck time or length;

The hypothesis of parallel universes estimates that one can find another identical copy of our Universe within the distance $10^{10^{118}}$ m.

Chapter 28 Distances in Applied Social Sciences

In this chapter we present selected distances used in real-world applications of Human Sciences. In this and next chapter, the expression of distances ranges from numeric (say, in meters) to ordinal (as a degree assigned according to some rule). Depending on the context, the distances are either practical ones, used in daily life and work outside of science, or those used as a metaphor for remoteness (the fact of being apart, being unknown, coldness of manner, etc.).

28.1 Distances in Psychology and Sociology

• Approximative human-scale distances

An arm's length is a distance (about 0.7 m, i.e., within personal distance) sufficient to exclude intimacy, i.e., discouraging familiarity or conflict; its analogs are: Italian braccio, Turkish pik, and Old Russian sazhen. The **reach distance** is the difference between maximum reach and arm's length distance.

The **striking distance** is a short distance (say, through which an object can be reached by striking).

The **spitting distance** is a very close distance.

The **shouting distance** is a short, easily reachable distance.

A stone's throw is a distance of about 25 fathoms (46 m).

The **hailing distance** is the distance within which the human voice can be heard.

The walking distance is the distance normally (depending on the context) reachable by walking. For example, some UK high schools define 2 and 3 miles as the statutory walking distance for children younger and older than 11 years. *Pace out* means to measure distance by *pacing* (walking with even steps).

The **acceptable commute distance**, in Real Estate, is the distance that can be covered in an acceptable travel time and increases with better connectivity.

• Optimal eye-to-eye distance

The **optimal eye-to-eye distance** between two persons was measured for some types of interaction. For example, such optimal viewing distance between a baby and its mother's face, with respect to immature motor and visual systems of the newborn, is 20–30 cm. Haynes, White, Held (1965) showed that during the first weeks of life the accommodation system does not yet function and the lens of the newborn is locked at the **focal distance** of about 19 cm.

• Distances between people

In [Hall69], four interpersonal bodily distances were introduced: the *intimate distance* for embracing, touching or whispering (15–45 cm), the *personal-casual distance* for conversations among good friends (45–120 cm), the *social-consultative distance* for conversations among acquaintances (1.2–3.6 m), and the *public distance* used for public speaking (over 3.6 m). Cf. **distances in Animal Behavior** in Chap. 23.

To each of those proxemics distances corresponds an intimacy/confidence degree and appropriated sound level. The distance which is appropriate for a given social situation depends on culture, gender and personal preference. For example, under Islamic law, proximity (being in the same room or secluded place) between a man and a woman is permitted only in the presence of their mahram (a spouse or anybody from the same sex or a pre-puberty one from the opposite sex). For an average westerner, personal space is about 70 cm in front, 40 cm behind and 60 cm on either side.

Example of other cues of nonverbal communication is given by angles of vision which individuals maintain while talking. The **people angular distance in a posture** is the spatial orientation, measured in degrees, of an individual's shoulders relative to those of another; the position of a speaker's upper body in relation to a listener's (for example, facing or angled away); the degree of body alignment between a speaker and a listener as measured in the coronal (vertical) plane which divides the body into front and back. This distance reveals how one feels about people nearby: the upper body unwittingly angles away from disliked persons and during disagreement.

Distancing behavior of people can be measured, for example, by the *stop distance* (when the subject stops an approach since she/he begins to feel uncomfortable), or by the *quotient of approach*, i.e., the percentage of moves made that reduce the interpersonal distance to all moves made.

• Psychological Size and Distance Scale

The CID (Comfortable Interpersonal Distance) scale by Duke and Nowicky (1972) consists of a center point 0 and eight equal lines emanating from it. Subjects are asked to imagine themselves on the point 0 and to respond to descriptions of imaginary persons by placing a mark at the point on a line at which they would like the imagined person to stop, that is, the point at which they would no longer feel comfortable. CID is then measured in mm from 0.

The GIPSDS (Psychological Size and Distance Scale) by Grashma and Ichiyama (1986) is a 22-item rating scale assessing interpersonal status and affect. Subjects are asked to draw circles, representing the drawer and other significant persons, so that the radii of the circles and the distances between them indicate the thoughts and feelings about their relationship. Then these distances and radii, measured in mm, represent psychological distance and status, respectively.

• Symbolic distance effect

In Psychology, the brain compares two concepts (or objects) with higher accuracy and faster reaction time if they differ more on the relevant dimension. For example, the performance of subjects when comparing a pair of positive numbers (x, y) decreases for smaller |x - y| (behavioral numerical distance effect).

The related magnitude effect (or Weber–Fechner law effect) is that performance decreases for larger $\min\{x,y\}$. Those effects are valid also for congenitally blind people; they learn spatial relation via tactile input (interpreting, say, numerical distance by placing pegs in a peg board).

A current explanation is that there exists a mental line of numbers, which is oriented from left to right (as 2, 3, 4) and non-linear (more mental space for smaller numbers). So, close numbers are easier to confuse since they are represented on the mental line at adjacent and not always precise locations. The quantity system providing a semantic representation of the size and distance relations between numbers is thought to be located in the parietal cortex.

Distancing

Distancing is any behavior or attitude causing to be or appearing to be at a distance.

In Martial Arts, **distancing** is the selection of an appropriate *combat range*, i.e., distance from the adversary. For other examples of spatial distancing, see **distances between people** and, in Chap. 29, **safe distancing** from a risk factor.

In *Mediation* (a form of alternative dispute resolution), **distancing** is the impartial and non-emotive attitude of the mediator versus the disputants and outcome.

In Psychoanalysis, **distancing** is the tendency to put persons and events at a distance. It concerns both the patient and the psychoanalyst.

In Developmental Psychology, **distancing** (Werner and Kaplan 1964, for deaf-blind patients) is the process of establishing a subject individuality as an essential phase (prior to symbolic cognition and linguistic communication) in learning to treat symbols and referential language. For Sigel (1970, for preschool children), **distancing** is the process of the development of cognitive representation: cognitive demands by the teacher or the parent help to generate a child's representational competence.

In books by Kantor, **distancing** refers to APD (Avoidant Personality Disorder): fear of intimacy and commitment (confirmed bachelors,

"femmes fatales," etc.) The **distancing language** is phrasing used by a person to avoid thinking about the subject or content of his own statement (for example, referring to death).

Cf. technology-related distancing and antinomy of distance.

• Distance education

Distance education is the process of providing instruction when students and instructors are separated by physical distance, and technology is used to bridge the gap. *Distance learning* is the desired outcome of distance education.

The transactional distance (Moore 1993) is a perceived degree of separation during interaction between students and teachers, and within each group. This distance decreases with *dialog* (a purposeful positive interaction meant to improve the understanding of the student), with larger autonomy of the learner, and with lesser predetermined structure of the instructional program.

Moral distance

The **moral distance** is a measure of moral indifference or empathy toward a person, group of people, or events.

The (moral) distancing is a separation in time or space that reduces the empathy that a person may have for the suffering of others, i.e., that increases moral distance. In particular, **distantiation** is the tendency to distance oneself (physically or socially, by segregation or congregation) from those that one does not value. Cf. **distanciation**.

On the other hand, the term *good distancing* (Sartre 1943, and Ricoeur 1995) means the process of deciding how long a given ethical link should be.

• Emotional distance

The **emotional distance** (or *psychic distance*) is the degree of emotional detachment (toward a person, group of people or events), aloofness, indifference by personal withdrawal, reserve.

The Bogardus Social Distance Scale (cf. social distance) measures, in fact, not social but this emotional distance; it offers following eight response items: would marry, would have as a guest in my household, would have as next door neighbor, would have in neighborhood, would keep in the same town, would keep out of my town, would exile, would kill. Dodd and Nehnevasja attached, in 1954, increasing distances of 10^t m, $0 \le t \le 7$, to these eight levels of the Bogardus scale.

The **propinquity effect** is the tendency for people to get emotionally involved, as to form friendships or romantic relationships, with those who have higher *propinquity* (physical/psychological proximity) with them, i.e., whom they encounter often. Walmsley (1978) proposed that emotional involvement decreases as $d^{-\frac{1}{2}}$ with increasing **subjective distance** d.

• Social distance

In Sociology, the **social distance** is the extent to which individuals or groups are removed or excluded from participating in one another's lives; a degree of understanding and intimacy which characterize personal and

social relations generally. This notion was originated by Simmel in 1903; in his view, the social forms are the stable outcomes of distances interposed between subject and object.

The Bogardus Social Distance Scale (cf. **emotional distance**) is scored so that the responses for each ethnic/racial group are averaged across all respondents which yields a RDQ (racial distance quotient) ranging from 1.00 to 8.00.

An example of relevant models: Akerlof [Aker97] defines an agent x as a pair (x_1, x_2) of numbers, where x_1 represents the initial, i.e., inherited, social position, and the position expected to be acquired, x_2 . The agent x chooses the value x_2 so as to maximize

$$f(x_1) + \sum_{y \neq x} \frac{e}{(h + |x_1 - y_1|)(g + |x_2 - y_1|)},$$

where e, h, g are parameters, $f(x_1)$ represents the intrinsic value of x, and $|x_1-y_1|$, $|x_2-y_1|$ are the inherited and acquired *social distances* of x from any agent y (with the social position y_1) of the given society.

• Rummel sociocultural distances

Rummel defined [Rumm76] the main sociocultural distances between two persons as follows:

- 1. **Personal distance:** one at which people begin to encroach on each other's territory of personal space.
- 2. Psychological distance: perceived difference in motivation, temperaments, abilities, moods, and states (subsuming *intellectual distance*).
- 3. **Interests-distance:** perceived difference in wants, means, and goals (including **ideological distance** on socio-political programs).
- Affine distance: degree of sympathy, liking or affection between the two.
- 5. Social attributes distance: differences in income, education, race, sex, occupation, etc.
- 6. **Status-distance:** differences in wealth, power, and prestige (including **power distance**).
- 7. Class-distance: degree to which one person is in general authoritatively superordinate to the other.
- Cultural distance: differences in meanings, values and norms reflected in differences in philosophy-religion, science, ethics-law, language, and fine arts.

• Cultural distance

The cultural distance between countries $x = (x_1, ..., x_5)$ and $y = (y_1, ..., y_5)$ (usually, US) is derived (in [KoSi88]) as the following composite index

$$\sum_{i=1}^{5} \frac{(x_i - y_i)^2}{5V_i},$$

where V_i is the variance of the index i, and the five indices represent [Hofs80]:

- 1. Power distance (preferences for equality)
- 2. Uncertainty avoidance (risk aversion)
- 3. Individualism versus collectivism
- 4. Masculinity versus femininity (gender specialization)
- 5. Confucian dynamism (long-term versus short-term orientation)

The above **power distance** measures the extent to which the less powerful members of institutions and organizations within a country expect and accept that power is distributed unequally, i.e., how much a culture has respect for authority. For example, Latin Europe and Japan fall in the middle range.

Wirsing (1973) defined *social distance* as a "symbolic gap" between rulers and ruled designed to set apart the political elite from the public. It consists of reinforced and validated ideologies (a formal constitution, a historical myth, etc.).

Davis (1999) theorized social movements (in Latin America) in terms of their shared distance from the state: geographically, institutionally, socially (class position and its economic counterpart, income level) and culturally. For example, the groups distanced from the state on all four dimensions are more likely to engage in revolutionary action.

Henrikson (2002) identified the following Political Geography distances between countries: attributional distance (according to cultural characteristics, say, democracy or not), gravitational distance (according to which political and other powers "decay") and topological distance (remoteness of countries increases when others are located in between them).

28.2 Distances in Economics and Human Geography

• Effective trade distance

The **effective trade distance** between countries x and y with populations x_1, \ldots, x_m and y_1, \ldots, y_n of their main agglomerations is defined in [HeMa02] as

$$\left(\sum_{1 \le i \le m} \frac{x_i}{\sum_{1 \le t \le m} x_t} \sum_{1 \le j \le n} \frac{y_j}{\sum_{1 \le t \le n} y_t} d_{ij}^r\right)^{\frac{1}{r}},$$

where d_{ij} is the bilateral distance (in kilometers) of corresponding agglomerations x_i, y_i , and r measures the sensitivity of trade flows to d_{ij} .

As an **internal distance of a country**, measuring the average distance between producers and consumers, Head and Mayer [HeMa02] proposed $0.67\sqrt{\frac{area}{\pi}}$.

Technology distances

The **technological distance** between two firms is a distance (usually, χ^2 - or **cosine distance**) between their *patent portfolios*, i.e., vectors of the number of patents granted in (usually, 36) technological sub-categories. Other measures are based on the number of patent citations, co-authorship networks, etc.

Granstrand's **cognitive distance** between two firms is the **Steinhaus distance** $\frac{\mu(A \triangle B)}{\mu(A \cup B)} = 1 - \frac{\mu(A \cap B)}{\mu(A \cup B)}$ between their technological profiles (sets of ideas) A and B seen as subsets of a measure space (Ω, A, μ) .

The economic model of Olsson (2000) defines the metric space (I, d) of all ideas (as in human thinking), $I \subset \mathbb{R}^n_+$, with some intellectual distance d. The closed, bounded and connected knowledge set $A_t \subset I$ extends with time t. New elements are, normally, convex combinations of previous ones: innovations within gradual technological progress. Exceptionally, breakthroughs (Kuhn's paradigm shifts) occur. The similar notion of thought space (an externalized mental space of ideas/knowledge and relationships among them in thinking) was used in Sumi, Hori and Ohsuga (1997) for computer-aided thinking with text; they proposed a system of mapping text-objects into metric spaces.

Introduced in Patel (1965) the **economic distance** between two countries is the time (in years) for a lagging country to catch up to the same per capita income level as the present one of an advanced country. Introduced in Fukuchi and Satoh (1999) the **technology distance** between countries is the time (in years) when a lagging country realizes a similar technological structure as the advanced one has now. The basic assumption of the *Convergence Hypothesis* is that the technology distance between two countries is smaller than the economic one.

• Production Economics distances

In quantitative Economics, a technology is modeled as a set of pairs (x, y), where $x \in \mathbb{R}_+^m$ is an input vector, $y \in \mathbb{R}_+^m$ is an output vector, and x can produce y. Such set T should satisfy standard economical regularity conditions.

The **technology directional distance function** of input/output x, y toward (projected and evaluated) a direction $(-d_x, d_y) \in \mathbb{R}^m_- \times \mathbb{R}^m_+$ is

$$\sup\{k \ge 0 : ((x - kd_x), (y + kd_y)) \in T\}.$$

The **Shephard output distance function** is $\sup\{k \geq 0 : (x, \frac{y}{k}) \in T\}$. The *frontier* $f_s(x)$ is the maximum feasible output of a given input x in a given system (or year) s. The **distance to frontier** (Färe, Crosskopf and Lovell 1994) of a production point (x, y), where $y = g_s(x)$, is $\frac{g_s(x)}{f_s(x)}$.

The Malmquist index measuring the change in TFP (total factor productivity) between periods s, s' (or comparing to another unit in the same period) is $\frac{g_s'(x)}{f_s(x)}$. The term distance to frontier is also used for the inverse of TFP in a given industry (or of GDP per worker in a given country) relative to the existing maximum (the frontier, usually, US).

Consider a production set $T \subset \mathbb{R}^{n_1} \times \mathbb{R}^{n_2}$ (input, output). The measure of the technical efficiency, given in Briec and Lemaire (1999) is the point-set distance $\inf_{y \in we(T)} ||x-y||$ (in a given norm ||.|| on $\mathbb{R}^{n_1+n_2}$) from $x \in T$ to the weakly efficient set we(T). It is the set of minimal elements of the poset (T, \preceq) where the partial order $\preceq (t_1 \preceq t_2)$ if and only if $t_2 - t_1 \in K$ is induced by the cone $K = int(\mathbb{R}^{n_1}_{>0} \times \mathbb{R}^{n_2}_{>0}) + \{0\}$.

• Action distance

The action distance is the distance between the set of information generated by the Active Business Intelligence system and the set of actions appropriate to a specific business situation. Action distance is the measure of the effort required to understand information and to effect action based on that information. It could be the physical distance between information displayed and action controlled.

• Death of Distance

Death of Distance is the title of the influential book [Cair01] arguing that the telecommunication revolution (the Internet, mobile telephones, digital television, etc.) initiated the "death of distance" implying fundamental changes: three-shift work, lower taxes, prominence of English, outsourcing, new ways of government control and citizens communication, etc. Physical distance (and so, Economic Geography) do not matter, we all live in a "global village." Thomas Friedman (2005) announced: "The World is flat." Bill Gates claimed (Financial Times 2006): "With the Internet having connected the world together, someone's opportunity is not determined by geography." The proportion of long-distance relationships in foreign relations increased.

Also the "death of distance" allows also both management-at-a-distance and concentration of elites within the "latte belt."

Similarly (see [Ferg03]), steam-powered ships and the telegraph (as rail-roads previously and cars later) led, via falling transportation costs, to the "annihilation of distance" in the nineteenth and twentieth centuries. Further in the past, archaeological evidence points out the appearance of systematic long-distance trade (\approx 140,000 years ago), and the innovation of projectile weapons (40,000 years ago) which allowed humans to kill large game (and other humans) from a safe distance.

But modern technology eclipsed distance only in that the time to reach a destination has shrunk (except where places previously well connected, say, by railroads have fallen off the beaten track). In fact, the distance (cultural, political, geographic, and economic) "still matters" for, say, a company's strategy on the emerging markets, for political legitimacy, etc. Bilateral trade decreases with distance; Disdier and Head (2004) report

a slight increase, over the last century, of this negative impact of distance. Webb (2007) claims that an average distance of trade in 1962 of 4,790 km changed only to 4,938 km in 2000. Partridge, Rickman, Ali and Olfert (2007) report that proximity to higher-tiered urban centers (with their higher-order services, urban amenities, higher-paying jobs, lower-cost products) is an increasingly important positive determinant of local job growth.

Moreover, increased access to services and knowledge exchange requires more face-to-face interaction and so, an increase in the role of distance. Despite globalization, new communication technologies and the dematerialization of economy, economic and innovation activity are highly localized spatially and tend to agglomerate more. Also, the social influence of individuals, measured by the frequency of memorable interactions, is heavily determined by distance.

In military affairs, Boulding (1965) and Bandow (2004) argued that twentieth-century technology reduced the value of proximity for the projection of military power because of "a very substantial diminution in the cost of transportation of armed forces" and "an enormous increase in the range of the deadly projectile" (say, strategic bombardment). It was used as partial justification for the withdrawal of US forces from overseas bases in 2004. But Webb (2007) counterargues that distance retains its importance: for example, any easing of transport is countered by increased strain put upon transport modes since both sides will take advantage of the falling costs to send more supplies. Also, by far the greatest movement of logistics continues to be conducted by sea, with little improvement in speed since 1900.

• Technology-related distancing

The Moral Distancing Hypothesis postulates that technology increases the propensity for unethical conduct by creating a moral distance between an act and the moral responsibility for it.

Print technologies divided people into separate communication systems and distanced them from face-to-face response, sound and touch. Television involved audile-tactile senses and made distance less inhibiting, but it exacerbated *cognitive distancing*: story and image are biased against space/place and time/memory. This distancing has not diminished with computers; interactivity has however increased. In terms of Hunter: technology only re-articulates *communication distance*, because it also must be regarded as the space between understanding and not. The collapsing of spatial barriers diminishes economic but not social and cognitive distance.

On the other hand, the *Psychological Distancing Model* in [Well86] relates the immediacy of communication to the number of information channels: sensory modalities decrease progressively as one moves from face-to-face to telephone, videophone, and e-mail. On-line settings tend to filter out social and relational cues. Also, the lack of instant feedback, because of e-mail communication, is asynchronous and can be isolating:

it and low bandwidth limit visual and aural cues. For example, moral and cognitive effects of distancing in on-line education are not known at present.

• Relational proximity

Economic Geography considers, as opposed to geographical proximity, different types of proximity (organizational, institutional, cognitive, etc.). In particular, **relational proximity** (or trust-based interaction between actors) is an inclusive concept of the benefits derived from spatially localized sets of economic activities. In particular, it generates relational capital through the dynamic exchange of locally produced knowledge.

The five dimensions of relational proximity are proximity: of contact (directness), through time (continuity, stability), in diversity (multiplicity, scope), in mutual respect and involvement (parity), of purpose (commonality).

Individuals are close to each other in a relational sense when they share the same interaction structure, make transactions or realize exchanges. They are *cognitively close* if they share the same conventions and have common values and representations (including knowledge and technological capabilities).

Bouba-Olga and Grossetti (2007) divide also socio-economic proximity into relational proximity (role of social networks) and *mediation proximity* (role of resources such as newspapers, directories, Internet, agencies, etc.)

• Commuting distance

The **commuting distance** is the distance (or travel time) separating work and residence when they are located in separated places (say, municipalities).

• Migration distance (in Economics)

The **migration distance**, in Economic Geography, is the distance between the geographical centers of the municipalities of origin and destination.

• Gravity models

The general **gravity model** for social interaction is given by the *gravity equation*

$$F_{ij} = a \frac{M_i M_j}{D_{ij}^b},$$

where F_{ij} is the "flow" (or "gravitational attraction," interaction, mass-distance function) from location i to location j (alternatively, between those locations), D_{ij} is the "distance" between i and j, M_i and M_j are relevant economic "masses" of i and j, and a, b are parameters. Cf. Newton's law of universal gravitation in Chap. 24, where b = 2. The first instances were formulated by Reilly (1929), Stewart (1948), Isard (1956) and Tinbergen (1962).

If F_{ij} is a monetary flow (say, export values), then M is GDP (gross domestic product), and D_{ij} is the distance (usually the **great circle**

distance between the centers of countries i and j). For trade, the true distances are different and selected by economic considerations. But the distance is a proxy for transportation cost, the time elapsed during shipment, cultural distance, and the costs of synchronization, communication, transaction. The **distance effect on trade** is measured by the parameter b; it is 0.94 in Head (2003) and 0.6 in Leamer and Levinsohn (1994).

If F_{ij} is a people (travel or migration) or message flow, then M is the population size, and D_{ij} is the travel or communication cost (distance, time, money).

If F_{ij} is the force of attraction from location i to location j (say, for a consumer, or for a criminal), then, usually b = 2. Reilly's law of retail gravitation is that, given a choice between two cities of sizes M_i, M_j and at distances D_i, D_j , a consumer tends to travel further to reach the larger city with the equilibrium point defined by

$$\frac{M_i}{D_i^2} = \frac{M_j}{D_i^2}.$$

• Distance decay (in Spatial Interaction)

In general, distance decay or distance effect (cf. Chap. 29) is the attenuation of a pattern or process with distance.

In Spatial Interaction, **distance decay** is the mathematical representation of the inverse ratio between the quantity of obtained substance and the distance from its source.

This decay measures the effect of distance on accessibility and number of interactions between locations. For example, it can reflect a reduction in demand due to the increasing travel cost. A more abstract example is provided by bid-rent distance decay: the cost of overcoming distance has a consequence in a class-based spatial arrangement around a city. In fact, with increasing distance (and so decreasing rent) commercial, industrial, residential and agricultural areas follow.

In location planning for a service facility (fire station, retail store, transportation terminal, etc.), the main concerns are *coverage standard* (the maximum distance, or travel time, a user is willing to overcome to utilize it) and distance decay (demand for service decays with distance).

Distance decay is related to **gravity models** and another "social physics" notion, **friction of distance**, which posits that distance usually requires some amount of effort, money, and/or energy to overcome.

• Nearness principle

The **nearness principle** (or Zipf's least effort principle, in Psychology) is the following basic geographical heuristic: given a choice, a person will select the route requiring the least expenditure of effort. Similarly, an information-seeking person will tend to use the most convenient search method, in the least exacting mode available (path of least resistance).

The geographical nearness principle is used in transportation planning and (Rossmo 2000) locating of serial criminals: they tend to commit their crimes fairly close to where they live.

The first law of geography (Tobler 1970) is: "Everything is related to everything else, but near things are more related than distant things."

• Consumer access distance

Consumer access distance is a distance measure between the consumer's residence and the nearest provider where he can get specific goods or services (say, a store, market or a health service).

Measures of geographic access and spatial behavior include distance measures (**map distance**, **road travel distance**, perceived travel time, etc.), **distance decay** (decreased access with increasing distance) effects, transportation availability and *activity space* (the area in km² of $\approx \frac{2}{3}$ of the consumer's routine activities).

For example, US consumers had access in 2007, within driving distance of their current pharmacy, to 30, 7 and 14 competing pharmacies in urban, suburban and rural areas, respectively. (By Medicare standards, this distance, corresponding to the residence of the vast majority of beneficiaries, is 2, 5 and 15 miles in urban, suburban and rural areas.) Also, $\approx 81\%$ of the population of Texas is within 25 miles of a medical oncologist and radiation therapy facility.

Similar studies for retailers revealed that the negative effect of distance on store choice behavior was (for all categories of retailers) much larger when this behavior was measured as "frequency" than when it was measured as "budget share."

• Distance selling

Distance selling, as opposed to face-to-face selling in shops, covers goods or services sold without face-to-face contact between supplier and consumer but through *distance communication means*: press adverts with order forms, catalog sales, telephone, tele-shopping, e-commerce (via Internet), m-commerce (via mobile phone). Examples of the relevant legislation are Consumer Protection (Distance Selling) Directive 97/7/EC and Regulations 2000 in EU.

The main provisions are: clear prior information before the purchase, its confirmation in a durable medium, delivery within 30 days, "cooling-off" period of 7 working days during which the consumer can cancel contract without any reason and penalty. Exemptions are: *Distance Marketing* (financial services sold at distance), business-to-business contracts and some purchases (of land, at an auction, from vending machines).

• Surname distance model

A surname distance model was used in [COR05] in order to estimate the preference transmission from parents to children by comparing, for 47 provinces of mainland Spain, the 47×47 distance matrices for surname distance with those of consumption distance and cultural distance. The distances were l_1 -distances $\sum_i |x_i - y_i|$ between the frequency vectors

 (x_i) , (y_i) of provinces x, y, where z_i is, for the province z, either the frequency of the i-th surname (**surname distance**), or the budget share of the i-th product (**consumption distance**), or the population rate for the i-th cultural issue, say, rate of weddings, newspaper readership, etc., (**cultural distance**), respectively.

Other (matrices of) distances considered there are:

- Geographical distance (in kilometers, between the capitals of two provinces)
- Income distance |m(x) m(y)|, where m(z) is mean income in the province z
- Climatic distance $\sum_{1 \le i \le 12} |x_i y_i|$, where z_i is the average temperature in the province z during the i-th month
- Migration distance $\sum_{1 \le i \le 47} |x_i y_i|$, where z_i is the percentage of people (living in the province z) born in the province i

Strong *vertical preference transmission*, i.e., correlation between surname and consumption distances, was detected only for food items.

• Distances in Criminology

The **geographic profiling** (or *geoforensic analysis*) aims to identify the spatial behavior (target selection and, especially, likely *point of origin*, i.e., the residence or workplace) of a serial criminal offender as it relates to the spatial distribution of linked crime sites.

The **offender's buffer zone** (or *coal-sack effect*) is an area surrounding the *offender's heaven* (point of origin) from which little or no criminal activity will be observed; usually, such a zone occurs for premeditated personal offenses. The primary streets and network arterials that lead into the buffer zone tend to intersect near the estimated offender's heaven. An 1 km buffer zone was found for UK serial rapists. Most personal offenses occur within about 2 km from the offender's heaven, while property thefts occur further away.

Given n crime sites $(x_i, y_i), 1 \leq i \leq n$ (where x_i and y_i are the latitude and longitude of i-th site), the Newton-Swoope Model predicts the offender's heaven to be within the circle around the point $(\frac{\sum_i x_i}{n}, \frac{\sum_i y_i}{n})$ with the search radius being

$$\sqrt{\frac{\max|x_{i_1} - x_{i_2}| \cdot \max|y_{i_1} - y_{i_2}|}{\pi(n-1)^2}},$$

where the maxima are over (i_1, i_2) , $1 \le i_1 < i_2 \le n$. The Ganter-Gregory Circle Model predicts the offender's heaven to be within a circle around the first offense crime site with diameter the maximum distance between crime sites.

The *centrographic models* estimate the offender's heaven as a *center*, i.e., a point from which a given function of travel distances to all crime sites is

minimized; the distances are the Euclidean distance, the Manhattan distance, the **wheel distance** (i.e., the actual travel path), perceived travel time, etc. Many of these models are the reverse of Location Theory models aiming to maximize the placement of distribution facilities in order to minimize travel costs. These models (*Voronoi polygons*, etc.) are based on the **nearness principle** (*least effort principle*).

The **journey-to-crime decay function** is a graphical **distance curve** used to represent how the number of offenses committed by an offender decreases as the distance from his/her residence increases. Such functions are variations of the center of gravity functions; cf. **gravity models**.

For detection of criminal, terrorist and other hidden networks, there are also used many data-mining techniques which extract latent associations (distances and **near-metrics** between people) from proximity graphs of their co-occurrence in relevant documents, events, etc.

Drop distance

In judicial hanging, the **drop distance** is the distance the executed is allowed to fall. In order to reduce the prisoner's physical suffering (to about a third of a second), this distance is pre-determined, depending on his/her weight, by special *drop tables*. For example, (US state) Delaware protocol prescribes, in pounds/feet, about 252, 183 and 152 cm for at most 55, 77 and at least 100 kg.

In Biosystems Engineering, a ventilation jet *drop distance* is defined as the horizontal distance from an air inlet to the point where the jet reaches the occupational zone. In Aviation, an airlift *drop distance* (or *drop height*) is the vertical distance between the aircraft and the drop zone over which the airdrop is executed.

28.3 Distances in Perception, Cognition and Language

• Oliva et al. perception distance

Let $\{s_1, \ldots, s_n\}$ be the set of stimuli, and let q_{ij} be the conditional probability that a subject will perceive a stimulus s_j , when the stimulus s_i was shown; so, $q_{ij} \geq 0$, and $\sum_{j=1}^n q_{ij} = 1$. Let q_i be the probability of presenting the stimulus s_i .

The Oliva at al. perception distance [OSLM04] between stimuli s_i and s_j is defined by

$$\frac{1}{q_i + q_j} \sum_{k=1}^n \left| \frac{q_{ik}}{q_i} - \frac{q_{jk}}{q_j} \right|.$$

Visual space

Visual space refers to a stable perception (internal representation) of the environment provided by vision, while haptic space (or tactile space) and auditory space refer to such representation provided by the senses of pressure perception and audition. The geometry of these spaces and the eventual mappings between them are unknown. The main observed kinds of distortion of vision and haptic spaces versus physical space follow; the first three were observed for auditory space also:

- 1. *Horopter lines*: perceived frontparallel (to the observer) lines are physically parallel only at a certain distance depending on subject and task.
- 2. Parallel-alleys: perceived parallel (to the medial plane of the observer) lines are, actually, some hyperbolic curves.
- 3. Distance-alleys: lines with corresponding points perceived as equidistant, are, actually, some hyperbolic curves. Usually, the parallel-alleys are lying within the distance-alleys and, for visual space, their difference is small at distances larger than 1.5 m.
- 4. Oblique effects: performance of certain tasks is worse when the orientation of stimuli is oblique rather than horizontal or vertical.
- 5. Equidistant circles: the **egocentric distance** is direction-dependent; the points perceived as equidistant from the subject lie on egg-like curves instead of circles.

The above effects and **size-distance phenomena** should be incorporated in a good model of visual space. In a visual space the distance d and direction are defined from self as the origin (the **egocentric distance**). There is evidence that visual space is almost affine and, if it admits a metric d, then d is a **projective metric**, i.e., d(x,y) + d(y,z) = d(x,z) for any three perceptually collinear points x, y, z. The main proposals for visual space are to see it as a Riemannian space of constant negative curvature (cf. **Riemannian color space** in Chap. 21), a general Riemannian/Finsler space, or an affinely connected (so not metric, in general) space [CKK03].

An *affine connection* is a linear map sending two vector fields into a third one. The expansion of perceived depth on near distances and its contraction at far distances indicate that the mapping between visual and physical space is not affine.

• Size-distance phenomena

Examples of size-distance phenomena of visual perception follow.

Emmert's size-distance law states that a retinal image is proportional in perceived size (apparent height) to the perceived distance of the surface it is projected upon. This law is based on the fact that the perceived size of an object doubles every time its perceived distance from the observer is cut in half and vice versa. Emmert's law accounts for *constancy scaling*, i.e., the fact that the size of an object is perceived to remain constant despite changes in the retinal image (as objects become more distant they begin, because of visual perspective, to appear smaller).

The **size-distance invariance hypothesis** posits that the ratio of perceived size and perceived distance is the tangent of the physical visual angle. In particular, objects which appear closer should also appear

smaller. But with the *moon illusion* it comes to a **size-distance paradox**. The Moon (and, similarly, the Sun) illusion is that, despite constancy of its visual angle (roughly, 0.52°), the moon at the horizon may appear to be about twice the diameter of the zenith moon. This illusion is still not understood completely; it is supposed to be cognitive: the size of the zenith moon is underestimated since it is perceived as approaching. The most common optical illusions distort size or length; for example, the Mueller-Lyer, Sander, and Ponzo illusions.

The **size-distance centration** is the overestimation of the size of objects located near the focus of attention and underestimation of it at the periphery.

• Probability-distance hypothesis

In Psychophysics, the **probability-distance hypothesis** is a hypothesis that the probability with which one stimulus is discriminated from another is a (continuously increasing) function of some subjective quasimetric between these stimuli (see [Dzha01]). Under this hypothesis, such a subjective metric is a **Finsler metric** if and only if it coincides in the small with the **intrinsic metric** (i.e., the infimum of the lengths of all paths connecting two stimuli).

• Distance ceptor

A distance ceptor is a nerve mechanism of one of the organs of special sense whereby the subject is brought into relation with his distant environment.

• Egocentric distance

In Psychophysiology, the **egocentric distance** is the perceived absolute distance from the self (observer or listener) to an object or a stimulus (such as a sound source); cf. **subjective distance**. Usually, the visual egocentric distance underestimates the actual physical distance to far objects, and overestimates it for near objects. Such distortion is direction-dependent: it decreases in a lateral direction.

In Visual Perception, the action space of a subject is 1–30 m; the smaller and larger spaces are called the personal space and vista space, respectively.

The **exocentric distance** is the perceived relative distance between objects.

• Distance cues

The distance cues are cues used to estimate the egocentric distance.

For a listener at a fixed location, the main auditory distance cues include: intensity (in open space it decreases by 5 dB for each doubling of the distance; cf. far field **acoustic distance** in Chap. 21), direct-to-reverberant energy ratio (in the presence of sound reflecting surfaces), spectrum, and binaural differences.

For an observer, the main visual distance cues include:

- Relative size, relative brightness, light and shade
- Height in the visual field (in the case of flat surfaces lying below the level of the eye, the more distant parts appear higher)

- Interposition (when one object partially occludes another from view)
- Binocular disparities, convergence (depending on the angle of the optical axes of the eyes), accommodation (the state of focus of the eyes)
- Aerial perspective (distant objects become bluer and paler), distance hazing (distant objects become decreased in contrast, more fuzzy)
- Motion perspective (stationary objects appear to a moving observer to glide past)

Examples of the techniques which use the above distance cues to create an optical illusion for the viewer, are:

- Distance fog: a 3D computer graphics technique so that objects further from the camera are progressively more blurred (obscured by haze).
 It is used, for example, to disguise too short draw distance, i.e., the maximal distance in a 3D scene that is still drawn by the rendering engine.
- Forced perspective: a film-making technique to make objects appear either far away, or nearer depending on their positions relative to the camera and to each other.

• Subjective distance

The **subjective distance** (or *cognitive distance*) is a mental representation of actual distance molded by an individual's social, cultural and general life experiences; cf. **egocentric distance**. Cognitive distance errors occur either because information about two points is not coded/stored in the same branch of memory, or because of errors in retrieval of this information. For example, the length of a route with many turns and landmarks is usually overestimated.

• Geographic distance biases

Sources of distance knowledge are either symbolic (maps, road signs, verbal directions) or directly perceived ones during locomotion: environmental features (visually-perceived turns, landmarks, intersections, etc.), travel time, and travel effort.

They relate mainly to the perception and cognition of the, **environmental distances** i.e., those that cannot be perceived in entirety from a single point of view but still can be apprehended through direct travel experience.

Examples of **geographic distance biases** (subjective distance judgments, location estimates) are:

- Observers are quicker to respond to locations preceded by locations that were either close in distance or were in the same region.
- Distances are overestimated when they are near to a reference point;
 for example, intercity distances from coastal cities are exaggerated.
- Subjective distances are often assymmetrical as the perspective varies with the reference object: a small village is considered to be close to a big city while the big city is likely to be seen as far away from it.

- Traveled routes segmented by features are subjectively longer than unsegmented routes; moreover, longer segments are relatively underestimated.
- Increasing the number of pathway features encountered and recalled by subjects leads to increased distance estimates.
- Structural features (such as turns and opaque barriers) breaking a
 pathway into separate vistas strongly increase subjective distance (suggesting that distance knowledge may result from a process of summing
 vista distances).
- Chicago-Rome illusion: belief that some European cities are located far to the south of their actual location; in fact, Chicago and Rome are at the same latitude (42°), as are Philadelphia and Madrid (40°), etc.
- Miami-Lima illusion: belief that cities on the east coast of US are located to the east of cities on the west coast of South America; in fact, Miami is 3° west of Lima.

Possible sources of such illusions could be perceptually based mental representations that have been distorted through normalization and/or conceptual non-spatial plausible-reasoning.

• Spatial reasoning

Spatial reasoning is the domain of spatial knowledge representation: spatial relations between spatial entities, and reasoning based on these entities and relations.

As a modality of human thought, spatial reasoning is a process of forming ideas through the spatial relationships between objects (as in Geometry), while verbal reasoning is the process of forming ideas by assembling symbols into meaningful sequences (as in Language, Algebra, Programming).

Spatial-temporal reasoning is the ability to visualize spatial patterns and mentally manipulate them over a time-ordered sequence of spatial transformations. In human-computer interaction, differences in such ability lead to certain users performing more efficiently than others at information search and information retrieval.

More specifically, *spatial visualization ability* is the ability to manipulate mentally two- and three-dimensional figures.

Visual thinking (or visual/spatial learning, picture thinking) is the common (about 60% of the general population) phenomenon of thinking through visual processing. Spatial-temporal reasoning is prominent among visual thinkers, as well as among kinesthetic learners (who learn through body mapping and physical patterning) and logical thinkers (mathematical/systems thinking) who think in patterns and relationships and may work diagrammatically without this necessary being pictorially.

In Computer Science, spatio-temporal reasoning aims at describing, using abstract relation algebras, the common-sense background knowledge on which human perspective of physical reality is based. It provides rather inexpensive reasoning about entities located in space and time.

• Spatial language

Spatial language consists of natural-language spatial relations used to indicate where things are, and so to identify or refer to them.

Among spatial relations there are *topological* (such as on, to, in, inside, at), *path-related* (such as across, through, along, around), *distance-related* and more complex ones (such as right/left, between, opposite, back of, south of, surround).

A distance relation is a spatial relation which specifies how far the object is away from the reference object: near, far, close, etc.

The distance concept of proximity (Pribbenow 1992) is the area around the RO (reference object) in which it can be used for localization of the LO (local object), so that there are visual access from RO and non-interruption of the spatial region between objects, while LO is less directly related to a different object. Such proximity can differ with physical distance as, for example, in "The Morning Star is to the left of the church." The area around RO, in which a particular relation is accepted as a valid description of the distance between objects, is called the *acceptance area*.

Pribbenow (1991) proposed five distance distinctions: inclusion (acceptance area restricted to projection of RO), contact/adjacency (immediate neighborhood of RO), proximity, geodistance (surroundings of RO) and remoteness (the complement of the proximal region around RO).

Jackendorff and Landau (1992) showed that in English there are mainly 3 degrees of distance distinctions: interior of RO (in, inside), exterior but in contact (on, against), proximate (near), plus corresponding negatives (such as outside, off of, far from).

Semantics of spatial language is considered in Spatial Cognition, Linguistics, Cognitive Psychology, Anatomy, Robotics, Artificial Intelligence and Computer vision. Cognitively based common-sense spatial ontology and metric details of spatial language are modeled for eventual interaction between Geographic Information Systems and users. An example of far-going applications is the Grove's Clean Space, a neuro-linguistic programming psychotherapy based on the spatial metaphors produced by (or extracted from) the client on his present and desired "space" (state).

• Language distance from English

There are many such measures based either on a typology (comparing formal similarities between languages), or language trees, or performance (mutual intelligibility and learnability of languages). Some examples of language distance from English follow.

Rutheford (1983) defined distance from English as the number of differences from English in the following three-way typological classification: subject/verb/object order, topic-prominence/subject-prominence and pragmatic word-order/grammatical word-order. It gives distances 1, 2, 3 for Spanish, Arabic/Mandarin, Japanese/Korean.

Borland (1983) compared several languages of immigrants by their acquisition of four areas of English syntax: copula, predicate complemen-

tation, negation and articles. The resulting ranking was English, Spanish, Russian, Arabic, Vietnamese.

Elder and Davies (1998) used ranking based on the following three main types of languages: isolating, analytic or root (as Chinese, Vietnamese), inflecting, synthetic or fusional (as Arabic, Latin, Greek), agglutinating (as Turkish, Japanese). It gave ranks 1, 2, 4, 5 for Romance, Slavic, Vietnamese/Khmer, Japanese/Korean, respectively, and the intermediate rank 3 for Chinese, Arabic, Indonesian, Malay.

The language distance index (Chiswick and Miller 1998) is the inverse of the language score of the average speaking proficiency (after 24 weeks of instruction) of English speakers learning this language (or, say, fluency in English of immigrants having it as native language). This score was measured by a standardized test at regular intervals by increments of 0.25; it ranges from 1.0 (hardest to learn) to 3.0 (easiest to learn). The score was, for example, 1.00, 1.25, 1.50, 1.75, 2.00, 2.25, 2.50, 2.75, 3.00 for Japanese, Cantonese, Mandarin, Hindi, Hebrew, Russian, French, Dutch, Afrikaans.

In addition to the above distances, based on syntax, and **linguistic distance**, based on pronunciation, see the lexical semantic distances in Sect. 22.2.

Cf. clarity similarity in Chap. 14, distances between rhythms in Chap. 21, Lasker distance in Chap. 23 and surname distance model.

• Editex distance

The main phonetic encoding algorithms are (based on English language pronunciation) *Soundex*, *Phonix* and *Phonex*, converting words into one-letter three-digits codes. The letter is the first one in the word and the three digits are derived using an assignment of numbers to other word letters. Soundex and Phonex assign:

0 to a, e, h, i, o, u, w, y; 1 to b, p, f, v; 2 to c, g, j, k, q, s, x, z; 3 to d, t; 4 to l; 5 to m, n; 6 to r.

Phonix assigns the same numbers, except for 7 (instead of 1) to f and v, and 8 (instead of 2) to s, x, z.

The **Editex distance** (Zobel and Dart 1996) between two words x and y is a cost-based **editing metric** (i.e., the minimal cost of transforming x into y by substitution, deletion and insertion of letters). For substitutions, the costs are 0 if two letters are the same, 1 if they are in the same letter group, and 2 otherwise.

The syllabic alignment distance (Gong and Chan 2006) between two words x and y is another cost-based **editing metric**. It is based on Phonix, the identification of syllable starting characters and seven edit operations.

• Phone distances

A *phone* is a sound segment that has distinct acoustic properties, and is the basic sound unit. Cf. *phoneme*, i.e., a family of phones that speakers usually hear as a single sound; the number of phonemes range, among

about 6,000 languages spoken now, from 11 in Rotokas to 112 in !Xóõ (languages spoken by about 4,000 people in Papua New Guinea and Botswana, respectively).

The two main classes of **phone distance** (distances between two phones x and y) are:

- 1. Spectrogram-based distances which are physical-acoustic distortion measures between the sound spectrograms of x and y
- 2. Feature-based phone distances which are usually the Manhattan distance $\sum_i |x_i y_i|$ between vectors (x_i) and (y_i) representing phones x and y with respect to a given inventory of phonetic features (for example, nasality, stricture, palatalization, rounding, sillability)

The Laver consonant distance refers to the improbability of confusing 22 consonantal phonemes of English, developed by Laver (1994) from subjective auditory impressions. The smallest distance, 15%, is between phonemes [p] and [k], the largest one, 95%, is, for example, between [p] and [z]. Laver also proposed a quasi-distance based on the likelihood that one consonant will be misheard as another by an automatic speech-recognition system.

Liljencrans and Lindlom (1972) developed a vowel space of 14 vowels. Each vowel, after a procedure maximizing contrast among them, is represented by a pair (x,y) of resonant frequencies of the vocal tract (first and second formants) in linear mel units with $350 \le x \le 850$ and $800 \le y \le 1,700$. Roughly, higher x values correspond to lower vowels and higher y values to less rounded or farther front vowels. For example, [u], [a], [i] are represented by (350, 800), (850, 1,150), (350, 1,700), respectively.

Phonetic word distance

The **phonetic word distance** (or pronunciation distance) between two words x and y is the **Levenstein metric** with costs, i.e., the minimal cost of transforming x into y by substitution, deletion and insertion of phones. A word is seen as a string of phones.

Given a **phone distance** r(u, v) on the International Phonetic Alphabet with the additional phone 0 (the silence), the cost of substitution of phone u by v is r(u, v), while r(u, 0) is the cost of insertion or deletion of u. Cf. in Sect. 23.3 distances on the set of 20 amino acids.

• Linguistic distance

The linguistic distance (or dialectology distance) between language varieties X and Y is the mean, for fixed sample S of notions, **phonetic** word distance between cognate (i.e., having the same meaning) words s_X and s_Y , representing the same notion $s \in S$ in X and Y, respectively. Usually, the **Levenshtein metric** defined in Chap. 11 is used (the minimum number of inserting, deleting or substituting sounds needed to recover the word pronunciation).

As an example of similar work, the **Stover distance** between phrases with the same key word is (Stover 2005) the sum $\sum_{-n \le i \le +n} a_i x_i$, where $0 < a_i < 1$, and x_i is the proportion of non-matched words between the

phrases within a moving window. Phrases are first aligned, by the common key word, to compare the uses of it in context; also, the rarest words are replaced with a common pseudo-token.

• Language distance effect

In Foreign Language Learning, Corder (1981) conjectured the existence of the following language distance effect: where the mother tongue (L1) is structurally similar to the target language, the learner will pass more rapidly along the developmental continuum (or some parts of it) than where it differs; moreover, all previous learned languages have a facilitating effect.

Ringbom (1987) added: the influence of the L1 is stronger at early stages of learning, at lower levels of proficiency and in more communicative tasks.

But such correlation could not be direct. For example, the written form of Chinese does not vary among the regions of China, but the spoken languages differ sharply. Alternatively, spoken German and Yiddish are close but have different alphabets.

• Long-distance dependence (in Language)

In Language, **long-distance dependence** is a construction – including diverse wh-questions (who, what, where, etc.), relative clauses and topicalization) – which permits an element in one position to fill the grammatical role associated with another position.

In Generative Linguistics, anaphora is a reciprocal (such as one another and each other) or reflexive (such as myself, herself, themselves, oneself, etc.) pronoun in English, or an analogous referential pattern in other language. In order to be interpreted, anaphora must get its content from an antecedent in the sentence, which is, usually, syntactically local as in "Mary excused herself." A long-distance anaphora is an anaphora with antecedent outside of its local domain, as in "The players told us stories about each other." Its resolution (finding what the anaphora refer to) is an unsolved linguistic problem of machine translation.

Cf. long range dependence in Chap. 29.

28.4 Distances in Philosophy, Religion and Art

• Kristeva non-metric space

Kristeva's (1980) basic psychoanalytic distinction is between pre-Oedipal and Oedipal aspects of personality development. Narcissistic identification and maternal dependency, anarchic component drives, polymorphic erotogenicism, and primary processes characterize the pre-Oedipal. Paternal competition and identification, specific drives, phallic erotogenicism, and secondary processes characterize Oedipal aspects. Kristeva describes the pre-Oedipal feminine phase by an enveloping, amorphous, **non-metric space** (Plato's *chora*) that both nourishes and threatens; it also defines

and limits self-identity. She characterizes the Oedipal male phase by a metric space (Aristotle's *topos*); the self and the self-to-space are more precise and well defined in topos. Kristeva posits also that the semiotic process is rooted in feminine libidinal, pre-Oedipal energy which needs channeling for social cohesion.

Deleuze and Guattari (1980) divided their *multiplicities* (networks, manifolds, spaces) into *striated* (metric, hierarchical, centered and numerical) and *smooth* ("non-metric, rhizomic and those that occupy space without counting it and can be explored only by legwork").

The above French post-structuralists use the metaphor non-metric in line with a systematic use of topological terms by the psychoanalyst Lacan. In particular, he sought the space J (of Jouissance, i.e., sexual relations) as a bounded metric space.

Back to Mathematics, the **non-metricity tensor** is the *covariant derivative* of a **metric tensor**. It can be non-zero for **pseudo-Riemannian metrics** and vanishes for **Riemannian metrics**.

Simone Weil distance

We call the **Simone Weil distance** a kind of moral radius of the Universe which the French philosopher, Christian mystic, social activist and self-hating Jew, Simone Weil (1909–1943) introduced in "The Distance," one of the philosophico-theological essays comprising her *Waiting for God* (posthumous English edition by Putnam, New York 1951).

She connects God's love to the distance; so, his absence can be interpreted as a presence: "every separation is a link" (Plato's metaxu). In her peculiar Christian theodicy, "evil is the form which God's mercy takes in this world," and the crucifixion of Christ (the greatest love/distance) was necessary "in order that we should realize the distance between ourselves and God... for we do not realize distance except in the downward direction." The Simone Weil God-cross distance (or Creator-creature distance) recalls the old question: can we equate distance from God with proximity to Evil? Her main drive, the purity, consisted of maximizing moral distance to Evil, embodied for her by "the social, Rome and Israel."

Cf. Pascal's God-nothing distance in Pensée, note 72: "For after all what is man in nature? A nothing in relation to infinity, all in relation to nothing, a central point between nothing and all and infinitely far from understanding either." Cf. also Tipler's (2007) Big Bang – Omega Point time/distance with Initial and Final Singularities seen as God-Father and God-Son. Omega point is (Teilhard de Chardin 1950) the supreme point of complexity and consciousness: the Logos, or Christ.

The Weil approach reminds also of the Lurian kabbalistic notions: tzimtzum (God's concealment, withdrawal of a part, creation by self-delimitation) and shattering of the vessels (evil as impure vitality of husks, produced whenever the force of separation loses its distancing function, and giving man the opportunity to choose between good and evil). The purpose is to bridge the distance between God (or Good) and the diversity of existence, without falling into the facility of dualism (as manicheanism

and gnosticism). It is done by postulating intermediate levels of being (and purity) during emanation (unfolding) within the divine and allowing humans to participate in the redemption of the Creation.

So, a possible individual response to the Creator is purification and ascent, i.e., a spiritual movement through the levels of emanation in which the coverings of impurity, that create distance from God, are removed progressively.

Meanwhile, a song "From a Distance," written by Julie Gold, is about how God is watching us and how, despite the distance (physical and emotional) distorting perceptions, there is still a little peace and love in this world.

• Distance to Heaven

Below are given examples of distances and lengths which old traditions related (sometimes as a metaphor) to such notions as God and Heaven.

In the early Hebrew mystical text *Shi'ur Qomah*, i.e., *The measure* of the (divine) body, the height of the Holy Blessed One is 236×10^7 parasangs, i.e., 14×10^{10} (divine) spans. In the Biblical verse "Who has measured the waters in the hollow of his hand and marked off the heaven with a span" (Isaiah 40:12), the size of the Universe is one such span.

The **cosmic light horizon** (or age/size of the Universe) is ≈ 13.7 billion light-years. Sefer HaTemunah (by Nehunia ben Hakane, first century) and Otzar HaChaim (by Yitzchok deMin Acco, thirteenth century) deduced that the World was created in thought 42,000 divine years, i.e., $42,000\times365,250\approx15.3$ billion human years, ago. This exegesis counts, using a 42-letter name at the start of Genesis, that now we are in the sixth cycle of the seven cosmic sh'mitah cycles, each one being 7,000 divine years long. Tohu and bohu followed and less than 6,000 years ago the creation of world in deed is posited.

In the Talmud (Pesahim, 94), the Holy Spirit points out to "impious Nebuchadnezzar" (planning "to ascend above the heights of the clouds like the Most High"): "The distance from earth to heaven is 500 year's journey alone, the thickness of the heaven again 500 years" This heaven is the firmament plate, and the journey is by walking. Seven other heavens, each 500 years thick, follow "and the feet of the holy Creatures are equal to the whole...." Their ankles, wings, necks, heads and horns are each consecutively equal to the whole. Finally, "upon them is the Throne of Glory which is equal to the whole." The resulting journey of 4,096,000 years amounts, at the rate of 80 miles per day, at about $\approx 2,600$ AU $\approx \frac{1}{100}$ of the distance to the nearest other star.

On the other hand, *Baraita de Massechet Gehinom* affirms in Sect. VII.2 that Hell consists of seven cubic regions of side 300 year's journey each; so, 6,300 years altogether.

Islamic tradition (Dawood, Book 40, Nr. 470) also attributes a journey of 71–500 years (by horse, camel or foot) between each *asmaa* (the ceiling containing one of the seven luminaries: Moon, Mercury, Venus, Sun, Mars, Jupiter, Saturn).

Vedic texts (Pancavimsa-Brahmana, circa $2000\,\mathrm{BC}$) states that the distance to Heaven is 1,000 Earth diameters and Sun (the middle one among seven luminaries) is halfway at 500 diameters. A similar ratio 500–600 was expected till the first scientific measurement of 1 AU (Earth–Sun distance) by Cassini in 1672. The actual ratio is $\approx 11,687$.

The sacred Hindu number $108 \ (=6^2+6^2+6^2=\prod_{1\leq i\leq 3}i^i)$, also connected to Golden Ratio as the interior angle 108^0 of a regular pentagon, is traced to following Vedic values: $108 \ \text{Sun's}$ diameters for the Earth–Sun distance and $108 \ \text{Moon's}$ diameters for the Earth–Moon distance. The actual values are ≈ 107.6 and (steadily increasing) ≈ 110.6 ; they could be computed without any instruments during an eclipse, since the angular size of Moon and Sun, viewed from Earth, is almost identical. Cf. the Metonic cycle (period of 19 tropical years, i.e., 6,939.60 days, that is equal to 235 synodic lunar months, $\approx 29.53 \ \text{days}$ each, plus about 2h) and the Saros cycle (period of 223 sinodic months $\approx 6,585.33 \ \text{days}$) that can be used to predict eclipses of the Sun and Moon.

Also, the ratio between Sun and Earth diameters is ≈ 108.6 , but it is unlikely that Vedic sages knew it. In Ayurveda, the devotee's distance to his "inner sun" (God within) consists of 108 steps; it corresponds to 108 beads of *japamala* (rosary): the devotee, while saying beads, does a symbolic journey from his body to Heaven.

• Swedenborg heaven distances

The Swedish scientist and visionary Emanuel Swedenborg (1688–1772), in Sect. 22 (Nos. 191–199, Space in Heaven) of his main work Heaven and Hell (1952, first edition in Latin, London 1758), posits: "distances and so, space, depend completely on interior state of angels." A move in heaven is just a change of such a state, the length of a way corresponds to the will of a walker, approaching reflects similarity of states. In the spiritual realm and afterlife, for him, "instead of distances and space, exist only states and their changes."

• Space (in Philosophy)

The present Newton–Einstein notion of **space** were preceded by the Aristotelian Cosmos (space is a finite system of relations between material objects) and earlier, in the same fourth century BC, by Democritus Void (space is the infinite container of objects).

For Newton, space was absolute: it existed permanently and independently of whether there is any matter in it. For Leibniz (in the same seventeenth century), space was a collection of relations between objects, given by their distance and direction from one another, i.e., an idealized abstraction from the relations between individual entities or their possible locations, which must therefore be discrete.

For Kant (eighteenth century), space and time are not objective features of the world, with substance or relation. Instead, they are part of an unavoidable systematic framework used by humans to organize their experiences.

Disagreement continues between philosophers over whether space is an entity, a relationship between entities, or part of a conceptual framework.

• Quotes on "near-far" distances

"Better is a nearby neighbor, than a far off brother." (the Bible)

"It is when suffering seems near to them that men have pity; as for disasters that are ten thousand years off in the past or the future, men cannot anticipate them, and either feel no pity for them, or at all events feel it in no comparable measure." (Aristotle)

"The path of duty lies in what is near, and man seeks for it in what is remote." (Mencius)

"Sight not what is near through aiming at what is far." (Euripides)

"Good government occurs when those who are near are made happy, and those who are far off are attracted." (Confucius)

"By what road," I asked a little boy, sitting at a cross-road, "do we go to the town?" – "This one," he replied, "is short but long and that one is long but short." I proceeded along the "short but long road." When I approached the town, I discovered that it was hedged in by gardens and orchards. Turning back I said to him, "My son, did you not tell me that this road was short?" – "And," he replied, "Did I not also tell you: "But long"? I kissed him upon his head and said to him, "Happy are you, O Israel, all of you are wise, both young and old." (Erubin, Talmud)

The Prophet Muhammad was heard saying: "The smallest reward for the people of paradise is an abode where there are 80,000 servants and 72 wives, over which stands a dome decorated with pearls, aquamarine, and ruby, as wide as the distance from Al-Jabiyyah [a Damascus suburb] to Sana'a [Yemen]." (Hadith, Islamic Tradition)

"There is no object so large ... that at great distance from the eye it does not appear smaller than a smaller object near." (Leonardo da Vinci)

"Nothing makes Earth seems so spacious as to have friends at a distance; they make the latitudes and longitudes." (Henri David Thoreau)

"In true love the smallest distance is too great, and the greatest distance can be bridged." (Hans Nouwens)

"There is an immeasurable distance between late and too late." (Og Mandino) $\,$

"Everything is related to everything else, but near things are more related than distant things." (Tobler's first law of Geography).

Cf. nearness principle in Chap. 28 and, in Chaps. 22 and 24, action at a distance.

• Antinomy of distance

The antinomy of distance, as introduced in [Bull12] for aesthetic experiences by beholder and artist, is that both should find the right amount of emotional distance, defined in Chap. 28 (neither too involved, nor too detached), in order to create or appreciate art. The fine line between objectivity and subjectivity can be crossed easily, and the amount of distance can fluctuate in time.

The **aesthetic distance** is a degree of emotional involvement of the individual, who undergoes experiences and objective reality of the art, in a work of art. It means the frame of reference that an artist creates, by the use of technical devices in and around the work of art, to differentiate it psychologically from reality; cf. **distanciation**. Some examples are: the perspective of a member of the audience in relation to the performance, the psychological and the emotional distance between the text and the reader, the *actor-character distance* in the Stanislavsky system of acting.

Morgan [Morg76] defines pastoral ecstasy as the experience of *role-distancing*, or the authentic self's supra-role suspension, i.e., the capacity of an individual to stand outside or above himself for purposes of critical reflexion. Morgan concludes: "The authentic self is an *ontological possibility*, the social self is an *operational inevitability*, and awareness of both selves and the creative coordination of both is the gift of ecstasy. Interplay of proximity and distance to the Other is central also in Levinas ethics."

A variation of the antinomy of distance appears in critical thinking: the need to put some emotional and intellectual distance between oneself and ideas, in order to better evaluate their validity. Another variation is detailed in *Paradox of Dominance: Distance and Connection* (posting on http://www.leatherpage.com) by Sprott.

The **historical distance**, in terms of [Tail04], is the position the historian adopts *vis-à-vis* his objects – whether far-removed, up-close, or somewhere in between; it is the fantasy through which the living mind of the historian, encountering the inert and unrecoverable, positions itself to make the material look alive. The antinomy of distance appears again because historians engage the past not just intellectually but morally and emotionally. Historical knowledge is always mediate/inferential, never empirical/perceptual. The formal properties of historical accounts are influenced by the affective, ideological and cognitive commitments of their authors.

A related problem is how much distance people must put between themselves and their pasts in order to remain psychologically viable; Freud showed that often there is no such distance with childhoods.

• Distanciation

In scenic art and literature, **distanciation** (Althusser 1968, on Brecht's alienation effect) consists of methods to disturb purposely (in order to challenge basic codes and conventions of spectator/reader) the narrative contract with him, i.e., implicit clauses defining logic behavior in a story. The purpose is to differentiate art psychologically from reality, i.e., to create some **aesthetic distance**.

One of distanciation devices is breaking of the fourth wall, when the actor/author addresses the spectators/readers directly through an imaginary screen separating them. The "fourth wall" is the conventional boundary between the fiction and the audience. It is a part of the suspension of disbelief between them: the audience tacitly agrees to provisionally suspend their judgment in exchange for the promise of entertainment.

Cf. distancing and distantiation.

• Far Near Distance

Far Near Distance is the name of the program of the House of World Cultures in Berlin which presents a panorama of contemporary positions of all artists of Iranian origin. Some examples of similar use of distance terms in modern popular culture follow.

"Some near distance" is the title of an art exhibition of Mark Lewis (Bilbao 2003), "A Near Distance" is a paper collage by Perle Fine (New York, 1961), "Quiet Distance" is a fine art print by Ed Mell, "Zero/Distance" is the title of an art exhibition of Jim Shrosbree (Des Moines, Iowa, 2007).

"Distance" is a Japanese film directed by Hirokazu Koreeda (2001) and an album of Utada Hikaru (her famous ballad is called "Final Distance"). It is also the stage name of a British musician Greg Sanders and the name of the late-1980 rock/funk band led by Bernard Edwards. "The Distance" is a film directed by Benjamin Busch (2000) and an album by American rock band "Silver Bullet" led by Bob Seger. "Near Distance" is a musical composition by Chen Yi (New York, 1988) and lyric by the Manchester quartet "Puressence."

The terms $near\ distance$ and $far\ distance$ are also used in Ophthalmology and for settings in some sensor devices.

Chapter 29 Other distances

In this chapter we group together distances and distance paradigms which do not fit in the previous chapters, being either too practical (as in equipment), or too general, or simply hard to classify.

29.1 Distances in Medicine, Antropometry and Sport

• Distances in Medicine

Some examples from this vast family of physical distances follow.

In Dentistry, the **inter-occlusal distance**: the distance between the occluding surfaces of the maxillary and mandibular teeth when the mandible is in a physiologic rest position.

The **inter-arch distance**: the vertical distance between the maxillary and mandibular arches. The **inter-ridge distance**: the vertical distance between the maxillary and mandibular ridges.

The **inter-proximal distance**: the **spacing distance** between adjacent teeth; *mesial drift* is the movement of the teeth slowly toward the front of the mouth with the decrease of the inter-proximal distance by wear.

The **inter-pediculate distance**: the distance between the vertebral pedicles as measured on the radiograph.

The **teardrop distance**: the distance from the lateral margin of the pelvic teardrop to the most medial aspect of the femoral head as seen on the anteposterior pelvic radiograph. A widening of at least 1 mm indicates excess hip joint fluid and so inflammation and other joint abnormalities.

The **source-skin distance**: the distance from the focal spot on the target of the X-ray tube to the skin of the subject as measured along the central ray.

The inter-aural distance: the distance between the ears. The inter-ocular distance: the distance between the eyes.

The **anogenital distance** (or AGD): the length of the *perineum*, i.e., the region between anus and genital area (the anterior base of the penis

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for a male). For a male it is normally twice what it is for a female; so this distance is a measure of physical masculinity (used in reproductive toxicity testing).

In Anesthesia, the **thyromental distance** (or TMD): the distance from the upper edge of the thyroid cartilage (laryngeal notch) to the mental prominence (tip of the chin) when the neck is extended fully. The **sternomental distance**: the distance from the upper border of the manubrium sterni to the tip of the chin, with the mouth closed and the head fully extended. When the above distances are less than 6.5 cm and 12.5 cm, respectively, a difficult intubation is indicated.

The **sedimentation distance** (or ESR, *erythrocyte sedimentation rate*): the distance red blood cells travel in 1 h in a sample of blood as they settle to the bottom of a test tube. ESR indicates inflammation and increases in many diseases.

The **stroke** distance: the distance a column of blood moves during each heart beat, from the aortic valve to a point on the arch of the aorta.

The distance between the lesion and the aortic valve being <6 mm, is an important predictor, available before surgical resection of DSS (discrete subaortic stenosis), of re-operation for recurrent DSS.

The **aortic diameter**: the maximum diameter of the outer contour of the aorta. It (as well as the cross sectional diameter of the left ventricle) varies between the ends of the *systole* (the time of ventricular contraction) and *diastole* (the time between those contractions); the *strain* is the ratio between the systolic and diastolic diameters.

The **dorsoventral interlead distance** of an implanted pacemaker or defibrillator: the horizontal separation of the right and left ventricular lead tips on the lateral chest radiograph, divided by the *cardiothoracic ratio* (ratio of the cardiac width to the thoracic width on the posteroanterior film).

In Laser Treatments, the **extinction length** and **absorption length** of the vaporizing beam is the distance into the tissue from the incident surface along the ray path over which 90% (or 99%) and 63%, respectively, of its radiant energy is absorbed.

In Ophthalmic Plastic Surgery, the marginal reflex distances MRD_1 and MRD_2 are the distances from the center of the pupil (identified by the corneal reflex created by shining a light on the pupil) to the margin of the upper or lower eyelid, while the **vertical palperbal fissure** is the distance between the upper and lower eyelid.

The main distances used in Ultrasound Biomicroscopy (for glaucoma treatment) are the **angle-opening distance** (from the corneal endothelium to the anterior iris) and the **trabecular ciliary process distance** (from a particular point on the *trabecular mesh-work* to the *ciliary process*).

Diffusion MRI is a modality of Magnetic Resonance Imaging producing noninvasively in vivo images of brain tissues weighted by their water diffusivity. The image-intensities at each position are attenuated proportionally to the strength of diffusion in the direction of its gradient. Diffusion in tissues, because of its *anisotropy* (dependence on direction), is described by a tensor instead of a diffusivity scalar. Tensor data are displayed, for each voxel, by ellipsoids; their length in any direction is the diffusion distance molecules cover in given time in this direction. The **diffusion tensor distance** is the length from the center to the surface of the diffusion tensor.

In brain MRI, the distances considered for *cortical maps* (i.e., outer layer regions of cerebral hemispheres representing sensory inputs or motor outputs) are: MRI distance map from the GW (gray/white matter) interface, cortical distance (say, between activation locations of spatially adjacent stimuli), cortical thickness (the shortest distance between the GW-boundary and the innermost surface of *pia matter* enveloping the brain) and *lateralization metrics*. Anderson (1996) found that the cortical thickness of Einstein's brain is 2.1 mm, while the average one is 2.6 mm; resulting closer packing of cortical neurons may speed up communication between them.

In Nerve Regeneration by transplantation of cultured stem cells, the **regeneration distance** is the distance between the point of insertion of the proximal stump and the tip of most distal regenerating axon.

• Distances in Oncology

In Oncology, the **tumor radius** is the mean radial distance R from the tumor origin (or its center of mass) to the tumor–host interface (the tumor/cell colony border); the cell proliferation along [0, R] is ≈ 0 up to some r_0 , then increases linearly in r for $r_0 < r \le r_1 < R$, and it happens mainly within the outermost band $[r_1, R]$.

The **tumor diameter** is the greatest vertical diameter of any section; the *tumor growth* is the geometric mean of its three perpendicular diameters.

In Oncological Surgery, the **margin distance**: the tumor-free surgical margin (after formalin fixation) of tumor resection, done in order to prevent local recurrence; Chan et al. (2007) assert that, for vulvar cancer, $\geq 8 \,\mathrm{mm}$ margin clearance is sufficient, instead of 2–3 cm recommended previously.

The **perfusion distance**: the shortest distance between the infusion outlet and the surface of the electrodes (during radiofrequency tumor ablation with internally cooled electrodes and saline infusion).

In Radiation Oncology, the maximum heart distance MHD is the maximum distance of the heart contour (as seen in the beam's eye view of the medial tangential field) to the medial field edge, and the central lung distance CLD is the distance from the dorsal field edge to the thoracic wall. A "L-bar" armrest, used to position the arm during breast cancer irradiation, decreases these distances, and so decreases the clinically relevant amount of heart and lung inside the treatment fields.

A distant cancer (or distant recurrence, relapse, metastasis) is a cancer that has spread from the original (primary) tumor to distant organs or distant lymph nodes. DDFS (Distant Disease-free Survival) – the time until such an event – is a parameter used in clinic trials.

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Tubiana (1986) claims that a critical tumor diameter and mass for metastatic spread exists, and that this threshold varies with the tumor type and may be reached before the primary tumor is detectable. For breast cancer, he found metastases in 50% of the women whose primary tumor had a diameter of $3.5 \, \mathrm{cm}$, i.e., a mass $\approx 22 \, \mathrm{g}$.

• Distances in Rheumatology

The main such distances (measured in cm to the nearest 0.1 cm) follow.

Occiput wall distance: the distance from the patient's occiput to the wall during maximal effort to touch the head to the wall, without raising the chin above its usually carrying level (when heels and, if possible, the back are against the wall).

Modified Schober test: the distance between two marked points (a point over the spinous process of L5 and the point directly 10 cm above) are measured when the patient is extending his lumbar spine in neutral position and then when he flexes forward as far as possible. Normally, the 10 cm distance increases to > 16 cm.

Lateral spinal flexion: the distance from the middle fingertip to the floor in full lateral flexion without flexing forward or bending the knees or lifting heels and attempting to keep the shoulders in the same place.

Chest expansion: the difference in cm between full expiration and full inspiration, measured at the nipples.

Intermalleolar distance: the distance between the medial malleoli when the patient (supine, the knees straight and the feet pointing straight up) is asked to separate the legs as far as possible.

• Distance healing

Distance (or *distant*, *remote*) **healing** is defined (Sicher and Targ 1998) as a conscious, dedicated act of mentation attempting to benefit another person's physical or emotional well-being at a distance.

It includes prayer (intercessory, supplicative and non-directed), spiritual/mental healing and strategies purporting to channel some supraphysical energy (Non-contact Therapeutic Touch, Reiki healing, External Qigong).

Distant healing is a part of popular alternative and complementary medicine but, in the absence of any plausible mechanism, it is highly controversial: some positive results of Therapeutic Touch and intercessory prayer are attributed to a placebo effect. Still, such rejection (as well as for Homeopathy) is also a matter of belief; cf. action at a distance (in Physics) in Chap. 24.

• Space-related phobias

Several **space-related phobias** have been identified: agoraphobia, astro-phobia, claustrophobia, and cenophobia, which are fear of open, celestial, enclosed and empty spaces, respectively. Balint's syndrome is unability to localize objects in space.

In Chap. 28, among applications of **spatial language** is mentioned Grove's Clean Space: a neuro-linguistic psychotherapy based on the

spatial metaphors produced by the client on his present and desired "space" (state).

• Neurons with spatial firing properties

Known types of neurons with spatial firing properties are listed below; cf. also spike train distances in Chap. 23.

Many mammals have in several brain areas *head direction cells* (Taube, Muller and Ranck 1990): neurons which fire only when the animal's head points in a specific direction within an environment.

Place cells (O'Keefe and Dostrovsky 1971) are principal neurons in the hippocampus that fire strongly whenever an animal is in a specific location (the cell's place field) in an environment.

Spatial view cells (Georges-Francois, Rolls and Robertson 1999) are neurons in hippocampus which fire when the animal views a specific part of an environment. They differ from head direction cells since they represent not a global orientation, but the direction towards a specific object. They also differ from place cells, since they are not localized in space.

Grid cells are neurons in the entorhinal cortex that fire strongly when an animal is in specific locations in an environment. Hafting, Fyhn, Molden, Moser, and Moser, who discovered them in 2005, conjecture that the network of these cells constitute a mental map of the spatial environment.

Border cells (Hafting, Fyhn, Molden, Moser and Moser 2008) are neurons in the entorhinal cortex that fire when a border is present in the proximal environment.

The smallest processing module of cortical neurons is a $\it minicolumn-a$ vertical column through the cortical layers of the brain, comprising 80–120 neurons that seem to work as a team. There are about 2×10^8 minicolumns in humans. The diameter of a minicolumn is about $28–40\,\mu m$. Smaller minicolumns (as observed in scientists and in people with autism) mean that there are more processing units within any given cortical area; it may allow for better signal detection and more focused attention.

• Visual Analogue Scales

In Psychophysics and Medicine, a **Visual Analogue Scale** (or *VAS*) is a self-report device used to measure the magnitude of internal states such as pain and mood (depression, anxiety, sadness, anger, fatigue, etc.) which range across a continuum and cannot easily be measured directly. Usually, VAS is a horizontal (or vertical, for Chinese subjects) 100 mm line anchored by word descriptors at each end. The *VAS score* is the distance, measured in millimeters, from the left hand (or lower) end of the line to the point marked by the subject.

The VAS tries to produce *ratio data*, i.e., ordered data with constant scale and a natural zero. It is more suitable when looking at change within individuals than comparing across a group.

Amongst scales used for pain-rating, the VAS is more sensitive than the simpler verbal scale (six descriptive or activity tolerance levels), the Wong–Baker facial scale (six grimaces) and the numerical scale (levels 542 29 Other distances

 $0, 1, 2, \ldots, 10$). On the other hand, the VAS is much simpler and less intrusive than full-length questionnaires for measuring internal states.

• Vision distances

The inter-pupillary distance (or *inter-ocular distance*): in Ophthalmology, the distance between the centers of the pupils of the two eyes when the visual axes are parallel. Typically, it is 2.5 inches (6.35 cm).

The *near acuity* is the eye's ability to distinguish an object's shape and details at a near distance such as 40 cm; the **distance acuity** is the eye's ability to do it at a far distance such as 6 m.

The *optical near devices* are designed for magnifying close objects and print; the **optical distance devices** are for magnifying things in the distance (from about 3 m to far away).

The *near distance*: in Ophthalmology, the distance between the object plane and the *spectacle* (eyeglasses) plane.

The *vertex distance*: the distance between a person's glasses (spectacles planes) and their eyes (the corneal).

The *infinite distance*: in Ophthalmology, the distance of 20 feet (6.1 m) or more; so called because rays entering the eye from an object at that distance are practically as parallel as if they came from a point at an infinite distance.

The **distance vision** is a vision for objects that are at least 6 m from the viewer.

The angular eye distance is the aperture of the angle made at the eye by lines drawn from the eye to two objects.

The RPV-distance (or resting point of vergence) is the distance at which the eyes are set to converge (turn inward toward the nose) when there is no close object to converge on. It averages about 45 inches (1.14 m) when looking straight ahead and comes in to about 35 inches (0.89 m) with a 30° downward gaze angle. Ergonomists recommend the RPV-distance as the eye-screen distance in sustained viewing, in order to minimize eyestrain.

The default accommodation distance (or resting point of accommodation, RPA-distance) is the distance at which the eyes focus when there is nothing to focus on.

In Painting, the *distance* is the part of a picture which contains the representation of those objects which are the farthest away, while the *middle distance* is the central part of a scene between the foreground and the background.

In a *Perspective Drawing* (when objects are drawn smaller at larger distances and distorted if viewed at an angle), a *point of distance* is the point where the visual rays meet. A *vanishing point* is the point at which parallel lines receding from an observer seem to converge; a picture can have several such points or none.

• Body distances in Antropometry

Besides weight and circumference, all the main standard measurements in Physical Anthropology and Human Osteology (including Forensic Anthropology and Paleoanthropology) are distances between some body landmark points or planes.

The main vertical distances from a standing surface are: stature (to the top of the head), nasion height (to the nasion, i.e., the top of the nose between the eyes), C7 level height (to the first palpable vertebra from the hairline down, C7), acromial height (to the acromion, i.e., the lateral tip of the shoulder), L5 level height (to the first palpable vertebra from the tailbone up, L5), knee height (to the patella, i.e., kneecap, plane). The similar vertical distances from a sitting surface are called sitting heights.

Examples of other body distances are the following breadths, lengths and depths:

Biacromial breadth: the horizontal distance between the right and left acromions

Hip breadth (seated): the lateral distance at the widest part of the hips Ankle distance (seated): the horizontal distance from L5 to the lateral malleolus (bony prominence at the distal end of the fibula)

Buttock-knee length: the distance from the buttocks to the patella

Total foot length: the maximum length of the right foot

Hand length: the length of the right hand between the stylion landmark on the wrist and the tip of the middle finger

 $Abdominal\ depth$ (seated): the maximum horizontal depth of the abdomen

Head and face measurement distances

The main linear dimensions of the cranium in Archeology are: lengths (of temporal bone, of tympanic plate, glabella-opistocranion), breadths (maximum cranial, minimum frontal, biauricular, mastoid), heights (of temporal bone, basion-bregma), thickness of tympanic plate, bifrontomolare-temporale distance.

Examples of the viscerocranium measurements in Craniofacial Anthropometry are:

 $\it Head\ width:$ the (horizontal) maximum breadth of the head above the ears

Head depth (or head length): the horizontal distance from the nasion to the opisthocranion (the most prominent point on the back of the head)

Inter-canthal distance (inner or outer): the distance between (inner or outer) canthi (corners of eyes)

Total face height (or face length): the distance from nasion to menton Morphological facial height: the distance between nasion and gnathion (the most inferior point of the mandible in the midline)

Face width (or bizygomatic width): the maximum distance between lateral surfaces of the zygomatic arches; the facial index is the ratio between face width and face length.

In Face Recognition, the sets of (vertical and horizontal) cephalofacial dimensions, i.e., distances between fiducial (standard of reference 544 29 Other distances

for measurement) facial points, are used. The distances are normalized, say, with respect to the **inter-pupillary distance** for horizontal ones. For example, the following five independent facial dimensions are derived in [Fell97] for facial gender recognition: distance E between outer eye corners, nostril-to-nostril width N, face width at cheek W and (vertical ones) eye to eyebrow distance B and distance L between eye midpoint and horizontal line of mouth. "Femaleness" relies on large E, B and small N, W, L.

• Gender-related body distance measures

The main gender-specific body configuration features are:

For females, WHR (waist-to-hip ratio), i.e., the ratio of the circumference of the waist to that of the hips at their widest part and BMI (body mass index), i.e., the ratio of the weight in kg and squared height in m^2

For males, SCR (waist-to-chest ratio), SHR (shoulder-to-hip ratio) and VHI (volume-to-height index), i.e., the ratio of the volume in liters and squared height in m²

Androgen equation (three times the shoulder width minus one times the pelvic width) which is higher for males

 $Second\mbox{-}to\mbox{-}forth\ digit\ (index\ to\ ring\ finger)\ ratio\ which is lower for\ males$ in the same population

Anogenital distance (length of the perineum, cf. distances in Medicine) which is larger for males

Also, a person's center of mass (slightly below the belly button) is lower for females, and women have lower mental rotation ability.

BMI and WHR indicate percentage of body fat and fat distribution, respectively; they are widely used in Medicine to assess risk factors. WHR is also seen as the main visual cue to female body attractiveness, with optimum ≈ 0.7 . But Rilling, Kaufman, Smith, Patel and Worthman (2009) claim that abdominal depth (the depth of the lower torso at the umbilicus) and WC (waist circumference) are stronger predictors. Fan, Dai, Liu and Wu (2005) claim that VHI is the main visual cue to male body attractiveness, with optimum 17.6 and 18.0 for female raters and male raters, respectively.

• Body distances for clothes

The European standard EN 13402 "Size designation of clothes" defined, in part EN 13402-1, a standard list of 13 body dimensions (measured in centimeters) together with a method for measuring each one on a person. These are: body mass, height, foot length, arm length, inside leg length, and girth for head, neck, chest, bust, under-bust, waist, hip, hand. Examples of these definitions follow.

Foot length: horizontal distance between perpendiculars in contact with the end of the most prominent toe and the most prominent part of the heel, measured with the subject standing barefoot and the weight of the body equally distributed on both legs.

Arm length: distance, measured using the tape-measure, from the armscye/shoulder line intersection (acromion), over the elbow, to the far end of the prominent wrist bone (ulna), with the subject's right fist clenched and placed on the hip, and with the arm bent at 90°.

Inside leg length: distance between the crotch and the soles of the feet, measured in a straight vertical line with the subject erect, feet slightly apart, and the weight of the body equally distributed on both feet.

The final part EN 13402-4 (a compact three-digits coding system for the size of clothes) of the standard is expected to became mandatory in EU after 2007. It should simplify the situation when, for a Miss Average $(88-72-96\,\mathrm{cm}, \mathrm{i.e.}, 34-28-37\,\mathrm{inch}, \mathrm{bust-waist-hips})$, her dress is (being 10 in US): 12 in the UK, C38 in Norway, Sweden and Finland, 38 in Germany and the Netherlands, 40 in Belgium and France, 44 in Italy, 44/46 in Portugal and Spain.

• Racing distances

In Racing, **length** is an informal unit of distance to measure the distance between competitors in a race; for example, in boat-racing it is the average length of a boat.

The **horse-racing distances** are measured in terms of the approximate length of a horse, i.e., about 8 feet (2.44 m).

Winning margins are measured in **lengths**, ranging from half the length to the **distance**, i.e., more than 20 lengths; the *length* is often interpreted as a unit of time equal to $\frac{1}{5}$ s. Smaller margins are: *short-head*, *head*, or *neck*. Also, the *hand*, i.e., 4 inches (10.2 cm), is used for measuring the height of horses.

• Triathlon race distances

The **Ironman distance** (started in Hawaii 1978) is a 3.86 km swim followed by a 180 km bike and a 42.2 km (marathon distance) run.

The international **Olympic distance** (started in Sydney, 2000) is 1.5 km (*metric mile*), 40 km and 10 km of swim, cycle and run, respectively.

Next to the Olympic distance are the *sprint distance* (0.75, 20, 5 km) and the *ITU long distance* (3, 80, 20 km).

DPS (distance per swim stroke) is a metric of swimming efficiency used in training triathletes; it is obtained by counting strokes on fixed pool distances. LSD (long slow distance) is a slang term for a training method for runners that involves running distances longer, and at a slower pace, than those of races. In Running, sprinting is divided into $100\,\mathrm{m}$, $200\,\mathrm{m}$, $400\,\mathrm{m}$, while middle distance mean races of various distance from $800\,\mathrm{m}$ to $5\,\mathrm{km}$.

Also, the distance is boxing slang for a match that lasts the maximum number (12 or 10) of scheduled rounds.

• Isometric muscle action

An **isometric muscle** action refers to exerting muscle strength and tension without producing an actual movement or a change in muscle length.

The technique of *isometric action training* is used mainly by weightlifters and bodybuilders. Examples of such *isometric exercises* include holding a weight at a certain position in the range of motion and pushing or pulling against an immovable external resistance.

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29.2 Equipment distances

Vehicle distances

The **perception-reaction distance** (or *thinking distance*): the distance a motor vehicle travels from the moment the driver sees a hazard until he applies the brakes (corresponding to human perception time plus human reaction time). Physiologically, it takes 1.3–1.5 s, and the brake action begins 0.5 s after application.

The safe following distance: the reglementary distance from the vehicle ahead of the driver. For reglementary perception-reaction time at least 2s (the 2 Second Rule), this distance (in meters) should be $0.56 \times v$, where v is the speed (in kilometer per hour). Sometimes the 3 Second Rule applied. The stricter rules are used for heavy vehicles (say, at least $50 \,\mathrm{m}$) and in tunnels (say, at least $150 \,\mathrm{m}$).

The **braking distance**: the distance a motor vehicle travels from the moment the brakes are applied until the vehicle completely stops.

The (total) **stopping distance**: the distance a motor vehicle travels from where the driver perceives the need to stop to the actual stopping point (corresponding to vehicle reaction time plus vehicle braking capability).

The **crash distance**: the amount of distance between the driver and the front end of a vehicle in a frontal impact (or, say, between the pilot and the first part of airplane to impact the ground). Crushworthy motor vehicles are designed with structural "crush zones" which collapse in a prescribed way at specified loads. They achieve, during crush milliseconds, a uniform deceleration (say, 25 g when measured in a fixed barrier frontal crush at $50 \,\mathrm{km}\,\mathrm{h}^{-1}$, where $g = 9.80665\,\mathrm{m}\,\mathrm{s}^{-2}$ is the standard gravity) of the passenger compartment and absorption of kinetic energy primarily outside of it. European regulation, for example, requires a $60 \,\mathrm{km}\,\mathrm{h}^{-1}$ impact test into a 40% deformable offset barrier.

The actual landing distance is the distance used in landing and braking to a complete stop (on a dry runway) after crossing the runway threshold at 50 feet (15.24 m); it can be affected by various operational factors.

The JAA/FAA **required landing distance** (used for dispatch purposes) is factor 1.67 of actual landing distance for a dry runway and factor 1.92 for a wet runway.

The accelerate-stop distance: the runway plus *stopway length* (able to support the airplane during an aborted takeoff) declared available and suitable for the acceleration and deceleration of an airplane aborting a takeoff.

The **endurance distance**: the total distance that a ground vehicle or ship can be self-propelled at any specified endurance speed.

The distance made good is a nautical term: the distance traveled after correction for current, *leeway* (the sideways movement of the boat away from the wind) and other errors that may not have been included in

the original distance measurement. Log is a device to measure the distance traveled through the water which is further corrected to a distance made good. Before log's introduction, sea distances were measured in units of a day's sail.

One of meanings of the term leg – a stage of a journey or course – includes a nautical term: the distance traveled by a sailing vessel on a single tack.

The GM-distance (or $metacyclic\ height$) of a ship is the distance between its center of gravity G and the metacenter, i.e., the projection of the $center\ of\ buoyancy$ (the center of gravity of the volume of water which the hull displaces) on the center line of the ship as it heels. This distance (usually, $1-2\ m$) determines the stability of the ship in water.

The **distance line**: in Diving, a temporary marker (typically, 50 m of thin polypropylene line) of the shortest route between two points. It is used, as a kind of Ariadne's thread, to navigate back to the start in poor visibility.

Bushell BackTrack GPS Device stores location in its memory so that one can be directed back to original starting point using directional arrows and GPS (Global Positioning System; see radio distance measurement) distance estimations.

• Buffer distance

In nuclear warfare, the **horizontal buffer distance** is the distance which should be added to the radius of safety in order to be sure that the specified degree of risk will not be exceeded. The **vertical buffer distance** is the distance which should be added to the fallout safe-height of burst, in order to determine a desired height of burst so that military significant fallout will not occur.

The term *buffer distance* is also used more generally as, for example, the buffer distance required between sister stores or from a high-voltage line. Cf. **clearance distance** and, in Chap. 25, **setback distance**.

• Offset distance

In nuclear warfare, the **offset distance** is the distance the desired (or actual) ground zero is offset from the center of the area (or point) target.

In Computation, *offset* is the distance from the beginning of a string to the end of the segment on that string. For a vehicle, **offset** of a wheel is the distance from its hub mounting surface to the centerline of the wheel.

The term *offset* is also used for the **displacement** vector (cf. Chap. 24) specifying the position of a point or particle in reference to an origin or to a previous position.

• Standoff distance

The **standoff distance** is the distance of an object from the source of an explosion (in warfare), or from the delivery point of a laser beam (in laser material processing). Also, in Mechanics and Electronics, it is the distance separating one part from another; for example, for insulating (cf. **clearance distance**), or the distance from a non-contact length gauge to measured material surface.

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• Distance in Military

In Military, the term **distance** usually has one the following meanings:

The space between adjacent individual ships or boats measured in any direction between foremasts

The space between adjacent men, animals, vehicles, or units in a formation measured from front to rear

The space between known reference points or a ground observer and a target, measured in meters (artillery), or in units specified by the observer

In amphibious operations, the *distant retirement area* is the sea area located to seaward of the landing area, and the *distant support area* is the area located in the vicinity of the landing area but at considerable distance seaward of it.

In military service, a **bad distance** of the troop means a temporary intention from the war service to extract itself. This passing was usually heavily punished and equated with that the *desertion* (an intention to extract itself durably).

• Proximity fuse

The **proximity fuse** is a fuse that is designed to detonate an explosive automatically when close enough to the target.

• Sensor network distances

The **stealth distance** (or *first contact distance*): the distance traveled by the moving object (or intruder) until detection by an active sensor of the network (cf. **contact quasi-distances** in Chap. 19); the *stealth time* is the corresponding time.

The first sink contact distance: the distance traveled by the moving object (or intruder) until the monitoring entity can be notified via a sensor network.

The **miss distance**: the distance between the lines of sight representing estimates from two sensor sites to the target (cf. **line-line distance** from Chap. 4).

The sensor tolerance distance: a range distance within which a localization error is acceptable to the application (cf. tolerance distance from Chap. 25).

• Proximity sensors

Proximity sensors are varieties of ultrasonic, laser, photoelectric and fiber optic sensors designed to measure the distance from itself to a target. For such laser range-finders, a special *distance filter* removes those sensor measurements which, with high probability, are shorter than expected, and which are therefore caused by an unmodeled object. Cf. **distances in Animal Behavior** in Chap. 23 (distance estimation by some insects).

The detection distance is the distance from the detecting surface of a sensor head to the point where a target approaching it is first detected. The maximum operating distance of a sensor is its maximum detection distance from a standard modeled target, disregarding accuracy. The stable

detection range is the detectable distance range in which a standard detected object can be stably detected with respect to variations in the operating ambient temperature and power supply voltage.

• Precise distance measurement

The resolution of TEM (transmission electronic microscope) is about $0.2 \,\mathrm{nm} \ (2 \times 10^{-10} \,\mathrm{m})$, i.e., the typical separation between two atoms in a solid. This resolution is 1,000 times greater than a light microscope and about 500,000 times greater than that of a human eye. However, only nanoparticles can fit in the vision field of an electronic microscope.

The methods, based on measuring the wavelength of laser light, are used to measure macroscopic distances non-treatable by an electronic microscope. However, the uncertainty of such methods is at least the wavelength of light, say, 633 nm.

The recent adaptation of Fabry–Perot metrology (measuring the frequency of light stored between two highly reflective mirrors) to laser light permits the measuring of relatively long (up to 5 cm) distances with uncertainty of only 0.01 nm.

• Radio distance measurement

DME (or **distance measuring equipment**) is an air navigation technology that measures distances by timing the propagation delay of UHF signals to a *transponder* (receiver-transmitter that will generate a reply signal upon proper interrogation) and back. DME is expected to be phased out by global satellite-based systems: GPS and, planned for 2009, Galileo (EU) and GLOSNASS (Russia).

The GPS (Global Positioning System) is a radio navigation system which permits one to get her/his exact position on the globe (anywhere, anytime). It consists of 24 satellites and a monitoring system operated by the US Department of Defense. The non-military part of GPS can be used just by the purchase of an adequate receiver and the accuracy is 10 m.

The **GPS** pseudo-distance (or *pseudo-range*) from a receiver to a satellite is the travel time of a satellite time signal to a receiver multiplied by the propagation time of the radio signal (about the speed of light). It is called *pseudo-distance* because of the error: the receiver clock is not so perfect as the ultra-precise clock of a satellite. The GPS receiver calculates its position (in latitude, longitude, altitude, etc.) by solving a system of equations using its pseudo-distances from at least four satellites and the knowledge of their positions. Cf. **radio distances** in Chap. 25.

• Transmission distance

The transmission distance is a range distance: for a given signal transmission system (fiber optic cable, wireless, etc.), it is the maximal distance the system can support within acceptable path loss level.

For a given network of contact that can transmit an infection (or, say, an idea with the belief system considered as the immune system), the **transmission distance** is the path metric of a graph (edges correspond to events of infection) via the most recent common ancestor, between (infectious agents isolated from) infected individuals.

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• Delay distance

The **delay distance** is a general term for the distance resulting from a given delay.

For example, in a meteorological sensor, the *delay distance* is the length of a column of air passing a wind vane, such that the vane will respond to 50% of a sudden angular change in wind direction.

When the energy of a neutron is measured by the delay (say, t) between its creation and detection, the *delay distance* is vt - D, where v is its velocity and D is the source-detector distance.

NASA's X-ray Observatory measures the distance to a very distant source via the delay of the halo of scattering interstellar dust between it and the Earth. Cf. also radio distance measurement.

In evaluations of visuospatial working memory (when subjects saw a dot, following a 10-, 20-, or 30-s delay, and then drew it on a blank sheet of paper), the *delay distance* is the distance between the stimulus and the drawn dot.

Master-slave distance

A master–slave system refers to a design in which one device (the master) fully controls one or more other devices (the slaves). It can be a remote manipulation system (say, a master controller with surgical end effector), a surveillance system (a stationary master camera with a wide field of view detects a moving person and turns a slave camera with a narrow field to that direction), a data transmission system and so on. The **master–slave distance** is a measure of distance between the master and slave devices. Cf. also Sect. 18.3.

• Instrument distances

The load distance: the distance (on a lever) from the fulcrum to the load. The effort distance (or resistance distance): the distance (on a lever) from the fulcrum to the resistance.

The **K-distance**: the distance from the outside fiber of a rolled steel beam to the web toe of the fillet of a rolled shape.

The **end distance**: the distance from a bolt, screw, or nail to the end of a (wood) structural member. The **edge distance**: the distance from a bolt, screw, or nail to the edge of a (wood) structural member.

The **calibration distance**: the standard distance used in the process of adjusting the output or indication on a measuring instrument.

• Sagging distance

The brazeability of brazing sheets materials is evaluated by their **sagging distance**, i.e., the deflection of the free end of the specimen sheet after brazing.

• Etch depth

Laser etching into a metal substrate produce craters. The **etch depth** is the central crater depth averaged over the apparent roughness of the metal surface.

• Feeding distance

Carbon steel shrinks during solidification and cooling. In order to avoid resulting porosity, a **riser** (cylindric liquid metal reservoir) provides liquid feed metal until the end of the solidification process.

A riser is evaluated by its **feeding distance**, which is the maximum distance over which a riser can supply feed metal to produce a radiographically sound (i.e., relatively free of internal porosity) casting. The **feeding length** is the distance between the riser and furthest point in the casting fed by it.

• Gear distances

Given two meshed gears, the distance between their centers is called the **center distance**. Examples of other distances used in basic gear formulas (such as b = a + c) follow.

Addendum (a): the radial distance between the *Pitch Circle* (the circle whose radius is equal to the distance from the center of the gear to the pitch point) and the top of the teeth.

Dedendum (b): the radial distance between the bottom of the space between the teeth and the top of the teeth.

Clearance (c): the distance between the top of a tooth and the bottom of the space into which it fits on the meshing gear.

Whole depth: the distance between the top of a tooth and the bottom of the space between teeth.

Backlash: the play between mating teeth.

• Creepage distance

The **creepage distance** is the shortest path distance along the surface of the insulation material between two conductive parts.

The shortest (straight-line) distance between two conductive parts is called the *clearance distance*; cf. the general term below.

• Clearance distance

A clearance distance (or separation distance, clearance) is, in Engineering and Safety, a physical distance or unobstructed space tolerance as, for example, the distance between the lowest point on the vehicle and the road (ground clearance).

For vehicles going in a tunnel or under a bridge, the *clearance* is the difference between the *structure gauge* (minimum size of tunnel or bridge) and the vehicles' *loading gauge* (maximum size). A clearance distance can be prescribed by a code or a standard between a piece of equipment containing potentially hazardous material (say, fuel) and other objects (buildings, equipment, etc.) and the public.

In general, *clearance* refers to the distance to the nearest "obstacle" as defined in a context. Cf. **buffer distance** and **setback distance** in Chap. 25.

Humidifier absorbtion distance

The absorbtion distance of a (water centrifugal atomizing) humidifier is the list of minimum clearance dimensions needed to avoid condensation.

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• Protective action distance

The **protective action distance** is the distance downwind from an incident (a spill involving dangerous goods which are considered toxic by inhalation) in which persons may become incapacitated.

The notion of the *mean distance between people and any hazardous event* operates also at a large scale: expanding the living area of human species (say, space colonization) will increase this distance and prevent many human extinction scenarios.

• Vertical separation distance

The **vertical separation distance** is the distance between the bottom of the drain field of a sewage septic system and the underlying water table. This separation distance allows pathogens (disease-causing bacteria, viruses, or protozoa) in the effluent to be removed by the soil before it comes in contact with the groundwater.

• Spray distance

The **spray** distance is the distance maintained between the nozzle tip of a thermal spraying gun and the surface of the workpiece during spraying.

• Fringe distance

Usually, the **fringe distance** is the spacing distance between *fringes*, for example:

dark and light regions in the interference pattern of light beams (cf., in Chap. 24, *Pendellösung fringes* in **dynamical diffraction distances**);

components into which a spectral line splits in the presence of an electric or magnetic field (*Stark* and *Zeeman effects*, respectively, in Physics).

For, say, a non-contact length gauge, the fringe distance is the value $\frac{\lambda}{2\sin\alpha}$, where λ is the laser wavelength and α is the beam angle.

In Image Analysis, there is also the *fringe distance* (Brown 1994) between binary images (cf. **pixel distance** in Chap. 21).

• Lens distances

The **focal distance** (effective focal length): the distance from the optical center of a lens (or a curved mirror) to the focus (to the image). Its reciprocal measured in meters is called the diopter and is used as a unit of measurement of the (refractive) power of a lens; roughly, the magnification power of a lens is $\frac{1}{4}$ of its diopter.

The back focal length is the distance between the rear surface of a lens and its image plane; the front focal length is the distance from the vertex of the first lens to the front focal point.

The **depth of field** (DoF): the distance in the object plane (in front of and behind the object) over which the system delivers an acceptably sharp image, i.e., the region where the blurring is tolerated at a particular resolution.

The depth of focus: the range of distance in the image plane (the eyepiece, camera, or photographic plate) over which the system delivers an acceptably sharp image. The **working distance**: the distance from the front end of a lens system to the object under inspection when the instrument is correctly focused. It can be used to modify the depth of the field.

The register distance (or flange distance): the distance between the flange (protruding rim) of the lens mount and the plane of the film image.

The conjugate image distance and the conjugate object distance: the distance along the optical axis of a lens from its principle plane to the image plane and the object plane, respectively. When a converging lens is placed between the object and the screen, the sum of the inverse conjugate image and inverse conjugate object distances is equal to the inverse focal distance.

A circle of confusion (CoC) is an optical spot caused by a cone of light rays from a lens not coming to a perfect focus; in Photography, it is defined usually as the largest blur circle that will still be perceived by the human eye as a point when viewed at a distance of 25 cm.

The *close* (or *minimum*, *near*) *focus distance*: the closest distance to which a lens can approach the subject and still achieve focus.

The **hyper-focal distance**: the distance from the lens to the nearest point (*hyper-focal point*) that is in focus when the lens is focused at infinity; beyond this point all objects are well defined and clear. It is the nearest distance at which the far end of the **depth of field** stretches to infinity (cf. **infinite distance**).

• Distances in Stereoscopy

A method of 3D imaging is to create a pair of 2D images by a two-camera system.

The inter-camera distance (or base line length, inter-ocular lens spacing) is the distance between the two cameras from which the left and right eye images are rendered.

The convergence distance is the distance between the base line of the camera center to the convergence point where the two lenses should converge for good stereoscopy. This distance should be 15–30 times the inter-camera distance.

The picture plane distance is the distance at which the object will appear on (but not behind or in front of) the picture plane (the apparent surface of the image). The window is a masking border of the screen frame such that objects, which appear at (but not behind or outside) it, appear to be at the same distance from the viewer as this frame. In human viewing, the picture plane distance is about 30 times the inter-camera distance.

Distance-related shots

A film *shot* is what is recorded between the time the camera starts (the director's call for "action") and the time it stops (the call to "cut").

The main **distance-related shots** (camera setups) are:

- Establishing shot: a shot, at the beginning of a sequence which establishes the location of the action and/or the time of day

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- Long shot: a shot taken from at least 50 yards (45.7 m) from the action

- Medium shot: a shot from 5–15 yards (4.6–13.7 m), including a small entire group, which shows group/objects in relation to the surroundings
- Close-up: a shot taking the actor from the neck upwards, or an object from a similarly close position
- Two-shot: a shot that features two persons in the foreground
- *Insert*: an inserted shot (usually a close up) used to reveal greater detail.

29.3 Miscellany

• Range distances

The range distances are practical distances emphasizing a maximum distance for effective operation such as vehicle travel without refueling, a bullet range, visibility, movement limit, home range of an animal, etc.

For example, the range of a risk factor (toxicity, blast, etc.) indicates the minimal **safe distancing**.

The **operating distance** (or *nominal sensing distance*) is the range of a device (for example, a remote control) which is specified by the manufacturer and used as a reference. The **activation distance** is the maximal distance allowed for activation of a sensor-operated switch. In order to stress the large range of a device, some manufacturers mention it in the product name; for example, *Ultimate Distance* golf balls (or softball bates, spinning reels, etc.).

• Spacing distances

The following examples illustrate this large family of practical distances emphasizing a minimum distance (cf. **minimum distance** in Chap. 16, nearest-neighbor **distance** in **Animal Behavior** in Chap. 28 and **first-neighbor distance** for atoms in a solid in Chap. 24).

The **miles in trail**: a specified minimum distance, in nautical miles, required to be maintained between airplanes. *Seat pitch* and *seat width* are airliner distances between, respectively, two rows of seats and the armrests of a single seat.

The **isolation distance**: a specified minimum distance required (because of pollination) to be maintained between variations of the same species of a crop in order to keep the seed pure (for example, $\approx 3 \,\mathrm{m}$ for rice).

The **stop-spacing distance**: the interval between bus stops; the mean stop-spacing distance in the US (for light rail systems) ranges from 500 m (Philadelphia) to 1,742 m (Los Angeles).

The **character spacing**: the interval between characters in a given computer font.

The **just noticeable difference** (JND): the smallest percent change in a dimension (for distance/position, etc.) that can be reliably perceived (cf. **tolerance distance** in Chap. 25).

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• Cutoff distances

Given a range of values (usually, a length, energy, or momentum scale in Physics), *cutoff* is the maximal or minimal value, as, for example, Planck units.

A **cutoff distance** is a cutoff in a length scale. For example, *infrared* cutoff and ultraviolet cutoff (the maximal and minimal wavelength that the human eye takes into account) are *long-distance* cutoff and short-distance cutoff, respectively, in the visible spectrum.

• Quality metrics

A quality metric (or, simply, *metric*) is a standard unit of measure or, more generally, part of a system of parameters, or systems of measurement. This vast family of measures (or standards of measure) concern different attributes of objects. In such a setting, our distances and similarities are rather "similarity metrics," i.e., metrics (measures) quantifying the extent of relatedness between two objects.

Examples include academic metrics, crime statistics, corporate investment metrics, economic metrics (indicators), education metrics, environmental metrics (indices), health metrics, market metrics, political metrics, properties of a route in computer networking, software metrics and vehicle metrics. For example, the site http://metripedia.wikidot.com/start aims to build an Encyclopedia of IT (Information Technology) performance metrics. Some examples of non-equipment quality metrics are detailed below.

A symmetry metric (Bhanji, Purchase, Cohen and James 1995) measures graph-drawing aesthetics by $\frac{\sum_{i=1}^{m}(a_{1i}+a_{2i}+a_{3i})}{a} \times \sum_{i=1}^{m}(a_{1i}+\frac{a_{2i}+n_i}{2})$, where a is the number of all arcs, m is the number of axes of symmetry and, for a given axis i, n_i is the number of vertices which are mirrored by another vertex about i, while a_{1i} , a_{2i} and a_{3i} are the number of arcs which are, respectively, bisected at right angles by i, mirrored by another arc about i and run along i. Axes of symmetry are taken as all straight lines i with n_i , a_{1i} , $a_{2i} \geq 1$.

Landscape metrics evaluate, for example, greenway patches in given landscape by *patch density* (the number of patches per km²), *edge density* (the total length of patch boundaries per hectare), *shape index* $\frac{E}{4\sqrt{A}}$ (where A is the total area, and E is the total length of edges), connectivity, diversity, etc.

Management metrics include: surveys (say, of market share, sales increase, customer satisfactions), forecast (say, of revenue, contingent sales, investment), R&D effectiveness, absenteeism, etc.

Risk metrics are used in Insurance and, in order to evaluate a portfolio, in Finance.

An **impact factor** is a quality metric ranking the relative influence as, for example:

PageRank of Google ranking web pages;

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ISI (now Thomson Scientific) *Impact Factor* of a journal measuring, for a given 2-year period, the number of times the average article in this journal is cited by some article published in the subsequent year;

Hirsch's h-index of a scholar which is the highest number of his/her published articles that have each received at least that number of citations.

• Distal and proximal

The antipodal notions near (close, nigh) and far (distant, remote) are also termed proximity and distality.

The adjective **distal** (or *peripheral*) is an anatomical term of location (on the body, the limbs, the jaw, etc.); corresponding adverbs are: distally, distad. For an *appendage* (any structure that extends from the main body), **proximal** means situated towards the point of attachment, while **distal** means situated around the furthest point from this point of attachment. More generally, as opposed to *proximal* (or *central*), *distal* means: situated away from, farther from a point of reference (origin, center, point of attachment, trunk). As opposed to *mesial* it means: situated or directed away from the midline or mesial plane of the body.

More abstractly: for example, the T-Vision (Earth visualization) project brings a *distal* perception of Earth, known before only to astronauts, down to Earth. In 2008 NASA's Deep Impact spacecraft has created a video of the Moon passing before Earth as seen from the spacecraft's distal (50 million km away) point of view.

Proximal and distal demonstratives are words indicating place deixis, i.e., a spatial location relative to the point of reference. Usually, they are two-way as this/that, these/those or here/there, i.e., in terms of the dichotomy near/far from the speaker. But, say, Korean, Japanese, Spanish, Portuguese and Thai make a three-way distinction: proximal (near to the speaker), medial (near to the addressee) and distal (far from both). English had the third form, yonder (at an indicated distance, usually within sight), still spoken in Southern US. Cf. spatial language in Chap. 28.

• Distance effect

The **distance effect** is a general term for the change of a pattern or process with distance. Usually, it means **distance decay**. For example, a static field attenuates proportionally to the inverse square of the distance from the source.

Another example of the distance effect is a periodic variation (instead of uniform decrease) in a certain direction, when a *standing wave* occurs in a time-varying field. It is a wave that remains in a constant position because either the medium is moving in the opposite direction, or two waves, traveling in opposite directions, interfer; cf. **Pendellösung length** in Chap. 24.

The distance effect, together with the size (source magnitude) effect determine many processes; cf. island distance effect, insecticide distance effect in Chap. 23 and symbolic distance effect, distance effect on trade in Chap. 28.

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• Distance decay

The distance decay is the attenuation of a pattern or process with distance. Cf. distance decay (in Spatial Interaction) in Chap. 28. It is the main case of distance effect.

Examples of distance-decay curves: Pareto model $\ln I_{ij} = a - b \ln d_{ij}$, and the model $\ln I_{ij} = a - b d_{ij}^p$ with $p = \frac{1}{2}$, 1, or 2 (here I_{ij} and d_{ij} are interaction and distance between points i, j, while a and b are parameters).

A mass-distance decay curve is a plot of "mass" decay when the distance to the center of "gravity" increases. Such curves are used, for example, to determine an *offender's heaven* (the point of origin; cf. distances in Criminology) or the galactic mass within a given radius from its center (using *rotation-distance curves*).

• Incremental distance

An **incremental distance** is a gradually increasing (usually, by a fixed amount) distance.

• Distance curve

A distance curve is a plot (or a graph) of a given parameter against a corresponding distance. Examples of distance curves, in terms of a process under consideration, are: time-distance curve (for the travel time of a wave-train, seismic signals, etc.), height-run distance curve (for the height of tsunami wave vs. wave propagation distance from the impact point), drawdown-distance curve, melting-distance curve and wear volume – distance curve.

A force-distance curve is, in SPM (Scanning Probe Microscopy), a plot of the vertical force that the tip of the probe applies to the sample surface, while a contact-AFM (Atomic Force Microscopy) image is being taken. Also, frequency-distance and amplitude-distance curves are used in SPM.

The term *distance curve* is also used for charting growth, for instance, a child's height or weight at each birthday. A plot of the rate of growth against age is called the **velocity-distance curve**. The last term is also used for the speed of aircraft.

• Propagation length

For a pattern or process attenuating with distance, the **propagation** length is the distance to decay by a factor of $\frac{1}{e}$.

Cf., for example, radiation length and the Beer-Lambert law in Chap. 24.

• Characteristic length

A characteristic length (or *scale*) is a convenient reference length (usually constant) of a given configuration, such as the overall length of an aircraft, the maximum diameter or radius of a body of revolution, or a chord or span of a lifting surface.

In general, it is a length that is representative of the system (or region) of interest, or the parameter which characterizes a given physical quantity.

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For example, the characteristic length of a rocket engine is the ratio of the volume of its combustion chamber to the area of the nozzle's throat, representing the average distance that the products of burned fuel must travel to escape.

• Path length

In general, a *path* is a line representing the course of actual, potential or abstract movement. In Topology, a *path* is a certain continuous function; cf. parametrized **metric curve** in Chap. 1.

In Physics, **path length** is the total distance an object travels, while *displacement* is the net distance it travels from a starting point. Cf mathematical **displacement inelastic mean free path**, **optical distance** and (in the item **dislocation distances**) **dislocation path length** in Chap. 24. In Chemistry, (cell) **path length** is the distance that light travels through a sample in an analytical cell.

In Graph Theory, **path length** is a discrete notion: the number of vertices in a sequence of vertices of a graph; cf. **path metric** in Chap. 1. Cf. **Internet IP metric** in Chap. 22 for **path length** in a computer network. Also, it means the total number of machine code instructions executed on a section of a program.

• Scale invariance

Scale invariance is a feature of laws or objects which do not change if **length scales** are multiplied by a common factor.

Examples of scale invariant phenomena are **fractals** (in Chap. 1) and power laws; cf. **scale-free network** in Chap. 22 and self-similarity in **long range dependence**. Scale invariance arising from a power law $y = Cx^k$, for a constant C and scale exponent k, amounts to linearity $\log y = \log C + k \log x$ for logarithms.

Much of scale invariant behavior (and complexity in nature) is explained (Bak, Tang and Wiesenfeld 1987) by *self-organized cruciality (SOC)* of many dynamical systems, i.e., the property to have the critical point of a phase transition as an attractor which can be attained spontaneously without any fine-tuning of control parameters.

Two moving systems are dynamically similar if the motion of one can be made identical to the motion of the other by multiplying all lengths by one scale factor, all forces by another one and all time periods by a third scale factor. Dynamic similarity can be formulated in terms of dimensionless parameters (where the units of measurement, and so scale factors, cancel out) as, for example, the Froude number $F = \frac{v}{\sqrt{gd}}$ (here v is the representative speed, d is the representative length, g is Earth standard gravity 9.80665 m s⁻²), and the Richardson number F^{-2} .

• Long range dependence

A (second-order stationary) stochastic process X_k , $k \in \mathbb{Z}$, is called **long** range dependent (or *long memory*) if there exist numbers $\alpha, 0 < \alpha < 1$, and $c_{\rho} > 0$ such that $\lim_{k \to \infty} c_{\rho} k^{\alpha} \rho(k) = 1$, where $\rho(k)$ is the

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autocorrelation function. So, correlations decay very slowly (asymptotically hyperbolic) to zero implying that $\sum_{k\in\mathbb{Z}} |\rho(k)| = \infty$, and that events so far apart are correlated (long memory). If above sum is finite and decay is exponential, then the process is *short range*.

Examples of such processes are the exponential, normal and Poisson processes, which are memoryless, and, in physical terms, systems in thermodynamic equilibrium. The above power law decay for correlations as a function of time translates into a power law decay of the Fourier spectrum as a function of frequency f and is called $\frac{1}{f}$ noise.

A process has a self-similarity exponent (or Hust parameter) H if X_k and $t^{-H}X_{tk}$ have the same finite-dimensional distributions for any positive t. The cases $H=\frac{1}{2}$ and H=1 correspond, respectively, to purely random process and to exact self-similarity: the same behavior on all scales (cf. fractal in Chap. 1, scale-free network in Chap. 22, and scale invariance). The processes with $\frac{1}{2} < H < 1$ are long range dependent with $\alpha = 2(1-H)$.

Long range dependence corresponds to heavy-tailed (or power law) distributions. The distribution function and tail of a non-negative random variable X are $F(x) = P(X \le x)$ and $\overline{F(x)} = P(X > x)$. A distribution F(X) is heavy-tailed if there exists a number $\alpha, 0 < \alpha < 1$, such that $\lim_{x\to\infty} x^{\alpha} \overline{F(x)} = 1$. Many such distributions occur in the real world (for example, in Physics, Economics, Internet) in both space (distances) and time (durations). A standard example is the Pareto distribution $\overline{F(x)} = x^{-k}$, x > 1, where x > 0 is a parameter (cf. distance decay).

Also, the random-copying model (the cultural analog of genetic drift) of the frequency distributions of various cultural traits (such as scientific paper citations, first names, dog breeds, pottery decorations) results (Bentley, Hahn and Shennan 2004) in a power law distribution $y = Cx^{-k}$, where y is the proportion of cultural traits that occur with frequency x in the population, and C and k are parameters.

Long-distance

The term **long-distance** refers usually telephone communication (long-distance call, long-distance operator) or to covering large distances by moving (long-distance trail, long-distance running, swimming, etc.) or, more abstractly: long-distance migration, commuting, supervision, relationship, etc. *Long-distance intercourse* (coupling at a distance) is found often in Native American folklore: Coyote, the Trickster, can reach a woman on the opposite bank of a lake.

Cf. long-distance dispersal, animal and plant long-distance communication, long range order, long range dependence, action at a distance (in Computing, Physics, along a DNA).

The term *short-distance* is rarely used. Instead, the adjective **short-range** means limited to (or designed for) short distances, or relating to the near future. Finally, **touching**, for two objects, is having (or getting) a zero distance between them.

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• Spatial analysis

In Statistics, **spatial analysis** (or *spatial statistics*) includes the formal techniques for studying entities using their topological, geometric, or geographic properties. More restrictively, it refers to *geostatistics* and *Human Geography* (techniques applied to structures at the human scale, especially in the analysis of geographic data).

Starting with mapping, surveying and geography initially, spatial analysis focuses now on computer-based techniques. It is applied in Geography, Biology, Epidemiology, Statistics, Geographic Information Science, Remote Sensing, Computer Science, Mathematics, and Scientific Modeling.

Also in Statistics, **spatial dependence** is a measure for the degree of associative dependence between independently measured values in an ordered (in situ or temporally) set, determined in samples selected at positions with different coordinates in a sample space. An example of such space–time dynamics: Gould (1997) showed that $\approx 80\%$ of the diffusion of HIV in US is highly correlated with the air passenger traffic (origin-destination) matrix for 102 major urban centers.

Spatial analysis considers spatially distributed data as *a priori* dependent one on another, unless proven otherwise. A significant degree of spatial dependence gives a higher degree of precision for the central value of a set of measured values; so, testing for spatial dependence is especially suitable in geostatistics (mineral exploration, mining, mineral processing, etc.).

• Space syntax

Space syntax is a set of theories and techniques (cf. Hiller and Hanson 1984) for the analysis of spatial configurations complementing traditional transport engineering and geographic accessibility analysis in Geographical Information System.

It breaks down space into components, analyzed as networks of choices, and then represents it by maps and graphs describing the relative connectivity and integration of parts. The basic notions of space syntax are, for a given space:

Isovist (or visibility polygon), i.e., the field of view from any fixed point Axial line, i.e., longest line of sight and access through open space

Convex space, i.e., the maximal inscribed convex polygon (all points within it are visible to all other points within it)

These components are used to quantify how easily a space is navigable, for the design of settings where wayfinding is a significant issue, such as museums, airports, hospitals, etc. Space syntax has also been applied to predict the correlation between spatial layouts and social effects such as crime, traffic flow, sales per unit area, etc.

• Go the distance

Go the (full) distance is a general distance idiom meaning to continue to do something until it is successfully completed.

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It is, for example, the name of a US Masters Swimming fitness event and a song from the 1997 Disney animated feature *Hercules*.

• Sabbath distance

The **Sabbath distance** (or *rabbinical mile*) is a **range distance**: 2,000 Talmudic cubits (1,120.4 m, cf. **cubit** in Chap. 27) which an observant Jew should not exceed in a public thoroughfare from any given private place on the Sabbath day.

Other Talmudic length units are: day's march, parsa, stadium (40, 4, $\frac{4}{5}$ of the rabbinical mile, respectively), and span, hasit, hand-breath, thumb, middle finger, little finger ($\frac{1}{2}$, $\frac{1}{3}$, $\frac{1}{6}$, $\frac{1}{24}$, $\frac{1}{30}$, $\frac{1}{36}$ of the Talmudic cubit, respectively).

• Galactocentric distance

A star's galactocentric distance (or galactocentric radius) is its range distance from the Galactic Center; it may also refer to a distance between two galaxies. The Sun's galactocentric distance is $\approx 8.5 \,\mathrm{kpc}$, i.e., $27,700 \,\mathrm{light}$ -years.

• Cosmic light horizon

The **cosmic light horizon** (or *age of the Universe*, or, in Chaps. 26 and 27, **Hubble distance**) is an increasing **range distance**: the maximum distance that a light signal could have traveled since the Big Bang, the beginning of the Universe.

At present, it is 13–14 \times 10^9 light-years, i.e., about 13 \times 10^{60} Planck lengths.

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